

# QCD corrections to charged-current decays with Heavy Sterile Neutrinos in initial or final state and their impact on $\tau$ decays

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**ABSTRACT:** Searches for a Heavy Sterile Neutrino  $N$  profit from precise predictions of inclusive decay rates, which enter predictions for branching fractions and lifetime. Once the decay channels into semi-hadronic final states are open, a reliable calculation of inclusive decay rates is only possible if  $N$  is heavy enough to permit a perturbative calculation, in analogy to the well-known case of semi-hadronic  $\tau$  lepton decays. We adopt the popular scenario in which  $N$  only interacts with the SM particles through  $N$ - $\nu_\ell$  mixing, where  $\ell = e, \mu, \tau$ . Using literature results for  $W$  boson correlators calculated to fourth order in the strong coupling  $\alpha_s$ , we study the quality of the perturbation series for  $N \rightarrow \ell + \text{hadrons}$  to determine the mass ranges for which inclusive decay widths can be predicted in a robust way. We present novel analytic results for the decay rate of  $N \rightarrow \tau + \text{hadrons}$  in terms of  $m_\tau/m_N$ . Our expressions equally apply to  $\tau \rightarrow N + \text{hadrons}$ , the width of which is found to be perturbatively calculable for  $m_N \lesssim 600$  MeV. Applying our result to the  $\tau$  lifetime, we determine the allowed parameter space for the  $N$ - $\nu_\tau$  mixing angle  $\theta$  and  $m_N$ . We find  $|\sin \theta| \leq 0.2$  for  $m_N = 600$  MeV and weaker bounds for a lighter  $N$ . In the mass region  $m_N \geq m_\tau$  we find constraints from the dependence of  $\tau$  decay rates on  $\cos \theta$ , with  $|\sin \theta| \leq 0.12$  inferred from the  $\tau$  lifetime. Combining  $\tau \rightarrow \pi^- \nu_\tau$  and  $\tau \rightarrow K^- \nu_\tau$  data gives  $|\sin \theta| = (9.1_{-7.8}^{+3.7}) \cdot 10^{-2}$  at  $1\sigma$  while  $N$ - $\nu_\tau$  mixing does not improve the agreement between theory and data for  $\tau \rightarrow \ell \bar{\nu}_\ell \nu_\tau$ . We find current data for the decay rate  $\Gamma(\tau \rightarrow \ell + \text{nothing})$  about  $1\sigma$  above the SM prediction for  $\Gamma(\tau \rightarrow \ell \bar{\nu}_\ell \nu_\tau)$ , which leads to useful constraints on  $\Gamma(\tau \rightarrow \ell X_{\text{dark}})$  or  $\Gamma(\tau \rightarrow \ell X_{\text{dark}} X_{\text{dark}})$  with dark-sector particles  $X_{\text{dark}}$  and might stimulate additional experimental effort on  $\tau \rightarrow \ell + \text{nothing}$ .

**KEYWORDS:** Sterile or Heavy Neutrinos, Semi-Leptonic Decays, Specific BSM Phenomenology

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## 1 Introduction

Heavy Sterile Neutrinos (HSNs) (a.k.a. Heavy Neutral Leptons) are a widely-studied extension to the Standard Model (SM), postulated in theories of Dark Matter [1, 2], leptogenesis [3, 4], or neutrino masses [5–9] (see ref. [10] for an overview). In standard scenarios it is assumed that HSNs interact with SM particles only through mixing with the active neutrinos  $\nu_e$ ,  $\nu_\mu$ , and  $\nu_\tau$ , so that different observables are highly correlated. In particular,  $W$ -mediated weak decays involving an HSN  $N$  and a charged lepton  $\ell$  in the initial or final state all involve the same element  $V_{N\ell}$  of the matrix describing the mixing of  $N$  with the active neutrinos. In the most simple case one assumes that a given HSN  $N$  only mixes with one active neutrino  $\nu_\ell$ , so that  $V_{N\ell}$  reduces to  $V_{N\ell} = \sin\theta$  with a neutrino mixing angle  $\theta$ . It is therefore desirable to study as many different decays as possible to first discover HSNs and subsequently corroborate or falsify the active-sterile mixing concept. In the former case one will aim at the precise determination of  $V_{N\ell}$  for  $\ell = e, \mu, \tau$ . In the latter case one will find apparent different values for  $V_{N\ell}$  in different decays and turn the attention to less minimal

models featuring new interactions between  $N$  and SM fermions, possibly mediated by new Higgs bosons or leptoquarks [11, 12].

