

Global peroxyacetyl nitrate (PAN) retrieval in the upper troposphere from limb emission spectra of the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS)

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Abstract. We use limb emission spectra of the Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) onboard the ENVIRONMENTAL SATellite (ENVISAT) to derive the first global distribution of peroxyacetyl nitrate (PAN) in the upper troposphere. PAN is generated in tropospheric air masses polluted by fuel combustion or biomass burning and acts as a reservoir and carrier of NO_x in the cold free troposphere. PAN exhibits continuum-like broadband structures in the mid-infrared region and was retrieved in a contiguous analysis window covering the wavenumber region 775–800 cm⁻¹. The interfering species CCl₄, HCFC-22, H₂O, ClONO₂, CH₃CCl₃ and C₂H₂ were fitted along with PAN, whereas pre-fitted profiles were used to model the contribution of other contaminants like ozone. Sensitivity tests consisting in retrieval without consideration of PAN demonstrated the existence of PAN signatures in MIPAS spectra obtained in polluted air masses. The analysed dataset consists of 10 days between 4 October and 1 December 2003. This period covers the end of the biomass burning season in South America and South and East Africa, in which generally large amounts of pollutants are produced and distributed over wide areas of the southern hemispheric free troposphere. Indeed, elevated PAN amounts of 200–700 pptv were measured in a large plume extending from Brasil over the Southern Atlantic, Central and South Africa, the South Indian Ocean as far as Australia at altitudes between 8 and 16 km. Enhanced PAN values were also found in a much more restricted area between northern subtropical Africa and India. The most significant northern midlatitude PAN signal was detected in an area at 8 km altitude extending from China into the Chinese Sea. The average mid and high latitude PAN amounts found at 8 km were around 125 pptv in the northern, but only be-

tween 50 and 75 pptv in the southern hemisphere. The PAN distribution found in the southern hemispheric tropics and subtropics is highly correlated with the jointly fitted acetylene (C₂H₂), which is another pollutant produced by biomass burning, and agrees reasonably well with the CO plume detected during end of September 2003 at the 275 hPa level (~10 km) by the Measurement of Pollution in the Troposphere (MOPITT) instrument on the Terra satellite. Similar southern hemispheric PAN amounts were also observed by previous airborne measurements performed in September/October 1992 and 1996 above the South Atlantic and the South Pacific, respectively.

1 Introduction

Peroxyacetyl nitrate (PAN) is generated in air masses polluted by fuel emissions or by biomass burning and is a possible indicator of photochemical smog. Its production is initialized by the reaction of hydrocarbons with the hydroxyl radical (OH) and subsequently with O₂ to form the peroxyacetyl radical CH₃C(O)OO as specific intermediate. The reversible reaction of the peroxyacetyl radical with nitrogen dioxide produces PAN

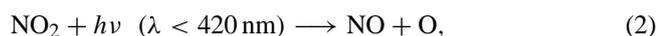


PAN was first detected in the Los Angeles area during smog episodes (Stephens et al., 1956). Phytotoxic episodes with PAN amounts of more than 15 ppbv were observed in Southern California in the 1980s (Temple et al., 1983). Even higher amounts of more than 30 ppbv were measured in 1997 in Mexico City (Gaffney et al., 1999). Since the 1980s PAN production has decreased in North America and Europe, but increased in Eastern Asia. The PAN amounts in clean air areas are generally substantially lower, namely between 50 and

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100 pptv (Singh et al., 2000a). Widespread southern hemispheric PAN pollution extending from South America over the South Atlantic, East and South Africa, the Southern Indian Ocean as far as Australia has been observed during the annual biomass burning period in South America and South Africa in September and October (Singh et al., 1996, 2000a).

The thermal decomposition rate of PAN is highly temperature dependent, resulting in lifetimes between 1 h at 298 K and about 5 months at 250 K (Singh, 1987). However, due to photolysis its mean lifetime is restricted to 3 months. Because of its temperature sensitivity PAN can only be transported over a few hundred kilometres in the lower troposphere, but over distances of more than 10 000 km in the cold upper troposphere (Singh, 1987). Thus it serves as a reservoir and carrier of NO_x in this height region. When these polluted air masses warm up, PAN is destroyed and NO_2 is released (Eq. 1), which is photolysed at wavelengths less than 420 nm. This leads to formation of nitric oxide (NO) and atomic oxygen:



Subsequently, atomic oxygen reacts with O_2 and produces tropospheric ozone:



Thus, PAN can contribute to the pollution of remote clean air areas, but has also strong influence on the oxidising power of the atmosphere. After photolytic destruction during daytime re-formation of PAN at night is possible. The fact, that contrary to ozone it has no important natural sources in the stratosphere, is an advantage of PAN over ozone as indicator of upper tropospheric hydrocarbon- NO_x photochemistry (Singh, 1987).

Up to now there exists no global dataset of upper-tropospheric PAN. Only a few airborne campaigns covering supra-regional areas have been performed in the free troposphere, e.g. the Global Tropospheric Experiment/Transport and Atmospheric Chemistry Near the Equator-Atlantic (GTE/TRACE A) experiment above the Tropical and South Atlantic (Singh et al., 1996) and the Pacific Exploratory Mission (PEM) Tropics-A campaign above the Pacific (Singh et al., 2000a). These missions were dedicated to study the extent and composition of the southern hemispheric biomass burning plume during September/October 1992 and 1996, respectively. Another airborne campaign off the west coast of North America, the Intercontinental Transport and Chemical Transformation 2002 (ITCT 2K2) project, was carried out to investigate the intercontinental PAN transport from East Asia (Roberts et al., 2004). Most of the other PAN data available are groundbased measurements in the boundary layer in highly polluted as well as very remote regions (Temple et al., 1983; Beine et al., 2000; Gaffney et al., 1999; Jacobi et al., 1999). Whereas all these airborne and groundbased

measurements employed in-situ techniques, only quite recently the first detection of upper-tropospheric PAN in infrared remote sensing spectra of MIPAS/ENVISAT and of the balloon-borne MIPAS instrument (MIPAS-B2) was reported (Allen, 2005a; Remedios et al., 2006, 2007). The MIPAS-B2 instrument is described in Fischer and Oelhaf (1996) and in Friedl-Vallon et al. (2004).

In this paper we present the first global distribution of PAN derived from spaceborne mid infra-red emission spectra. In detail, we will describe the retrieval method, give an error estimation, discuss sensitivity tests, and present the global PAN distribution observed between 4 October and 1 December 2003. This period covers the end of the biomass burning season in South America, Central and South Africa and in Australia. Thus, a widespread distribution of elevated PAN amounts and of other pollutants was expected in the southern hemispheric tropical and subtropical upper troposphere. While this paper is focused on the description of the PAN retrieval scheme and on the presentation of the retrieved PAN distribution, a more detailed investigation of the biomass burning plume by combined analysis of several pollutants and by trajectory calculations will be presented in a subsequent paper.

