



# On the utility function of experiments in fundamental science

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## ABSTRACT

The majority of experiments in fundamental science today are designed to be multi-purpose: their aim is not simply to measure a single physical quantity or process, but rather to enable increased precision in the measurement of a number of different observable quantities of a natural system, to extend the search for new phenomena, or to exclude a larger phase space of candidate theories. Most of the time, a combination of the above goals is pursued; this breadth of scope adds a layer of complexity to the already demanding task of designing the measurement apparatus in an optimal way, by defining suitable geometries and choosing the most advantageous materials and appropriate detection technologies. The precise definition of a global optimality criterion may then require experimentalists to find a consensus on the relative scientific worth of those goals.

In this work we discuss the problem of formulating a utility function for multipurpose experiments, as an enabling step to employ artificial intelligence tools to explore the design space and assist humans in finding solutions at the Pareto front. For that purpose, we consider two use cases in particle physics research and one in astro-particle physics; in the latter case we show, using a recently developed optimization software, how the precise definition of a multi-target utility function may enable a significant increase of its value above that offered by human-designed detector layouts.

## 1. Introduction

Research in fundamental science has a long history of synergy with computer science developments. Some of the applications required to study subnuclear matter and interactions, to model galaxy formation and interactions, to describe matter distribution in the universe, and to investigate similar complex systems often require access to extensive computing resources, innovative algorithms, and specialized computing hardware. A glaring example of the interplay between fundamental science and high-performance computing comes from the field of lattice Quantum Chromo-Dynamics (QCD). Lattice QCD calculations involve discretizing spacetime into a grid, or lattice, to solve the equations

governing the strong interaction between quarks and gluons. The complexity of those calculations grows exponentially with the resolution of the lattice and the physical parameters being studied, such as temperature or quark masses. To tackle these challenges, researchers have been relying on supercomputers equipped with specialized architectures, and developed innovative algorithms for parallel computation. Over the past few decades the growth in computing performance has enabled lattice QCD to turn from sideline research into an oracle capable of providing precise predictions of hadron masses, decay rates, and interaction cross-sections [1].

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In light of the above, it should come as no surprise that in recent years experiments in particle and astro-particle physics – the focus of this work, from which we draw a few examples – have extensively adopted deep learning technologies, integrating them in their data analysis workflows [2]. In parallel, significant efforts have begun to be directed toward the integration of neural networks and other machine learning models into online data acquisition systems and processes [3]. Quantum Computing (QC) developments are also now closely followed by the community of experimental physicists, with a view to identifying specific use cases where that new technology may provide effective solutions to very hard problems that are amenable to be treated by it, through a successful encoding and implementation of their underlying computing tasks [4].

What the above landscape of computer science-driven advancements has largely left behind – at striking variance with the general embracing of deep learning that the community of fundamental science researchers has otherwise been displaying – is the task of optimizing the design of the instruments that collect the data ultimately needed for the extraction of information on the studied phenomena<sup>1</sup>. What the community has so far been referring to as “optimization” of particle detectors and other similar apparatus is typically constituted by the appraisal of a quite limited set of design parameters that modify in non-radical ways a pre-defined, experience-driven design concept through the discrete sampling of their possible values. The crucial hindrance to a more extensive investigation of the design space has in the past been the stochastic nature of the physical reactions at the basis of data-generating processes, which makes the likelihood function intractable [10]. Further, because of the absence of a quantitative recipe to combine the merits of different properties of the apparatus, an evaluation of the relative worth of different choices is usually performed on a narrow subset of the scientific results achievable by the experiment. The typical *modus operandi* includes the generation of large high-fidelity simulations of the relevant physical phenomena, followed by careful modeling of the resulting interaction of particles with the apparatus. Those data constitute the basis of the evaluation of a limited number of proxies of the overall experimental goals. *E.g.*, in the design of a detector for a particle collider targeting precision Higgs physics, those proxies could consist in the statistical uncertainty on the self-couplings of the Higgs boson achievable with a given amount of data; while for a gamma-ray observatory such as the one we discuss in Section 4 *infra*, the considered proxy could be the discovery sensitivity to point sources in the sky as a function of energy. The significant limitations connected to the large computing cost of high-fidelity GEANT4 [11] or CORSIKA [12] simulations provide a rationale for the procedure but do not justify it any longer today: progress in deep learning techniques and development of automatic differentiation tools are enablers of a completely different approach to the problem, which the community cannot any longer afford to ignore. In particular, the success of today’s generative models offers effective bypass solutions to the likelihood intractability.

A motivation for investigating the potential of a more holistic approach to the design of experiments is offered by observing that “experience-driven” design choices may actually fall very far from optimality even in the simplest use cases. A proof comes from the study of the relevant figure of merit of MuOnE, an experiment that aims to determine with high precision the differential cross section of elastic scattering of muons off electrons at high energy [13]. A reduction of a full factor of two in the relative resolution in event  $q^2$ , a quantity closely tracking the objective of the intended measurement, was demonstrated by a simple exploration of the geometry of the apparatus, when crucial co-design elements were included in the investigation [14] (see Fig. 1).

<sup>1</sup> Recently the community has been trying to organize efforts toward the co-design of hardware and software of future experiments; examples come from the MODE Collaboration [5–8] and the new EUCAIF initiative [9].

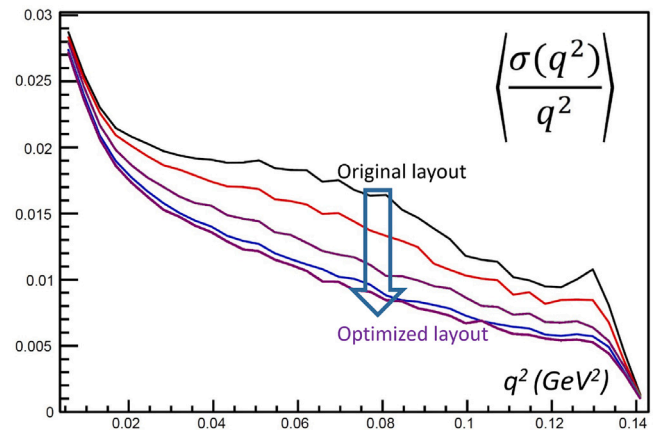


Fig. 1. Reduction in the estimated relative uncertainty in event  $q^2$  as a function of event  $q^2$  for elastic muon scattering, for progressively improved layouts of the detection apparatus.

Source: Figure reprinted from [14].

A gain of a factor of two in sensitivity is an off-the-scale performance improvement in particle physics applications<sup>2</sup>. Yet similar potential improvements may be there, waiting to be achieved at no significant increase in cost or complexity even in experiments that are admittedly not at the high end in the complexity scale of similar endeavors in high-energy physics (HEP).

Two factors require to be carefully appraised for their potential as show-stoppers in the way to artificial-intelligence-powered experiment optimization: on one side the high dimensionality of the design space, and on the other the multi-purpose nature of the utility function of typical large-scale experiments. The MuOnE example mentioned *supra*, in truth, fails to deliver as representative on both counts. The dimensionality of the parameter space considered by its optimization task lays in the  $O(10)$  range. Furthermore, the experiment is uncharacteristically focused toward a single experimental goal: reducing as much as possible the uncertainty in the differential cross section of muon-electron scattering at high scattering  $q^2$ , as an enabling step toward reducing the leading systematic uncertainty in the theoretical prediction for the muon  $g - 2$  anomaly, which constitutes the sole reason for building the experiment. We briefly comment on the two mentioned factors *infra*, before we proceed to examine the latter in more detail in the remainder of this work.

### 1.1. The design space

The complexity of multipurpose detectors and similar instruments employed in fundamental science experimentation implies that a full investigation of the most effective choices for the materials, the detection technology, and the geometrical arrangement of the system may correspond to the exploration of  $O(1000-10,000)$ -dimensional design spaces. Such a large phase space of available options is problematic to explore with discrete sampling or evolutionary optimization techniques. This large dimensionality is also multi-faceted and heterogeneous, making automatic exploration challenging; it stems primarily from two distinct sources that have grown in importance in recent years: on one side, emerging technologies enabling complex three-dimensional arrangements of detection elements; and on the other, the expanding variety of available materials, components, and detection technologies. We can ascribe to the first of these sources, for instance,

<sup>2</sup> The typical scale could be set by the intense and long-term effort spent by the CMS collaboration to precisely measure the Higgs boson decay to photon pairs. Here, machine learning methods were deployed to increase photon energy resolution, obtaining improvements ranging from 5 to 30% [15].

the three-dimensional printing of scintillator elements [16], which allows highly non-trivial arrangements within detector volumes; similarly, advances in manufacturing silicon pixel detectors have expanded design possibilities considerably [17]. The second source is contributed by advancements in materials science and manufacturing options, and it presents a specific challenge to gradient-based optimization techniques [6] due to the discrete nature of the space of solutions, which does not directly lend itself to continuous scanning.

Only little more than a decade ago, the simultaneous optimization of 10,000 parameters, heterogeneous in nature and connected in non-trivial ways to the data-generation mechanisms at the basis of the extraction of information on the studied physical phenomena, would be considered beyond the state-of-the-art of computing tasks. Yet today, profit-driven optimization tasks routinely involve orders of magnitude higher complexity; e.g., while the exact parameter count of GPT-4 is undisclosed by OpenAI, recent estimates suggest it is likely in the hundreds of billions, potentially reaching 1.7 trillion parameters [18]. However, that task was significantly expensive — the sheer cost of electricity required for it is estimated at several hundred million US dollars. Although particle physics is not an inexpensive area of fundamental science, even large experiments are unlikely to be able to invest even a fraction of that sum in R&D studies any time soon.

