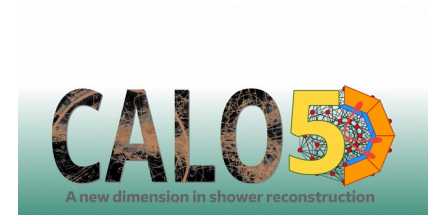


5D Calorimetry: Recent Results

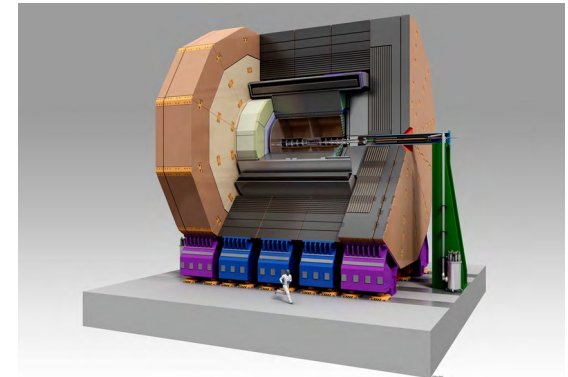
Uli Einhaus, Bohdan Dudar
LCWS Valencia
15.10.2025



- 5 dimensions of calorimeter data: x , y , z , E , t
 - Aim: study of implementation and utilisation of ps-timing capability in high-granularity calorimeter to improve performance
 - Calo5D project working on various aspects from hardware over reconstruction simulation to physics impact
-
- This talk: using neural networks / ML and timing to enhance calorimeter performance
 - Cooperation with Jan Kieseler (KIT) on ML part



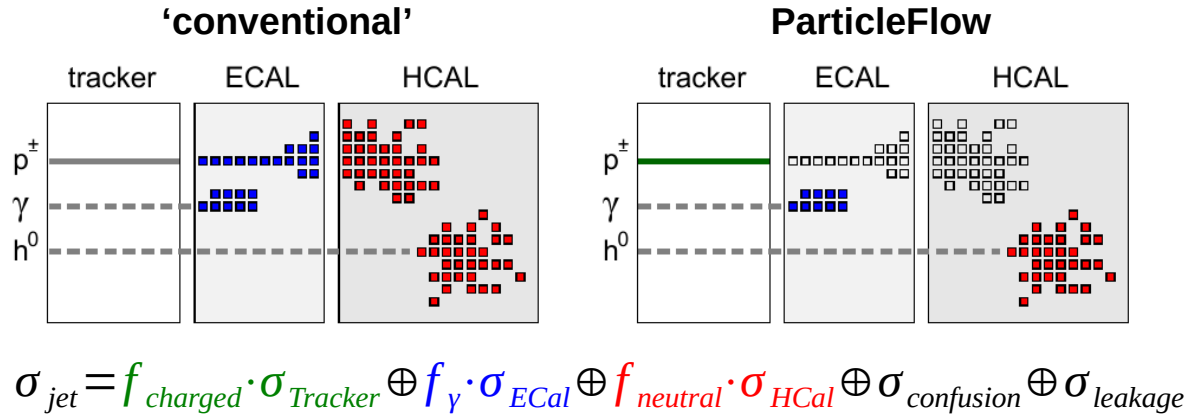
- International Large Detector, developed for ILC, now also proposed for FCC-ee
- Full-sim in DD4HEP, simulation and reconstruction chain ‘battle-proven’ in large MC productions
- Setup:
 - Si vertex + gaseous tracker, calorimeter inside 3.5 T magnet
 - ECal: $5 \times 5 \text{ mm}^2$ in 30 layers
 - HCal: $3 \times 3 \text{ cm}^2$ in 48 layers
- Designed for ParticleFlow with low-material tracker and high-granularity calorimeter



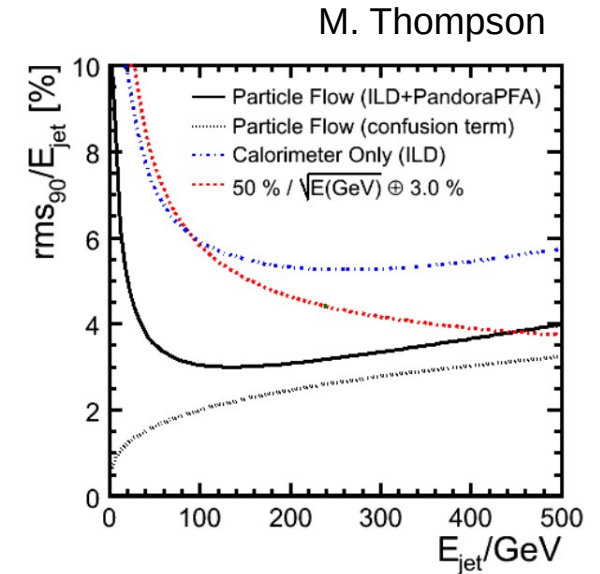
[arxiv:2003.01116](https://arxiv.org/abs/2003.01116)

ParticleFlow overview

- Idea: utilise the best subdetector for each particle / jet constituent
 - work horse: highly granular calorimeter
 - future collider detectors (e.g. ILD) developed with PFlow in mind, also applied to existing detectors (e.g. CMS HGCAL for HL-LHC)

