



# Determination of the atmospheric lifetime and global warming potential of sulfur hexafluoride using a three-dimensional model

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**Abstract.** We have used the Whole Atmosphere Community Climate Model (WACCM), with an updated treatment of loss processes, to determine the atmospheric lifetime of sulfur hexafluoride (SF<sub>6</sub>). The model includes the following SF<sub>6</sub> removal processes: photolysis, electron attachment and reaction with mesospheric metal atoms. The Sodankylä Ion Chemistry (SIC) model is incorporated into the standard version of WACCM to produce a new version with a detailed *D* region ion chemistry with cluster ions and negative ions. This is used to determine a latitude- and altitude-dependent scaling factor for the electron density in the standard WACCM in order to carry out multi-year SF<sub>6</sub> simulations. The model gives a mean SF<sub>6</sub> lifetime over an 11-year solar cycle ( $\tau$ ) of 1278 years (with a range from 1120 to 1475 years), which is much shorter than the currently widely used value of 3200 years, due to the larger contribution (97.4 %) of the modelled electron density to the total atmospheric loss. The loss of SF<sub>6</sub> by reaction with mesospheric metal atoms (Na and K) is far too slow to affect the lifetime. We investigate how this shorter atmospheric lifetime impacts the use of SF<sub>6</sub> to derive stratospheric age of air. The age of air derived from this shorter lifetime SF<sub>6</sub> tracer is longer by 9 % in polar latitudes at 20 km compared to a passive SF<sub>6</sub> tracer. We also present laboratory measurements of the infrared spectrum of SF<sub>6</sub> and find good agreement with previous studies. We calculate the resulting radiative forcings and efficiencies to be, on average, very similar to those re-

ported previously. Our values for the 20-, 100- and 500-year global warming potentials are 18 000, 23 800 and 31 300, respectively.

## 1 Introduction

Sulfur hexafluoride (SF<sub>6</sub>) is an anthropogenic greenhouse gas which is mainly used as an electrical insulator, with other applications as a quasi-inert gas. Although its main sources are in the Northern Hemisphere, its atmospheric abundance is increasing globally in response to these emissions and its long atmospheric lifetime (defined as the ratio of the atmospheric burden of a trace gas to its rate of loss from the atmosphere). SF<sub>6</sub> is characterized by large absorption cross sections for terrestrial infrared radiation such that the presently increasing SF<sub>6</sub> abundance will contribute a positive radiative forcing over many centuries. The important known removal sources are electron attachment and photolysis. Recently, Totterdill et al. (2015) have also measured bimolecular rate constants for the reaction of SF<sub>6</sub> with mesospheric metals.

Harnisch and Eisenhauer (1998) reported that SF<sub>6</sub> is naturally present in fluorites, and outgassing from these materials leads to a natural background atmospheric abundance of 0.01 pptv. However, at present the anthropogenic emissions of SF<sub>6</sub> exceed the natural ones by a factor of 1000 or more

and are responsible for the rapid increase in its atmospheric abundance. Surface measurements show that atmospheric SF<sub>6</sub> has increased steadily since the 1980s (Geller et al., 1997; Maiss and Brenninkmeijer, 1998) and is still increasing fairly linearly. The current global mean surface mixing ratio is 8.60 pptv, which is increasing by about 0.33 pptv yr<sup>-1</sup> (~ 4 % yr<sup>-1</sup>; Dlugokencky et al., 2016).

SF<sub>6</sub> provides a useful tracer of atmospheric transport in both the troposphere and stratosphere. Rates for transport of pollutants into, within and out of the stratosphere are important parameters that regulate stratospheric composition. The basic characteristics of the stratospheric Brewer–Dobson circulation (BDC) are known from observations of trace gases such as SF<sub>6</sub>: air enters the stratosphere at the tropical tropopause, rises at tropical latitudes, and descends at middle and high latitudes to return to the troposphere. Understanding the rate of this transport on a global scale is crucial in order to predict the response of stratospheric ozone to climatic or chemical change. SF<sub>6</sub> is essentially inert in the troposphere to middle stratosphere and is removed by electron attachment and photolysis in the upper stratosphere and mesosphere (Ravishankara et al., 1993). This tracer therefore provides an ideal probe of transport on timescales of importance in the stratospheric circulation and quantitative information on mean air mass age for the lower and middle stratosphere.

The mean age of air (AoA) is the interval between the time when the volume mixing ratio of a linearly increasing atmospheric tracer reaches a certain value at a given location in the stratosphere and an earlier time when this mixing ratio was reached at a reference location. Mean AoA is expressed as (Hall and Plumb, 1994; Waugh and Hall, 2002)

$$\text{AoA} = t(\chi, l, z) - t(\chi, l_0, z_0), \quad (1)$$

where  $t$  is time,  $\chi$  is the volume mixing ratio,  $l$  and  $z$  are latitude and altitude, and the 0 subscripts denote the reference latitude and altitude, which are chosen to be the upper tropical troposphere (for the purposes of our simulations we chose the grid point latitude = 1° N, altitude = 13.9 km). In principle the trend of AoA can be used to diagnose changes in the strength of the Brewer–Dobson circulation (BDC); in practice, however, it is very difficult to obtain unambiguous results on trends from this or any other trace gas (Garcia et al., 2011). Ideally, AoA should be determined experimentally using a tracer with a very small (or zero) chemical sink in the stratosphere or mesosphere. Otherwise, a correction must be applied to account for this loss. A correction would also be necessary for any non-linear tropospheric growth. However, for the period considered for diagnosing age of air in this paper (2002–2007), the growth of SF<sub>6</sub> is approximately linear, so we can reasonably neglect such a correction for SF<sub>6</sub>-derived AoA (Hall and Plumb, 1994).

Ravishankara et al. (1993) reported the atmospheric lifetime of SF<sub>6</sub> to be 3200 years by considering electron attachment and vacuum ultraviolet (VUV) photolysis. They also

studied the loss of SF<sub>6</sub> by reaction with O(<sup>1</sup>D) but found the rate too slow to be important. They deduced that electron attachment was the dominant loss process and quantified this process using a 2-D model, wherein they assumed that all SF<sub>6</sub> molecules are destroyed after attachment of an electron (with a rate constant of 10<sup>-9</sup> cm<sup>3</sup> molecule<sup>-1</sup> s<sup>-1</sup>). They therefore argued that their lifetime of 3200 years could be a lower limit, but clearly this result depends on the accuracy of the 2-D electron density, which was calculated using only photochemistry. Morris et al. (1995) subsequently extended the work of Ravishankara et al. (1993) by including an ion chemistry module in the same 2-D model. They also made other assumptions to maximize the impact of electron attachment on SF<sub>6</sub> loss and derived a lifetime as short as 800 years (which could be further sporadically decreased by large solar proton events). Using a 3-D middle atmosphere model, Redmann et al. (2001) estimated the lifetime to be 472 years when SF<sub>6</sub> is irreversibly destroyed purely by direct electron attachment and to be 9379 years when SF<sub>6</sub> loss is assumed to occur only via indirect loss (via the formation of SF<sub>6</sub><sup>-</sup>) and ionization via the reactions with O<sub>2</sub><sup>+</sup> and N<sub>2</sub><sup>+</sup>. They concluded that the estimated lifetime depends strongly on the electron attachment mechanism because the efficiency of this process as a permanent removal process of SF<sub>6</sub> depends on the competition between the reaction of SF<sub>6</sub><sup>-</sup> with H and HCl and the photodetachment and reaction with O and O<sub>3</sub>. Here we extend the above studies and investigate the SF<sub>6</sub> lifetime using a state-of-the-art 3-D chemistry climate model with a domain from the surface to 140 km. Our modelled electron density is based on results of a detailed ion chemistry model, and we use a detailed methodology for treating the atmospheric background electrons, which is based on Troe's formalism (Troe et al., 2007a, b; Viggiano et al., 2007).

In addition to determining the SF<sub>6</sub> lifetime, in this study we report new measurements of the infrared absorption cross sections for SF<sub>6</sub> and input these into a line-by-line radiative transfer model in order to obtain radiative forcings and efficiencies. These values are then used to determine more accurate values of global warming potentials (GWPs) based on their cloudy-sky adjusted radiative efficiencies. GWP is the metric used by the World Meteorological Organisation (WMO) and Intergovernmental Panel on Climate Change (IPCC) to compare the potency of a greenhouse gas relative to an equivalent emission of CO<sub>2</sub> over a set time period. The definitions of these radiative terms are given in Appendix A.