The search for HSNs produced at colliders and the interpretation of experimental limits requires accurate predictions of branching fractions. In addition, it is important to know the  $N$  lifetime to understand whether  $N$  decays outside or within the detector. Especially interesting is the parameter space for which  $N$  decays with a displaced decay vertex, which corresponds to an ideal, background-free signal. To predict branching fractions or lifetimes one must calculate the total decay width  $\Gamma_{\text{tot}}(N)$ . In this paper we consider semi-hadronic decays of  $N$  and present precise predictions for  $\Gamma(N \rightarrow \ell + \text{hadrons})$  for the case that the HSN mass  $m_N$  is large enough that accurate perturbative QCD predictions are possible. To this end we exploit literature results on the correlator of two charged quark currents to five-loop order of QCD [13] and add the appropriate phase-space integration to predict  $\Gamma(N \rightarrow \ell + \text{hadrons})$ . The main purpose of this analysis is the determination of the range for  $m_N$  for which the perturbative calculation is applicable, by requiring that with higher orders in  $\alpha_s$  the size of the perturbative correction and the dependence on the unphysical renormalization scale  $\mu$  decrease. Specifically, if the invariant mass of the hadronic system governed by  $m_N - m_\ell$  is too small, perturbation theory breaks down, and the same is true for certain larger values of  $m_N$  which are closely above the masses of the  $D$  or  $B$  mesons. For  $\ell = e$  or  $\mu$  we can simply use the formulae for  $\tau \rightarrow \nu_\tau + \text{hadrons}$  decays from the literature [13, 14], but for  $\Gamma(N \rightarrow \tau + \text{hadrons})$  we perform a new calculation to account for  $m_\tau \neq 0$  and find closed results in terms of  $m_\tau/m_N$  with polylogarithms (up to  $\alpha_s^3$ ) or a series representation (at order  $\alpha_s^4$ ). With this result we can further predict  $\Gamma(\tau \rightarrow N + \text{hadrons})$  for the case that  $m_N$  is too large to be neglected and analyse the impact of  $N$ - $\nu_\tau$  mixing on  $\tau$  lifetime and branching fractions. While our calculations are complete for  $\Gamma(\tau \rightarrow N + \text{hadrons})$  and the impact of these rates on the  $\tau$  lifetime, the calculated  $N \rightarrow \ell + \text{hadrons}$  decay rates do not fully account for the  $N$  lifetime, which receives further important contributions from  $Z$ -mediated  $N \rightarrow \nu_\ell + \text{hadrons}$  decay modes. These calculations are relegated to a future publication.

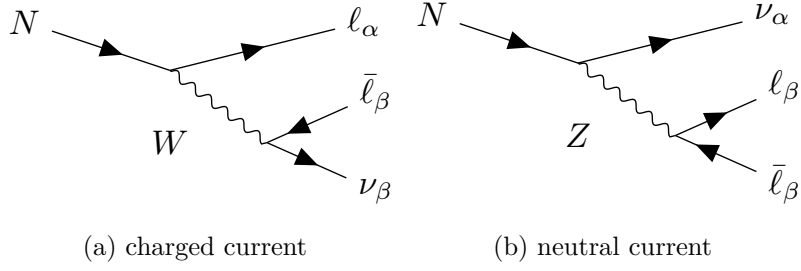
Our paper is organised as follows: in section 2 we recapitulate the known results entering the calculation of the inclusive hadronic HSN decay width, in section 3 we show the details of the actual calculation followed by a discussion of the phenomenology of our results in section 4. We conclude in section 5.