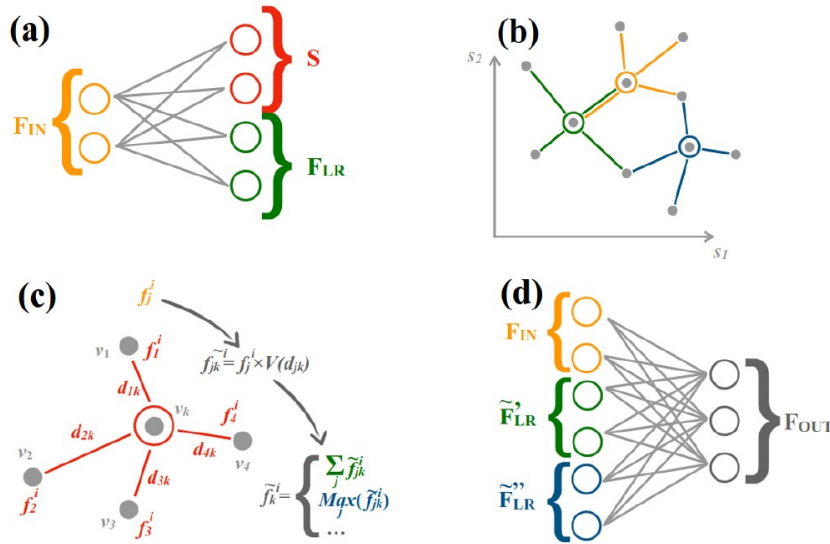


- Use neural networks and timing to reduce confusion term
- Issue with number of channels for fixed-geometry NN (e.g. CNN) → use GNN



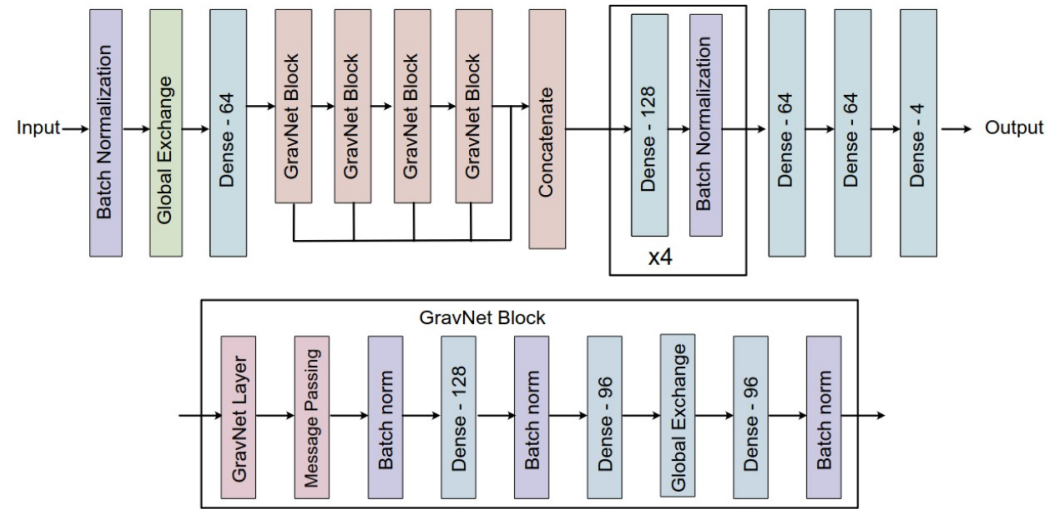
Overview GravNet + Object Condensation

- ‘Holistic’ approach with end-2-end reconstruction - clustering and features reconstruction (position, energy, PID) in one integrated Graph Neural Network
- Calorimeter hits are clustered in (low dimension) latent space S
- Feature properties F_{LR} of nearest neighbours in S are combined with different aggregator function to predict object features, weighted with $\exp(-d^2) \rightarrow$ GravNet



visualisation of the GravNet steps

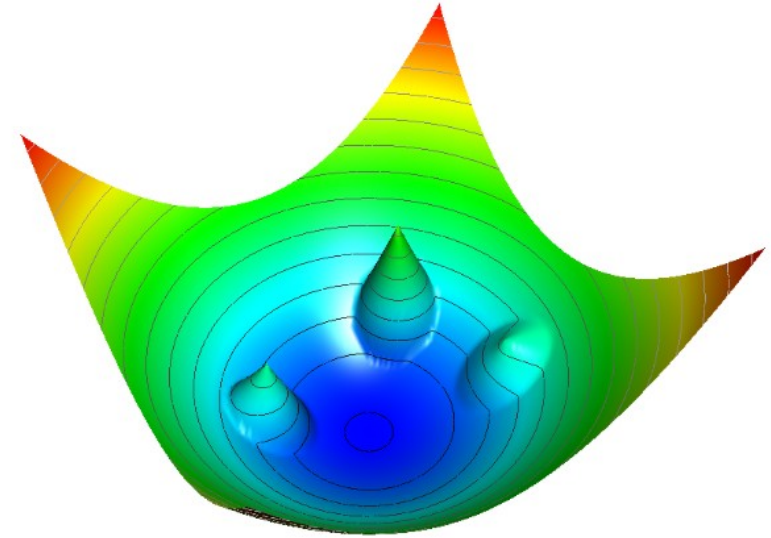
[arxiv:1902.07987](https://arxiv.org/abs/1902.07987)