2 Michelson Interferometer for Passive Atmospheric Sounding

The Michelson Interferometer for Passive Atmospheric Sounding (MIPAS) was launched onboard the Sun-synchronous polar-orbiting European ENVIRONMENTAL SATellite (ENVISAT) on 1 March 2002. It is a limb-viewing Fourier transform infrared (FTIR) emission spectrometer covering the mid-infrared spectral region from 685 to 2410 cm^{-1} . Its wide spectral coverage and high spectral resolution, which is 0.035 cm^{-1} (unapodised) for spectra of the original full resolution measurement mode evaluated here, enables simultaneous observation of numerous trace gases (Fischer and Oelhaf, 1996; European Space Agency (ESA), 2000). This measurement mode consists of rearward limb-scans covering tangent altitudes between about 6 and 68 km within 17 altitude steps. The step-width is 3 km up to 42 km and increases up to 8 km above 52 km. The level-1B radiance spectra used for PAN retrieval are data versions 4.61–4.62 provided by the European Space Agency (ESA) (Nett et al., 2002). Several data processors, amongst others one at the Institut für Meteorologie und Klimaforschung (IMK), have been developed for near-real time and off-line retrieval of profiles of atmospheric trace species from level-1B MIPAS spectra (von Clarmann et al., 2003a). The PAN distribution presented in this paper was produced with the retrieval processor of the IMK, which was developed to produce consistent datasets containing considerably more trace species than included in the operational dataset provided under ESA responsibility.

3 Retrieval method and error estimation

The retrievals performed with the IMK data processor consist of inversion of MIPAS level-1B spectra to vertical profiles of atmospheric state parameters by constrained non-linear least squares fitting (e.g., Rodgers, 2000, and references therein) in a global-fit approach:

$$\mathbf{x}_{i+1} = \mathbf{x}_i + (\mathbf{K}^T \mathbf{S}_y^{-1} \mathbf{K} + \mathbf{R} + \lambda \mathbf{I})^{-1} \times (\mathbf{K}^T \mathbf{S}_y^{-1} (\mathbf{y}_{\text{meas}} - \mathbf{y}(\mathbf{x}_i)) - \mathbf{R}(\mathbf{x}_i - \mathbf{x}_a)), \quad (4)$$

where \mathbf{x}_{i+1} is the vector of the unknown state parameters calculated in iteration $i+1$, \mathbf{y}_{meas} is the measurement vector, and $\mathbf{y}(\mathbf{x}_i)$ is the result of the radiative transfer model using the parameters \mathbf{x}_i of iteration number i . \mathbf{S}_y is the measurement noise covariance matrix, \mathbf{K} is the Jacobian matrix containing the partial derivatives $\partial \mathbf{y}(\mathbf{x}_i) / \partial \mathbf{x}_i$, \mathbf{R} is a regularization matrix, $\lambda \mathbf{I}$ (scalar times unity matrix) is a damping term as proposed by Levenberg (1944) and Marquardt (1963), and \mathbf{x}_a is an a-priori profile used to constrain the retrieval result. To avoid any influence on the shape of the retrieved PAN profiles, an all-zero a-priori profile is chosen, whereas a midlatitude PAN profile of the Model for Ozone And Related chemical Tracers (MOZART) (Horowitz et al., 2003) is used as first-guess. Application of a smoothing constraint is necessary to attenuate instabilities, because the retrieval grid used has an altitude spacing of 1 km up to 44 km and of 2 km between 44 and 70 km, which means oversampling in comparison with the height distance between the tangent altitudes. Similar as in other retrievals of MIPAS data performed at IMK, Tikhonov's regularization scheme is used with a first derivative operator as constraint (Steck, 2002, and references therein).

Radiative transfer calculations are performed with the Karlsruhe Optimized and Precise Radiative Algorithm (KOPRA) (Stiller, 2000). Prior to PAN many other trace species are evaluated in the retrieval chain. First, spectral shift, the temperature profile and the tangent heights are fitted. Thereafter various trace gas profiles are retrieved in a successive manner, in which the major species H_2O , O_3 , CH_4 , N_2O , and HNO_3 are analysed first. The fitted profiles of all these species are used to model their radiative contribution in the PAN retrieval unless they are joint-fitted again. Beside each target species, one continuum profiles and one height-independent zero level calibration corrections are jointly fitted for each microwindow. The retrieval strategy is described in more detail in, e.g., von Clarmann et al. (2003b), Glatthor et al. (2004) and Höpfner et al. (2004) and references therein. Cloud-flagging, i.e. removal of cloud-contaminated spectra prior to retrieval, is performed following the standard method applied at IMK, using a threshold of 4.0 (Glatthor et al., 2006).

Since PAN is a relatively heavy and complex molecule (141 atomic mass units), the line positions are very dense and the line wings of its single rotational-vibrational transition lines overlap, which leads to broadband continuum-like

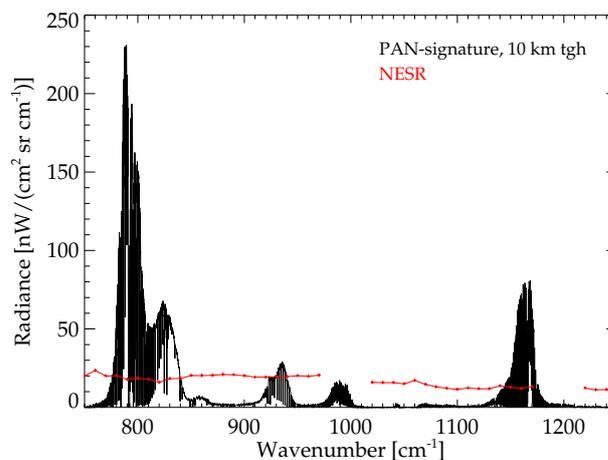


Fig. 1. PAN signatures at 10 km altitude (black), displayed as difference spectrum between model calculations for a tropical atmosphere containing a PAN amount of ~ 200 pptv (at 10 km) and for a tropical PAN-free atmosphere. The main features are the bands centred at 794, 930, 991 and 1163 cm^{-1} . The red curves show the typical noise-equivalent spectral radiance (NESR) in MIPAS spectra from channels A, AB and B (only low-wavenumber edge).

