



# Preferential adsorption of para and ortho water molecules on charged nanoparticles in planetary ice clouds

Johanna Weidelt<sup>1</sup>, Thomas Dresch<sup>2</sup>, Denis Duft<sup>2</sup>, and Thomas Leisner<sup>2,3</sup>

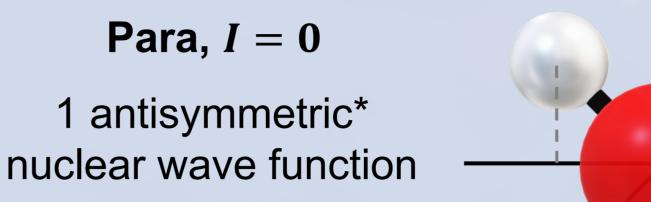
<sup>1</sup>Faculty of Physics, Bielefeld University, Germany <sup>2</sup>Institute of Meteorology and Climate Research, Karlsruhe Institute of Technology, Germany <sup>3</sup>Institute of Environmental Physics, University of Heidelberg, Germany

jweidelt@physik.uni-bielefeld.de, thomas.dresch@kit.edu

# Motivation

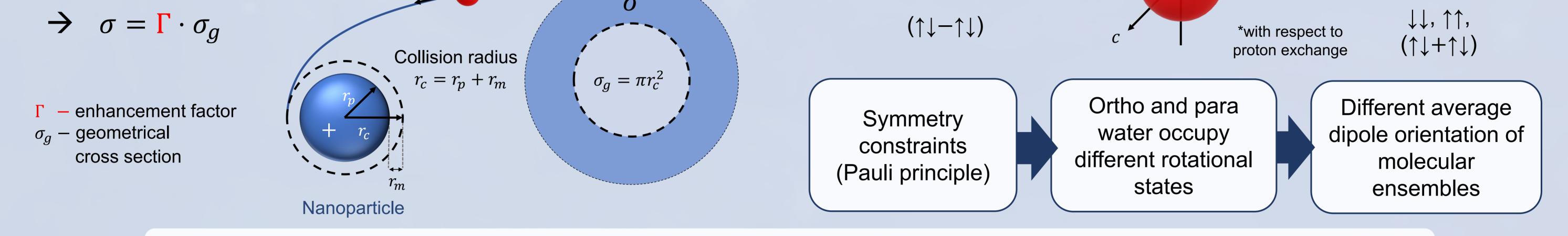
In the Earth mesopause and other extraterrestrial settings, **nanometer**sized singly charged particles form by condensation of evaporated meteorite material. They exhibit an **enhanced water adsorption cross** section due to the strong charge-dipole-interaction [1]. The enhancement factor depends on the average dipole orientation of the water molecules in their Stark-shifted rotational states.

In water, rotational states are strongly linked to the nuclear spin. The spins of the two indistinguishable protons in water can be either parallel (ortho) or antiparallel (para) [2]:



Ortho, I = 1

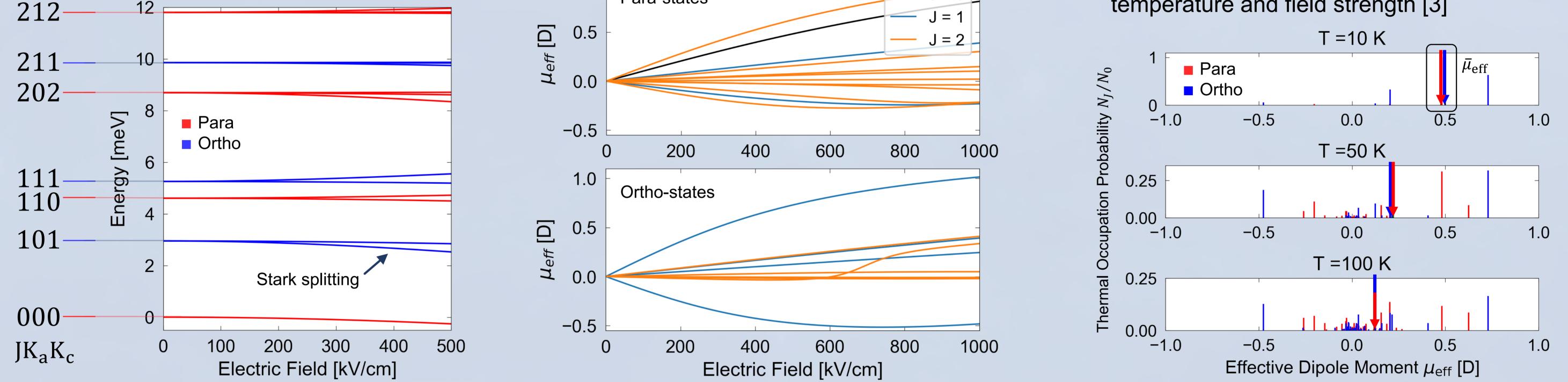
3 symmetric\* nuclear wave functions



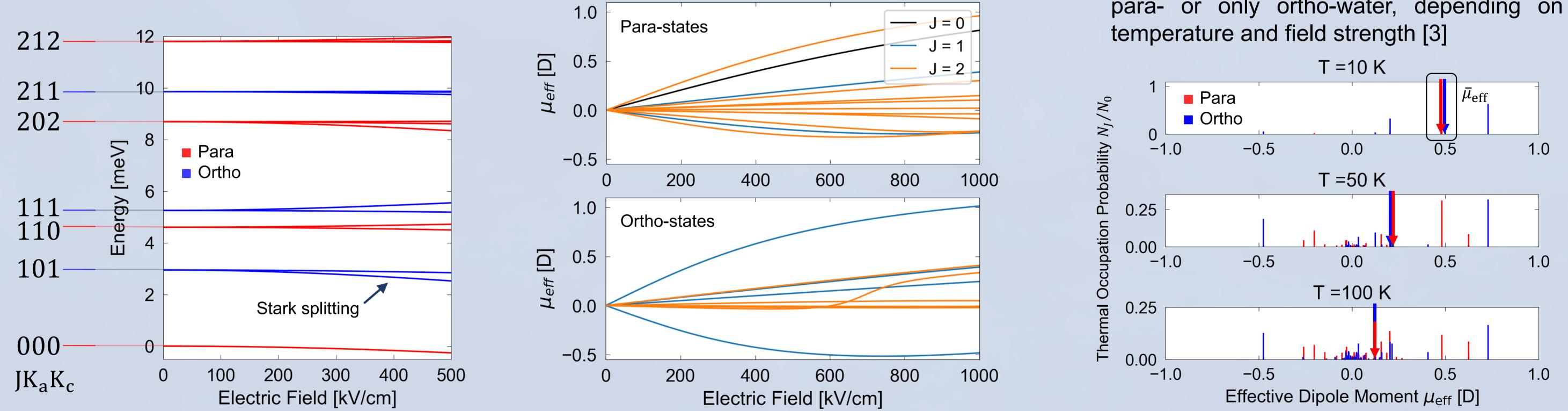
How does the adsorption enhancement factor of the water molecule depend on its nuclear spin state?

## Method

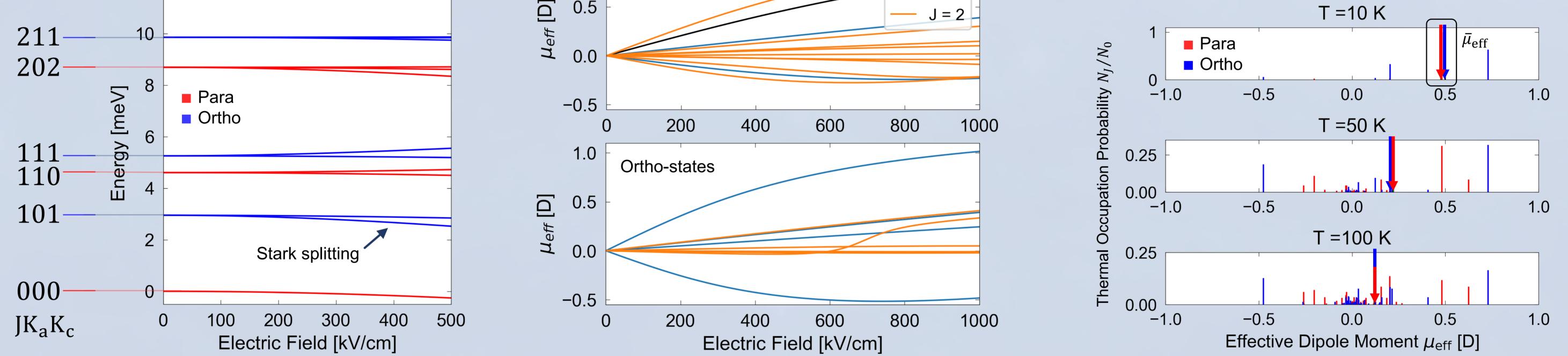
Calculation of the rotational states of water and division into ortho- and para states using the program CMIStark [4]



Calculation of the **effective dipole** moment of each energy state [4]