In addition to the substantial computational costs (and significant carbon footprint) associated with thoroughly exploring high-dimensional design spaces, the required software tools must also be developed from scratch. Even the small-scale end-to-end optimization of MuOnE required one person-year of software development and analysis; larger and more complex systems may involve costs difficult to allocate, especially at early design stages when funding is uncertain.

To the rescue in this situation comes the relative non-specific nature of some of the ingredients of the problem. For example, the optimization of a ground-based gamma-ray detector array such as the one we discuss in Section 4 *infra* may require a fast simulation of extended atmospheric showers that may produce a good representation of the secondary particles flux on the ground, given the primary particle identity, energy, and incidence angle. The effort to develop a differentiable model of the secondary particles flux can be significant, but it may be worth spending it if the resulting code is reusable for future optimization tasks of other instruments. Similarly, the development of a differentiable model of the pattern recognition and regression procedures necessary to extract information on the particle flow and incident energy of hadrons in a calorimeter can be a very demanding task, but it becomes an attractive option if the code is designed in a way that ensures flexibility and adaptability to different software pipelines, such that it may be reused to handle detectors of different geometry, components, and materials. In other words, there is an intrinsic modularity in the software tasks performing the modeling of the data-generating processes, and the reconstruction of the resulting observable features in an instrument; leveraging that modularity makes the construction of differentiable programming models of hardware and software possible, enabling co-design optimization. The MODE Collaboration<sup>3</sup> has undertaken the ambitious task of developing reusable and modular software pipelines for gradient-based experiment optimization, and is in the process of building a library of software solutions to the optimization of experimental endeavors of low- to mid-scale complexity. This approach may provide in the long run the tools necessary to faster solutions to more and more complex optimization questions.

## 1.2. Multi-purposedness

The question that remains on the table is the one of coping with the ambitious nature of large experiments in fundamental science, that usually target not one, but a number of different scientific goals at the

same time. From a machine learning perspective, there is a quantum leap in the complexity scale when going from a single, well-defined optimization task (such as that of the MuOnE experiment we discussed *supra*) to multi-target optimization. The reason is connected to the mathematical specification of the loss function, or more conveniently the negative of the loss, which is sometimes called objective function, but which we will rather call *utility* in this work. In a multi-objective scenario, the utility function should be written as the sum of different terms, each of which gets multiplied by some constant to provide the means to specify our relative appraisal of the different components:

$$U = \lambda_1 U_1 + \lambda_2 U_2 + \dots \quad (1)$$

The above definition may induce non-convexity of the total utility function  $U$  in the parameter space we wish to explore, even if the two components  $U_1$  and  $U_2$  are separately convex, when there are interactions between the parameters of the two terms. The set of parameters which provides the maximum of  $U$  may also change in a non-linear manner following a modification of the ratio of the  $\lambda$  factors. This has two key implications: on one side, the solution becomes highly sensitive to the exact definition we choose for the utility; and on the other side, the maximization task is made considerably more complex, due to the appearance of local minima and saddle points. In such cases, the concept of *Pareto front* becomes relevant, as it helps organize combinations of parameter values that provide good utility outcomes using non-dominated sorting. The Pareto front consists of parameter values where any change to improve one component of the utility will necessarily decrease at least one other component. Non-dominated sorting enables the systematic exploration of the Pareto front, identifying suitable compromises between the often conflicting utility components. However, investigating the Pareto front may, in typical situations, be computationally quite expensive.

While Pareto front optimization ups the ante considerably, it is a solvable problem. From a pragmatic point of view, it is a negligible complication when we compare it to the much more problematic task of coming up with an agreed-upon list of utility components to be jointly optimized, and corresponding agreed-upon values of the  $\lambda$  multipliers: one may imagine that large scientific collaborations could be unwilling or unable to pull that off. The main purpose of this work is to examine that problem in some detail, to assess whether it is really a show-stopper in the way of the automated exploration of the configuration space of large experiments.

The present document is structured as follows. In Section 2 and Section 3 we discuss how an experiment-wide utility function can in principle be defined for experiments in particle physics, and in fact we argue how in some cases they are implicitly agreed upon by the researchers as they define their rules for data collection. In Section 4 we discuss a use case from particle astrophysics, where we may focus on other distinctive features of similar situations. We offer some concluding observations in Section 5.

## 2. Case study 1: Hadron collider physics with CDF

As a case study centered on experimentation in elementary particle physics, let us focus on the Collider Detector Facility (CDF) experiment at the Fermilab Tevatron collider. CDF was commissioned in the early eighties. It started operations with a brief demonstrative data-taking in 1985, then collected data in three main running periods from 1987 to 1988 (Run 0), from 1992 to 1996 (Run 1), and from 2000 to 2012 (Run 2, when the collaboration changed name to “CDF 2 collaboration” after a major upgrade). For the purpose of our investigation, we will focus on the first two data-taking runs of the experiment, which collected the data that eventually allowed the collaboration to claim discovery of the top quark [19].

<sup>3</sup> <https://mode-collaboration.github.io>.

## 2.1. The CDF experiment

The CDF experiment is an appropriate testbed to examine how the formulation of an experiment-wide utility function can be expressed in a HEP use case, because of its characteristics and the special role it played within subnuclear physics in the 1990ies.

CDF was conceived between the end of the 1970s and the start of the 1980s by a small collaboration of American, Italian, and Japanese physicists who set out to build a detector which would study the proton–antiproton collisions produced by the Tevatron collider, a machine which would allow collisions at a center-of-mass energy three times higher than that offered by the CERN  $S\bar{p}\bar{p}S$  machine, which was by then already in its commissioning phase. It was too late to challenge the CERN facility on the discovery of the  $W$  and  $Z$  bosons, so the main goal of CDF was the discovery of the last missing quark, top.

It is interesting to note that already in its early design stage CDF was understood to need a strong solenoidal field to measure with high precision the momentum of charged leptons that were expected from top quark decays; the competitor of CDF, the  $D\bar{O}$  experiment, would just a few years later be designed with the main goal of producing the best determination of the  $W$  boson mass, by constructing a high-performance electromagnetic calorimeter which, unhindered from any solenoid material, could achieve high precision in the measurement of electrons energy. This early focus on all-important single goals of the two collaborations (which later both turned out to benefit highly from the multi-purposedness and wide range of measurements and searches enabled by their redundant detection systems and by the Tevatron's unprecedented beam energy) probably has more to do with the competition between the United States and Europe at the HEP frontier than on the perceived worth of different scientific investigations. Still, for the sake of reflecting on the utility function of these experiments, we note that its definition in the early eighties was not really difficult, so that important decisions on the resources allocated to the construction of the detector subsystems could be straightforward to take: for CDF, anything that contributed to catching the top quark took priority. Incidentally, this focused approach made the decision to invest a conspicuous effort in the risky, unprecedented challenge of constructing and operating a silicon microvertex detector (SVX) in the core of CDF an easier one to take [20]. The SVX turned out to be instrumental in characterizing top pair-production events, through the identification of secondary vertices from  $b$ -hadron decays.

After the initial engineering run of 1985–86 and the first data-taking period of Run 0 had proven the good performance of the Tevatron collider, as well as demonstrated the quality of event reconstruction that the CDF detector could achieve, from the outset Run 1 of the Tevatron was expected to deliver a sufficient amount of data and the necessary precision of their reconstruction to warrant success in the discovery of top, if that particle was at reach of the collider energy. But in addition to the top quark, many CDF collaborators were hoping they would manage to make other fundamental, groundbreaking discoveries, by finding evidence of some new physics signal predicted by one of a number of scenarios that extended the Standard Model of particle physics. In fact, despite the incredible successes that the theory of elementary particles had until then achieved, it was generally believed that new phenomena would become accessible by studying physical reactions at the electroweak scale. In particular, some new physics scenarios that extended the Standard Model had just become popular among theorists; most notably, Supersymmetric models [21], which had as a natural mass scale for their particles the one of electroweak symmetry breaking, one which was then fully at reach of the 1.8 TeV collisions that the Tevatron collider was delivering; Technicolor theories [22] were also considered quite promising.

Further, the SVX was installed in CDF just before the start of Run 1, giving for the first time access to the determination of trajectories of charged tracks with micrometric precision to an experiment operating in the harsh radiation-riddled environment of hadron collisions.

Given the very large cross section for production of B hadron states at a proton–antiproton machine, a wealth of entirely new B physics measurements were for the first time at reach of the experimentalists, provided that enough data could be collected by track-based triggers.<sup>4</sup>

Finally, the possibility to test perturbative calculations of quantum chromodynamics (QCD) in a fully new energy regime, and up to several hundred GeV in the momenta of outgoing bodies in two-to-two processes, had spurred a lot of interest in QCD phenomenologists. Diffractive physics was also a very hot topic, and CDF included a small number of experimentalists who were highly motivated to shed more light on the complex low-energy phenomena involving the exchange of colorless quanta between the colliding particles.

To the above *pot-pourri* of areas of research where the state of the art of human knowledge could be pushed outwards by new CDF data analyses, one must add the development of new particle detectors for its own sake. A significant number of physicists in CDF were heavily involved in developing improvements in the detection techniques and in the design of new electronic chips to increase the potential of triggering or online data processing. CDF offered to be the ideal testing ground for those new hardware concepts. These ideas need to be included in any consideration of the mix of desiderata that motivated the CDF members in their research.