signatures. For this reason spectroscopic data are available as tabulated absorption cross sections only. In our retrievals PAN is modelled using a cross section dataset provided by the University of Leicester, which consists of measurements at three temperatures, 250, 273 and 295 K, with a nominal unapodised spectral resolution of 0.25 cm^{-1} (Allen et al., 2005b, c). Based on model calculations for a climatological tropical atmosphere, the most prominent atmospheric PAN band is that centred at 794 cm^{-1} (the ν_{16} NO_2 stretch), followed by a second band at 1163 cm^{-1} (the ν_{10} C-O stretch) and other considerably weaker bands. Figure 1 shows the difference spectrum for a tangent altitude of 10 km between an atmosphere containing a PAN amount of ~ 200 pptv (at 10 km) and a PAN-free atmosphere. Comparison with the noise-equivalent spectral radiance (NESR) of MIPAS indicates that for this PAN amount a signal-to-noise ratio of larger than 10 can be obtained in the band at 794 cm^{-1} . After test of different sizes of the analysis window, PAN retrievals were finally performed in one contiguous microwindow in the wavenumber region 775–800 cm^{-1} , which covers the strongest signatures of the band at 794 cm^{-1} . Due to the large amount of MIPAS data to be analysed, the weaker part of this band between 800 and 830 cm^{-1} as well as the band at 1163 cm^{-1} were not used for retrieval to save computation time. However some sample consistency tests were performed, in which the whole band at 794 cm^{-1} as well as the band at 1163 cm^{-1} were evaluated. For correct modelling of the CO_2 Q-branch at 792 cm^{-1} radiative transfer calculations were performed with inclusion of CO_2 line-mixing. Because of its relatively weak signal and broadband structure, PAN requires a careful retrieval to separate its signature from the

atmospheric continuum and from other broadband species, namely CCl_4 , HCFC-22, ClONO_2 and CH_3CCl_3 , which also contribute to the atmospheric signal in this spectral interval. Thus, beside PAN these species as well as H_2O and C_2H_2 , one continuum profile and one height-independent offset (one microwindow only) were also fitted. The results of preceding retrievals were used to model the temperature profile and the contribution of other interfering species like ozone, because test calculations had demonstrated that joint-fit of these species had only little influence on the retrieved PAN profile.

Figure 2 shows a PAN profile obtained on 21 October, 2003, over Eastern Africa (top left) and another profile from the same orbit obtained above northern midlatitudes (top right). Over East Africa a PAN volume mixing ratio (vmr) as high as 700 pptv was detected at the lowermost height useable for retrieval at this geolocation (11 km, due to clouds). Towards higher altitudes PAN decreases to 65 pptv at the tropopause (17 km) and to 50 pptv at 20 km. The midlatitude profile exhibits PAN values of ~ 175 pptv in the troposphere, a rapid decrease at the tropopause and also values around 50 pptv in the lowermost stratosphere. The corresponding averaging kernels (middle panels) indicate, that at the tropical geolocation a PAN signal is obtainable between 11 and 23 km and at the midlatitude geolocation, which is not contaminated by clouds, between 5 and 23 km. The calculated number of degrees of freedom is 2.4 for the tropical and 4.6 for the midlatitude geolocation. The higher midlatitude value is due to higher temperatures between 11 and 25 km and further extension of the scan into the troposphere. In both cases the height resolution is 3.5–6 km (degrading with altitude). The averaging kernels show that the PAN amounts retrieved in the lowermost stratosphere have a certain tropospheric contribution. Nevertheless, these PAN values are in good agreement with PAN amounts reported by Singh et al. (2000b) for the lower stratosphere above the Northern Atlantic (solid red squares).

The diagrams at the bottom of Fig. 2 show the total estimated PAN retrieval error, measurement noise and the major error components for both geolocations, in which all error components are 1-sigma uncertainties. The error estimation is based on the actual retrieved temperatures, tangent heights, PAN profiles and simulated spectra and Jacobians of the final iteration (cf. Glatthor et al., 2004). The error contributions displayed for the tropical (left) and midlatitude geolocation (right) are not exactly the same, because due to different atmospheric conditions slightly different components proved important (only components larger than 2% or 4%, respectively, are shown). The meaning of the various error sources listed in the legend is the following: total: total retrieval error; noise: measurement noise; param: total contribution of all error components except of noise; o3, hno3, cof2, co2: errors due to uncertain knowledge on the respective trace gases; tem, los, tgrra: errors due to uncertainty in the temperature profile, in instrumental pointing and in the

horizontal temperature gradient (1 K, 150 m and 1 K/100 km assumed, respectively); spectro, shift, gain, ils: errors caused by uncertain knowledge on PAN spectroscopic data, frequency calibration, gain calibration and on instrumental line shape (3.2%, 0.0005 cm^{-1} , 1% and 3% assumed). The errors caused by temperature and instrumental pointing are somewhat overestimated, because the basing uncertainties of 1 K and 150 m are assumed uncorrelated and rather conservative in comparison with the estimated standard deviation resulting from the preceding retrieval of temperature and line-of-sight. The error contribution of the joint-fitted gases CCl_4 , HCFC-22, ClONO_2 , CH_3CCl_3 , H_2O and C_2H_2 is assumed zero.

For the tropical geolocation the total error is minimal (5%) at the lowermost useable tangent height (13 km), because here the sensitivity of the retrieval is maximal, expressed by the maximum of the averaging kernel at this altitude. Due to the considerably smaller averaging kernel the error is larger at 11 km. Towards higher altitudes, where the PAN amount declines rapidly, the total error increases to 40% at the tropopause (17 km) and remains above 30% in the stratosphere. The stratospheric error estimates do not approach or exceed 100%, because, among other things, the smoothing error is not taken into account in our error assessment. Measurement noise increases from 3% at 13 km to more than 25% at the tropopause. Except for 13 km, uncertain knowledge on the instrumental line-of-sight (LOS) is the dominant error contribution in the altitude range 10–17 km. Further important error sources in this altitude range are the uncertainties in spectroscopic data, instrumental line shape (ILS), temperature and, above 15 km, in the ozone vmr. The spectroscopic uncertainty is around 3% at most altitudes, which corresponds to the assumed 1-sigma error of 3.2% of the PAN cross sections at 250 K (cf. Allen et al., 2005b, c). Uncertain knowledge on ILS and temperature causes errors of up to 3 and 4%, respectively, in the altitude range 10–17 km. The error due to unknown knowledge on ozone, as calculated from the estimated standard deviation (ESD) of the pre-fitted ozone profile, is 7% at 17 km, but decreases rapidly towards lower altitudes. At the northern midlatitude geolocation (bottom right), where considerably lower PAN amounts were detected, the total error amounts to 14–20% between 7 and 12 km. Towards lower altitudes the total error increases strongly up to 75% at 5 km because of decreasing sensitivity of the retrieval (cf. Fig. 2, middle right). Due to declining PAN amounts the error increases up to 30% in the lower stratosphere. The main error component below 9 km is the uncertainty in LOS and above 9 km the noise error, the latter ranging from 5% at 8 km to 25% in the stratosphere. Further major error components are the same as at the East African location. The upper tropospheric total errors derived here are considerably smaller than the natural variability of PAN, which, e.g., in a compilation given by Roberts et al. (2004) ranges from 20 to 700 pptv.