## 2 Preliminaries

### 2.1 Sterile neutrino decays at tree level

Several decays of HSN or of leptons and hadrons into HSN have been calculated at tree level [15–20]. Here we briefly collect the tree-level results relevant to our study and introduce our notation.

An HSN  $N$  can decay via a  $W$  boson into a charged lepton and a quark-antiquark pair  $(u', \bar{d}')$ . Here, we only consider cases in which  $N$  is heavy enough to permit  $N \rightarrow \ell u \bar{d}$  and  $N \rightarrow \ell u \bar{s}$  decays, but for large enough values of  $m_N$  also final states with  $u' = c$  or  $d' = b$



**Figure 1.** Tree diagrams for an HSN decay. For  $\alpha = \beta$  both diagrams interfere. Diagrams were drawn with the help of TikZ-Feynman [21].

can be accessible. The decay width reads

$$\Gamma(N \rightarrow \ell^- u' \bar{d}') = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2 |V_{u'd'}|^2}{192\pi^3} \times 12 \int_{(x_{u'}+x_{d'})^2}^{(1-x_\ell)^2} \frac{dx}{x} (1+x_\ell^2-x)(x-x_{u'}^2-x_{d'}^2) \sqrt{\lambda(1,x,x_\ell^2)\lambda(x,x_{u'}^2,x_{d'}^2)}, \quad (2.1)$$

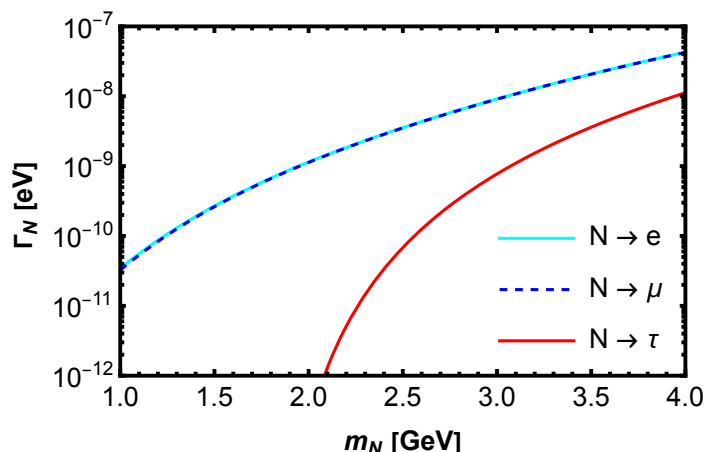
here  $G_F$  is the Fermi constant,  $m_N$  is the mass of the sterile neutrino,  $V_{N\ell}$  parametrises the mixing between  $N$  and  $\nu_\ell$ ,  $V_{u'd'}$  is the relevant CKM matrix element,  $N_c = 3$  is the number of colours,  $x_i = m_i/m_N$ , and  $\lambda(a,b,c) = (a-b-c)^2 - 4bc$  is the Källén function.

Decay rates of an HSN into leptons  $N \rightarrow \ell_i \bar{\ell}_j \nu_j$  are found from eq. (2.1) by setting  $N_c = 1$  and  $V_{u'd'} = 1$ :

$$\Gamma(N \rightarrow \ell_\alpha \bar{\ell}_\beta \nu_\beta) = \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \times 12 \int_{x_\beta^2}^{(1-x_\alpha)^2} \frac{dx}{x} (1+x_\alpha^2-x)(x-x_\beta^2) \sqrt{\lambda(1,x,x_\alpha^2)\lambda(x,x_\beta^2,0)}, \quad (2.2)$$

where  $x_\alpha = m_{\ell_\alpha}/m_N$ . This relation only holds for  $\alpha \neq \beta$ . We dub the lepton directly attached to the HSN the tagging lepton (see figure 1). For  $\alpha = \beta$  there are additional contributions to the leptonic width due to interference with the neutral current decay. The decay width becomes