## 2.2. The CDF trigger system

At the start of Run 0 (1987), the CDF author list included 227 physicists from 22 institutions in 6 countries; these numbers kept growing during Run 0 and Run 1, and reached 481 from 47 institutions in 12 countries in the year 2002. The experiment was led by a pair of spokespersons, whose job included the search for a consensual view in the collaboration on important decisions concerning the operation of the detector and the scientific output of the experiment. They did so by consulting with leaders of several working groups, which included ones dealing with specific hardware subsystems, and ones dealing with the different physics analyses. Of special relevance was a Trigger working group, wherein decisions were taken on the allocation of data collection bandwidth to a number of different rules for the storage of interesting events during the runs of the machine.

During Run 0 and Run 1, the Tevatron yielded proton–antiproton collisions inside the core of the CDF detector at a rate of 280 kHz, yet the data acquisition system could only sustain a rate of about 50–75 events per second that could be fully stored on tape. The CDF trigger was the system responsible for online reconstruction and selection that progressively reduced the data rate, allowing for the best possible impedance matching of those two very different numbers. The trigger was therefore a crucial step turning detector readouts into analyzable data on tape, which the experiment relied upon to preserve its discovery and measurement potential.

The trigger system was organized in three levels. Level 1, operated by dedicated hardware boards, used pre-set selection cuts on fast primitives read in from specific detector components (calorimeter towers and muon chambers) to bring the initial inflow of data at the bunch crossing rate of 280 kHz to an output of about 1 kHz. Level 2 was a hybrid of dedicated hardware boards and optimized software algorithms, which could use a more fine-grained information from all detector components to produce a fast reconstruction of the main physics objects present in the event, as a preliminary step to apply a selection which could further reduce the throughput to a few hundred Hz. Level 3, which was run on optimized software in a commercial PC farm, finally performed a speed-optimized reconstruction of the

<sup>4</sup> Proper secondary vertex-sensitive track triggers were only developed for Run 2 of the Tevatron [23], but CDF in Run 1 could still rely on low-momentum electron and muon triggers for the collection of large data samples enriched with B hadron decays.



full events using the same algorithms used offline. The application of further selection requirements which could now be based on the measured physics objects brought the throughput further down to about 50–75 Hz, which was a rate at which the data acquisition system could sustain the writing of the full event information, properly reduced by previous processing, into data tapes. It is worth noting how the crucial action of Level 2 and Level 3 was performed in parallel. Level 2 in particular used four buffers acquiring the input data from Level 1; employing a simple queuing logic, these buffers distributed the input load to hardware processors on a “first come, first serve” basis. When all four buffers were busy processing events, data from any further input event accepted by Level 1 would be lost. Level 3 processors would also ignore any input from subsequent events when they were all busy.

The question posed by the large mismatch between input and output data rates could be simply put as how to decide what data to keep and what data to throw away. A trigger menu was devised, to set minimum energy or momentum thresholds on interesting reconstructed objects (electrons, muons, jets, missing transverse energy). The sum of the accept rates of the various trigger conditions had to not exceed 50 Hz or so, lest the data acquisition system would drown in its own dirty water, and generate dead time. Effectively, the detector would then be blind for a fraction of the time. For example, if an accept signal were given to 80 events per second, the first 50 would get written, yet the remaining 30 would get dropped, being overrun by new ones coming in; even if the dropped events were the most valuable to store, their relative value would never get to be appraised. Dead time was a democratic, but highly ineffective way of coping with a higher rate of proton–antiproton collisions than the trigger could handle.

The matter was made worse by the fact that, during its operation, the Tevatron collider always strove – and nobody would object, as this maximized the discovery potential of the experiments! – to achieve the highest density of protons and antiprotons in its beams. This caused each bunch crossing to generate a number variable from a few to a few tens of independent proton–antiproton collisions, depending on the run conditions. Every time a new run would start, its conditions would be different. Perhaps a trigger menu that had worked well the other day, when the Tevatron delivered a lower instantaneous luminosity to CDF, would today generate a huge dead time, because the larger number of collisions produced by the higher luminosity would generate a higher accept rate by the trigger; further, the relationship was non linear, as the pile-up of several low-energy proton–antiproton collisions that individually would not manage to fire the trigger conditions could conspire to collectively make it. The weekly “trigger meeting”, aptly held in a meeting room adjacent to the one hosting the Level 2 trigger hardware, was the venue where CDF physicists took informed decisions on which thresholds to raise from which trigger condition, in order to keep the dead time below a physiological level of about 10%.

To make a concrete example, there were so-called “high- $p_t$ ” electron and muon triggers. These were constituted by sets of conditions that ensured that the experiment would collect events with high-purity charged lepton candidates of high transverse momentum ( $p_t$ ), typically produced by the decay of  $W$  or  $Z$  bosons. Due to their significant acceptance of  $W$  boson decays, those triggers were ideal for collecting events which could be the result of the production of a pair of top quarks. The full functionality of high- $p_t$  electron and muon triggers had therefore to be preserved, hence their thresholds were never tweaked to cope with variations in instantaneous luminosity. This was a relatively easy decision, as the accept rate of the resulting data streams was relatively small, of the order of 1 Hz. But then there were jet triggers: in principle, jets are the commonest thing one can see in a proton–antiproton collision, and their high cross-section poses a direct threat to a fixed-throughput trigger system: so one could accept to collect only a small fraction of events with high-energy jets. However, many CDF scientists would have loved to collect as many jet events as possible. Perhaps 20 Hz of jet events could be stored? Well, that could be a sound decision, but then there were scientists who wanted to study B physics,

which relied on collecting low-momentum electrons and muons. Those, too, were relatively frequent. Should CDF collect more low-momentum leptons or more high-energy jets?

In truth, above the problem has been massively simplified, as the trigger menu in CDF included over a hundred different sets of requirements on the energy and type of objects measured by the online reconstruction. In addition, a dynamical prescaling system was in place to automatically reduce or increase the acceptance rate of specific trigger conditions, by discarding a varying fraction of the events accepted by those conditions. Dynamic prescaling was applied to “expendable” datasets, whose collection was in other words conditional to the avoidance of dead time. But the bottomline remains clear: deciding what kind of datasets to collect, and how many events to acquire from each, was equivalent to deciding what was the overall goal of the experiment. This means that in a way, the trigger of a limited-bandwidth hadron collider experiment is tantamount to an operative definition of its utility function.

### 2.3. A proxy to the scientific value of different analysis targets

Eventually, CDF did discover the top quark, as it did produce groundbreaking measurement in a number of other measurements and searches. From the vantage point of a *a posteriori* examination, we have a way to judge on the relative merit of the various physics results that the experiment produced, which correlates with the impact that those measurements had in the science CDF advanced. This is constituted by the number of citations of the publications that the experiment published based on those measurements. While such a metric may not be objectively the best one to assess the scientific worth of different results, it has the merit of being simple to define and relatively unbiased, as long as it compares articles from the same experiment.

Examining the 20 most cited works published by the CDF collaboration based on data collected before the start of Run 2 of the Tevatron, we get the following picture for the number of citations as of January 10, 2025 (see Table 1 for more detail):

1. Observation of top quark production, and related publications [24–26]: **6840** citations
2.  $J/\psi$  and  $\psi(2S)$  production, plus related publications [27–29]: **1359** citations
3. Detector descriptions [20,30,31]: **2026** citations
4. QCD results on jet properties [32–34]: **1327** citations
5. Measurement of total p-antip cross section [35,36]: **869** citations
6.  $B_c$  observation and a related publication [20,37]: **827** citations
7. Double parton scattering [38,39]: **770** citations
8. W boson mass measurement [40,41]: **658** citations

Leaving aside the three detector-related articles, after observing that the citations they got (14.5% of the total) give us a scale of the relative importance of detector developments *per se*, we are left with the task of aggregating the seven remaining classes of publications together based on the triggers that produced the respective analyses. First of all, the top and  $W$  mass publications were mainly produced by analyzing data collected by high-transverse-momentum lepton triggers, and brought CDF authors collectively 6840 citations; the  $J/\psi$  and  $b$  physics papers, mainly produced through the study of low- $p_t$  lepton triggers, brought 2186 more citations; and the QCD-based measurements, brought in by study of single and multiple jets and minimum bias triggers, a total of 2966 citations. By normalizing these numbers, we may finally assess that the “relative value” of the high- $p_t$  lepton triggers, low- $p_t$  lepton triggers, and hadronic activity-based triggers in CDF during Run 1 (the period of data taking which produced the publications above discussed) is respectively of 57.0%, 18.2%, and 24.7%, with sub-percent statistical uncertainties (albeit admittedly also carrying a much more significant additional error from their ill-defined nature).

A small refinement of the above numbers comes if we observe that three of the publications referring to the broad category of QCD

**Table 1**

The twenty CDF publications based on Run 0 and Run 1 data which collected the largest numbers of citations as of January 10 2025. The Category column indicates the type of physics goals addressed by the study: category 0 refers to detector descriptions; category 1 includes top quark searches and measurements, and electroweak studies; category 2 includes studies of heavy-flavored hadron properties; and category 3 includes QCD-related studies. See the text for detail.