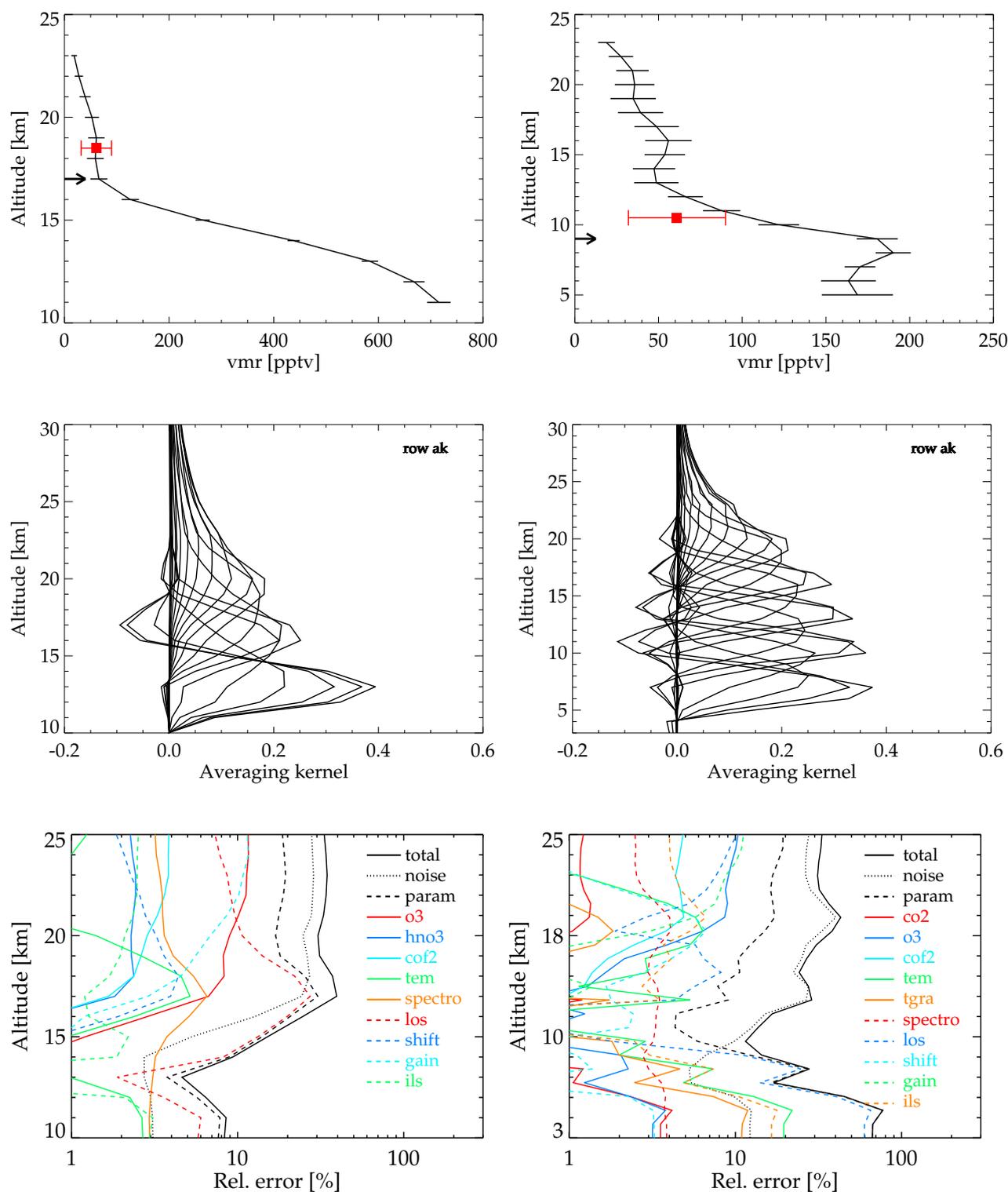


Fig. 2. Left: (top) PAN vmr-profile measured by MIPAS on 21 October, 2003, in polluted air masses over Eastern Africa (12.9° S, 37.2° E); due to cloud-contamination no PAN data are available from below 11 km; the error bars represent the noise error; the black arrow indicates the tropopause height and the red square the average PAN amount in the lowermost stratosphere given by Singh et al. (2000b) along with the scatter of the dataset; (middle) rows of the respective averaging kernel plotted versus altitude; (bottom) total retrieval error (solid black line), noise error (dotted black line), total parameter error (dashed black line), and contributions of various other error sources (see legend, detailed explanation in Sect. 3). Right: Same as left but for a geolocation from the same orbit at northern midlatitudes (51.4° N, 153.7° W).

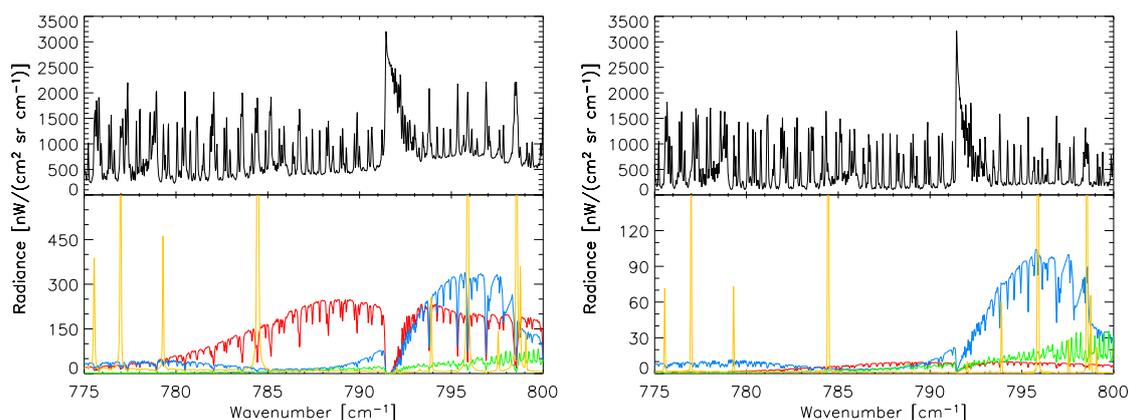


Fig. 3. Left: Modelled spectrum (black) and contributions of fitted PAN (red), CCl_4 (blue), HCFC-22 (green) and H_2O (brown) for 13 km altitude at an East-African geolocation (12.9°S , 37.2°E), where high PAN amounts of 600 pptv were retrieved. The contributions of PAN, CCl_4 , HCFC-22 and H_2O are the differences between model spectra with and without the respective gas. The top ends of the H_2O lines are cut to zoom the broadband contributors. Right: Same as left, but for 19 km altitude.

An additional error source is the fact that the lowest laboratory temperature, for which PAN cross sections have been determined is 250 K, whereas the upper tropospheric temperature often is considerably lower, e.g. 210 K at 13 km altitude at the East African geolocation. To estimate the possible bias resulting from this discrepancy, test retrievals were performed using the PAN cross section data sets for 273 K or 295 K only. Thereby the retrieved PAN amount at this geolocation, which due to the low temperatures had been evaluated using the cross sections for 250 K in the standard retrieval, increased by 13% and 26% at 13 km altitude, respectively. Provided that this linear dependence on temperature maps to laboratory PAN cross sections actually required for colder upper tropospheric temperatures, this would mean that the retrieval overestimates the “true” PAN amount at 13 km at this geolocation by 23%. For a general estimation of this effect, average temperatures were calculated for the southern hemispheric area of high PAN amounts (cf. Sect. 5), which were 247, 236, 221, 209 and 199 K for the altitudes of 8, 10, 12, 14 and 16 km. Under the assumption of linearity this means that the retrieved PAN amounts at these altitudes would have to be reduced by 2, 8, 14, 19 and 23%, respectively.