$$\begin{aligned} \Gamma(N \rightarrow \nu_\alpha \ell_\alpha \bar{\ell}_\alpha) & \quad (2.3) \\ &= \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \times \left[ \sqrt{1-4x_\ell^2} \left( \left( \frac{g_V^2}{2} + g_V \right) (1-10x_\ell^2 + 18x_\ell^4 - 36x_\ell^6) \right. \right. \\ &+ \left. \left( \frac{g_A^2}{2} + g_A \right) (1-18x_\ell^2 - 22x_\ell^4 + 12x_\ell^6) + 1 - 14x_\ell^2 - 2x_\ell^4 - 12x_\ell^6 \right) \\ &- 24 \left( \left( g_V^2 + 2g_V \right) x_\ell^6 (2-3x_\ell^2) \right. \\ &+ \left. \left( g_A^2 + 2g_A \right) x_\ell^4 (2-2x_\ell^2 + x_\ell^4) + x_\ell^4 (2-2x_\ell^4) \right) \ln \left( \frac{2x_\ell}{1 + \sqrt{1-4x_\ell^2}} \right) \Big], \end{aligned}$$



**Figure 2.** Decay width of an HSN into leptons. Here the notation  $N \rightarrow \ell$  means the plotted line is the sum of all partial widths associated with the tagging lepton i.e.  $\Gamma(N \rightarrow e) = \sum_{\ell} \Gamma(N \rightarrow e\bar{\ell}\nu_{\ell})$ . For all three channels the mixing angle is chosen as  $V_{N\ell} = 10^{-3}$ .

here  $g_A = -1/2$ ,  $g_V = -1/2 + 2s_w^2$  with  $s_w = \sin \theta_w$  and the Weinberg angle  $\theta_w$ . Our results agree with those of Bodarenko et al. in ref. [10]. In figure 2 we show the decay widths as a function of  $m_N$ . The widths for  $e$  or  $\mu$  as tagging leptons are almost indistinguishable.

## 2.2 QCD correlators

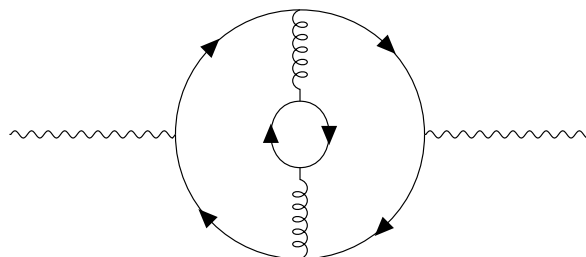
The QCD dynamics of decays of HSNs into hadrons is governed by the decays of virtual electroweak gauge bosons into hadrons. The decays of  $W$  and  $Z$  bosons into hadrons have been calculated up to  $\mathcal{O}(\alpha_s^4)$  [13, 22]. Here we use these results to calculate the charged-current contribution to the decay of an HSN.

The optical theorem relates any inclusive decay width to the imaginary part of the self-energy of the decaying particle. The non-trivial part of the  $W$  or  $Z$  boson self-energy is the correlator of two charged quark currents, namely the time ordered product of the vector and axial vector currents of the quarks  $j_{\mu,ij}^{V/A} = \bar{q}_i \gamma_{\mu} (\gamma_5) q_j$ . It may be parametrized as follows [23, 24]

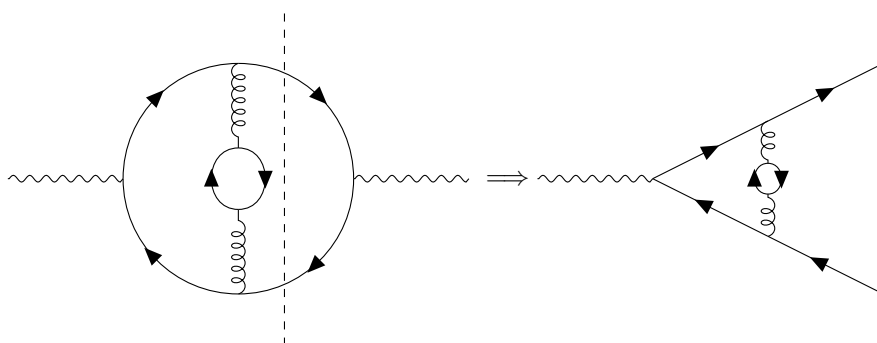
$$\begin{aligned} \Pi_{\mu\nu,ij}^{V/A}(q, m_i, m_j, \mu, \alpha_s) &= i \int dx e^{iqx} \langle 0 | \hat{T} \{ j_{\mu,ij}^{V/A}(x) j_{\nu,ij}^{V/A\dagger}(0) \} | 0 \rangle \\ &= g_{\mu\nu} \Pi_{ij,V/A}^{[1]}(q^2) + q_{\mu} q_{\nu} \Pi_{ij,V/A}^{[2]}(q^2) \\ &= (-g_{\mu\nu} q^2 + q_{\mu} q_{\nu}) \Pi_{ij,V/A}^{(1)}(q^2) + q_{\mu} q_{\nu} \Pi_{ij,V/A}^{(0)}(q^2). \end{aligned} \quad (2.4)$$