GravNet model setup we used

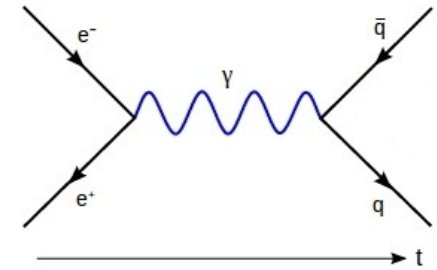
Object Condensation

- For clustering introduce attractive potential $\sim d^2$ for all hits belonging to an object towards each other, and a repulsive potential $\sim -d$ with short range towards all other hits
- To reduce computational effort, introduce β quantifying similarity to object properties, only use hit with highest β per object to generate combined potential
- Additional loss per object of $1-\beta$ of the condensation point encourages one hit per object to become representative of object and ‘gather’ related hits into a cluster



Object condensation attractive + repulsive potential visualisation
[arxiv:2002.03605](https://arxiv.org/abs/2002.03605)

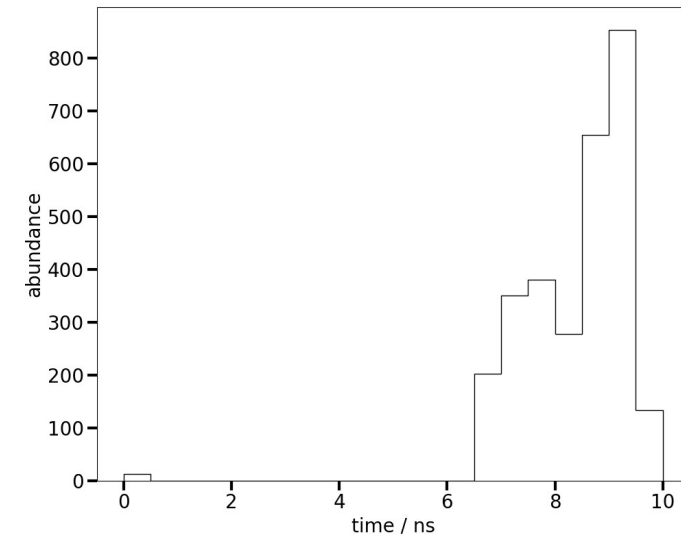
- Use ILD MC production [data set](#) of $e^+e^- \rightarrow (\gamma)q\bar{q}$ at 250 GeV (2fZhad), 45k events
- $O(\text{a few } 1000)$ hits per event
- Truth: properties of MCParticle that entered calo and caused energy depositions
 - full MCParticle chain before calo from geant4 is available; in case of multiple contributions to the same hit, largest energy fraction is chosen
 - index, end point in calo (x,y,z), (energy, PID)
 - issue: particles that backscatter from the calorimeter sometimes are not registered, and hits are associated with the original particle that entered the calorimeter
- Simplification cut: all hits with $r_{\varphi,\theta} > .5$ m from true object are ommitted



Properties used

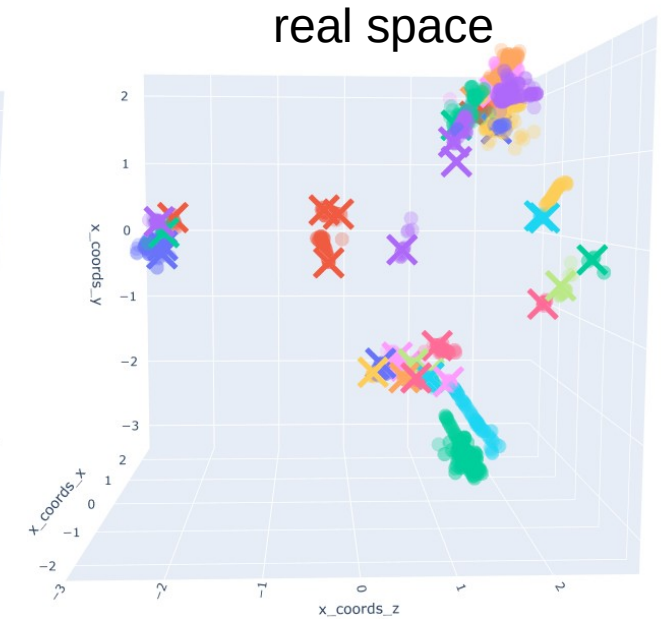
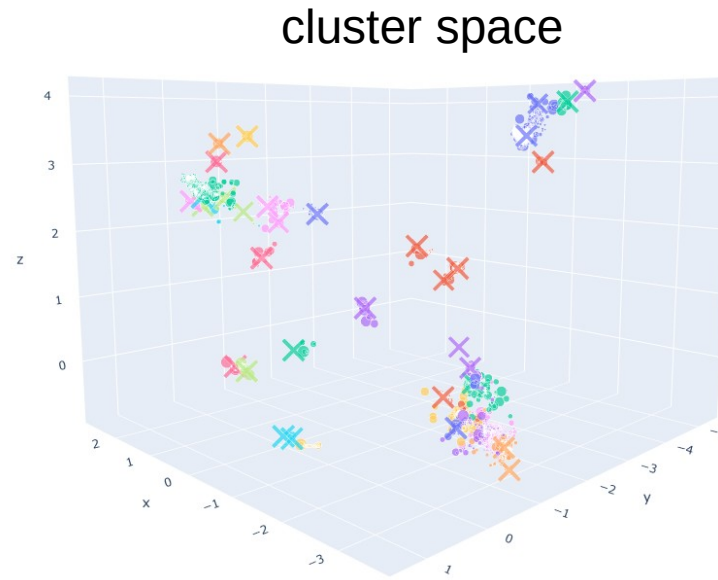
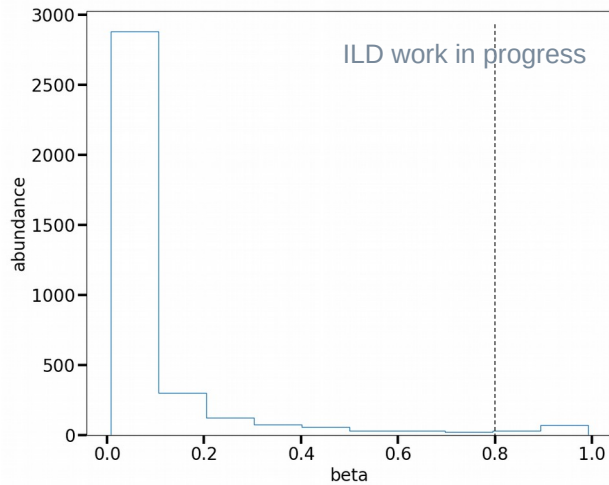
- Input: tracks, ECal and HCal hits with 9 properties each
 - $x, y, z, r_{x,y}, \theta, \eta, E$ (p for a track), time, ID (isTrack)
 - three timing 'resolutions' are compared: perfect, smeared with 100 ps, none ($t=0$)
 - time tail cut: $t [t > 20\text{ns}] := -1$
- Output: 12
 - target: β , cluster coordinates (x,y,z), position (x,y,z), energy, PID (4 channels)
- Loss:
 - object condensation $\rightarrow \beta$, cluster coordinates
 - hit loss \rightarrow position (energy and class not used atm)