Figure 3 shows the modelled spectrum and the retrieved signatures of PAN, CCl_4 , HCFC-22 and H_2O for the East African geolocation at tangent altitudes of 13 km (left) and 19 km (right). To maintain clarity the weak signatures of the other joint-fit parameters ClONO_2 , CH_3CCl_3 and C_2H_2 are not shown. At 13 km PAN exhibits a broadband signature of up to more than $200 \text{ nW}/(\text{cm}^2 \text{ sr cm}^{-1})$ (bottom left). The broadband structures of CCl_4 and HCFC-22 are mainly interfering between 792 and 800 cm^{-1} , but only little at lower wavenumbers, which enables a separation of the PAN signal. At 19 km the signature of PAN has decreased to $10 \text{ nW}/(\text{cm}^2 \text{ sr cm}^{-1})$ (bottom right), which is a factor of

2 lower than the instrumental NESR and also lower than the contributions of CCl_4 and HCFC-22. Although the signal-to-noise ratio is enhanced by the extension of the low PAN signature over a considerable number of gridpoints, this shows that the retrieved stratospheric PAN amounts have a large uncertainty.

4 Sensitivity tests

To check the retrieved PAN amounts we investigated, if the measured MIPAS spectra can be modelled just as well by complete neglect of PAN. Figure 4 (left) shows the MIPAS spectrum from 13 km altitude obtained on 21 October 2003 on orbit 8585 at the above presented East African geolocation (12.9°S , 37.2°E), where a high PAN vmr of 600 pptv was retrieved. Further it contains fitted spectra and residuals resulting from retrieval with and without inclusion of PAN, i.e. assuming a PAN-free atmosphere in the latter case and retrieval of the joint-fit gases only. The residuals clearly show that the fit without consideration of PAN is much worse than the adjustment with modelling of PAN, confirming the large PAN amount retrieved at this altitude. This is also expressed by the RMS deviation between measured and modelled spectra, which is reduced from $41.2 \text{ nW}/(\text{cm}^2 \text{ sr cm}^{-1})$ for the retrieval without PAN to $26.9 \text{ nW}/(\text{cm}^2 \text{ sr cm}^{-1})$, i.e. by 35%, for the retrieval with inclusion of PAN. For comparison the result of the same test for 19 km altitude at the same geolocation, where a very low PAN amount of 50 pptv was retrieved, is also shown (Fig. 4, right). In this case the residuals are nearly equal (RMS of 16.82 and $16.76 \text{ nW}/(\text{cm}^2 \text{ sr cm}^{-1})$), indicating that retrieval without inclusion of PAN is technically almost just as good. However the comparably good fit for the PAN-free atmosphere is obtained with a slightly negative, i.e. unphysical, amount of CH_3CCl_3 . This phenomenon

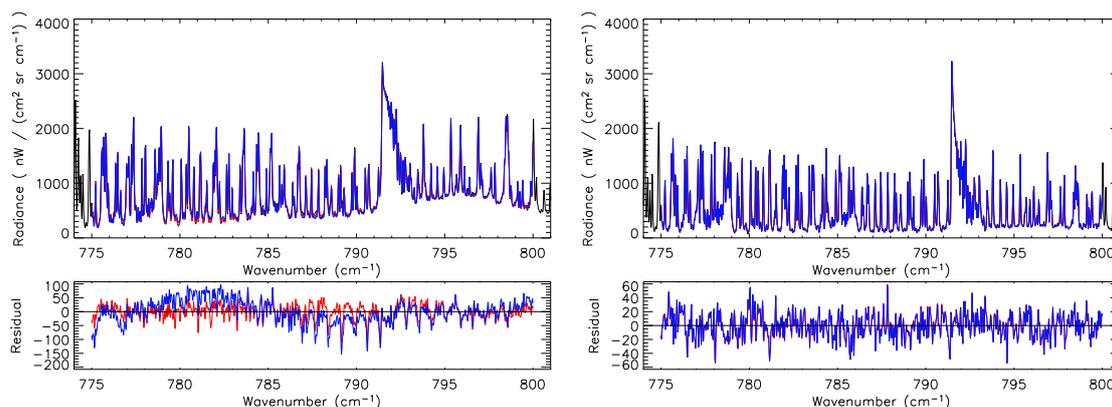


Fig. 4. Left: MIPAS spectrum from 13 km altitude, obtained on 21 October 2003 over Eastern Africa (12.9° S, 37.2° E) (black), as well as fitted spectra and residuals with (red) and without (blue) inclusion of PAN. Residuals are in the same units as the radiance plots themselves. Right: Same as left, but for 19 km altitude.

becomes even much stronger at 13 km, where neglect of PAN leads to large negative CH₃CCl₃ and ClONO₂ amounts.

The result of this sensitivity test for the whole orbit 8585 is displayed in Fig. 5, where the ratio of the RMS-values from retrieval without and with inclusion of PAN is plotted against the retrieved PAN amount. The colour code indicates the tangent height of the respective spectra. For data points from altitudes of 16 km or higher the RMS-ratio is very close to unity for values below 100 pptv, indicating no sensitivity to PAN. These altitudes are at or above the tropopause, where indeed only low PAN amounts are expected. However, at lower altitudes of 7–14 km there are other points with PAN amounts below 100 pptv and larger RMS-ratios of up to 1.1. This indicates an altitude-dependent PAN significance-threshold of ~100 pptv at altitudes of 16 km or higher and of ~60 pptv for lower altitudes. In addition, there is a subset of datapoints, which show a linear correlation of the RMS-ratio with the retrieved PAN amounts, meaning an increasingly worse fit without consideration of PAN. For these data points, which are from altitudes between 7 and 13 km, the RMS-ratio increases up to more than 1.6 for fitted PAN amounts of 200–600 pptv.

PAN detection in MIPAS/ENVISAT spectra, but following a somewhat different method, has also been reported by Allen (2005a) and by Remedios et al. (2006).

Additional tests performed in our investigation consisted in correlation between the retrieved PAN and the jointly fitted CCl₄, HCFC-22 and CH₃CCl₃ amounts and showed no correlation or anticorrelation. This ensured that the broadband signatures of these species were not erroneously interpreted as PAN.

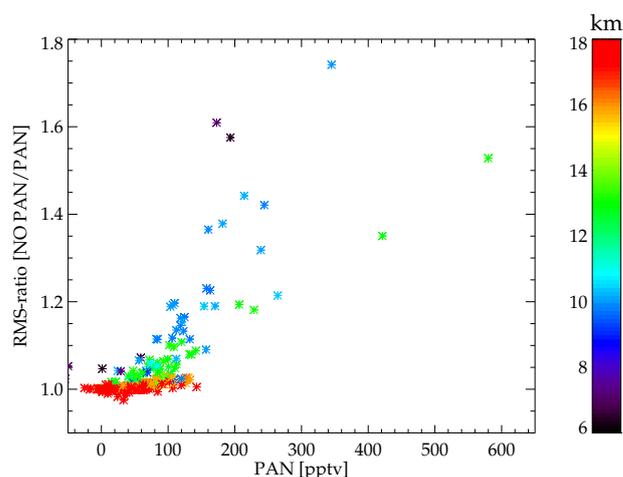


Fig. 5. Ratio of RMS-values resulting from retrievals without and with inclusion of PAN plotted versus retrieved PAN amounts for orbit 8585 of 21 October 2003. Data are color-coded with the tangent height; tangent heights higher than 18 km are also coded red.