Item	Abbreviated title	Year	Category	Citations	Notes
1	Observation of top quark production [...] [24]	1995	1	4288	
2	The CDF Detector: an Overview [30]	1988	0	1988	Detector description
3	Evidence for top quark production [...] (PRD) [25]	1994	1	1062	
4	Evidence for top quark production [...] (PRL) [26]	1994	1	832	
5	The Topology of 3-jet events [32]	1992	3	571	
6	$J/\psi$ and $\psi(2S)$ production [...] [27]	1997	2	515	
7	Observation of the $B_c$ meson [...] [37]	1998	2	469	
8	Measurement of the $p\bar{p}$ Total Cross-Section [...] [35]	1993	3	452	Special run
9	Inclusive $J/\psi$ , $\psi(2S)$ and b-quark production [...] [28]	1992	2	448	
10	Double parton scattering in $p\bar{p}$ collisions [...] [38]	1997	3	448	
11	Inclusive jet cross section [...] [36]	1996	3	417	
12	CDF central and end-wall calorimeter [31]	1988	0	403	Detector description
13	Production of $J/\psi$ mesons from $\chi_c$ meson decays [...] [29]	1997	2	396	
14	Charged Jet Evolution and the Underlying event [...] [33]	2002	3	395	Minimum bias data
15	Transverse Momentum Distributions of Charged Particles [...] [34]	1988	3	361	Special run/MB data
16	Observation of $B_c$ mesons [...] [42]	1998	2	358	
17	The Silicon Vertex Detector [...] [20]	1994	0	354	Detector description
18	A Measurement of the $W$ boson mass [40]	1990	1	336	
19	Study of four jet events and evidence for double parton [...] [39]	1993	3	322	
20	A Measurement of the $W$ boson mass [...] [41]	1990	1	322	

measurements (numbers 8, 14, and 15 in the list shown in Table 1) were produced through the analysis of data collected by special runs, where the experiment could afford to acquire a significant number of so-called “minimum bias” events that produced very little detected signals in the detector, and/or when the Tevatron collider was run at a lower energy corresponding to that previously studied by the  $S\bar{p}\bar{p}S$  collider at CERN. It is an interesting fact *per se* that three of the 20 most cited publications by CDF in Run 0 and Run 1 were based on the analysis of those very quick runs, which subtracted an irrelevant amount of integrated luminosity to the total bounty collected by the experiment! The choice of devoting run time to these QCD measurements was never controversial; the above observation not only justifies the special procedure *a posteriori*, but it allows us to note how the existence of a facility that sheds light on previously unexplored territory (the highest center-of-mass energy ever achieved in hadron collisions) offers scientific value through the diversity of studies it enables, some of which may not be a major driver of the decision to build the facility in the first place. In any case, by excluding also those three publications we remain with 14 which break down into these fractions of citations: 63.4% to top and electroweak related analyses, brought in by high- $p_t$  lepton triggers; 20.3% to b-quark analyses relying on low-transverse-momentum leptons; and 16.3% to QCD analyses relying on regular jet trigger data.

And what were the overall rates of the triggers producing the three classes of events used by the above analysis categories in CDF during Run 1? Alas, those numbers are not easy to determine with precision, as they varied significantly over time; apparently, they were not reported in detail in published documents. Furthermore, the CDF collaboration made the rather upsetting move of not making public their internal CDF notes archive after the experiment stopped its data-taking phase.<sup>5</sup> What we may note here is the high- $p_t$  lepton triggers took way less than 57.0% of the Level-2 output bandwidth, primarily because of the rarity of the corresponding Level-1 accept rates of data streams containing lepton-rich primitives. Even if the momentum thresholds of the trigger objects of interest (electromagnetic clusters in the calorimeter, and muon stubs in the muon chambers) as well as the objects’ quality criteria had been further loosened from their operation values to try and increase the data statistics, the discovery potential of the collected data would not have changed significantly, as

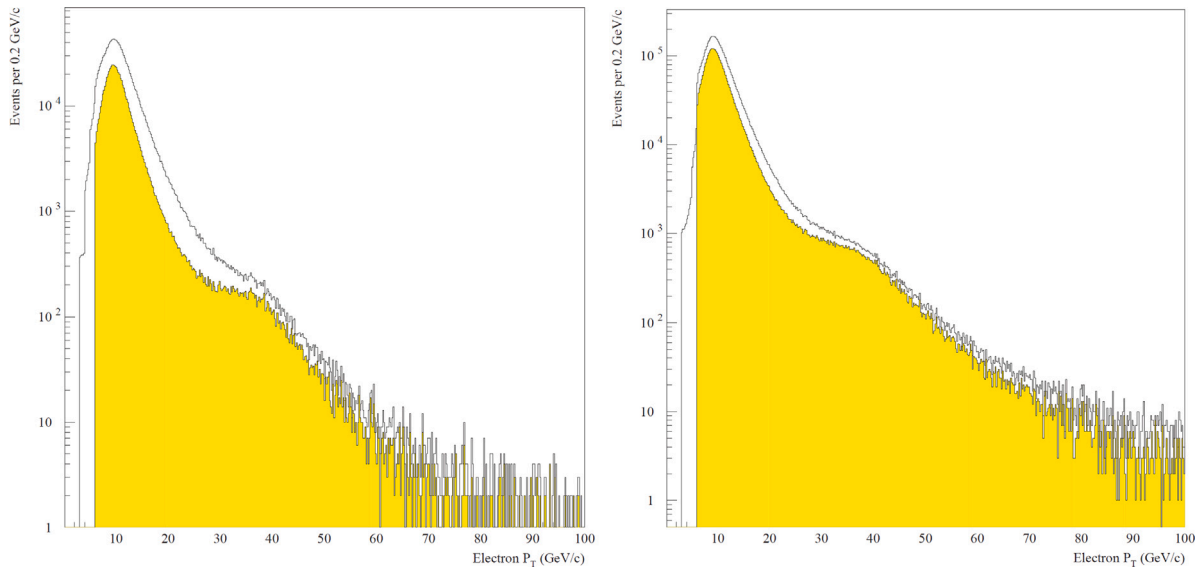
most of the reconstructible top quark events would not require looser thresholds to be acquired (see Fig. 2). In this sense, the bandwidth occupied by top-discovery-related datasets is not a good gauge of the relative value of the resulting physics output by the experiment. At first sight, this also suggests that a top discovery could in principle have been produced by the experiment even if the data acquisition system had been significantly reduced in capacity. However, massive use of low-transverse-momentum lepton triggers as well as of jet triggers was made by the analyses which proved the existence of top quark decays in CDF Run 1 data, because the resulting data samples were instrumental in providing crucial cross-checks [24,25] and, in the case of jet triggers, a data-driven parametrization of the rate of fake secondary vertex  $b$ -quark tags in jets [25]. What we can conclude from the above is that even for the single goal of discovering a new particle at a hadron collider, the optimal choice of selection requirements to store collision data is by no means trivial, and it is a problem only amenable by co-design techniques: *e.g.*, a larger amount of low-momentum leptons and jet data would have allowed CDF to develop more performant B-tagging algorithms, and to estimate with higher precision their efficiency; but the resulting bandwidth increase of the corresponding triggers would then have had to be balanced carefully, something which can only be done by software tools available since the last decade.

Analyses which could have benefited from a loosening of the energy, momentum, and quality thresholds of low-transverse-momentum electron and muon candidates clearly include those of heavy-flavored hadrons, which as we showed above brought an overall 20.3% of the citations to prominent CDF Run 1 publications. Here again, the rate of low-transverse momentum electron and muon triggers were much lower than that fraction: *e.g.*, the number of events collected in the “inclusive electron” data stream in Run 1A (1B) amounted to 1,874,289 (6,125,629) events (see Fig. 2), which corresponds to far less than 1% of the output bandwidth of the experiment during typical Run 1 conditions<sup>6</sup>; the number of inclusive muon events was even smaller.

Instead, it is much more interesting to observe how QCD triggers, which were almost invariably the ones whose energy thresholds were tweaked upward by experimenters when dead-time had to be reduced, ended up producing a whopping 24.7% of the high-impact scientific results of CDF in Run 1. Such a datum highlights how the actual

<sup>5</sup> Formally, the CDF Collaboration still exists, and a trickle of publications continues to appear on scientific journals to this day, a testimony of the immense value of the data the experiment produced.

<sup>6</sup> As the integrated luminosity of the mentioned datasets is of  $19.4 \pm 0.8 \text{ pb}^{-1}$  and  $81.5 \pm 9.0 \text{ pb}^{-1}$  [44], the datasets correspond to cross sections in the 70–100 nb range. This, noted that the instantaneous luminosity during Run 1 averaged to less than  $10^{31} \text{ cm}^{-2} \text{ s}^{-1}$ , indicates a corresponding Level-3 accept rate of the inclusive electron triggers of less than 1 Hz.



**Fig. 2.** Transverse momentum distribution of electron candidates collected by inclusive electron triggers in Run 1A (left) and Run 1B (right) before (empty histogram) and after standard quality cuts (yellow histogram) [43]. The contribution from  $W$  and  $Z$  decays is evident in both graphs; it is not significantly affected by the momentum-dependent selection operated by the trigger, which causes the turn-on shape of the distributions at transverse momenta of about 10 GeV. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)

scientific impact of a multi-purpose experiment in fundamental physics may turn out to be quite different from the intended one. It also pays justice to those CDF members who indefatigably kept fighting during trigger meetings to avert the prescaling of jet-based trigger selections, or their plain removal from the trigger list.<sup>7</sup>

The bottom-line of the above discussion of CDF triggers in Run 1, for what concerns the focus of this work, is that at hadron colliders the limited bandwidth of the data acquisition system demands a quantitative appraisal of the different goals of the experiment; yet the defined working point of the trigger rates and the size of the resulting datasets can only be taken as a qualitative proxy to the overall utility function.