- $$L_{pos} = \text{Huber}(\sqrt{\Delta x^2 + \Delta y^2 + \Delta z^2 + 10^{-6}}, 0.1)$$



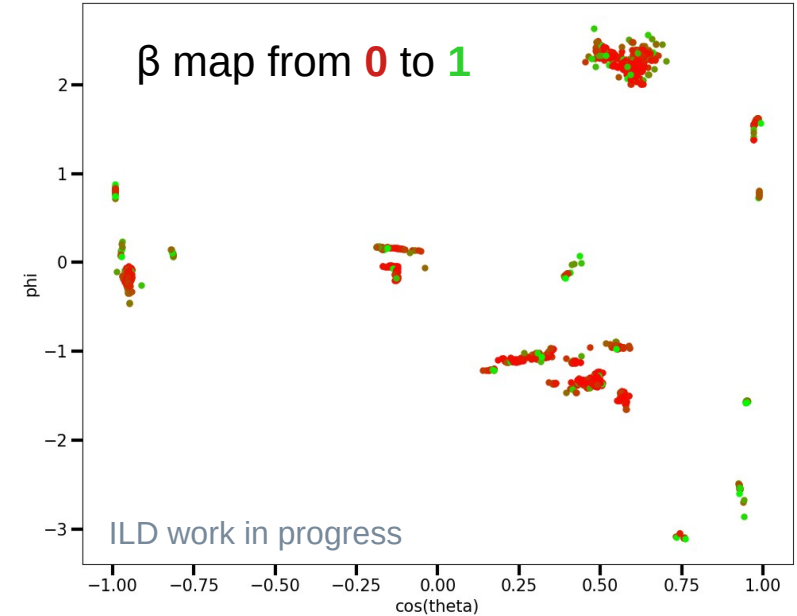
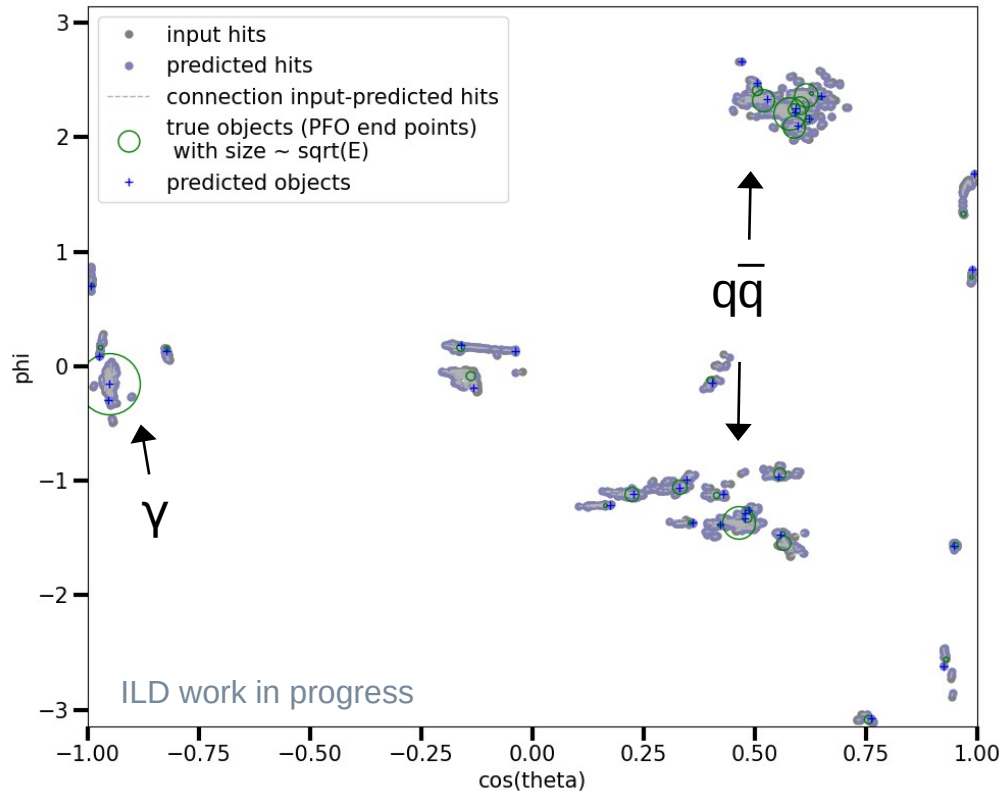
Post-inference thresholding