## 2 Methodology

### 2.1 WACCM 3-D model

To simulate atmospheric SF<sub>6</sub> we have used the Whole Atmosphere Community Climate Model (WACCM). Here we use WACCM 4 (Marsh et al., 2013), which is part of the NCAR Community Earth System Model (CESM; Lamarque et al.,

**Table 1.** SF<sub>6</sub> loss reactions included in WACCM.

Loss process	Rate constant	Reference and comments
Na + SF <sub>6</sub>	$k = 1.80 \times 10^{-11} \exp(-590.5/T)$	From Totterdill et al. (2015); refitted for mesospheric temperatures 215–300 K.
K + SF <sub>6</sub>	$k = 13.4 \times 10^{-11} \exp(-860.6/T)$	From Totterdill et al. (2015); refitted for mesospheric temperatures 215–300 K.
Electron attachment	Associative attachment: $k_{\text{EA,ass}} = k_{\text{at}} \times (k_{(\text{SF}_6^- + \text{H})}[\text{H}] + k_{(\text{SF}_6^- + \text{HCl})}[\text{HCl}]) / (j_{\text{PD}} + k_{(\text{SF}_6^- + \text{H})}[\text{H}] + k_{(\text{SF}_6^- + \text{HCl})}[\text{HCl}] + k_{(\text{SF}_6^- + \text{O}_3)}[\text{O}_3] + k_{(\text{SF}_6^- + \text{O})}[\text{O}])$ Dissociative attachment: $k_{\text{EA,diss}} = k_{\text{at}} \times \beta$ , where $\beta$ is the fraction of SF <sub>6</sub> <sup>−</sup> that dissociates into SF <sub>5</sub> <sup>−</sup> .	Totterdill et al. (2015).
Photolysis	Lyman $\alpha$ : $\sigma$ (121.6 nm) = $1.37 \times 10^{-18} \text{ cm}^2$ Parameterized expression over the range of 115–180 nm, based on previous measurements.	Totterdill et al. (2015).

2012), configured to have 88 pressure levels from the surface to the lower thermosphere ( $5.96 \times 10^{-6}$  Pa, 140 km) and a horizontal resolution of  $1.9^\circ \times 2.5^\circ$  (latitude  $\times$  longitude). The model contains a detailed treatment of middle atmosphere chemistry including interactive treatments of Na and K (Plane et al., 2015). We use the specified dynamics (SD) version of the model to allow comparison with observations (see Garcia et al., 2014, for details). The SF<sub>6</sub> surface emission flux and initial global vertical profiles were taken from a CCM1 (Chemistry Climate Model Initiative) simulation using the same version of SD-WACCM with the same nudging analyses (D. Kinnison, personal communication, 2013).

Lyman- $\alpha$  photolysis is the only SF<sub>6</sub> loss reaction in the standard version of WACCM, and in this work we have added the additional processes given in Table 1. The rate constants for the SF<sub>6</sub> + metal reactions have been measured in our laboratory for mesospheric conditions (Totterdill et al., 2015); here we use the experimental values for the reactions with Na and K. For the photolysis of SF<sub>6</sub> we used the standard WACCM methodology but with the updated Lyman- $\alpha$  cross section from our laboratory of  $1.37 \times 10^{-18} \text{ cm}^2 \text{ molecule}^{-1}$  (Totterdill et al., 2015). The WACCM Lyman- $\alpha$  flux is taken from Chabrilat and Kockarts (1997).

Electron attachment to SF<sub>6</sub> plays a major role in its atmospheric removal and so both dissociative and non-dissociative attachment are considered in this study. The detailed method is described in a recent paper (Totterdill et al., 2015) and here only a brief summary is given. The removal process by the attachment of low-energy electrons to SF<sub>6</sub> can be described using Troe's theory (Troe et al., 2007a, b; Viggiano et al., 2007). In the middle and lower mesosphere, electrons are mostly attached to neutral species in the form of anions. However, above 80 km the concentration of free electrons increases and the direct electron attachment to SF<sub>6</sub> becomes more likely. This can happen either by associative attachment forming the SF<sub>6</sub><sup>−</sup> anion, which can then undergo

chemical reactions with H, O, O<sub>3</sub> and HCl, or by dissociative attachment forming the SF<sub>5</sub><sup>−</sup> anion fragment. The probability  $\beta$  of dissociative attachment when an electron is captured by SF<sub>6</sub> is given by

$$\beta(p, T) = \frac{k_{\text{dis}}}{k_{\text{at}} + k_{\text{dis}}}, \quad (2)$$

where  $k_{\text{dis}}$  is the rate constant for dissociative attachment and  $k_{\text{at}}$  is the rate constant for associative attachment.  $\beta$  can be expressed as

$$\beta(p, T) = \exp(-4587T + 7.74) \times 10^{[4.362 - 0.582 \log_{10}(p/\text{Torr}) - 0.0203(\log_{10}(p/\text{Torr}))^2 / 5.26 \times 10^{-4}]}, \quad (3)$$

where  $T$  is the temperature in K and  $p$  is the pressure in Torr (Totterdill et al., 2015).

We include both associative and dissociative electron attachment using WACCM-predicted electron concentrations (see Table 1). Note that the SF<sub>6</sub><sup>−</sup> anion is not modelled directly. As it is short-lived, we can treat it as being in steady state and consider the relative rates of its destruction pathways. Therefore, the net SF<sub>6</sub> attachment loss rate is calculated by multiplying  $k_{\text{at}}$  by the ratio of permanent destruction of the resulting SF<sub>6</sub><sup>−</sup> (reactions of SF<sub>6</sub><sup>−</sup> with H and HCl) to the total sum of these reactions plus processes which recycle SF<sub>6</sub><sup>−</sup> to SF<sub>6</sub> (reactions with O and O<sub>3</sub>, and photodetachment). The SF<sub>6</sub><sup>−</sup> photodetachment rate of  $0.11 \text{ s}^{-1}$  at midday was estimated using photodetachment cross sections from Eisfeld (2011; 280–376 nm) and Bopp et al. (2007; 376–707 nm). Note there is a typographical error in Totterdill et al. (2015) where the rate is stated as  $1.1 \text{ s}^{-1}$ .

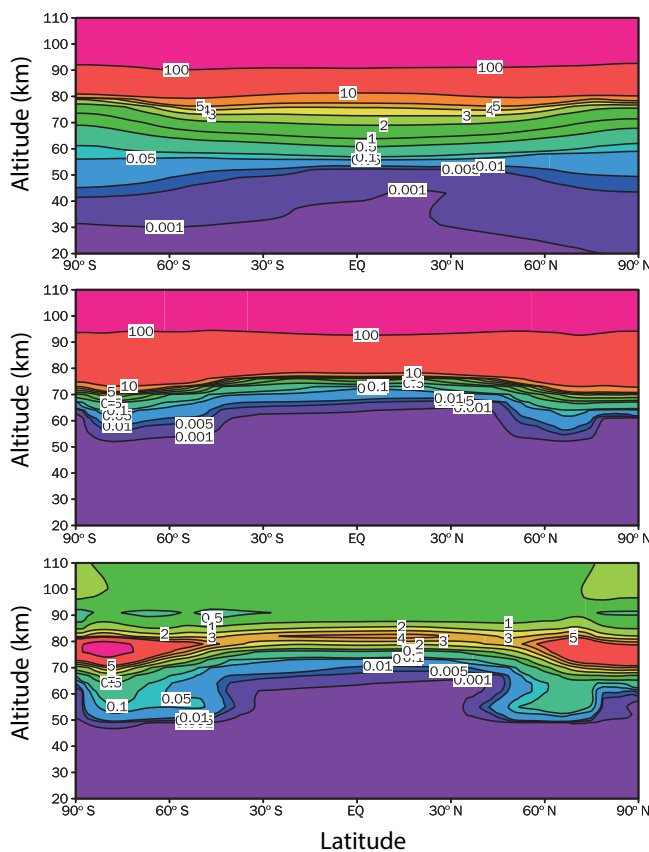
In order to use a realistic electron concentration, the role of negative ions in the  $D$  region must be considered. The  $D$  region is the lowest part of the ionosphere, extending from about 60 to 85 km. It is characterized by the appearance of cluster ions (e.g. proton hydrates  $\text{H}^+ \cdot (\text{H}_2\text{O})_n$ , where

**Table 2.** Positive and negative ions included in WACCM-SIC.

Positive ions	$O_2^+$ , $O_4^+$ , $NO^+$ , $NO^+(H_2O)$ , $O_2^+(H_2O)$ , $H^+(H_2O)$ , $H^+(H_2O)_2$ , $H^+(H_2O)_3$ , $H^+(H_2O)_4$ , $H^+(H_2O)_5$ , $H^+(H_2O)_6$ , $H_3O^+(H_2O)_2(CO_2)$ , $H_3O^+(OH)$ , $O_2^+(CO_2)$ , $H_3O^+(OH)(CO_2)$ , $H_3O^+(OH)(H_2O)$ , $O_2^+(H_2O)(CO_2)$ , $O_2^+(H_2O)_2$ , $O_2^+(N_2)$ , $NO^+(H_2O)_2$ , $H^+(H_2O)(CO_2)$ , $O^+$ , $N^+$ , $N_2^+$ , $NO^+(H_2O)_3$ , $O_4^+$ , $H^+(H_2O)_2(CO_2)$ , $H^+(H_2O)_2(N_2)$
Negative ions	$O_3^-$ , $O^-$ , $O_2^-$ , $OH^-$ , $O_2^-(H_2O)$ , $O_2^-(H_2O)_2$ , $O_4^-$ , $CO_3^-$ , $CO_3^-(H_2O)$ , $CO_4^-$ , $HCO_3^-$ , $NO_2^-$ , $NO_3^-$ , $NO_3^-(H_2O)$ , $NO_3^-(H_2O)_2$ , $NO_3^-(HNO_3)$ , $NO_3^-(HNO_3)_2$ , $Cl^-$ , $ClO^-$ , $NO_2^-(H_2O)$ , $Cl^-(H_2O)$ , $Cl^-(CO_2)$ , $Cl^-(HCl)$

$n \leq 6$ ) and negative ions (e.g.  $O_2^-$ ,  $CO_3^-$  and  $NO_3^-$ ) rather than free electrons. These species predominate because the atmospheric pressure is high enough to facilitate the three-body attachment of ligand species like  $H_2O$  to positive ions and electrons to neutral molecules. Therefore, a scaling factor was introduced that converts the standard WACCM electron concentrations, which are calculated from charge balance with the five major positive  $E$  region ions ( $N^+$ ,  $N_2^+$ ,  $O^+$ ,  $O_2^+$  and  $NO^+$ ), to more realistic electron concentrations. We have recently incorporated the Sodankylä Ion Chemistry (SIC) model into the standard version of WACCM to produce a new version (WACCM-SIC) containing a detailed  $D$  region ion chemistry with cluster ions and negative ions (Kovács et al., 2016). The mesospheric positive and negative ions in WACCM-SIC are listed in Table 2. The electron scaling factor in each grid box of WACCM was then defined as the annually averaged ratio of  $[e]_{\text{WACCM-SIC}} / [e]_{\text{WACCM}}$  for the year 2013, where  $[e]_{\text{WACCM-SIC}}$  is the electron density calculated from WACCM-SIC and  $[e]_{\text{WACCM}}$  from the standard WACCM.