#### 2.4. Assessing the weight of sensitivity to potential new discoveries

One final word concerns the “missing in action” category of CDF articles describing searches for new physics beyond the Standard Model. It is indeed quite striking to observe how no CDF publication of this kind appears among the twenty most-cited works in Table 1. The experiment invested significant effort in designing and maintaining triggers that collected, *e.g.*, events with photons, missing transverse energy, isolated tracks, or other non-mainstream trigger primitives potentially sensitive to exotic phenomena; collectively, these data streams had a substantial acceptance rate, perhaps of the order of 10 to 20 Hz. Additionally, an “express stream” specifically designed to quickly identify exotic signatures potentially signaling new physics operated throughout data taking, enabling rapid turnaround for possible groundbreaking discoveries. To facilitate immediate flagging of potentially anomalous events, an automated software named *PhysMon* (designed by David Saltzberg before the start of Run 1) scanned express stream data in real-time, alerting collaborators daily via a dedicated email list. Unfortunately, despite occasional excitement and even the production of theoretical papers triggered by intriguing yet never fully explained events (see, *e.g.*, [46]), CDF never produced a solid claim of physics beyond the Standard Model.<sup>8</sup>

Such mismatches between significant resource investment and the eventual scientific return are intrinsically impossible to avoid, as success depends crucially on the unfathomable existence or absence of new physics processes. Even a single clear signal beyond the Standard Model, had it emerged, would very likely have become the highest-cited CDF publication today. Without the benefit of hindsight, experiments must therefore continuously navigate difficult choices about allocating effort, resources, and bandwidth to “fishing expeditions”. These decisions inevitably reflect not only scientific considerations and prior beliefs but also factors related to funding dynamics and political pressures. Such external considerations substantially influence the relative weights given to searches for unknown phenomena versus guaranteed incremental advancements provided by precision measurements. Ultimately, a significant degree of subjectivity appears impossible to eliminate entirely from the utility functions guiding HEP experiments.

### 3. Case study 2: Definition of a multi-target utility at lepton colliders

In the previous section we have seen how the definition of an experiment-wide utility function in hadron collider physics experiments may be implicitly related to the issue of reducing the input data flow, which demands crucial decisions in the settings of the trigger system. Other HEP experiments that deal with smaller data rates may not need to take such draconian measures in real-time, but they are often still confronted with defining proxies of an experiment-wide utility in the definition of their data collection schedule. For example, electron–positron colliders of low energy, such as the Beijing Electron Synchrotron (BES) or the Super-KEK-B accelerator in Tsukuba, need to decide how much data to deliver to the detector at different energy points. The choice directly impacts the precision of different measurements of particle properties, so the run schedule tightly correlates with the relative value given by experimental groups to the different scientific outputs potentially at reach by their detectors.

<sup>7</sup> In the book “*Anomaly - Collider Physics and the Search for New Phenomena at Fermilab*” [45], T. Dorigo describes such a situation in detail.

<sup>8</sup> A long and tormented Run 1 analysis did evidence some unexplained anomalies in events containing  $W$  bosons and multiple jets [47], and an

analysis of inclusive lepton data in Run 2 mysteriously containing muon-rich jets also briefly made headlines around the world [48], but these claims were never confirmed.

**Table 2**

Publications by the BES-III collaborations listed in decreasing order of number of citations, with mention of datasets used for the analysis and the center-of-mass energy of the collisions. For some publications the exact integrated luminosity is not reported by the experiment in the article.

Item	Abbreviated title	Year	Cit.	L (fb <sup>-1</sup> )	$\sqrt{s}$ (GeV)
1	Design and Construction of the BES III Detector [49]	2010	1269	N/A	N/A
2	Observation of a Charged Charmonium-like Structure in $e^+e^- \rightarrow \pi^+\pi^-J/\psi$ [50]	2013	1156	0.525	4.260
3	Future Physics Programme of BES III [51]	2020	594	N/A	N/A
4	Observation of a [...] $Z_c(4020)$ and Search for the $Z_c(3900)$ [52]	2013	533	3.400	3.900–4.420
5	Observation of [...] Structure in $e^+e^- \rightarrow (D^*\bar{D}^*)\pi^\pm$ at $\sqrt{s} = 4.26$ GeV [53]	2014	413	0.830	4.260
6	Observation of a Charged $(D\bar{D}^*)^\pm$ Mass Peak in $e^+e^- \rightarrow \pi D\bar{D}^*$ at $\sqrt{s} = 4.26$ GeV [54]	2014	391	0.525	4.260
7	Measurement of the $e^+e^-\pi^+\pi^-$ cross section [...] using initial state radiation [55]	2016	341	2.930	3.773
8	Precise measurement of the $e^+e^- \rightarrow \pi^+\pi^-J/\psi$ cross section [...] 3.77 to 4.60 GeV [56]	2017	285	9.000	3.770–4.600
9	Observation of $e^+e^-\gamma X(3872)$ at BESIII [57]	2014	251	N/R	4.009–4.420
10	Observation of a Near-Threshold Structure [...] in $e^+e^- \rightarrow K^+(D_s^-D^{*0} + D_s^{*-}D^0)$ [58]	2021	233	3.700	4.628–4.698
11	Evidence of Two Resonant Structures in $e^+e^- \rightarrow \pi^+\pi^-h_c$ [59]	2017	205	5.600	3.896–4.600
12	Precision measurement of the integrated luminosity [...] [60]	2015	203	N/R	3.097
13	Confirmation of [...] and observation of [...] $X(2370)$ in $J/\psi \rightarrow \gamma\pi^+\pi^-\eta'$ [61]	2011	200	N/R	3.097
14	Polarization and Entanglement in Baryon–Antibaryon Pair Production [...] [62]	2019	198	N/R	3.097
15	Search for hadronic transition $\chi_{cJ} \rightarrow \eta_c\pi^+\pi^-$ and observation of $\chi_{cJ} \rightarrow K\bar{K}\pi\pi$ [63]	2013	195	N/R	3.686
16	Measurements of absolute hadronic branching fractions of $\Lambda_c^+$ baryon [64]	2016	191	0.567	4.599
17	Observation of $Z_c(3900)^0$ in $e^+e^- \rightarrow \pi^0\pi^0J/\psi$ [65]	2015	188	2.810	4.190–4.420
18	Measurement of the integrated luminosities [...] at $\sqrt{s} = 3.650$ and $3.773$ GeV [66]	2013	184	2.917	3.650–3.773
19	Branching fraction measurements of $\chi_{c0}$ and $\chi_{c2}$ to $\pi^0\pi^0$ and $\eta\eta$ [67]	2010	179	N/R	3.686
20	Observation of $e^+e^- \rightarrow \pi^0\pi^0h_c$ and a Neutral Charmoniumlike Structure $Z_c(4020)^0$ [68]	2014	166	2.461	4.230–4.360

The BES-III experiment is a clear example of how the precise run schedule defines an experiment-wide utility function. BES-III studies electro-positron collisions in an energy range where tau physics, charm physics, and exotic hadron resonances may be studied in high detail, as are the properties of QCD at low energy. The number of interesting reactions, bound states, and phenomena that the experiment is sensitive to is very large, but each of these targets demands the collection of a different dataset. Wide-range energy scans allow the study of the  $R$  evolution and cross-section dependencies on center-of-mass energy, as well as determine form factors in hadron production, while runs at the peak of resonances such as the  $\psi(2S)$  enable precision measurements of charm states. Further, narrow scans at threshold are needed for measurements of the tau lepton mass, as well as for the estimate of hadron resonances parameters and branching fractions; while for searches of narrow resonances, fine-grained energy scans are needed where at each energy point a sufficient amount of data is collected to warrant observation of potential new states. In such a situation, the experiment must carefully design a run schedule which may optimize the scientific throughput by allocating days of running to different center-of-mass energy values.

A look at the 20 most cited among the 659 publications produced by the BES-III collaboration to date (see Table 2) shows how articles reporting the observation of new hadron states are among those of highest interest for the HEP community; they collectively brought 71.3% of the citations of analysis articles (i.e., the 20 listed except the first and third one, which deal with the detector and the future plan of the experiment, respectively) in the reported set. Indeed, in this energy regime hadronic physics is to date still concealing many ill-understood phenomena, and empirical classification remains a prime tool of investigation. A larger statistics of collisions at finely spaced energy points might allow BES-III to shed more light on the complex cross-section behavior above the  $\psi(2S)$  region – a problem known as the “Y” problem; the scientific worth of data enabling other measurements of charmonium properties, form factors, and tau lepton mass is easier to assess.

The decision on how to divide the run time of the accelerator between fine-grained, wide-range energy scans and careful measurements at threshold or at the peak of known resonances is a complex one to make, as it involves a measure of arbitrariness. Yet the definition of a precise plan is still possible: indeed, in their 2020 study [51] the BES-III collaboration does not shy away from detailing the luminosity desiderata of BES-III for the near future, after appraising the various goals of the experiment. We report in Table 3 a concise version of a table (Table 7.1) presented in the Summary section of that document,

**Table 3**

Declared goals of future BES-III studies of  $e^+e^-$  collisions, collision energy, and required integrated luminosity and run time for the current (C) or an upgraded (U) BES-II detector; N/R stands for non reported.

Source: Excerpt from [51].