5 Discussion

5.1 Observed PAN distribution

The analysed dataset consists of 150 orbits from 10 days between 4 October and 1 December, 2003, namely 4, 13, 21 and 26 October, 1, 4, 11, 21 and 24 November, and 1 December. This period covers the end of the biomass burning season in the southern hemisphere. Figure 6 exemplarily shows the global PAN distribution retrieved from single spectra for 21 October 2003 at 12 km altitude. The white areas, which are mainly located in the tropics and subtropics, are data gaps caused by cloud contamination. Nevertheless a contiguous area of high PAN values (250–600 pptv) extending from the

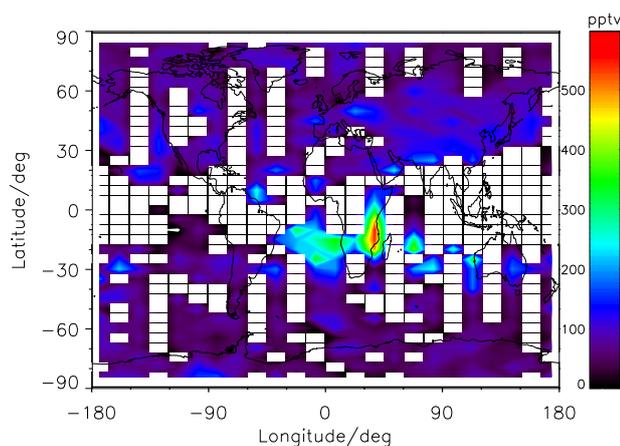


Fig. 6. Global PAN distribution on 21 October 2003 at 12 km altitude, obtained from MIPAS data. The white areas are data gaps due to cloud contamination.

Southern Atlantic to East Africa is visible, which we attribute to extensive biomass burning. Interrupted by data gaps, this area seems to be continued above the Southern Indian Ocean as far as Australia. To enhance the spatial data coverage we show the PAN distribution averaged over the whole period in the next two plots.

The zonal average of the whole dataset (Fig. 7) exhibits clear differences between the northern and southern hemispheric PAN distribution. In the southern hemisphere there is a large plume containing elevated PAN amounts of more than 100 pptv between the equator and about 50° S in the altitude range 7–15 km, which is obviously caused by biomass burning. Maximum values of more than 200 pptv are between 20° S and 40° S at altitudes from 7 to 10 km. There is another, but much more restricted, region of elevated PAN amounts in the northern hemispheric tropics, which contains maximum PAN amounts of 100–150 pptv. No data are available in the northern tropics below 8 km due to permanent cloud contamination of the MIPAS spectra. Further there are also elevated PAN amounts throughout the northern hemispheric mid and high latitudes, decreasing from 150 pptv at 7 km to 75 pptv at 11 km. The major reason for these values is probably industrial pollution. However, we cannot exclude any other pollution source like boreal forest fires. In comparison to the northern hemisphere, the upper tropospheric PAN values at southern high latitudes, which are much less industrially polluted, are much lower, namely between 75 pptv at high midlatitudes and 25 pptv above the Antarctic. At stratospheric altitudes above 15 km the situation is reversed. Here, obviously due to beginning subsidence over the Arctic, the northern hemispheric mid and high latitude PAN amounts are lower than those from the southern hemisphere.

Figure 8 shows the averaged global PAN distribution at 8, 10, 12 and 16 km altitude. At 8 km there are still large data gaps in the tropics due to cloud-contamination. The

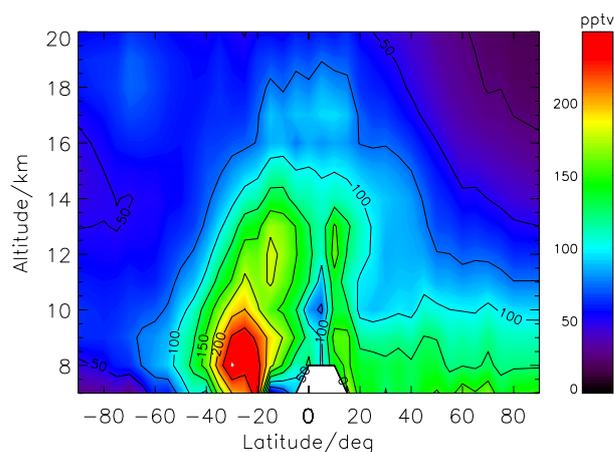


Fig. 7. Latitudinal PAN distribution, averaged zonally and over 10 days between 4 October and 1 December, 2003.

most prominent feature is a contiguous band of high PAN amounts in the southern hemispheric tropics and subtropics, which extends from the tropical and subtropical Atlantic over Southern Africa, the Southern Indian Ocean and Australia as far as to the West Pacific. The PAN values in this band are generally between 200 and 300 pptv and increase up to 500 pptv above Southern Africa and the adjacent Indian Ocean. The high PAN amounts above the Western Indian Ocean are represented by 15 individual profiles. As already mentioned above, we attribute this feature to southern hemispheric biomass burning. In the northern hemispheric subtropics, a more restricted area of elevated PAN is visible at 8 km, extending from the African Sahel zone to India and, possibly strengthened by additional sources of PAN, further northeast above China and the Chinese Sea. At 10 km high PAN values appear above the whole southern and eastern Africa. At 12 km the area of high southern hemispheric PAN amounts does not extend so far east and ends at the West Australian coast. However, high PAN amounts of 400–500 pptv are now visible above North-East Brasil, i.e. in an area which was masked by clouds below. At 16 km altitude, where deterioration by clouds has mostly disappeared and the global coverage is nearly complete, the highest PAN amounts of up to 200 pptv cover an area extending from tropical Brasil above the Southern Atlantic and tropical Africa to the Southern Indian Ocean. Like in Fig. 7, the northern mid and high latitude PAN values are lower than those from the southern hemisphere.