Here  $m_{i,j}$  are quark masses and  $q$  is the momentum of the gauge boson. In our case of the hadronic charged current  $q_i$  and  $q_j$  are up-type and down-type quarks, respectively. In the last line above the correlators have been decomposed into their transverse and longitudinal components

$$\Pi_{ij,V/A}^{(1)}(q^2) = -\frac{\Pi_{ij,V/A}^{[1]}(q^2)}{q^2}, \quad \Pi_{ij,V/A}^{(0)}(q^2) = \Pi_{ij,V/A}^{[2]}(q^2) + \frac{\Pi_{ij,V/A}^{[1]}(q^2)}{q^2}. \quad (2.5)$$



(a) Sample non-singlet diagram to  $\mathcal{O}(\alpha_s^2)$ .



(b) One possible cut of the loop diagram in figure 3(a). This corresponds to the decay amplitude  $W \rightarrow q\bar{q}$  at  $\mathcal{O}(\alpha_s^2)$  interfering with the tree-level amplitude.

**Figure 3.**  $W$  boson self-energy diagram contributing to the correlator.

In studies of  $\tau$  decays one commonly adopts the definition [23]

$$\Pi_{ij, V/A}^{(1+0)}(q^2) \equiv \Pi_{ij, V/A}^{(1)}(q^2) + \Pi_{ij, V/A}^{(0)}(q^2). \quad (2.6)$$

The correlators  $\Pi^{[1]}$  and  $\Pi^{[2]}$  are related to each other via a Ward identity [25, 26]

$$q^\mu q^\nu \Pi_{\mu\nu, ij}^{V/A} = q^4 \Pi_{ij, V/A}^{(0)} = (m_j \mp m_i)^2 \Pi_{ij}^{S/P} + (m_j \mp m_i) \langle 0 | \bar{q}_j q_j \mp \bar{q}_i q_i | 0 \rangle. \quad (2.7)$$

This identity connects the longitudinal part of the  $V$  or  $A$  correlator to the scalar or pseudo-scalar correlators  $\Pi_{ij}^{S/P}$ , respectively, where

$$\Pi_{ij}^{S/P}(q) = i \int dx e^{iqx} \langle 0 | \hat{T} \{ j_{ij}^{S/P}(x) j_{ij}^{S/P\dagger}(0) \} | 0 \rangle, \quad j_{ij}^{S/P} = \bar{q}_i (\gamma_5) q_j, \quad (2.8)$$

and the last term in eq. (2.7) is the quark condensate. Thus in the chiral limit

$$\Pi_{ij, V/A}^{(0)} = 0. \quad (2.9)$$

For the  $W$  boson there are only contributions to the correlator by so-called non-singlet diagrams as seen in figure 3(a). Non-singlet diagrams only have cuts through at least two fermion lines. Cuts indicate individual contributions to the imaginary part of a loop diagram with the particles on the cut line being on-shell (see figure 3(b)). From now on we adopt the chiral limit with zero quark masses. The applicability of perturbative QCD requires  $q^2 \gg m_i^2$ , which is certainly fulfilled for the three lightest quarks for  $m_N \gtrsim$  a few times  $\Lambda_{\text{QCD}}$ , which must be fulfilled in any perturbative QCD calculation. If one of the final state quarks is

$c$  or  $b$ , there is a range of  $q^2$  for which the  $c$  or  $b$  quark mass is non-negligible and can be included in an expansion in terms of  $m_{c,b}^2/q^2$ . These mass corrections are beyond the scope of this paper, i.e. for heavy hadrons our formulae are only valid for HSN decays with  $q^2$  large enough that  $m_c$  or  $m_b$  can be set to zero i.e.  $m_b^2 > q^2 \gg m_c^2$  or  $q^2 \gg m_b^2$ .