Energy (GeV)	Physics motivations	Final L (fb <sup>-1</sup> )	Run days C	Run days U
1.8–2.0	R values, nucleon cross sections	0.1	60	50
2.0–3.1	R values, cross sections	N/R	250	80
3.686	Light hadron and glueball charmonium decays	4.5	150	90
3.770	D meson decays	20.0	610	360
4.180	$D_s$ decay, XYZ states	6.0	140	50
4.0–4.6	XYZ, higher charmonia cross sections	30.0	770	310
4.6–4.9	Charmed baryon and XYZ cross sections	15.0	1490	600
4.74	$\Sigma_c\bar{\Lambda}_c$ cross sections	1.0	100	40
4.91	$\Sigma_c\bar{\Sigma}_c$ cross sections	1.0	120	50
4.95	$\Xi_c$ decays	1.0	130	50

to show how BES-III explicitly quotes the exact number of days to run at each energy point, separately for a current or upgraded version of the detector. We believe the table closely tracks the scientific value that the collaboration appraises to the measurements and searches that justify the collection of those datasets and is, therefore, a good illustration of how a multi-purpose HEP experiment can indeed formulate a global utility function.

#### 4. Case study 3: Gamma-Ray astrophysics

Here we examine the case of a ground-based array of water Cherenkov detectors measuring high-energy gamma rays in the TeV–PeV range, in order to study how the different scientific goals of an astro-particle physics experiment may impose conflicting requirements on the detector geometry.

##### 4.1. The SWGO Gamma-Ray array

We consider a number  $N_{det}$  of idealized detector units placed on the ground at high altitude (4800 m above sea level), corresponding to the conditions of the site (Pampa la Bola, in the Chilean Andes) chosen by



the SWGO Collaboration [69] for a future detector of this kind. SWGO will detect radiation, electrons and muons of extensive atmospheric particle showers generated in the atmosphere by cosmic gamma-rays (the signal) and distinguish them from the large background from charged particles (mostly protons). A global goal of the experiment is to correctly tag and reconstruct the rare gamma-ray events in the haystack of cosmic rays: 1 every  $10^4$ – $10^5$ . The particle shower hits the ground within an ellipse spanning thousands to millions of square meters around the shower development axis. The science case of SWGO is broad, encompassing studies of PeV emission morphology near the galactic center, cosmic ray transport in galactic sources like pulsar wind nebulae, variability in active galactic nuclei, and searches for signals from dark matter-dominated objects [69]. The instrument's figure of merit is therefore defined by its energy and angular resolution, as well as by the gamma-ray acceptance and background discrimination capability; these ingredients play in differently for each of the scientific use cases mentioned above.

In order to meaningfully sample the area on the ground hit by shower particles, an array of about 6000 water Cherenkov detectors is foreseen, with each detector composed by a water-filled tank with a diameter of a few meters. The most advantageous ground distribution of such an array is debatable: Is it better to strive for a homogeneous sampling, or to realize a decreasing fill factor from the center? Are radially symmetric layouts better than asymmetric ones? And does this depend on the specific science case considered?

For the sake of studying the dependence of the instrument performance on the layout geometry, it is possible to abstract away from the details of the detection process, and assume that each unit measures with 100% efficiency all secondary particles traversing it with an energy above some minimal threshold (e.g.,  $E > 10$  MeV), with perfect discrimination between soft (electrons, positrons, gamma rays) and hard component (muons). Such an idealized detector performance is deliberately chosen to eliminate confounding effects arising from realistic detector imperfections. This allows us to isolate how the geometry alone impacts the ultimate performance of the instrument. Effects of realistic detector performance and their coupling with detector layout details can be investigated at a later refinement stage.

In order to study the extraction of meaningful physical measurements from cosmic showers, that include as we mentioned above a large background from primary protons along with the gamma-ray signal component, we use a closed-form parametrization of the radial distribution  $dN/dR$  of the number of secondary particles of different kinds around the shower core, expressed as a function of primary particle identity (proton or gamma ray), energy  $E$ , and angles of incidence (polar angle  $\theta$  and azimuthal angle  $\phi$ ), which we have developed to study the optimization of the SWGO array in [70]. Following that work, we consider the problem of defining an array of  $N_{det}$  detector units laid on flat ground as a  $(2N_{det} - 3)$ -parameter optimization problem, with the  $i^{th}$  detector contributing with the  $x_i$  and  $y_i$  coordinates of its center, and with three degrees of freedom being defined by the center of the array – set to  $(0, 0)$  – and the choice of the azimuthal direction of the  $x$  axis in the plane. The problem can be simplified by considering “macro-tank” clusters of  $M$  tightly packed detector units, which reduces the dimensionality by a corresponding factor; following [70],  $M$  is here chosen to be 19 (for a three-ring hexagonal cluster around a central tank at position  $x_i, y_i$ , with now  $i = 1, \dots, N_{det}/M$ ); a further reduction of dimensionality by a factor 3 is operated by enforcing 120-degree symmetry around the center of the array. Starting from an arbitrary default configuration, the array may be modified by letting the position of each triplet of macro-tanks change following the gradient of a utility function which we wish to maximize:

$$x_i \rightarrow x_i + \eta_i \frac{dU}{dx_i} \quad y_i \rightarrow y_i + \eta_i \frac{dU}{dy_i} \quad (2)$$

where the  $\eta_i$  parameters are dynamically adjusted to dampen oscillating behavior and rewarding consistent movements of each detector in one specific direction on the plane, and where a vectorial average of

gradients over the three macro-tanks laying in a symmetric position around the center is implicit. The utility function and its derivatives with respect to a change of each detector's position on the ground are computed using at each iteration a batch of simulated gamma and proton showers, whose parameters ( $E$ ,  $\theta$ ,  $\phi$ , and shower core position on the ground  $X_0, Y_0$ ) are reconstructed with a likelihood maximization separately under each of the two primary hypotheses. The log-likelihood ratio for the two hypotheses serves as a test statistic to discriminate the signal component. For more details, we refer the interested reader to the stem article [70].

#### 4.2. Choices of the utility

While Ref. [70] discusses details of the optimization task and the developed software, exemplifying its functionality and providing evidence for significant potential improvements on two different choices of a utility function, here we wish to focus our attention on the definition of a multi-target utility function and on the related issues that arise in a specific use case not considered there, working our way to examine the resulting three-dimensional Pareto front.

The detection and precise measurement of high-energy gamma rays opens the way to a number of studies of astrophysical sources and phenomena in the cosmos. As discussed previously for particle collider detectors, it is challenging to precisely define the relative importance of gamma-ray flux accuracy, energy resolution, and angular resolution. Further, the discrimination of proton from gamma primaries affects directly the overall performance of the instrument at the very high-end of the energy spectrum, where the gamma signal component is drowned in hadronic showers backgrounds that exceed it by four to five orders of magnitude.

We rewrite the utility function discussed in [70] below:

$$U_1 = \lambda_{GF} U_{GF} + \lambda_{IR} U_{IR} + \lambda_{PR} U_{PR} \quad (3)$$

where the utility  $U_1$  is composed of three partly independent contributions:  $U_{GF}$ , a term proportional to the precision in the determination of the gamma-ray flux for a given integration time of the observatory (defined as the estimated flux divided by its uncertainty);  $U_{IR}$ , a term inversely proportional to the expectation value of the uncertainty in the energy of detected gamma rays integrated over the full considered energy spectrum; and  $U_{PR}$ , a term inversely proportional to the expected uncertainty in the direction of incidence of the gamma rays. This definition allows us to exemplify the significant challenges posed by the optimization of the detector layout. To optimize the detector's positions in the plane, defined by their  $x_i, y_i$  coordinates (with  $i = 1, \dots, N_{det}$ , or if macro-units consisting of several detector units are used,  $i = 1, \dots, N_{det}/M$  with  $M$  the number of units per macro-unit), we need to obtain the gradient updates of Eq. (2). This involves computing the following expressions involving derivatives of the  $U_1$  components:

$$\frac{dU}{dx_i} = w_{GF} \lambda_{GF} \frac{dU_{GF}}{dx_i} + w_{IR} \lambda_{IR} \frac{dU_{IR}}{dx_i} + w_{PR} \lambda_{PR} \frac{dU_{PR}}{dx_i} \quad (4)$$

$$\frac{dU}{dy_i} = w_{GF} \lambda_{GF} \frac{dU_{GF}}{dy_i} + w_{IR} \lambda_{IR} \frac{dU_{IR}}{dy_i} + w_{PR} \lambda_{PR} \frac{dU_{PR}}{dy_i}. \quad (5)$$

Here, the introduced parameters  $w_{GF}$ ,  $w_{IR}$ , and  $w_{PR}$  are gradient-scaling factors introduced to balance the relative contributions of the gradients of the three utility components. While the coefficients  $\lambda_{GF}, \lambda_{IR}, \lambda_{PR}$  reflect the scientific priorities of the experiment, the scaling factors  $w$  need to be adjusted dynamically during optimization to account for the disparity in the magnitudes of the gradient components. This ensures that no single term dominates the updates to the detector positions, thereby improving convergence and maintaining a balanced optimization. The need for such scaling becomes apparent when considering the sensitivity of each term to detector movements. The  $U_{GF}$  term, for example, can experience large changes due to the outward displacement of a single detector located on the array's outer rim, which allows it to detect showers previously outside its range. In

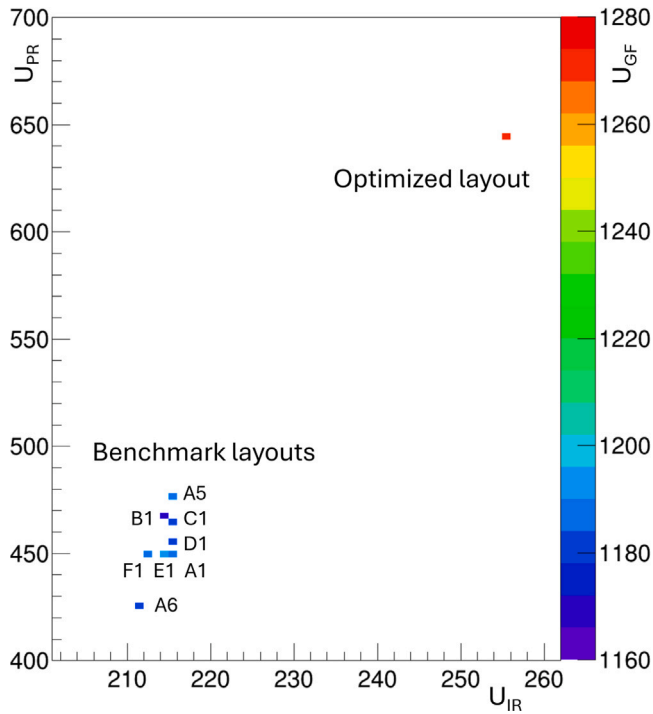


Fig. 3. Value of the three components of the utility for 330 19-tank macro-units. The 8 benchmark layouts A1, A5, A6, B1, C1, D1, E1, F1 [70] are compared to an optimized layout. See the text for detail.

contrast, the  $U_{IR}$  and  $U_{PR}$  terms exhibit much smaller variations with individual detector movements, as their definitions involve averages over all detected showers, where only a subset of detectors contribute to the reconstruction of any given event.