In order to check on a qualitative basis, if the observed southern hemispheric tropical and subtropical PAN distribution can be caused by biomass burning, fire counts of the Tropical Rainfall Measuring Mission (TRMM) satellite (Giglio et al., 2000) were inspected for the time period 25 September until 1 December 2003. Indeed, numerous fire count clusters were found in South America, Central and

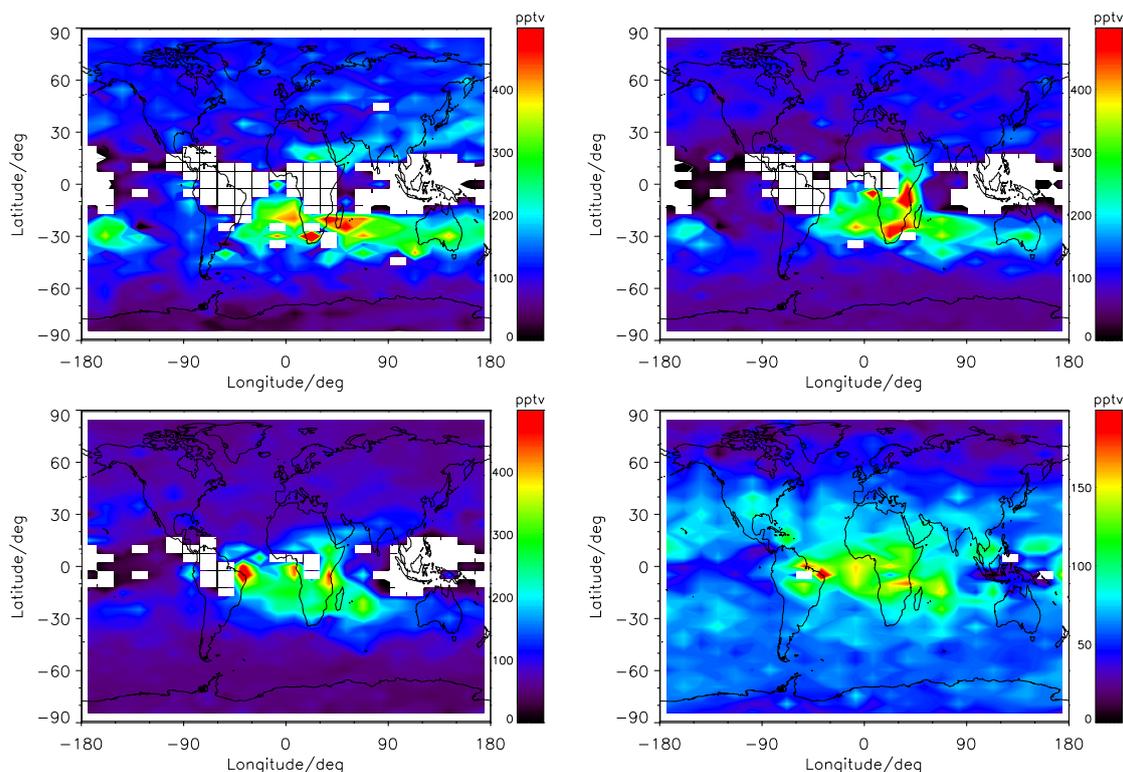


Fig. 8. Global PAN distribution at 8, 10, 12 and 16 km altitude (top left to bottom right) averaged over 10 days between 4 October and 1 December, 2003. The white areas are data gaps due to cloud contamination.

Southern Africa and in Australia over the whole period, indicating intensive biomass burning (<http://tsdis.gsfc.nasa.gov/tsdis/Fire/fireintro.html>). This is also confirmed by fire counts of the Moderate Resolution Imaging Spectroradiometer (MODIS) experiment on the Terra satellite published by Edwards et al. (2006, Fig. 5a). As mentioned in Sect. 1, a more detailed and quantitative investigation of the biomass burning plume by combined analysis of several pollutants and trajectory calculations starting above the observed fire count clusters will be presented in a subsequent paper.

Figure 9 shows daily averages of the PAN values in an area containing the southern hemispheric plume (left) and from the same latitude band but from a much less polluted area above the Central and Eastern Pacific (right) at 8, 12, 16 and 20 km altitude. Comparison of both plots shows that the PAN values in the area containing the biomass burning plume are strongly enhanced at 8 and 12 km and moderately elevated at 16 km throughout the observation period, which basically justifies the temporal averaging performed above. At 8 km the PAN amounts in the plume range from 200 to 370 pptv, whereas they are around 200 and 80 pptv at 12 and 16 km. The corresponding values above the Pacific are much lower, about 130, 90 and 60 pptv, respectively. At 8 km altitude these PAN values seem also to be moderately enhanced. As mentioned above, the stratospheric PAN amounts from

20 km are only little significant, but exhibit approximately the same values in both areas.

5.2 Correlation of southern hemispheric PAN with joint-fitted C_2H_2

Contrary to PAN the joint-fitted trace gas acetylene (C_2H_2) exhibits narrow transition lines in MIPAS spectra (e.g. a well separated line at 776.1 cm^{-1}), which can not be misinterpreted by other continuum-like structures. Since C_2H_2 is another, although not unique, product of biomass burning (Singh et al., 1996), it is suited for a cross-check of the PAN amounts retrieved in the southern hemispheric tropics and subtropics. In Fig. 10 the PAN and C_2H_2 data from the latitude band 0° to 40° S, i.e. from the biomass burning zone and the adjacent background region, are plotted against each other for the altitude region 10–16 km. The high correlation between both datasets ($r=0.79$) confirms the measured southern hemispheric PAN distribution. The slightly negative C_2H_2 values are mainly associated with PAN values below 100 pptv from altitudes of 15–16 km, which are anyway only little significant (cf. Sect. 4).

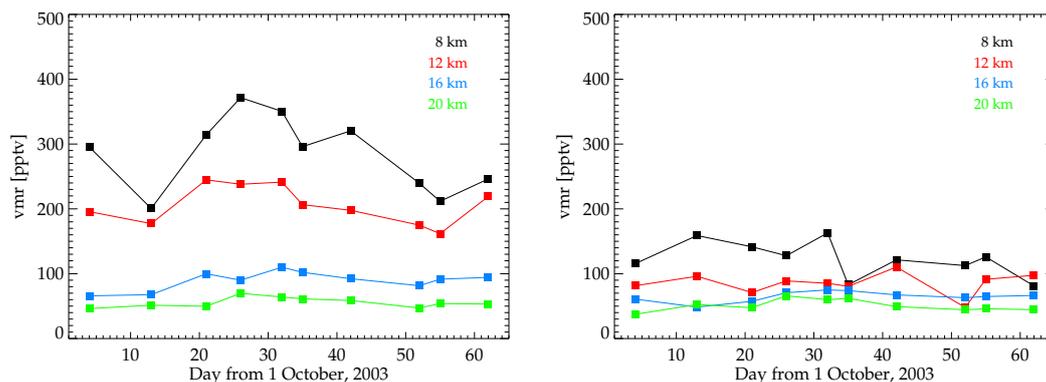


Fig. 9. Left: Daily averages of PAN amounts measured in the biomass burning zone (0° – 30° S and 60° W– 135° E) at 8 km (black), 12 km (red), 16 km (blue) and 20 km altitude (green). Right: Same as left, but for an adjacent background zone above the South Pacific (0° – 30° S, 90° W– 180° W).