The correlators may be expressed as a series in

$$a_\mu = \frac{\alpha_s(\mu)}{\pi}, \quad (2.10)$$

where the coefficients are functions of the logarithm  $\ln(-s/\mu^2)$  with  $s = q^2$  the invariant mass of the hadronic state and the logarithm evaluated below the branch cut according to the Feynman prescription  $s \rightarrow s + i\delta$ , so that  $\text{Im} \ln(-s/\mu^2 - i\delta) = \ln(s/\mu^2) - i\pi$ . Since the longitudinal part of the correlator is proportional to the square of the quark mass it will not contribute in the chiral limit (see eq. (2.7)). The sum of the longitudinal and transverse part of the correlator is given by [14]

$$\Pi_{ij, V/A}^{(1+0)}(s) = -\frac{1}{12\pi^2} \sum_{n=0}^{\infty} a_\mu^n \sum_{k=1}^{n+1} c_{n,k} \ln^k \left( \frac{-s}{\mu^2} \right), \quad (2.11)$$

and the coefficients  $c_{n,1}$  are given by [13, 27–32]

$$c_{0,1} = c_{1,1} = 1, \quad (2.12)$$

$$c_{2,1} = -\left( \frac{11}{12} - \frac{2}{3}\zeta_3 \right) n_f + \frac{365}{24} - 11\zeta_3, \quad (2.13)$$

$$c_{3,1} = \left( \frac{151}{162} - \frac{19}{27}\zeta_3 \right) n_f^2 - \left( \frac{7847}{216} - \frac{262}{9}\zeta_3 + \frac{25}{9}\zeta_5 \right) n_f + \frac{87029}{288} - \frac{1103}{4}\zeta_3 + \frac{275}{6}\zeta_5, \quad (2.14)$$

$$\begin{aligned} c_{4,1} = & \left( \frac{203}{324}\zeta_3 + \frac{5}{18}\zeta_5 - \frac{6131}{5832} \right) n_f^3 + \left( -\frac{40655}{864}\zeta_3 + \frac{5}{6}\zeta_3^2 - \frac{260}{27}\zeta_5 + \frac{1045381}{15552} \right) n_f^2 \\ & + \left( \frac{12205}{12}\zeta_3 - 55\zeta(3)^2 + \frac{29675}{432}\zeta_5 + \frac{665}{72}\zeta_7 - \frac{13044007}{10368} \right) n_f \\ & - \frac{7315}{48}\zeta_7 + \frac{65945}{288}\zeta_5 + \frac{5445}{8}\zeta_3^2 - \frac{5693495}{864}\zeta_3 + \frac{144939499}{20736}. \end{aligned} \quad (2.15)$$

Here  $n_f$  denotes the number of active flavours and  $\zeta_n \equiv \zeta(n)$  are values of the Riemann  $\zeta$ -function. To reconstruct the remaining coefficients  $c_{n,k}$  we can use the renormalization group (RG) equation. To this end, one considers the Adler function [33]

$$D^{(1+0)}(s) = -s \frac{d}{ds} \Pi_{ij, V/A}^{(1+0)}(s), \quad (2.16)$$

which is a physical object and  $\mu$ -independent. Employing

$$\mu^2 \frac{d}{d\mu^2} D_{ij, V/A}^{(1+0)}(s) = 0, \quad (2.17)$$

one finds [14]

$$c_{2,2} = -\frac{\beta_0}{2} c_{1,1} \tag{2.18}$$

$$c_{3,3} = \frac{\beta_0^2}{3} c_{1,1}, \quad c_{3,2} = -\frac{\beta_1}{2} c_{1,1} - \beta_0 c_{2,1} \tag{2.19}$$