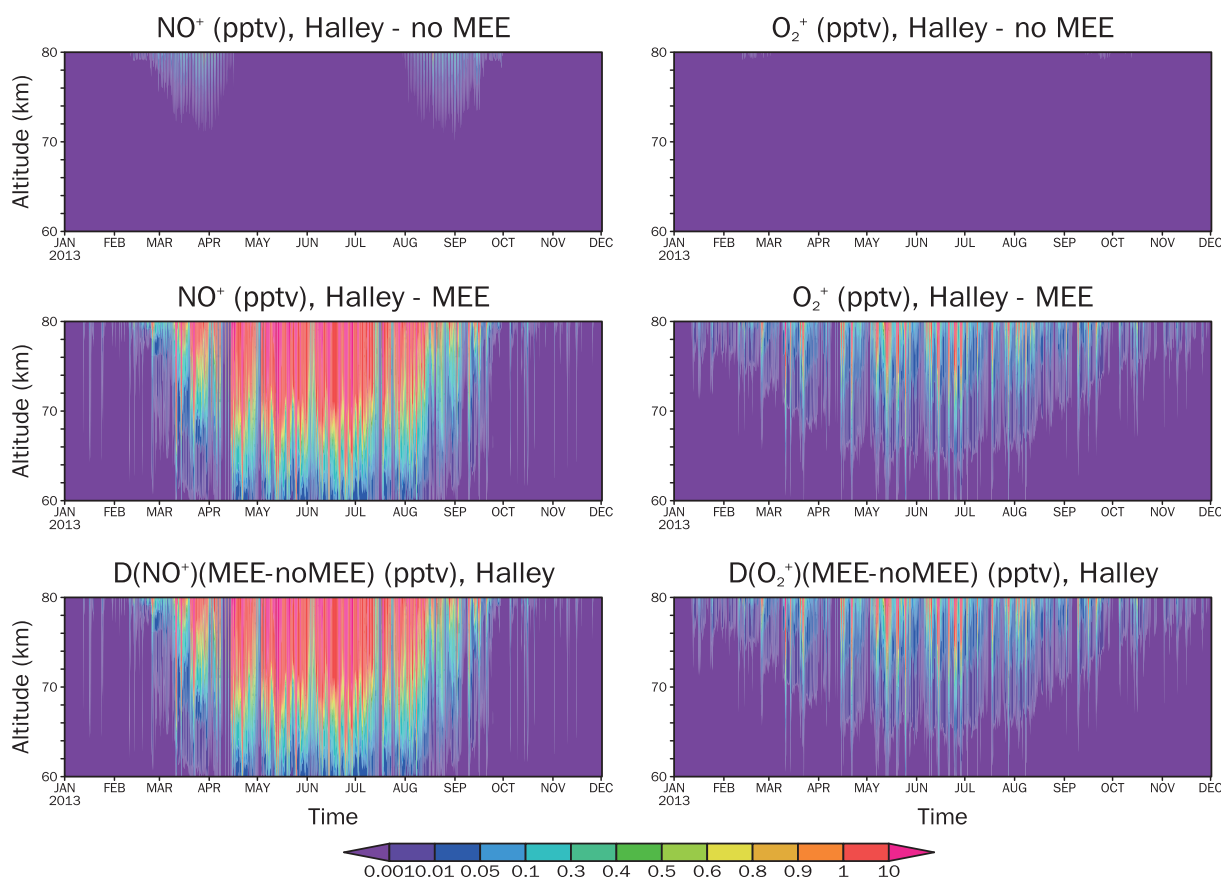
The scaling factor, which varies with altitude and latitude, is shown in Fig. 1 (bottom panel) together with the electron densities from the standard WACCM (top panel) and WACCM-SIC (middle panel) models. The annually averaged electron concentration in the WACCM-SIC model is significantly smaller in the lower and middle mesosphere than in the standard WACCM, which is expected because of negative ion formation. In the polar regions the scaling factor is around 0.01 at 50 km and 0.5 at 70 km. Note that in the upper mesosphere (70–80 km), the electron density in WACCM-SIC is larger than WACCM, by over a factor of 5 near the poles. This results from the inclusion of medium-energy electrons (MEE) (electrons with an energy between 30 keV and 2 MeV) in WACCM-SIC. Figure 2 shows the



**Figure 1.** Top: annual average electron concentration for 2013 from the standard WACCM model (in  $10^2$  electrons  $cm^{-3}$ ). Middle: annual average electron concentration for 2013 from WACCM-SIC model (in  $10^2$  electrons  $cm^{-3}$ ). Bottom: annually averaged electron scaling factor for 2013.

effect of MEE by comparing WACCM-SIC runs with and without this source of ionization in the upper mesosphere included. To describe the effect of ionization, WACCM-SIC uses ionization rates ( $I$ ) as a function of time and pressure, which were calculated from the spectra based on the proton energy-range measurements in standard air as described by Verronen et al. (2005). According to Fig. 3 of Meredith et al. (2015), the annually averaged medium-energy electron flux for 2013 approximately corresponds to the long-term, 20-year average. This allows us to assume that the annually averaged electron density of 2013 from WACCM-SIC can be used to scale the long-term simulations using the standard WACCM aimed at determining the atmospheric lifetime of  $SF_6$ .

The WACCM simulation included five different  $SF_6$  tracers in order to quantify the importance of different loss processes. All of these tracers used the same emissions but differed in their treatment of  $SF_6$  loss reactions. One  $SF_6$  tracer included no atmospheric loss (i.e. a passive tracer). Three tracers included one of the following loss processes for  $SF_6$ : (i) reaction with mesospheric metals (Na, K), (ii) electron



**Figure 2.** Time series of volume mixing ratio profiles (pptv) of  $\text{NO}^+$  (left panels) and  $\text{O}_2^+$  (right panels) above Halley ( $76^\circ\text{S}$ ) from two WACCM-SIC simulations. Top panels show the values obtained from the model run without medium-energy electrons (MEE); the middle panels show the run with medium-energy electrons; and the bottom panels show the absolute differences between the two model runs.

attachment and (iii) UV photolysis. Finally, one “total reactive”  $\text{SF}_6$  tracer included all three loss processes. This total reactive tracer should be the most realistic and was used in the radiative forcing calculations. WACCM was run for the period 1990–2007, and the first 5 years were treated as spin-up. For the analysis the monthly mean model outputs were saved and later globally averaged for the lifetime calculations.

## 2.2 Infrared absorption spectrum and radiative forcing

Previous quantitative infrared absorption spectra of  $\text{SF}_6$  have been compared in Hodnebrog et al. (2013; their Table 12). There are differences of  $\sim 10\%$  between existing integrated cross section estimates, and the measurements cover different spectral ranges. We therefore performed a more complete set of measurements over a wider spectral range, in order to reduce uncertainty in the absorption spectrum and hence the radiative efficiency of  $\text{SF}_6$ . Measurements were taken using an experimental configuration consisting of a Bruker Fourier transform spectrometer (Model IFS/66), which was fitted with a mid-infrared (MIR) source used to generate radiation

which passed through an evacuable gas cell with optical path length 15.9 cm. The cell was fitted with KBr windows, which allow excellent transmission between 400 and  $40\,000\text{cm}^{-1}$ . The choice of source and window were selected so as to admit radiation across the mid-IR range where bands of interest are known to occur. Room temperature ( $296 \pm 2\text{K}$ ) measurements were carried out between 400 and  $2000\text{cm}^{-1}$  at a spectral resolution of  $0.1\text{cm}^{-1}$  and compiled from the averaged total of 128 scans to 32 background scans at a scanner velocity of 1.6 kHz. Gas mixtures were made using between 8 and 675 Torr of  $\text{SF}_6$  diluted up to an atmosphere using  $\text{N}_2$ , according to the method described in Totterdill et al. (2016).

Radiative forcing calculations were made using the Reference Forward Model (RFM; Dudhia, 2013), which is a line-by-line radiative transfer model based on the previous GENLN2 model (Edwards, 1987). Results obtained from this model were validated against the DISORT radiative transfer solver (Stamnes et al., 2000) included within the libRadtran (Library for Radiative Transfer) package (Mayer and Kylling, 2005). A full description of these models and parameters used alongside discussion of the treatment of clouds and model comparison is also given in Totterdill et al. (2016).