- For predicted objects, i.e. PFOs, thresholds of min beta and min dist to next PFO are chosen after running inference
- beta: few entries with values close to 1 represent the well reconstructed PFOs
- In space projections: matching colours belong to the same MCParticle, X mark predicted PFOs, here for $\beta > 0.8$



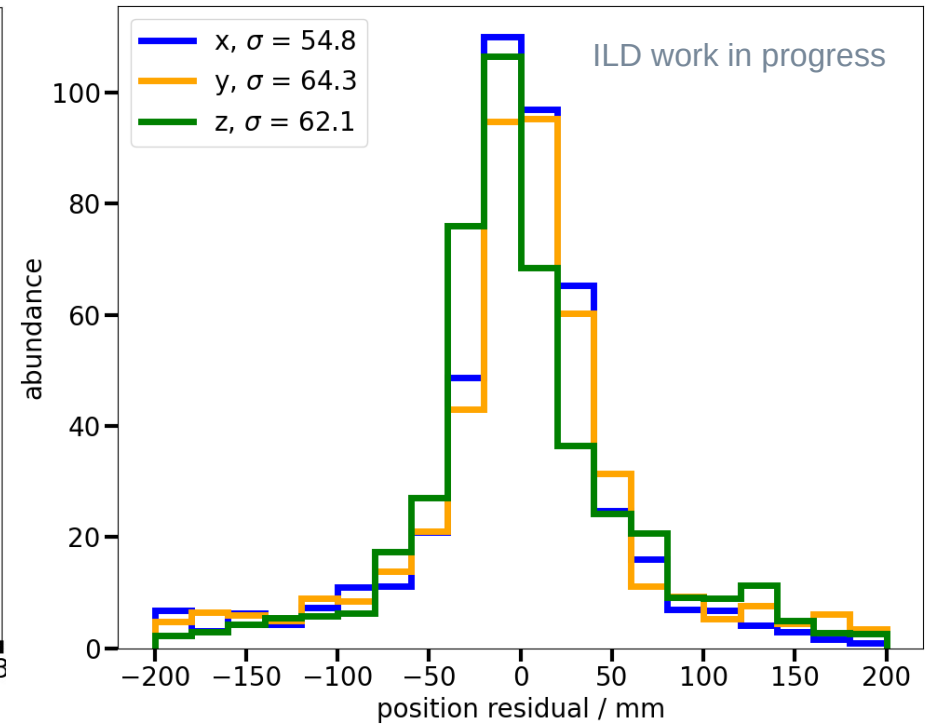
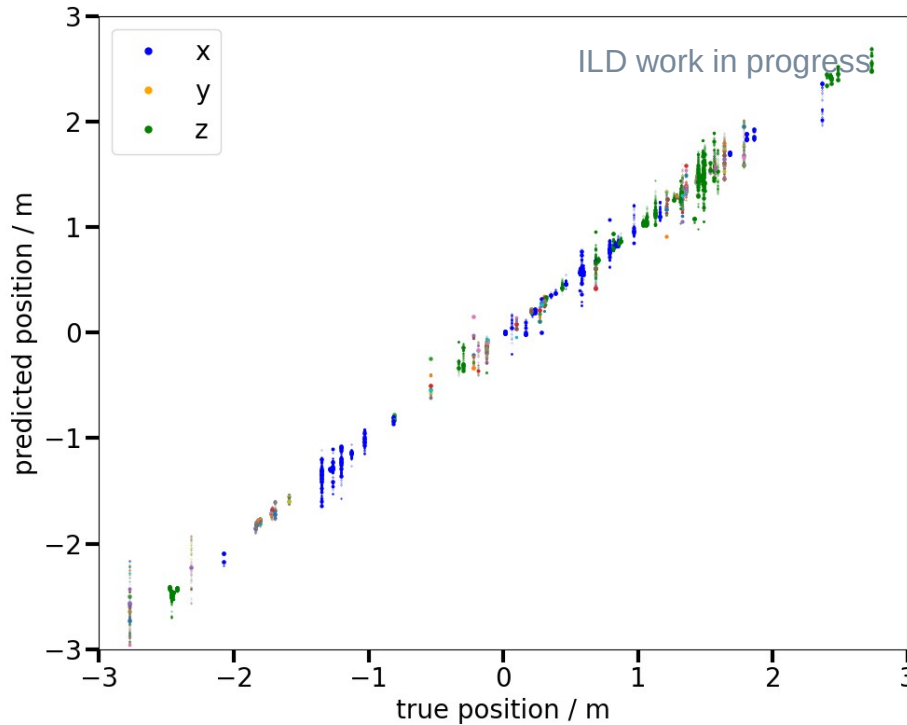
Angular projection of example event