$$c_{4,4} = -\frac{\beta_0^3}{4} c_{1,1}, \quad c_{4,3} = \frac{5}{6} \beta_1 \beta_0 c_{1,1} + \beta_0^2 c_{2,1}, \tag{2.20}$$

$$c_{4,2} = -\frac{1}{2} (\beta_2 c_{1,1} + 2\beta_1 c_{2,1} + 3\beta_0 c_{3,1}),$$

where  $c_{n,k} = 0$  if  $k > n$  and  $n \neq 0$  and the coefficients  $\beta_i$  of the QCD beta function are listed in appendix A.

### 3 QCD corrections to charged current HSN decays

Using the correlator in eq. (2.11) the differential inclusive decay rate of a HSN via a charged current reads

$$\frac{d\Gamma(N \rightarrow \ell X)}{ds} = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \times \frac{12\pi}{m_N^2} \left(1 + x_\ell^2 - \frac{s}{m_N^2}\right) \sqrt{\lambda\left(1, \frac{s}{m_N^2}, x_\ell^2\right)} \tag{3.1}$$

$$\times \left[ \left(1 + 2\frac{s}{m_N^2} + x_\ell^2 - \frac{4x_\ell^2}{1 + x_\ell^2 - \frac{s}{m_N^2}}\right) \text{Im}\Pi^{(1+0)}(s) - 2\frac{s}{m_N^2} \text{Im}\Pi^{(0)}(s) \right],$$

integrating over  $s$ , or the dimensionless variable  $x = s/m_N^2$ , leads to [23]

$$\Gamma(N \rightarrow \ell X) = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \times 12\pi \int_0^{(1-x_\ell)^2} dx (1 + x_\ell^2 - x) \sqrt{\lambda(1, x, x_\ell^2)} \tag{3.2}$$

$$\times \left[ \left(1 + 2x + x_\ell^2 - \frac{4x_\ell^2}{1 + x_\ell^2 - x}\right) \text{Im}\Pi^{(1+0)}(m_N^2 x) - 2x \text{Im}\Pi^{(0)}(m_N^2 x) \right],$$

with  $x_\ell = m_\ell/m_N$  and where we used

$$\Pi^{(J)} = |V_{ud}|^2 \left( \Pi_{ud,V}^{(J)} + \Pi_{ud,A}^{(J)} \right) + |V_{us}|^2 \left( \Pi_{us,V}^{(J)} + \Pi_{us,A}^{(J)} \right) + \dots \tag{3.3}$$

$$= 2(|V_{ud}|^2 + |V_{us}|^2 + \dots) \Pi_V^{(J)},$$

since the vector and axial vector contribution are equal in magnitude. Here the dots indicate to the inclusion of heavier mesons i.e.  $D$  or  $B$  mesons. Since  $\Pi^{(0)} \sim m_q^2$  we neglect it. The integration region with small  $x$  involves hadronic resonances while the integrand in eq. (3.2) is smooth. Eq. (3.2) nevertheless reproduces the correct result as can be derived from analyticity properties of  $\Pi^{(J)}$ , see appendix B for details.

Choosing  $\mu = m_N$  the imaginary part of the correlator reads

$$\begin{aligned}
 \text{Im } \Pi_V^{(1+0)}(m_N^2 x) &= -\frac{1}{12\pi^2} \sum_{n=0}^{\infty} a_{m_N}^n \sum_{k=1}^{n+1} c_{n,k} \text{Im } \ln^k(-x) \\
 &= \frac{1}{12\pi} \left\{ 1 + a_{m_N} + a_{m_N}^2 (c_{2,1} + 2c_{2,2} \ln(x)) \right. \\
 &\quad + a_{m_N}^3 (c_{3,1} + 2c_{3,2} \ln(x) - (\pi^2 - 3 \ln^2(x))c_{3,3}) \\
 &\quad + a_{m_N}^4 (c_{4,1} + 2c_{4,2} \ln(x) - (\pi^2 - 3 \ln^2(x))c_{4,3} \\
 &\quad \left. - (4\pi^2 \ln(x) - 4 \ln^3(x))c_{4,4} \right\}
 \end{aligned} \tag{3.4}$$

and leads to integrals of the form

$$I_k = \int_0^{(1-x_\ell)^2} dx \left( (1+x_\ell^2-x)(1+2x+x_\ell^2) - 4x_\ell^2 \right) \sqrt{\lambda(1, x, x_\ell^2)} \ln^k(x). \tag{3.5}$$