Without the  $w$  scaling factors, the gradient contributions from  $U_{GF}$  often dominate, leading to updates that prioritize flux-related improvements at the expense of energy and pointing resolutions. This can result in a configuration where  $U_{GF}$  increases significantly, but the overall utility  $U_1$  may decrease due to a substantial degradation in  $U_{IR}$  and  $U_{PR}$ . These imbalances are amplified by simultaneous updates of many detector positions, introducing complex dependencies that complicate the optimization landscape. The non-convex nature of the utility  $U_1$ , arising from the interplay of its components, further exacerbates this complexity, as it creates a landscape with potentially many local optima, saddle points, and flat regions. Consequently, the precise rule for dynamically varying the  $w$  parameters may be difficult to define and is liable to affect unpredictably the capability of the gradient descent algorithm to converge to solutions at the Pareto front.

We note that the optimization model from [70] significantly simplifies the full physics scenario of SWGO due to various approximations. However, these simplifications do not alter the fundamental conclusions on the complexity of the optimization task.

#### 4.3. An example

To illustrate how, in spite of the difficulties mentioned above, the search for an optimal point in the full-dimensional configuration space can be advantageous even in the presence of conflicting multi-target optimization criteria, we compare in Fig. 3 the three components of the utility function (Eq. (3)), for arrays of 330 19-unit macro-tanks laid down in configurations that mimic the A1, A5, A6, B1, C1, D1, E1, F1 SWGO benchmark layouts of 300-m radius described in [70]. We use for this purpose batches of 2000 showers sampled from a uniform energy distribution in  $[0.1 - 10]$  PeV. The figure shows how the eight benchmark configurations do not differ significantly in the

value of the three utility components  $U_{GF}$ ,  $U_{IR}$ , and  $U_{PR}$ ; by only examining the utility values of those eight solutions one might be led to believe that they lay close to the Pareto front. However, a 1000-epoch gradient descent loop searching for the best configuration of the same number of units, initially deployed within a uniformly populated 300-meter-radius circle, and targeting the optimization of the sum  $U_1^* = U_{IR} + U_{PR}$  alone, achieves quite significantly better performance than the highest achiever of the 8 benchmarks on both components of the maximized utility ( $+19 \pm 5\%$  on  $U_{IR}$ , and  $+35 \pm 4\%$  on  $U_{PR}$ ), while also marginally improving (by  $(6 \pm 4\%)$ ) the non-considered  $U_{GF}$  component. Further examples are discussed in [70].

Let us now return to the full definition of the  $U_1$  utility, and consider how the optimal definition of the layout of detectors may be strongly dependent on the goals of the experiment, by simply comparing optimized layouts that considered atmospheric showers sampled from different energy distributions: the assumed relative frequency of higher- or lower-energy showers in the training datasets acts as a weight in determining the relative value of showers of different energy in the utility. We control the distribution of showers energy by a coefficient in the power law defining their density function:

$$f(E) = ke^{-SE} \quad (6)$$

where the constant  $k$  is a normalization factor and where energy  $E$  is considered, in the considered setup, to vary from 100 TeV to 10 PeV. In practice, by changing the  $S$  parameter from 1.0 to  $-1.0$ , we effectively mimic an emphasis of the experiment on the physics of showers in the few hundred TeV range, to an emphasis in the several PeV range. By running optimization jobs that rely on data sampled from those different distributions we may therefore observe what is the corresponding effect on the layout; further, we may then compare the performance of the optimized layouts when tasked with reconstructing showers following a different energy distribution, to verify what price is paid in choosing layouts optimized with different energy priors.

Fig. 4 shows 10 optimized configurations of 1140 detection units, where the units are joined into 60 19-unit aggregates whose position is separately optimized. As mentioned *supra*, a triangular symmetry is enforced by the algorithm to reduce the dimensionality of the problem, such that the pattern on the ground is made up by three equal 120-degree sections. All in all, this corresponds to a 37-dimensional design problem which does not pose very significant computing demands; 1000-epoch updates are performed on the detector layouts, with a starting configuration packed in a circle around the origin.

We may observe that the overall appearance of the shown configurations is similar, with the strikingly common feature of six macro-tanks deployed at large distance from the center in a hexagonal arrangement, and a more complex central arrangement. The outer macro-tanks help in the reconstruction of angular and energy parameters of the most extended, highest-energy showers while guaranteeing that they get accepted by the trigger selection, which in these simulations filters for reconstruction showers that yield a signal in  $N \geq 50$  detectors.<sup>9</sup> Other than that, there are important differences: in particular, optimization runs targeting an energy spectrum with  $S > 0$  produce less extended layouts than ones using  $S < 0$ , as higher-energy showers more frequent in data used by the latter have a larger footprint on the ground, and thus benefit from wider arrays.

If we examine the graphs shown in Fig. 5, we see an example of the complexity of an optimization task with a multi-target utility. While the relative multipliers  $w$  of the three utility gradients are automatically equalized during the optimization, some of the utility components may still not increase as much as others (second panel from the left in the

<sup>9</sup> Given that each macro-unit is an aggregate of 19 detector tanks, in principle a signal in all detectors of a single macro-unit might suffice to meet the trigger criterion, but the reconstruction would then return very ill-measured shower parameters.

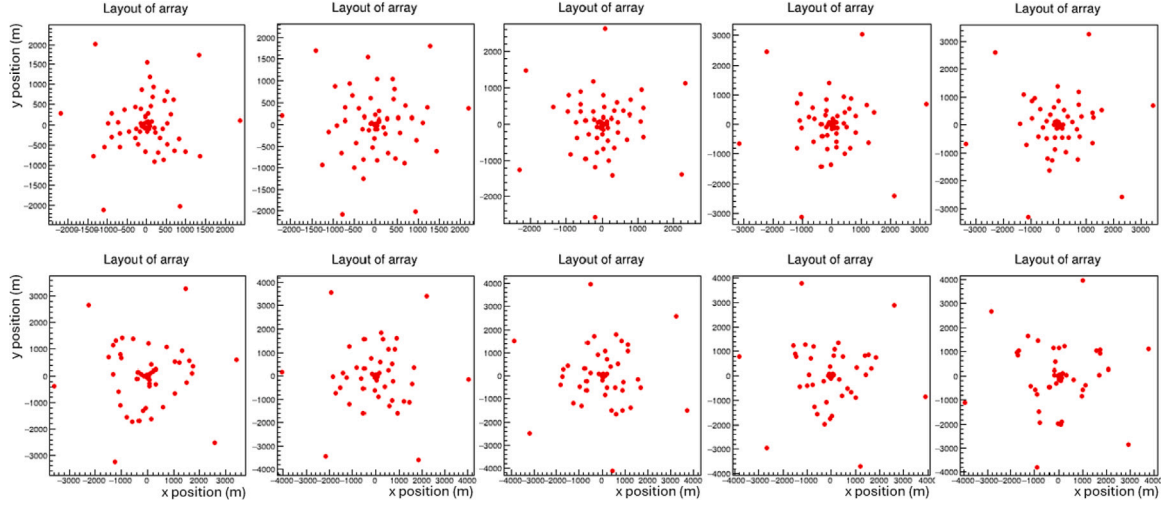


Fig. 4. Distribution on the ground of 60 macro-units resulting from the optimization of the utility function described in the text, for gamma rays following a power law with coefficients  $S$  varying from +1.0 (top left) to +0.2 (top right) and from 0.0 (bottom left) to  $-0.8$  (bottom right). See the text for more detail.