- Predicted positions excellently clustered around targets
- In many cases one predicted PFO (+) per MCParticle (○), here: 35 and 34



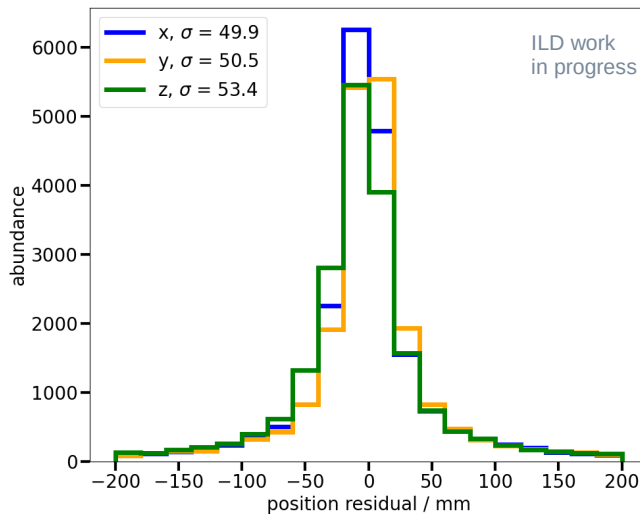
Position reconstruction

- Position per hit, weighted with beta
- Good correlation, hit position prediction uncertainty about 6 cm with current training setup

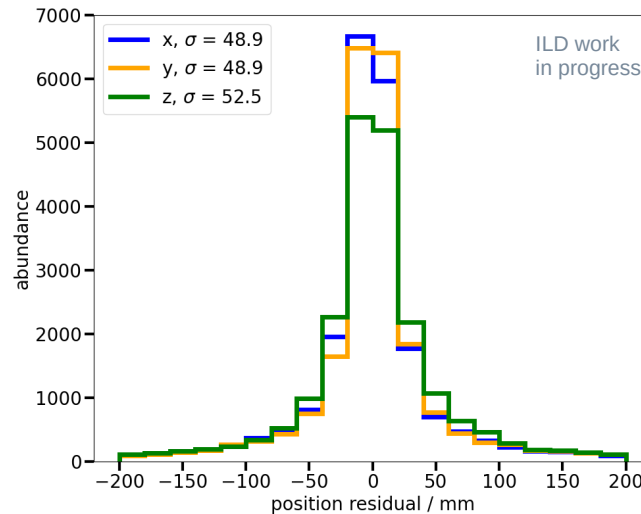


- PFO position prediction, statistics for 400 test events, ~ 50 mm position uncertainty
- Slight improvement (~ 1 mm; $\Delta_\sigma \approx 0.3$ mm) with adding time