**Table 3.** Partial (reactions with electrons, photolysis and metals; K, Na) and total atmospheric lifetimes (years) of SF<sub>6</sub> from different studies. Numbers in parentheses show relative percentage contributions of loss due to the different processes.

Study	Lifetime/years			Model dimensions
	Photolysis	Electron attachment	Total	
Ravishankara et al. (1993)	13 000 (24 %)	4210 (76 %)	3200	2-D
Morris et al. (1995)	N/A	N/A	800	2-D
This work	48 000 (2.6 %)	1339 (97.4 %)	1278	3-D

### 3 Results

#### 3.1 Global distributions of SF<sub>6</sub> from WACCM simulations

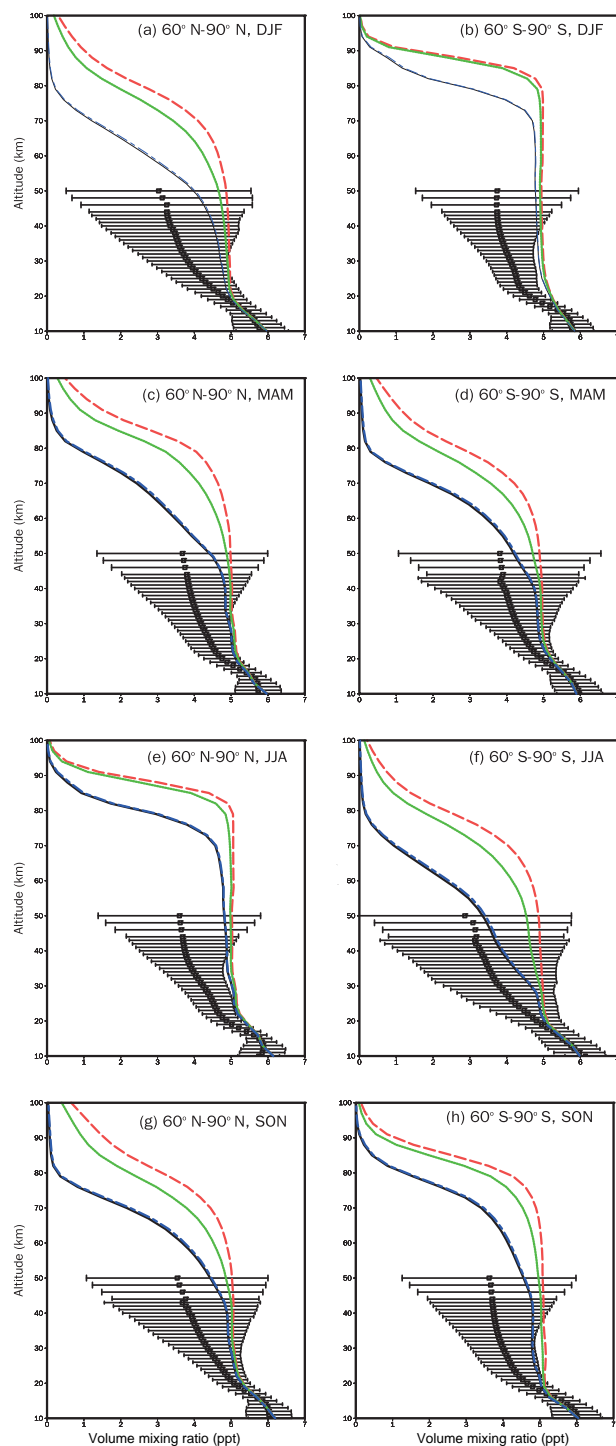
Figure 3 shows typical zonal mean profiles of the WACCM SF<sub>6</sub> tracers in the north and south polar regions for different seasons, compared to MIPAS (Michelson Interferometer for Passive Atmospheric Sounding) observations for the year 2007 (Haenel et al., 2015). Although the MIPAS SF<sub>6</sub> data provide much more coverage horizontally and vertically compared to in situ aircraft and balloon data, they have only been validated up to 35 km (Stiller et al., 2008). Validation at higher altitudes is not possible due to the lack of suitable reference data. Details of the validation of the MIPAS data version used here (V5h\_SF6\_20 for the full-resolution product from 2004 and earlier; V5r\_SF6\_222 and V5r\_SF6\_223 for the reduced-resolution period of 2005 and later) can be found in Haenel et al. (2015), including Fig. S2 of their Supplement. The WACCM passive SF<sub>6</sub> tracer has a mixing ratio profile that is almost constant with altitude in the stratosphere and lower mesosphere below 60 km and then decreases by about 10 % by 70 km, after which it decreases more rapidly. Comparison of the tracers that include loss processes show that the removal of SF<sub>6</sub> is dominated by electron attachment, with a small contribution (< 3 %) directly from photolysis. The mesospheric metals make a negligible contribution because the Na and K layers occur in the upper mesosphere above 80 km (with peaks around 90 km), and the concentrations of these metal atoms are too low. Figure 3 shows that the model mean profile agrees well with the MIPAS mean in the polar lower stratosphere (around 20 km) but exhibits a positive bias of around 1 pptv at higher altitudes in the middle stratosphere. However, the figure also shows that the variability in the observed SF<sub>6</sub> at high latitudes is large. The time variation of modelled SF<sub>6</sub> shown in Fig. 4 corresponds to an emission rate (slope) of  $6.5 \times 10^{-3} \text{ Tg yr}^{-1}$ , i.e. a  $0.29 \text{ pptv yr}^{-1}$  increase in global mean volume mixing ratio, and a volume mixing ratio of 6.4 pptv by the end of 2007.

Figure 5 shows the zonal mean annual mean SF<sub>6</sub> distribution from the five WACCM tracers and MIPAS observations for 2007. Figure 5a (and Fig. 3) shows that there is a rapid decrease in SF<sub>6</sub> above 75 km even for the inert tracer. This can be explained by diffusive separation, which becomes pro-

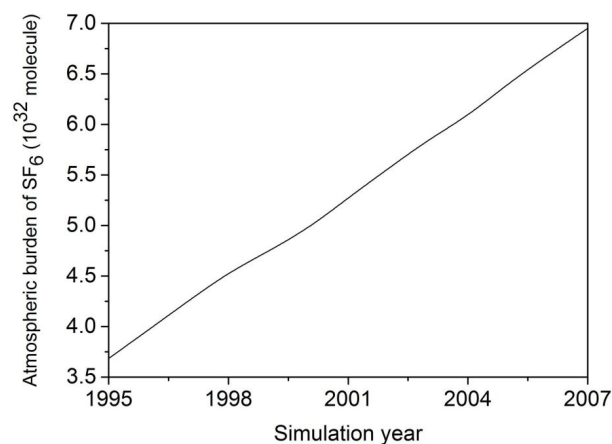
nounced in the upper mesosphere because SF<sub>6</sub> is a relatively heavy molecule (molar mass 0.146 kg) compared to the mean mass of air molecules (mean molar mass 0.0288 kg; cf. Garcia et al. 2014, where similar behaviour is seen for CO<sub>2</sub>, another relatively heavy molecule). Panels a–c of the figure all show SF<sub>6</sub> decreasing above  $\sim 80$  km, and panels a and c are almost identical, while in panel b the decrease begins about 4 km lower. This is all consistent with the notion that metals do not affect SF<sub>6</sub> and photolysis contributes only slightly. The fact that diffusive separation prevents SF<sub>6</sub> from reaching altitudes where photolysis is faster must be contributing to the very long lifetime (48 000 years, Table 3) found when photolysis is the only loss considered. By contrast, in Fig. 5d SF<sub>6</sub> decreases rapidly above 70 km, which is related to the fact that loss via electron attachment is important at these lower altitudes. Thus, in this case, SF<sub>6</sub> loss occurs below the altitudes ( $\sim 90$  km) where diffusive separation is important (and where air density is higher), which makes it a much more effective loss mechanism. The WACCM SF<sub>6</sub> tracer that includes all loss processes (Fig. 5e) has a very similar distribution to that which only treats loss due to electron attachment (Fig. 5d), which emphasizes how this process dominates SF<sub>6</sub> loss in the model. This model tracer can be compared to the MIPAS observations in Fig. 5f, which shows that WACCM appears to reproduce the general features of the MIPAS distribution (note the smaller altitude range in panels e and f of Fig. 5). However, it is also clear that WACCM SF<sub>6</sub>, even with all losses considered, decreases with altitude much more slowly at all latitudes than MIPAS SF<sub>6</sub>. This could indicate a problem with the model's meridional transport. However, a BDC that is too fast would tend to produce low levels of SF<sub>6</sub> at middle and high latitudes in the descending branch, which does not seem to be the case. Therefore, at least two other possible scenarios could be responsible for the discrepancy: SF<sub>6</sub> loss in WACCM is still somewhat underestimated despite the inclusion of the electron attachment, or MIPAS SF<sub>6</sub> is biased low above  $\sim 20$  km.