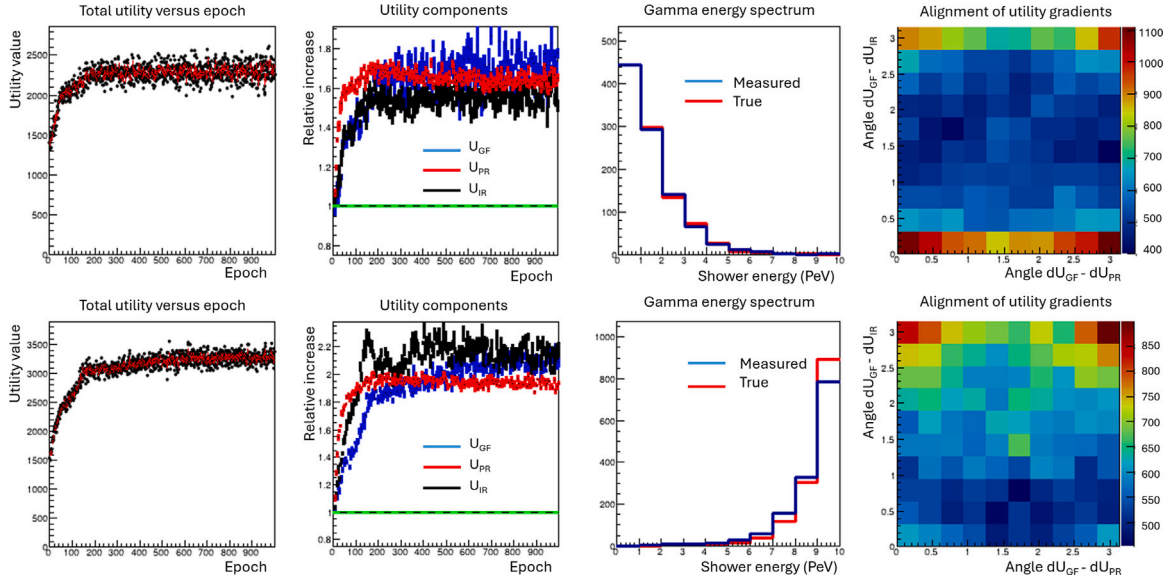


Fig. 5. From left to right, the graphs show: evolution of the total utility  $U_I$  as a function of epoch number; evolution of the three components of the utility versus epoch; generated and reconstructed spectrum of gamma-ray energies; and angle between the gradient vectors of pairs of utility components in the  $x, y$  plane during the optimization. The top line refers to the optimization based on a gamma-ray spectrum with  $S = 1.0$ ; the bottom line refers to the optimization based on a gamma-ray spectrum with  $S = -1.0$ . See the text for more details.

two rows). Also, the rightmost panels of the figure show how the three two-dimensional gradient vectors in the detector plane are usually not aligned with one another: for the high-energy  $S = 1.0$  case the gradient of the flux utility  $U_{GF}$  is mostly pointing in the opposite direction of the gradient of the integrated energy resolution piece  $U_{IR}$ ; the opposite behavior is apparent for the  $S = -1.0$  case. A more mixed behavior is seen for the alignment between the gradients of the  $U_{GF}$  and pointing resolution utility component  $U_{PR}$ .

If we compare the utility values corresponding to gamma-ray spectra of low-energy ( $S = 1.0$ ) or high-energy ( $S = -1.0$ ) for the 11 layouts optimized with fluxes of gamma rays generated with spectra corresponding to values of  $S$  varying from +1.0 to  $-1.0$  (see Table 4), we see how the choice of a target energy spectrum at optimization stage affects the potential of the experiment in the study of gamma rays of different energy: an array optimized for  $S = +1$  will lose 8%–10% of its effectiveness when studying gamma rays coming from a harder spectrum, and an array optimized for  $S = -1.0$  will similarly lose about 10% in terms of its utility.

Table 4

Comparison of the total utility  $U_I$  estimated for 11 different layouts (each previously optimized based on gamma rays with energy spectra corresponding to 11 values of  $S_{opt}$  from +1.0 to  $-1.0$ ) using fluxes of low-energy-rich ( $S = 1.0$ ) or high-energy-rich ( $S = -1.0$ ) gamma rays. Utility values should therefore be compared only within columns.

$S_{opt}$	$U_{S=1.0}$	$U_{S=-1.0}$
1.0	$2360 \pm 21$	$2880 \pm 7$
0.8	$2384 \pm 9$	$3072 \pm 8$
0.6	$2368 \pm 10$	$2901 \pm 7$
0.4	$2358 \pm 10$	$2979 \pm 10$
0.2	$2371 \pm 11$	$3070 \pm 9$
0.0	$2220 \pm 13$	$3153 \pm 9$
-0.2	$2255 \pm 11$	$3226 \pm 10$
-0.4	$2271 \pm 12$	$3187 \pm 9$
-0.6	$2247 \pm 12$	$3152 \pm 10$
-0.8	$2098 \pm 14$	$3049 \pm 9$
-1.0	$2234 \pm 12$	$3239 \pm 49$



It should be clear how the typical staged approach of construction of extended arrays such as SWGO – that may decide to build first an array with a limited number of detection units, and later add more units if more funding becomes available – may pose a challenge: *e.g.*, the collaboration might decide to target a lower energy range in the first stage, and thus build the array in a configuration optimized for that purpose; when adding more detectors to then target a higher energy, the first part of the detector remains however fixed (as changing the position of the initial units would constitute a very significant additional cost); the resulting extended array cannot then reach the same performance it would have if the position of all units had been optimized for the higher energy range in the first place.

#### 4.4. A word on flexible configurations

It is useful to mention here that in some cases it may be possible to design an experiment in a way that allows for seamless reconfiguration of its elements to fulfill different tasks, and reach optimality for each. That is the case, *e.g.*, of the ALMA array of radio-telescopes in northern Chile [71]. To achieve this kind of ultimate flexibility, the individual elements of the array are built such that they are movable with custom trucks over the high-altitude plain where the array is deployed. ALMA can thus arrange its radio-telescopes in a tight cluster to achieve high performance for large-aperture surveys, or over a wide area to maximize its angular resolution. The spectacular scientific output of the experiment is a testament to the success of this strategy.

It is necessary to point out that reconfigurable arrangements impose significant constraints on the weight and dimensions of the re-deployable elements, and that in particular, they cause a significant budget overhead over that of a static array. For an array such as SWGO, where each detector unit needs water, power, and signal cable connections to a central hub and counting house, and when the number of units is eventually foreseen to reach several thousand, the option is completely impractical. One possible workaround in this context is to construct and position a larger number of detector tanks on the ground than can initially be instrumented and operated. In the first phase, only a subset of these tanks would be filled with water and equipped with detection instruments, based on an optimized configuration for the initial scientific goals. In a second phase, water and detection instruments could be relocated to some of the initially inactive tanks, allowing the array configuration to be adapted to optimally study a different set of goals; or, if more funding became available, all tanks could be made operational. In any case, when such solutions can be implemented, budget remains the main driver of a final decision.

## 5. Conclusions

Machine learning tools today offer the possibility to explore the very large space of design configurations for fundamental science experiments in a continuous way, through the generation of differentiable surrogate models of the stochastic elements that may be present in the data generation processes. Such explorations may identify groundbreaking solutions by exploiting the subtle correlations and interplay of the many construction parameters of particle detectors and similar apparatus. In such a situation, the exact definition of an experiment-wide utility function, which may summarize the relative value of the various scientific goals of an experiment, becomes a crucial ingredient to enable an end-to-end optimization of its hardware and software.

In this work, we have argued that the definition of a global utility function is not only possible, but implicitly done already in many experimental situations, and therefore it should not be considered a significant hindrance in the way of global optimization programs. The case of the BES-III experiment mentioned in Section 3 is a glaring demonstration of the fact that, when it is important for a large scientific collaboration to quantitatively define the relative worth of different

scientific goals, the corresponding decisions are promptly and precisely taken.

As briefly discussed in Section 2.4, defining a global utility function becomes particularly challenging when experiments are potentially sensitive to groundbreaking new phenomena, and the subjective belief that the acquirable data might contain signals of fundamentally unknown processes (“unknown unknowns”) is not unanimously shared by collaborators. In such cases, experiments must inevitably strike compromises regarding the relative value assigned to precision measurements and exploratory searches. Decisions made during the experiment’s design, construction, and operation clearly reflect these inherent trade-offs; these are influenced not only by scientific judgments but also by external factors such as funding constraints and institutional priorities. Nonetheless, the very existence of such implicit choices suggests that explicitly formalizing these trade-offs in a mathematically defined utility function remains feasible, albeit with an unavoidable component of subjectivity.

The conflict between different targets of an exploratory instrument such as a ground-based array for cosmic ray studies has been discussed in Section 4. In that case, using an optimization program developed for the study of the layout of the SWGO observatory, we showed in practice how the design can be modified by changes in the preferred energy range of studied atmospheric showers. When possible, configurations that can be modified after construction would allow the experiment to retain optimality for a variable set of scientific goals during its lifetime.

## CRediT authorship contribution statement

**Tommaso Dorigo:** Writing – review & editing, Visualization, Software, Methodology, Investigation, Formal analysis, Data curation, Conceptualization. **Michele Doro:** Writing – review & editing, Supervision, Resources, Project administration, Conceptualization. **Max Aehle:** Writing – review & editing, Validation, Methodology. **Muhammad Awais:** Writing – review & editing, Visualization, Validation, Software. **Nicolas R. Gauger:** Writing – review & editing, Validation, Software. **Rafael Izbicki:** Writing – review & editing, Conceptualization. **Jan Kieseler:** Writing – review & editing, Validation, Investigation, Conceptualization. **Ann B. Lee:** Writing – review & editing, Supervision, Investigation, Conceptualization. **Luca Masserano:** Writing – review & editing, Supervision, Software, Conceptualization. **Federico Nardi:** Writing – review & editing, Validation, Software, Investigation, Formal analysis. **Alexander Shen:** Writing – review & editing, Validation, Software, Investigation, Data curation. **Luis Recabarren Vergara:** Writing – review & editing, Validation, Software, Data curation.

## Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

The author is an Editorial Board Member/Editor-in-Chief/Associate Editor/Guest Editor for this journal and was not involved in the editorial review or the decision to publish this article.

## Data availability

Data from publications are publicly available; code and data for studies of astrophysics use case is available on github at <https://github.com/tdorigo/SWGOLO>.



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