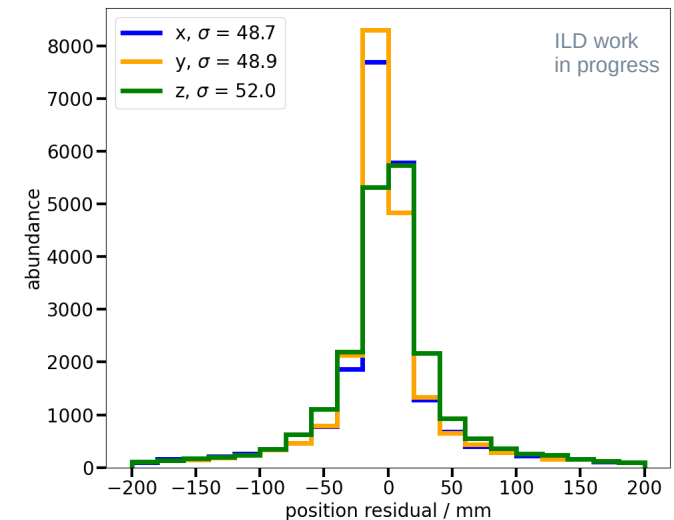
no timing



smearing with 100 ps

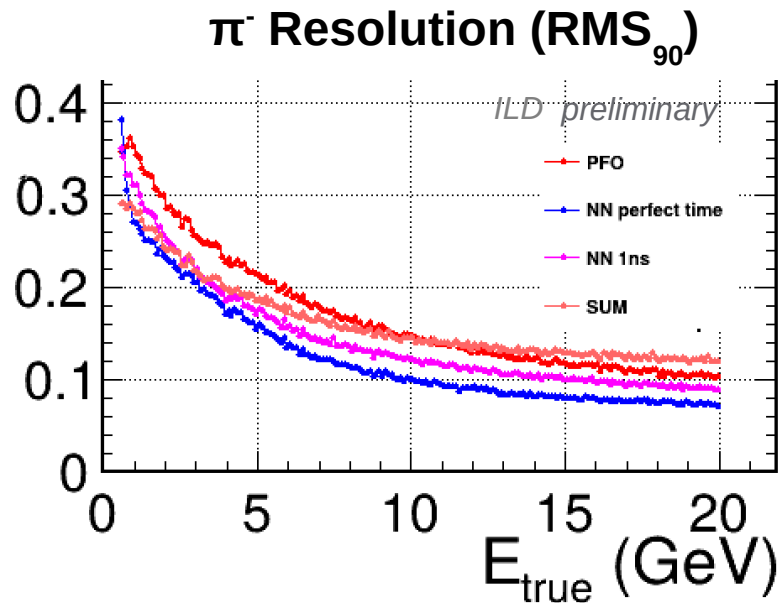


perfect timing



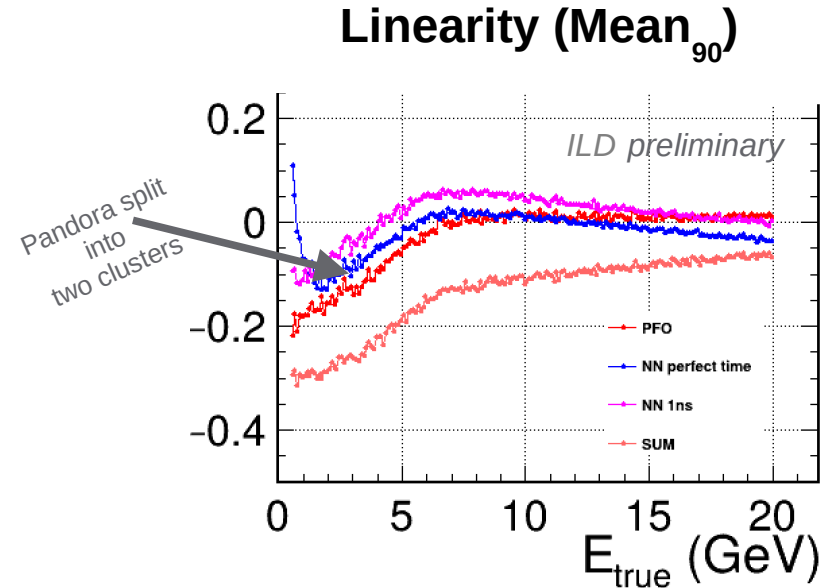
Optimising software compensation with ML

- Starting point: single particles – ML software compensation works well here!



Results

- NN outperforms Pandora on single pions
- Better timing \rightarrow better resolution!
- Could be further optimized.



Two big buts...

- Pandora is optimized for physics, not single particles!
- Migrating “single-particle” trained NN on physics events is challenging due to confusions

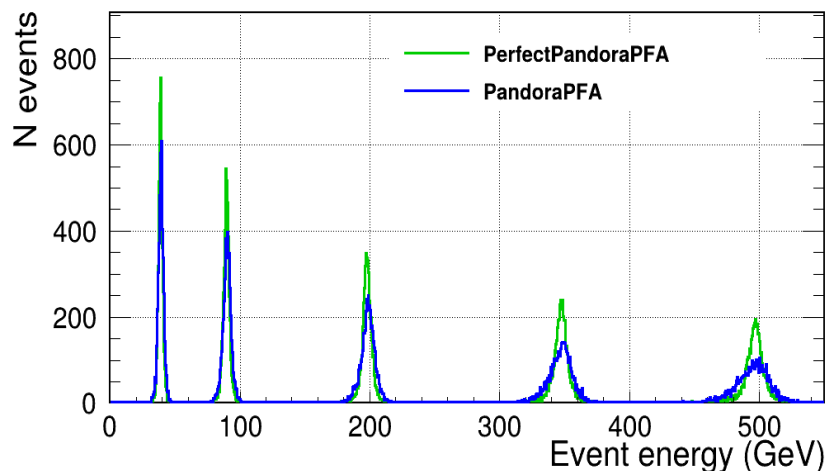
Application to full $q\bar{q}$ events

Setup:

- ❖ $Z' \rightarrow q\bar{q}$ (u, d, s)
- ❖ ILD full reconstruction
- ❖ E_{CM} : 40, 91, 200, 350, 500 GeV

Challenge:

- ❖ NN that outperforms Pandora PF
- ❖ Estimate role of timing



Current approach

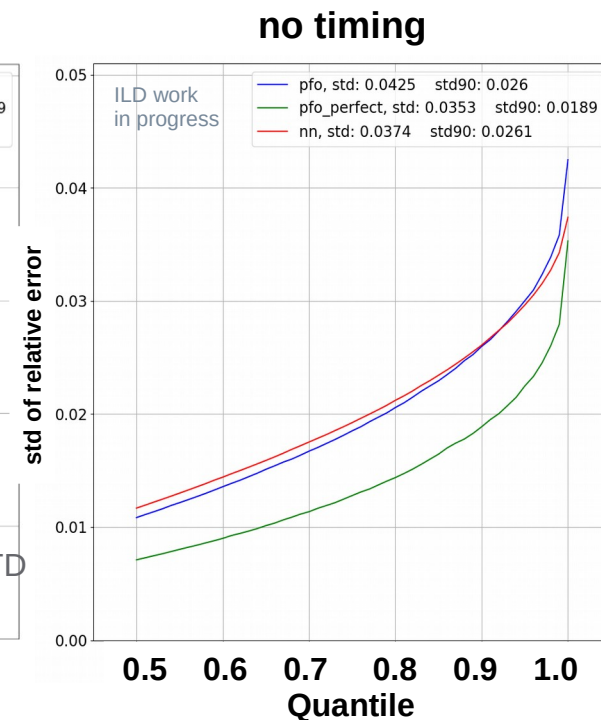
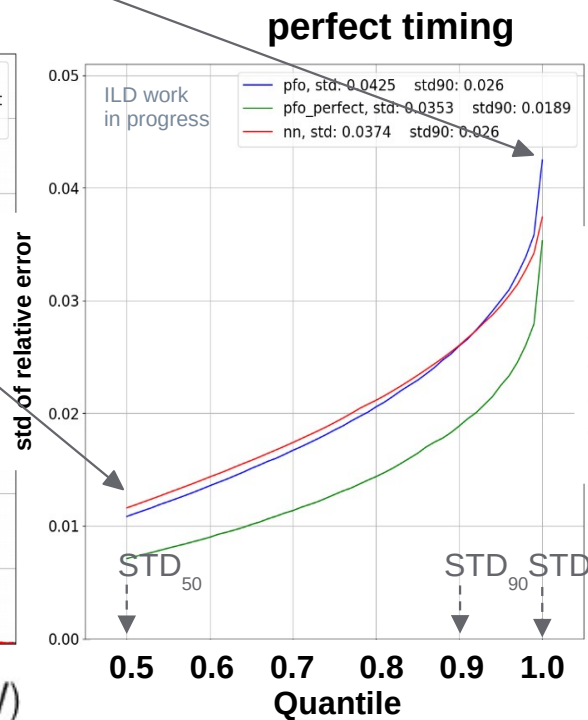
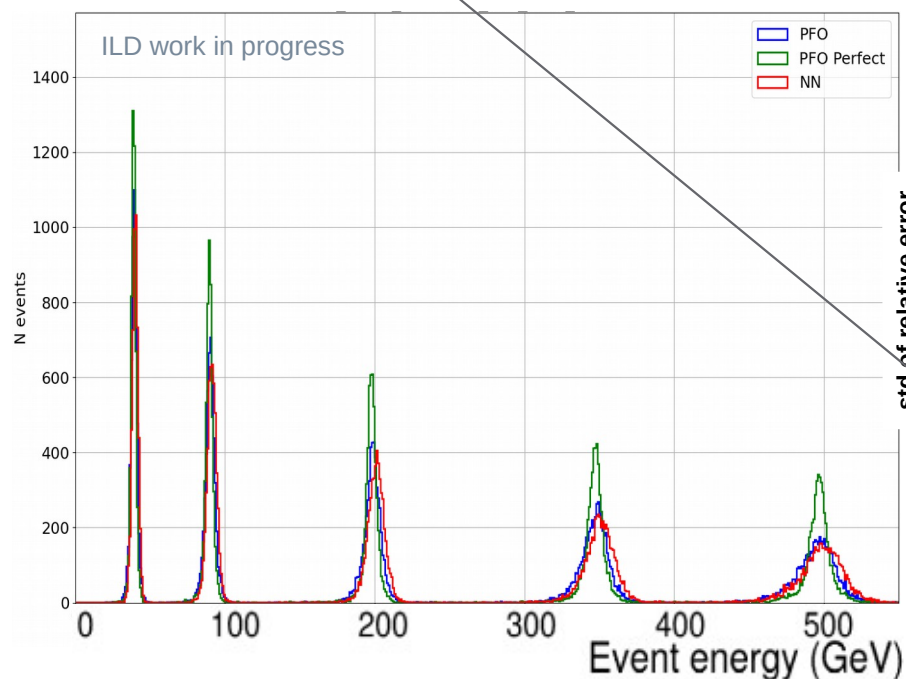
Input	Hit features (details in the back-up): positional, energy, time (perfect time res.)
Model	<u>Dynamic Graph CNN (DGCNN)</u>
Output	Corrected per-hit energies (E_{hit})
Target	Generator level $E_{\text{CM}} = \sum_i E_{\text{hit}, i}$
Loss	MSRE (Mean Squared Relative Error)
Train data	E_{CM} : 40, 91 , 200, 350 , 500 GeV (total 25k evts)
Val data	E_{CM} : 40, 91, 200, 350, 500 GeV (total 25k evts)

First results: slight improvement with NN compared to Pandora

Results

- ❖ NN is significantly better than Pandora on the full distribution (fixes bias+outliers)
- ❖ NN is similar to Pandora on the core of the distribution (still far from ideal)
- ❖ So far, timing shows no impact on performance, but it is too early to draw conclusions

Work in progress!



- **Promising results** with GravNet + Object Condensation for **ParticleFlow**: clustering and position prediction work well, optimisation ongoing
- **Additional features** energy and PID to be added next to go for full jet energy resolution
- In **software compensation** DGCNN performs similar to Pandora; with some improvement for the full distribution → working well to correct bias and outliers
- **Timing** is of small impact in both cases
- Work ongoing, lots of space to still explore!

Input features of DGCNN for SWC

Position	x, y, z,	<i>Cartesian detector coordinates</i>
hit	rho, r, phi, theta,	<i>Spherical detector coordinates</i>
features	d_long, d_perp,	<i>Distance from hit to the shower CoG</i>
	is_charged, is_neutral, is_undefined,	<i>Charged of Pandora PFO</i>
	is_ecal, is_hcal, is_yoke, is_lcal, is_lhcal, is_bcal,	<i>ILD subdetectors specific</i>
	is_barrel, is_endcap, is_ring,	
	layer,	
Energy	e,	<i>Hit's energy (note: for hits associated to tracks, ignore actual</i>
hit		<i>hit's energy and split track's energy evenly among all Pandora</i>
features		<i>cluster hits. To be improved.)</i>
	e_frac	<i>Hit energy fraction from the total energy of all hits within</i>
		<i>Pandora cluster</i>
Time	t,	<i>Absolute hit time</i>
hit	dt,	<i>Time relative to the earliest time in the cluster</i>
features	t_minus_c	<i>Absolte hit time – r/c</i>
		<i>(For now perfect time resolution assumed)</i>