#### 3.2 Atmospheric lifetime

The atmospheric lifetime is defined as the ratio of the atmospheric burden to the atmospheric loss rate. This definition was used to calculate annual mean lifetime values from the WACCM output containing the individual rates for the



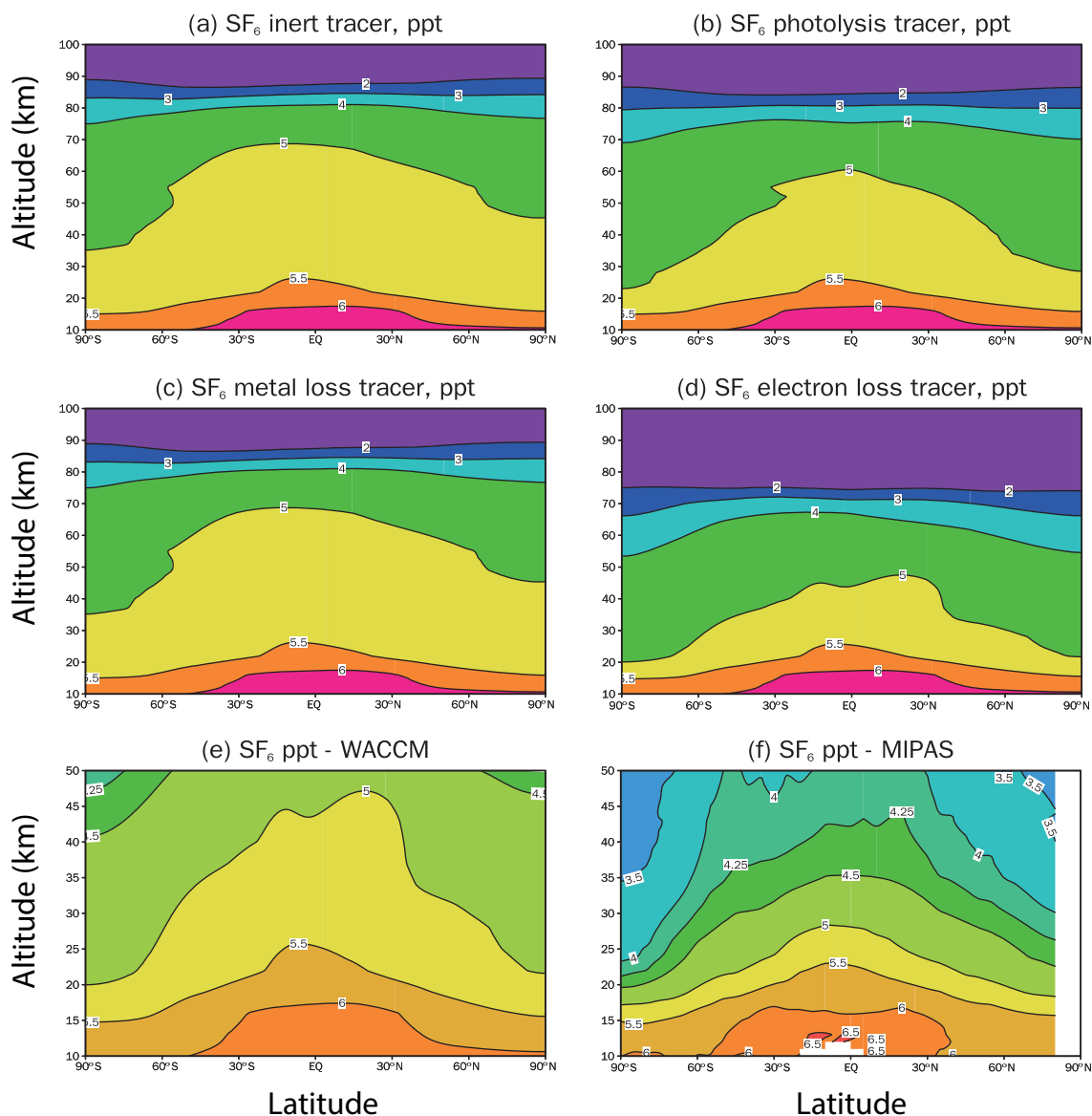
**Figure 3.** Seasonal mean volume mixing ratios (pptv) of the different SF<sub>6</sub> tracers for the polar regions (60–90° N and 60–90° S latitudes) in 2007 as a function of altitude for MIPAS (Michelson Interferometer for Passive Atmospheric Sounding) satellite-observed SF<sub>6</sub> (black symbols with standard deviations for  $\pm 1\sigma$ ; Stiller et al., 2012), the total WACCM-SF<sub>6</sub> (blue solid line), the photolysis WACCM-SF<sub>6</sub> tracer (green solid line) and the inert WACCM SF<sub>6</sub> tracer (red dashed line).



**Figure 4.** Variation of the total annual atmospheric burden of SF<sub>6</sub> during the simulation from 1995 to 2007. According to this the emission rate (slope) was determined to be  $6.5 \times 10^{-3} \text{ Tg yr}^{-1}$ , corresponding to  $0.29 \text{ pptv yr}^{-1}$ .

different loss processes. During the simulation the total atmospheric burden of SF<sub>6</sub> increased linearly as expected (see Fig. 4) from  $3.4 \times 10^{32}$  molecules with an annual increment of  $2.3 \times 10^{31}$  molecules yr<sup>-1</sup>. Figure 6 shows the variation in SF<sub>6</sub> lifetime from 1995 to 2007, corresponding to a full solar cycle (the solar minima occurred in May 1996 and January 2008). The figure demonstrates that the lifetime has a strong dependence on solar activity, being anti-correlated with solar activity, for example as measured by the radio flux at 10.7 cm (2800 MHz; Tapping, 2013), which ranges over  $(72\text{--}183) \times 10^{-22} \text{ W m}^{-2} \text{ Hz}^{-1}$ , with an average value of  $90.3 \times 10^{-22} \text{ W m}^{-2} \text{ Hz}^{-1}$ . The mean SF<sub>6</sub> lifetime over the same solar cycle period is 1278 years, with a range from 1120 to 1475 years. The annual averaged electron number density in the polar regions is also plotted in Fig. 6; as expected, it is correlated with the 10.7 cm radio emission (Tapping, 2013).

As noted in Sect. 1, the SF<sub>6</sub> lifetime has been reported to be 3200 years by Ravishankara et al. (1993). For this they used a total electron attachment rate constant of  $k_{\text{EA}} = 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ . In Morris et al. (1995) the calculated lifetime decreased to 800 years by considering ion chemistry and assuming that the associative attachment forming SF<sub>6</sub><sup>-</sup> does not regenerate the parent molecule, thereby obtaining a lower limit for the lifetime. Reddmann et al. (2001) estimated the lifetime to be 472 years when SF<sub>6</sub> is irreversibly destroyed purely by direct electron attachment and to be 9379 years when SF<sub>6</sub> loss is assumed to occur only via indirect loss (via the formation of SF<sub>6</sub><sup>-</sup>) and ionization via the reactions with O<sub>2</sub><sup>+</sup> and N<sub>2</sub><sup>+</sup>. In the present study we have directly applied Troe's theory (Troe et al., 2007a, b; Viggiano et al., 2007) to determine the efficiency of electron attachment as a function of temperature and pressure and the branching



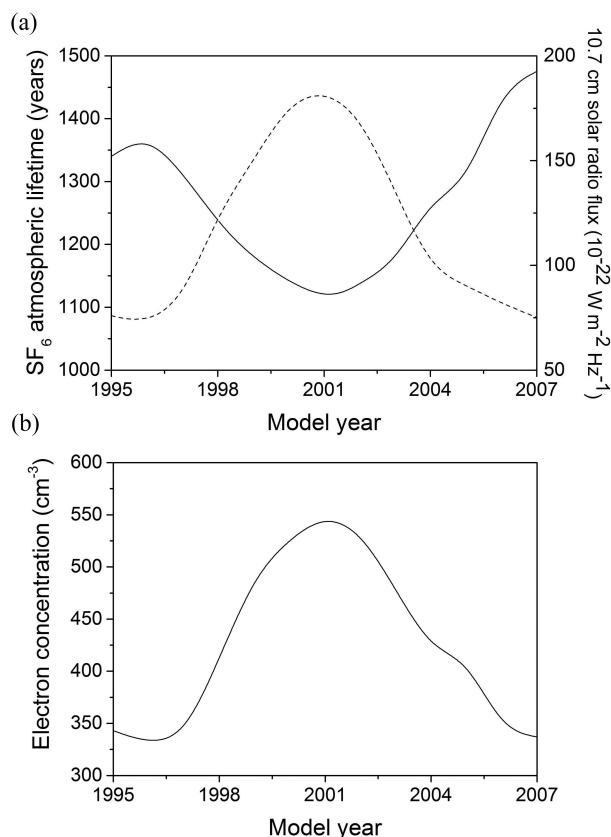
**Figure 5.** Annual zonal mean latitude–height volume mixing ratios (pptv) of the different WACCM SF<sub>6</sub> tracers in 2007: **(a)** inert SF<sub>6</sub> tracer; **(b)** SF<sub>6</sub> tracer removed by photolysis only; **(c)** SF<sub>6</sub> tracer removed by mesospheric metals only; **(d)** SF<sub>6</sub> tracer removed by electron attachment only; and **(e)** total reactive SF<sub>6</sub>. Panel **(f)** shows the SF<sub>6</sub> volume mixing ratio for 2007 from MIPAS observations. Note the different altitude ranges and contour intervals for panels **(a)–(d)** versus panels **(e)–(f)**.

ratio for dissociative attachment (Eq. 2), which we extrapolated to mesospheric conditions (Totterdill et al., 2015).