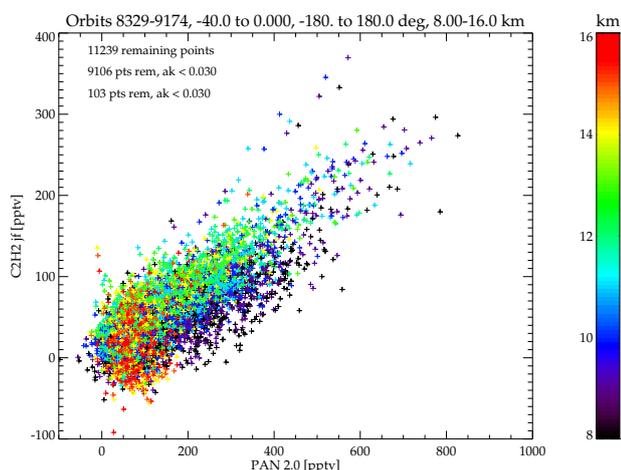


Fig. 10. Correlation between PAN and joint-fitted C_2H_2 amounts, measured by MIPAS on 10 days between 4 October and 1 December, 2003, in the latitude band 0° to 40° S between 8 and 16 km altitude; correlation coefficient $r=0.79$.

5.3 Comparison with other space- and airborne measurements

The biomass burning plume detected in the MIPAS PAN distribution (cf. Fig. 8) is in reasonably good agreement with the area of enhanced CO observed by the Measurement of Pollution in the Troposphere (MOPITT) instrument on the Terra satellite (Edwards et al., 2006, Figs. 1b and 4a), which these authors also attribute to biomass burning in South America, South Africa and Australia. During the last week of September 2003 MOPITT measured a plume of elevated CO (up to ~ 160 ppbv), which on the 275 hPa (~ 10 km) level mainly covered tropical Brazil, the Southern tropical and subtropical Atlantic, Southern Africa as well as the Western Indian Ocean, and less distinct extended until Western Australia. During the subsequent period 1–9 October 2003, which cor-

responds to the beginning of our measurement period, MOPITT observed a band of high CO total columns at 30° S, which even circumscribed the globe.

A qualitative comparison can also be made with airborne in-situ PAN measurements performed by the NASA Global Tropospheric Experiment (GTE) in September/October 1992 over the South Atlantic and the Pacific Exploratory Mission (PEM) in September/October 1996 over the remote South Pacific (Singh et al., 1996, 2000a). These campaigns were dedicated to study the distribution of species in pollution plumes originating from biomass burning regions in South America and South Africa. Over the South Atlantic average PAN amounts of about 300 pptv were found in the altitude region 3–11 km. This agrees well with the average of MIPAS Southern Atlantic (10° S to 30° S) PAN data, which is between 330 pptv at 8 km and 230 pptv at 11 km (cf. Fig. 8). For the remote South Pacific these authors report average PAN values of about 60 pptv between 5 and 11 km altitude, whereas the MIPAS values in this region (0° to 30° S and 90° W to 150° W) are between 115 pptv at 8 km and 80 pptv at 11 km.

Enhanced northern hemispheric PAN as displayed in Fig. 7 and Fig. 8 (top left) has also been found by Kotchenruther et al. (2001) and Roberts et al. (2004), who report on PAN amounts of about 200 pptv in the altitude region 6–8 km off the westcoast of the United States and attribute these measurements to long-range transport of industrial pollution from East-Asia. Further, Remedios et al. (2007) retrieved PAN amounts of 260–180 pptv in the altitude region 7.5–10.5 km from MIPAS-B2 spectra obtained over the Western Mediterranean region. Elevated PAN values at high northern latitudes as shown in Fig. 7 have also been measured by e.g. Bottenheim and Gallant (1989), who report on PAN-vmrs of up to 300 pptv at 7 km altitude at 70° N and attribute their observations to arctic haze episodes caused by industrial pollution.

6 Summary and conclusions

We have used spaceborne mid-infrared measurements of MIPAS to derive the first global upper tropospheric PAN distribution. The evaluated dataset consists of ten days covering the period 4 October to 1 December 2003. PAN was retrieved in the wavenumber region 775–800 cm^{-1} , which covers its most prominent mid-infrared signature. Beside PAN the interfering broadband species CCl_4 , HCFC-22, CH_3CCl_3 and ClONO_2 as well as H_2O and C_2H_2 were joint-fitted. PAN profiles were derived for the upper troposphere and lower stratosphere, with about 2.5–4.5 degrees of freedom (depending on useable tangent heights) and a height resolution of 3.5–6 km. The total retrieval error, obtained from an analytical error analysis, ranges from 5% for strongly enhanced tropical PAN amounts (600 pptv at 13 km) to 14–20% for moderate midlatitude PAN amounts (175 pptv) near the tropopause. The error increases up to 75% towards midlatitude lower tropospheric altitudes and to more than 30% in the lower stratosphere. An additional uncertainty results from incomplete coverage of the upper tropospheric temperatures by the PAN cross section data set used. Due to this discrepancy the “true” PAN amounts in the cold tropical upper troposphere might be overestimated by up to $\sim 30\%$. A sensitivity test to check the retrieved PAN data showed a degradation of the fit quality for retrievals without consideration of PAN, which increased with the PAN amount obtained before. Further, this test indicated a PAN significance-threshold of ~ 60 pptv for altitudes up to ~ 15 km and of ~ 100 pptv above.

The analysed period covers the end of the biomass burning season in the southern hemispheric tropics and subtropics, in which large amounts of pollutants inclusive PAN are produced, uplifted into the free troposphere and distributed over large areas. Indeed, high PAN amounts were found in the southern hemispheric tropics and subtropics in the altitude range 8 to 16 km. At 8 km elevated PAN values formed a nearly worldwide band extending from Brasil over the Southern Atlantic, Central and South Africa, the Southern Indian Ocean and Australia as far as to the South Pacific, whereas they were more confined to the area between Brasil and the Western Indian Ocean at higher altitudes. The PAN amounts in this biomass burning area ranged from 200 to 500 pptv at 8 km and from 100 to 200 pptv at 16 km. Enhanced PAN amounts were also found in a much more restricted area between northern subtropical Africa and India. The most significant PAN signal at northern hemispheric midlatitudes with values of up to 250 pptv was detected in an adjacent area extending from China into the Chinese Sea. The average mid and high latitude PAN values at 8 km altitude were 125 pptv in the northern and between 75 and 50 pptv in the southern hemisphere.

During the analysed period numerous fire count clusters were observed in the southern hemispheric tropics and subtropics by the TRMM satellite. Further, the southern hemi-

spheric PAN distribution is well correlated with the joint-fitted C_2H_2 , which is another pollutant produced by biomass burning, and there is good agreement with the CO distribution measured by the MOPITT experiment in late September and early October 2003. On average, the retrieved PAN amounts are also consistent to airborne PAN measurements performed during the southern hemispheric biomass burning periods of 1992 and 1996. These correspondencies confirm the southern hemispheric PAN distribution observed by MIPAS and its characteristics as biomass burning plume. A more detailed investigation of this plume by combined analysis of several pollutants will be given in a subsequent paper.

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