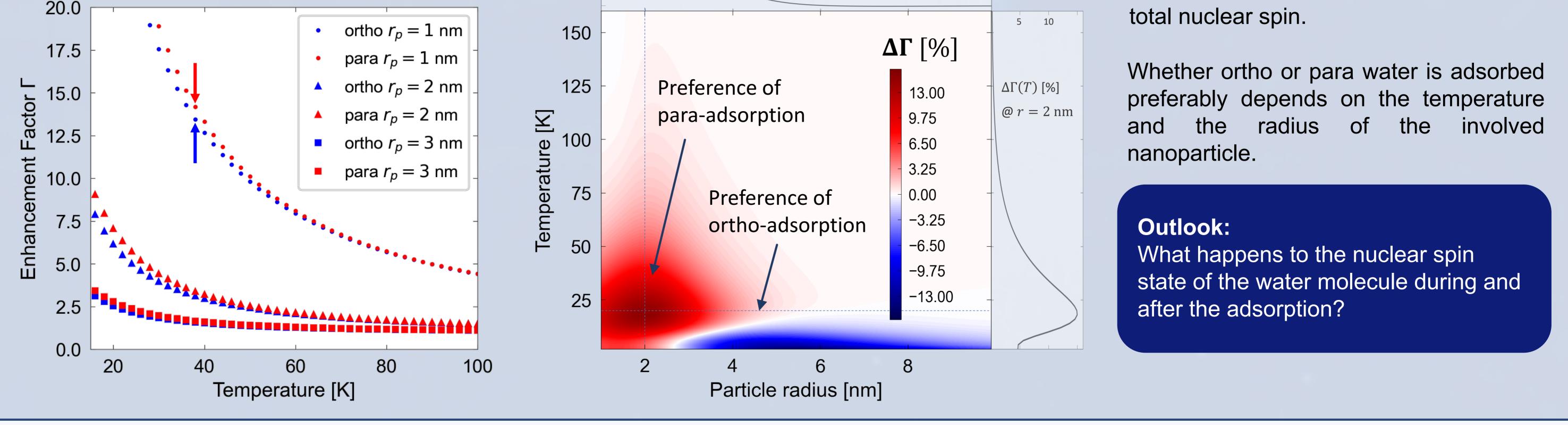
Calculation of the **average effective dipole moment** of a molecular ensemble (weighted by thermal occupation) consisting of only para- or only ortho-water, depending on

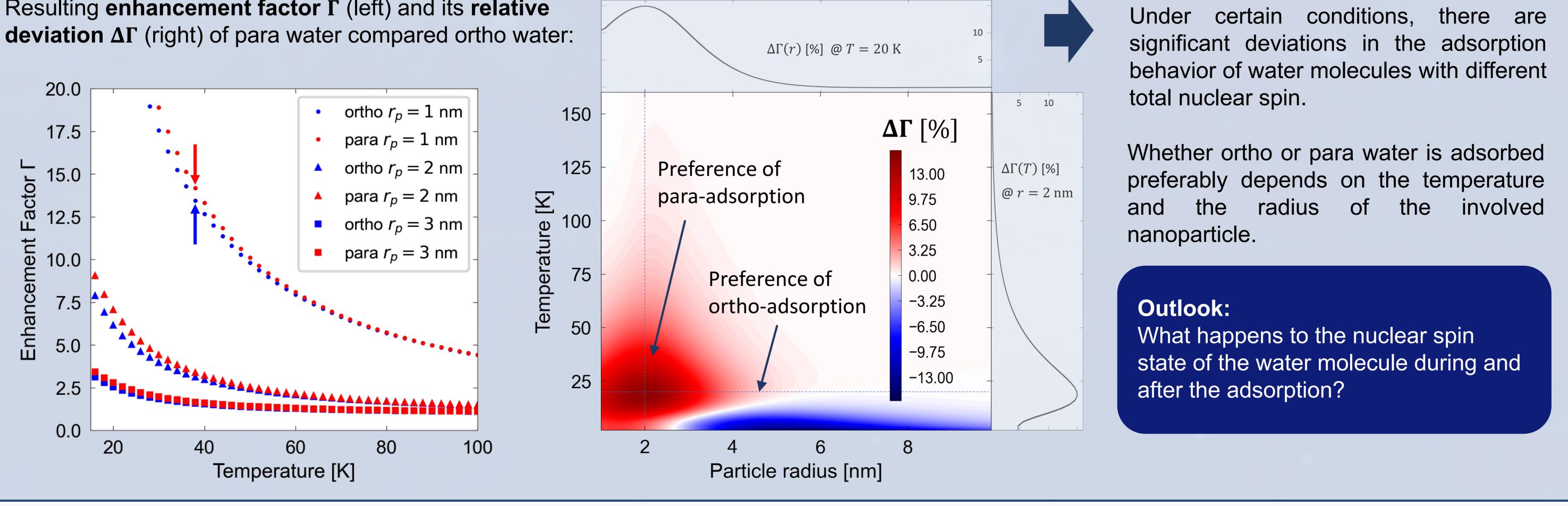


Having calculated the average effective dipole moment of a molecular ensemble of ortho- and para molecules, each, we can obtain their respective enhancement factor [1]:

### **Results: Collision Cross Section Enhancement**

Resulting enhancement factor  $\Gamma$  (left) and its relative





#### **References:**

1) M. Nachbar et al., DPG Annual Conference UP6.1 (2021)

3) R. Moro et al., Phys. Rev. A **75**, 013415 (2007) 2) D. Horke et al., Ang. Chem. **126**, 12159 (2014) 4) Y. P. Chang et al., Comput Phys Commun **185**, 339–349 (2014) Background image: Orion Nebula James Webb Telescope NASA

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