Our estimated partial lifetime of SF<sub>6</sub> due to photolysis (i.e. the lifetime calculated considering photolysis as the only atmospheric loss process) for the SF<sub>6</sub> tracer which includes all loss processes is 48 000 years, which is considerably longer than that the 13 000 years determined by Ravishankara et al. (1993) despite our Lyman- $\alpha$  cross section ( $1.37 \times 10^{-18} \text{ cm}^2$ , Table 1) being only  $\sim 22\%$  smaller than the value measured by Ravishankara et al. ( $1.76 \times 10^{-18} \text{ cm}^2$ ). One reason why our photolysis-related partial lifetime is longer is that WACCM includes dif-

fusive separation, which was not described in the earlier 2-D model study. The inclusion of diffusive separation reduces sharply the abundance of SF<sub>6</sub> at high altitudes, where photolysis is most effective. Another contributing factor could be that the VUV photolysis is important only above 80 km, while in our model runs SF<sub>6</sub> is mostly destroyed by electron attachment, which results in less being transported into this upper-mesospheric region. When we analyse our WACCM SF<sub>6</sub> tracer which is subject to photolysis loss only, the resulting steady-state overall lifetime (i.e. lifetime calculated using the rates of all loss processes) for the last model year (2007) is 17 200 years, which is only 32 % larger than the





**Figure 6.** (a) Variation in atmospheric lifetime of SF<sub>6</sub> (solid line) and 10.7 cm solar radio flux (dashed line) during the WACCM simulation. (b) Variation of the WACCM electron concentration (cm<sup>-3</sup>) at 80 km, averaged over polar latitudes (60–90° N and 60–90° S).

value of Ravishankara et al. (1993) and thus more consistent with the difference in the Lyman- $\alpha$  cross sections. Finally, if we do not include the electron scaling factor to reduce the electron density below 80 km due to negative ion formation, then the SF<sub>6</sub> lifetime decreases to 776 years (not shown), which is similar to the value of 800 years obtained by Morris et al. (1995).

### 3.3 Impact of SF<sub>6</sub> loss on mean age of stratospheric air

As SF<sub>6</sub> is a chemically stable molecule in the stratosphere and troposphere and has an almost linearly increasing tropospheric abundance, its atmospheric mixing ratio is often used to determine the mean age of stratospheric air. This is an important metric in atmospheric science as the distribution of ozone and other greenhouse gases depends significantly on the transport of air into, within, and out of the stratosphere. WACCM contains an idealized, linearly increasing age of air tracer (AOA1) that provides model age values for model experiments (Garcia et al., 2011).

Age of air has generally been derived from observations by treating SF<sub>6</sub> as a passive (non-reactive) tracer. The assump-

tion is that the global loss rate is too slow to significantly affect the lifetime. This was confirmed by Garcia et al. (2011) when only photolysis was included. However, when loss via electron attachment is also considered, the lifetime may become short enough that this assumption is no longer valid, in which case the stratospheric mixing ratio would appear to correspond to an earlier tropospheric mixing ratio than in reality. We have compared the passive WACCM SF<sub>6</sub> tracer with that subject to all loss processes, which yields the new lifetime of 1278 years. The difference between these two tracers indicates the error in the derived age of air that would arise in the real atmosphere if SF<sub>6</sub> is assumed to be a passive tracer. The error caused by chemical removal can be expressed as

$$\Delta(\text{AoA}) = \text{AoA}(\text{reactive tracer}) - \text{AoA}(\text{passive tracer}), \quad (4)$$

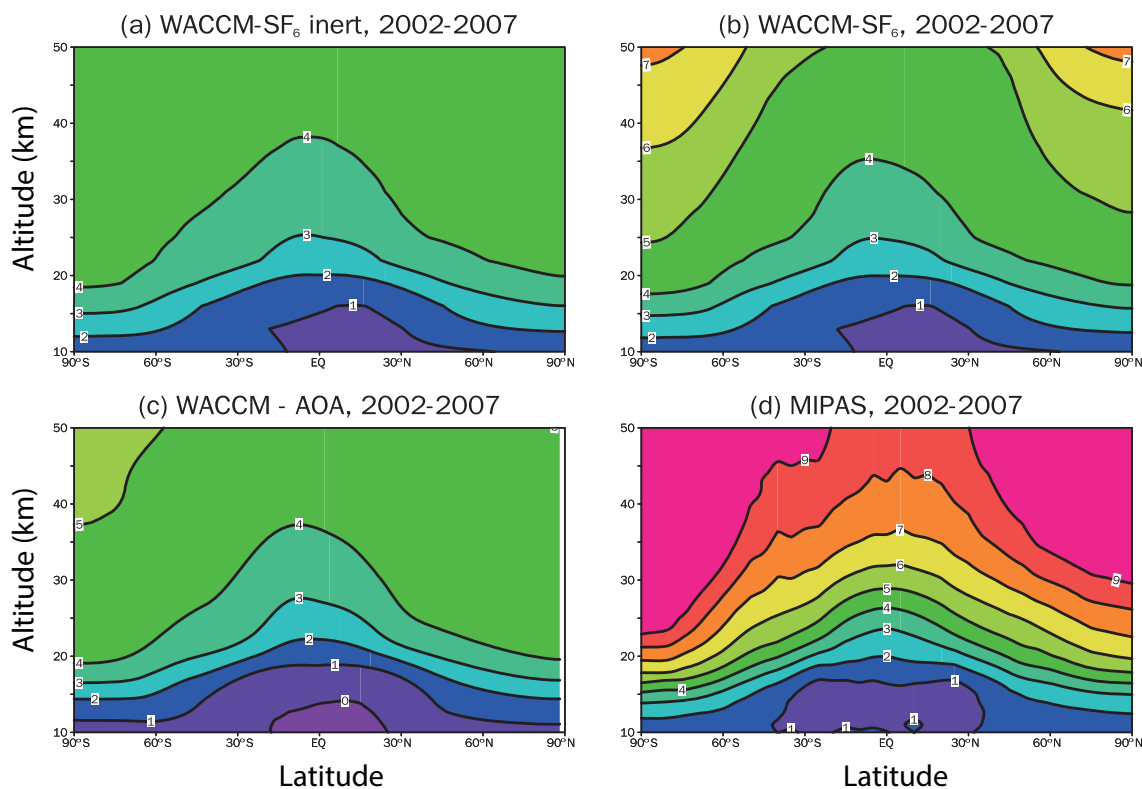
where  $\Delta(\text{AoA})$  is the difference in the age of air value caused by chemical loss,  $\text{AoA}(\text{reactive tracer})$  is the calculated age of air considering the chemical removal and  $\text{AoA}(\text{passive tracer})$  is the value obtained from a non-reactive tracer. The expression for the age of air at any point in the stratosphere can be obtained from a simplified version of (Eq. 1) that is derived from a Taylor series expansion, retaining only the linear term. It is then expressed as

$$\text{AoA} = [(\chi_0(\text{SF}_6) - \chi(\text{SF}_6))/r(\text{SF}_6)], \quad (5)$$

where  $\chi(\text{SF}_6)$  and  $\chi_0(\text{SF}_6)$  are the SF<sub>6</sub> volume mixing ratios at the actual and the reference (tropical tropopause) points, respectively, while  $r(\text{SF}_6)$  is the rate of increase in tropospheric SF<sub>6</sub>. In our simulations  $r(\text{SF}_6)$  is 0.29 pptv yr<sup>-1</sup> (Fig. 4), which is an approximation as the growth rate is not constant in reality. Stiller et al. (2012) report a value of 0.24 pptv yr<sup>-1</sup> based on observations over the MIPAS period. These two simplifications will lead to deviations between WACCM and MIPAS age data. If Eq. (5) is substituted into Eq. (4) then the error in age will be

$$\Delta(\text{AoA}) = (\chi(\text{SF}_6, \text{passive}) - \chi(\text{SF}_6, \text{reactive}))/r(\text{SF}_6). \quad (6)$$

This error, along with the mean age itself was calculated from WACCM output for 2007. Figure 7 shows the annual mean ages determined from the WACCM simulation from 2002 to 2007 using the total reactive and the inert SF<sub>6</sub> tracers and the idealized AOA1 age tracer. There is a clear difference between the age values derived from the passive SF<sub>6</sub> and the idealized AoA tracer. If Eq. (5) is used to determine the age values, there is no guarantee that the age values derived from the two tracers will be identical; the rate was determined from the increase in the SF<sub>6</sub> burden (0.29 pptv yr<sup>-1</sup>) and this was provided by the linear fit (Fig. 4), which can misrepresent the growth rate at a specific time. Figure 7 also shows



**Figure 7.** Annual mean age of stratospheric air (years) for the period of 2002–2007 determined from a WACCM simulation using (a) the inert SF<sub>6</sub> tracer; (b) the total reactive SF<sub>6</sub> tracer; (c) the idealized AOA1 tracer. Panel (d) shows the age values derived for the same period from our analysis of MIPAS SF<sub>6</sub> observations.

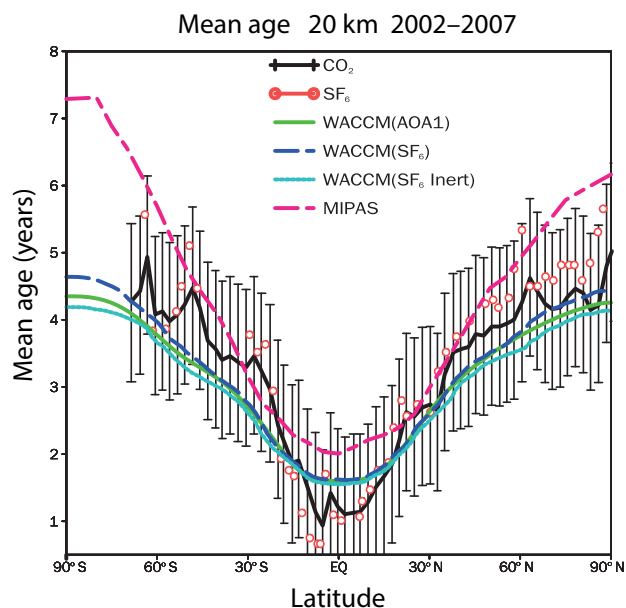
**Table 4.** Integrated absorption cross sections for SF<sub>6</sub> measured in this work and ratios with values obtained by GEISA (Gestion et Etude des Informations Spectroscopiques Atmosphériques; Jacquinet-Husson et al., 2011; Hurley, 2003; Varanasi, 2001) and HITRAN (High-resolution TRANsmision molecular absorption database; Rothman et al., 2012).

Band limits (cm <sup>-1</sup> )	Integrated cross section (10 <sup>-16</sup> cm <sup>2</sup> molec <sup>-1</sup> cm <sup>-1</sup> )	Ratio of integrated cross sections in this work to previous studies		
		Hurley (2003)	Varanasi (2001)	HITRAN
925–955	2.02	1.07	1.01	0.99
650–2000	2.40	–	1.09	–

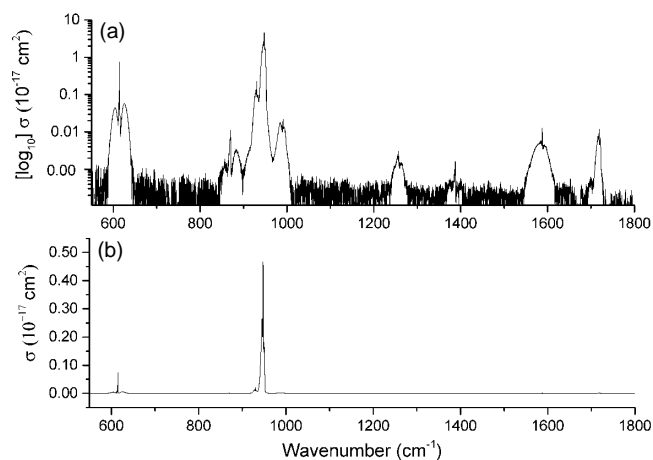
the difference between the age values obtained from the reactive and inert SF<sub>6</sub> tracers. It can be seen that consideration of the reactive SF<sub>6</sub> tracer does indeed affect the determined mean age values, mostly where electron attachment dominates. The age estimates at high latitudes are most sensitive to chemical loss because the air that reaches these locations has descended from the high altitudes where SF<sub>6</sub> loss predominantly occurs. According to the MIPAS satellite observations

(Stiller et al., 2012; Haenel et al., 2015.), the derived age value over the tropical lower stratosphere at 25 km is slightly more than 3 years, while the WACCM simulations with the reactive SF<sub>6</sub> tracer predicts 3 years. Comparing Fig. 7a and b, the effect of chemical removal in this region is minor (0.01-year or 0.5% change), and therefore it does not have much impact on the inferred atmospheric transport. At the poles the effect is much more significant; the difference at 25 km between the reactive and inert SF<sub>6</sub> tracers is up to 0.55 years (9%). In summary, in the troposphere–stratosphere at low latitudes, the effect of chemical removal is not very significant and the error on the estimated mean age caused by the assumption of SF<sub>6</sub> being a passive tracer is not important. However, the effect of chemical removal becomes more significant at high latitudes.

We can also compare modelled and observed mean age values in the lower stratosphere (20 km). Figure 8 shows the mean age profiles from WACCM tracers, ER-2 observations (Hall et al., 2009) and our analysis of MIPAS SF<sub>6</sub> satellite data at 20 km. From this it can be seen that in the tropical region the mean age values are similar between the idealized age tracer and the inert and reactive SF<sub>6</sub> tracers. This is consistent with no loss of SF<sub>6</sub> having occurred in air parcels in the deep tropics. At high latitudes there is an up to 0.5-year

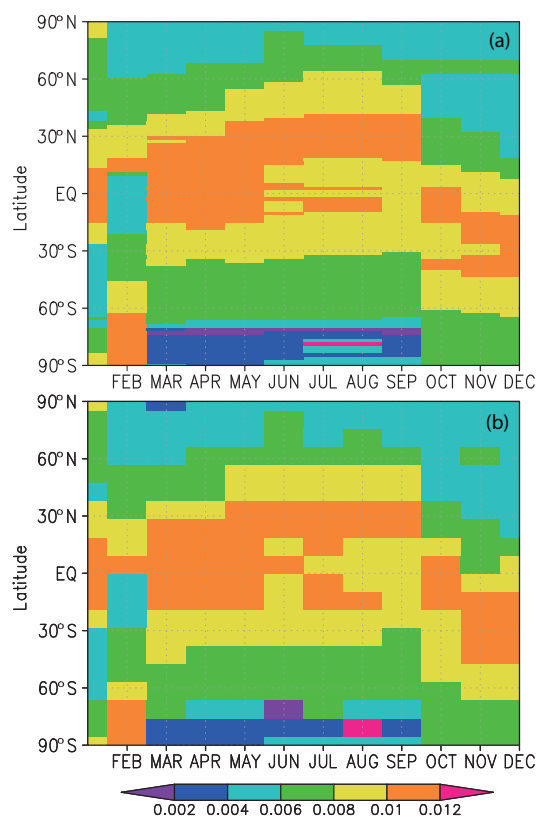


**Figure 8.** Mean age values at 20 km altitude derived from MIPAS satellite (dashed magenta line) and ER-2 aircraft observations ( $\text{SF}_6$  red open circles,  $\text{CO}_2$  black crosses; Hall et al., 1999). The error bars apply to the age derived from the ER-2 observations. Also shown is the mean age derived from WACCM tracers: reactive  $\text{SF}_6$  (dashed blue line), passive  $\text{SF}_6$  (light blue line) and AoA tracer (AOA1, solid green line).



**Figure 9.** Infrared absorption spectrum of  $\text{SF}_6$  at  $\sim 295$  K on (a) a logarithmic y axis and (b) a linear y axis. The logarithmic scale in panel (a) highlights the relative positions of the minor bands.

difference in the modelled mean ages, with the reactive  $\text{SF}_6$  tracer producing the oldest apparent age. The differences in mean age between the tracers is larger in the SH polar region than in the NH because the polar region is less well mixed. The tendency is very similar when we compare the WACCM mean ages to the MIPAS observations. Note that the satellite



**Figure 10.** Latitude–time plots for instantaneous radiative forcing ( $\text{W m}^{-2}$ ) by  $\text{SF}_6$  as a function of latitude and month at (a) high-latitude resolution ( $1.5^\circ$  spacing) and (b) low-latitude resolution ( $9^\circ$  spacing).

observations show more seasonal variability in the middle and high latitudes than in the tropics.

### 3.4 Radiative efficiency and forcing

To determine the radiative efficiency and global warming potential of  $\text{SF}_6$ , integrated cross sections were taken from two public molecular spectroscopic databases: (i) the GEISA-2009/2011 (Gestion et Etude des Informations Spectroscopiques Atmosphériques) spectroscopic database (Jacquinet-Husson et al., 2011), which uses the data of Varanasi (2001) and Hurley (2003), and (ii) the HITRAN-2012 (HIGH-resolution TRANsmission molecular absorption database) molecular spectroscopic database (Rothman et al., 2012), which uses the data of the Pacific Northwest National Laboratory Infra-Red Database (Sharpe et al., 2004). Values were also measured in this study. The literature values are presented in Table 4 for comparison with our experimentally determined values, and the full  $\text{SF}_6$  spectrum obtained in this study is given in Fig. 9. In our study the spectrometer error is  $\pm 1.0\%$  for all experiments, and the uncertainty in the sample concentrations of  $\text{SF}_6$  was calculated to be  $0.7\%$ . Spectral noise was averaged at  $\pm 5 \times 10^{-21} \text{ cm}^2 \text{ molecule}^{-1}$

**Table 5.** Calculated instantaneous and stratospheric adjusted radiative forcings and radiative efficiencies of SF<sub>6</sub> in clear and all-sky conditions.\*

	Instantaneous		Stratospheric adjusted	
	Clear	All-sky	Clear	All-sky
Radiative forcing (m W m <sup>-2</sup> )	76.43	48.91	81.81	56.01
Radiative efficiency (W m <sup>-2</sup> ppbv <sup>-1</sup> )	0.77	0.50	0.85	0.59

\* Based on present day atmospheric SF<sub>6</sub> surface concentration of 9.3 pptv.

**Table 6.** Comparison of 20-, 100- and 500-year global warming potentials for SF<sub>6</sub> from this work with values from IPCC (2013).

	Global warming potential		
	GWP <sub>20</sub>	GWP <sub>100</sub>	GWP <sub>500</sub>
This work <sup>a</sup>	18 000	23 700	31 300
IPCC (2013) <sup>b</sup>	17 500	23 500	32 600 <sup>c</sup>
Difference (%) (This work – IPCC)	+3 %	+1 %	–4 %

<sup>a</sup> Based on our atmospheric lifetime of 1278 years and radiative efficiency (RE) of 0.59 W m<sup>-2</sup> ppbv<sup>-1</sup>. <sup>b</sup> Based on an atmospheric lifetime of 3200 years and RE of 0.57 W m<sup>-2</sup> ppbv<sup>-1</sup>. <sup>c</sup> Based on an atmospheric lifetime of 3200 years and RE of 0.52 W m<sup>-2</sup> ppbv<sup>-1</sup> from IPCC AR4 (Forster et al., 2007).

per 1 cm<sup>-1</sup> band. However, at wavenumbers < 550 cm<sup>-1</sup>, towards the edge of the mid-infrared where opacity of the KBr optics increases, this value was 1 × 10<sup>-20</sup> cm<sup>2</sup> molecule<sup>-1</sup> per 1 cm<sup>-1</sup> band. The error from determining the scaling cross section was 5 %. This results in an average overall error of ±5 % in the cross sections.

The intensities of the main SF<sub>6</sub> absorption bands (925–955 cm<sup>-1</sup>) measured in this study are 7 % greater than those reported by Hurley (2003), 1 % greater than Varanasi (2001) and 1 % lower than those given in HITRAN (Rothman et al., 2012; Table 4). Comparison of our results to Varanasi (2001) between 650 and 2000 cm<sup>-1</sup> gives an agreement within 9 %. Note that these differences are within the combined error of both experiments.

The instantaneous and stratospheric adjusted SF<sub>6</sub> radiative efficiencies in clear and cloudy-sky conditions are given in Table 5. These are also presented as present-day radiative forcings employing a current surface concentration of 9.3 pptv (NOAA, 2016; see Fig. 4). The radiative efficiency was calculated in the RFM for each month between 90° S and 90° N at latitudinal resolutions (on which the data were averaged to obtain the zonal mean vertical profile) of 1.5 and 9.0°. The tropopause used the standard WMO lapse rate definition (see Totterdill et al., 2016). Figure 10 shows the seasonal–latitudinal variation of the instantaneous clear-sky radiative forcing for SF<sub>6</sub> on the high- (1.5°) and low- (9°) resolution grids. Employing profiles averaged over the lower-resolution grid gives an average forcing within 1 % of the

higher-resolution grid. Using only a single annually averaged global mean profile led to a 10 % error in radiative forcing when compared to our monthly resolved high-resolution profile, supporting the findings of Freckleton et al. (1998) and Totterdill et al. (2016).

A selection of experiments were carried out over a range of months and latitudes to investigate the sensitivity of the forcing calculations to the bands used. The average contributions from the main bands were compared to the calculation with the full measured spectrum. The results showed that the 580–640 and 925–955 cm<sup>-1</sup> bands contribute almost 99 % to the instantaneous radiative forcing. Our forcing calculations suggest that the SF<sub>6</sub> minor bands contribute only a small amount to the final value. This means that deviations between our experimentally determined spectra and those in the literature only result in a significant change to previously published radiative forcings and efficiencies when that deviation occurs over a major band.

The SF<sub>6</sub> adjusted cloudy-sky radiative efficiency published by the IPCC AR5 report and used to determine its GWP values is 0.57 W m<sup>-2</sup> ppbv<sup>-1</sup> (Myhre et al., 2013). This compares to our adjusted cloudy-sky radiative efficiency of 0.59 W m<sup>-2</sup> ppbv<sup>-1</sup> (Table 5). A review on radiative efficiencies and global warming potentials by Hodnebrog et al. (2013) provides a comprehensive list of all published values for these parameters for many species including SF<sub>6</sub>. They established the range of published radiative efficiencies for SF<sub>6</sub> to be 0.49–0.68 W m<sup>-2</sup> ppbv<sup>-1</sup>, with a mean value of 0.56 W m<sup>-2</sup> ppbv<sup>-1</sup>. They also made their own revised estimate using an average of the HITRAN (Rothman et al., 2012) and GEISA (Jacquinet-Husson et al., 2011) spectral databases and found a best estimate of (0.565 ± 0.025) W m<sup>-2</sup> ppbv<sup>-1</sup>. Their mean value for radiative efficiency is very close to that determined in this study using similar conditions (0.59 W m<sup>-2</sup> ppbv<sup>-1</sup>).

### 3.5 Global warming potential

Table 6 gives our estimates of the 20-, 100- and 500-year GWPs based on cloudy-sky adjusted radiative efficiencies of SF<sub>6</sub> compared with IPCC AR5 values (IPCC, 2013). Our 20-, 100- and 500-year global warming potentials for SF<sub>6</sub> are 18 000, 23 800 and 31 300, respectively. The 20- and 100-year values are 3 % greater and 1 % greater, respectively,

than their IPCC counterparts, and the 500-year GWP is 4 % smaller than its AR4 counterpart (Forster et al., 2007). The forcing efficiencies determined in this study are somewhat higher than the previously published values (see Sect. 3.4), which would imply a higher GWP value. However, our shorter atmospheric lifetimes would lead to a smaller GWP estimate, with a larger effect on longer time horizons when atmospheric loss becomes relevant. The trade-off between these competing effects is apparent in Table 6, where SF<sub>6</sub> exhibits a 20-year GWP that is slightly larger than the IPCC value, while the 500-year GWP is slightly smaller. The radiative efficiency effect is most obvious for the case of the 20-year GWP where, because the atmospheric lifetime of SF<sub>6</sub> is 1278 years, the species does not have time for any significant loss to occur.

#### 4 Conclusions

The 3-D Whole Atmosphere Community Climate Model was used to simulate the SF<sub>6</sub> atmospheric distribution over the period of 1995–2007. From the concentrations and the knowledge of the electron attachment, photolysis and metal reaction rates we determined the atmospheric lifetime, which shows a significant dependence (–12 to +15 %) on the solar cycle due to varying electron density. The mean SF<sub>6</sub> atmospheric lifetime over a solar cycle was determined to be 1278 years (ranging from 1120 to 1475 years), which is different to previously reported literature values and much shorter than the widely quoted value of 3200 years. The reason is our more detailed treatment of electron attachment using a new formalism to describe both associative and dissociative attachment and the use of a detailed model of *D* region ion chemistry to evaluate the partitioning of electrons and negative ions below 80 km. Further refinement to this lifetime estimate from modelling studies depends on decadal length simulations of the model with detailed ion chemistry for the realistic long-term estimation of electron density in the upper atmosphere without the need for scaling factors as employed here.

Based on this new estimate of the SF<sub>6</sub> lifetime, we find that the derived mean age of stratospheric air from observations can be slightly affected by the atmospheric removal of SF<sub>6</sub>. In the polar region the age of air values differ by up to 9 % when the values from inert and reactive model tracers are compared, suggesting that SF<sub>6</sub> loss does not have a large influence on the age values but that it should be included in detailed analyses.

We also reinvestigated the radiative efficiency and global warming potential of SF<sub>6</sub>. Our radiative efficiency value reported here,  $0.59 \pm 0.045 \text{ W m}^{-2} \text{ ppbv}^{-1}$ , is slightly higher than the IPCC AR5 estimate of  $0.57 \text{ W m}^{-2} \text{ ppbv}^{-1}$ . The global warming potentials of SF<sub>6</sub> for 20, 100 and 500 years have been determined to be 18 000, 23 800 and 31 300, respectively. We find that our revised lifetime and efficiency values cancel each other out somewhat, so overall do not play a significant role in modifying the GWP estimates on these time horizons.

#### 5 Data availability

The output from the WACCM experiments can be obtained by emailing Martyn Chipperfield.

## Appendix A

The climate metrics used in this paper are defined in the following ways.

Radiative forcing (RF) is the change in the net irradiance at the tropopause after allowing stratospheric temperatures to adjust to radiative equilibrium (expressed in watts per square metre;  $\text{W m}^{-2}$ ) due to an internal or external change in the forcing of the climate system, such as a change in the concentration of  $\text{CO}_2$  or the output of the Sun.

Radiative efficiency (RE) is the radiative forcing per kilogram of compound.

Global warming potential (GWP) is defined by the expression

$$\text{GWP} = \frac{\int_0^{\text{TH}} a_\chi [\chi(t)] dt}{\int_0^{\text{TH}} a_r [r(t)] dt}, \quad (\text{A1})$$

where TH is time horizon;  $a_\chi$  is radiative forcing due to a unit increase in atmospheric abundance of the compound ( $\text{W m}^{-2} \text{ kg}^{-1}$ ) and  $[\chi(t)]$  is its time-dependent decay in concentration following its instantaneous release at time  $t = 0$ . The denominator contains the corresponding quantities for  $\text{CO}_2$  as a reference gas (Myhre et al., 2013). GWP is the most common metric used by the WMO and IPCC to compare the potency of a greenhouse gas relative to an equivalent emission of  $\text{CO}_2$  over a set time period. GWP takes into account species lifetime. This means a species with a very high radiative forcing may still have a low GWP if it also possesses a short atmospheric lifetime. However, GWP does not account for factors such as changes in emission or the introduction of replacement species. Criticisms of GWP are discussed in greater detail by Myhre et al. (2013).

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