Up to  $k = 2$  the integrals are analytically calculable in terms of dilogarithm and trilogarithm functions:

$$I_0 = \frac{1}{2} \left( 1 - 8x_\ell^2 + 8x_\ell^6 - x_\ell^8 - 12x_\ell^4 \ln(x_\ell^2) \right), \tag{3.6}$$

$$\begin{aligned}
 I_1 &= -\frac{19}{24} + \frac{13}{3}x_\ell^2 - 12\zeta_2 x_\ell^4 - \frac{13}{3}x_\ell^6 + \frac{19}{24}x_\ell^8 \\
 &\quad + \left( -3x_\ell^4 - 4x_\ell^6 + \frac{1}{2}x_\ell^8 \right) \ln(x_\ell^2) \\
 &\quad + \left( 1 - 8x_\ell^2 + 8x_\ell^6 - x_\ell^8 \right) \ln(1-x_\ell^2) + 12x_\ell^4 \text{Li}_2(x_\ell^2),
 \end{aligned} \tag{3.7}$$

$$\begin{aligned}
 I_2 &= \frac{265}{144} - \frac{151}{18}x_\ell^2 - 12(\zeta_2 + 2\zeta_3)x_\ell^4 + \left( \frac{151}{18} - 16\zeta_2 \right) x_\ell^6 + \left( -\frac{265}{144} + 2\zeta_2 \right) x_\ell^8 \\
 &\quad + \left( -\frac{19}{6} + \frac{52}{3}x_\ell^2 - 48\zeta_2 x_\ell^4 - \frac{52}{3}x_\ell^6 + \frac{19}{6}x_\ell^8 + 48x_\ell^4 \text{Li}_2(x_\ell^2) \right) \ln(1-x_\ell^2) \\
 &\quad + (2 - 16x_\ell^2 + 16x_\ell^6 - 2x_\ell^8) \ln^2(1-x_\ell^2) \\
 &\quad + \left( -x_\ell^2 + \frac{15}{2}x_\ell^4 + \frac{23}{3}x_\ell^6 - \frac{19}{12}x_\ell^8 \right. \\
 &\quad \left. + (-1 + 8x_\ell^2 - 8x_\ell^6 + x_\ell^8) \ln(1-x_\ell^2) + 24x_\ell^4 \ln^2(1-x_\ell^2) - 12x_\ell^4 \text{Li}_2(x_\ell^2) \right) \ln(x_\ell^2) \\
 &\quad + (-1 + 8x_\ell^2 + 12x_\ell^4 + 8x_\ell^6 - x_\ell^8) \text{Li}_2(x_\ell^2) \\
 &\quad + 24x_\ell^4 \text{Li}_3(x_\ell^2) + 48x_\ell^4 \text{Li}_3(1-x_\ell^2).
 \end{aligned} \tag{3.8}$$

In appendix C we give approximations to these integrals for easier use.

We next derive a new analytical result holding as well for  $k \geq 3$ , defined in terms of a well converging sum. To do this we first use the following series representation of the

square root of the Källén function,

$$\frac{d^n}{dx^n} \sqrt{\lambda(1, x, x_\ell^2)} = \sum_{j=0}^{\lfloor \frac{n}{2} \rfloor} (-1)^{n-j} \left(-\frac{1}{2}\right)_{n-j} (2j-1)!! \binom{n}{2j} \quad (3.9)$$

$$\times \frac{\left(\lambda'(1, x, x_\ell^2)\right)^{n-2j} \left(\lambda''(1, x, x_\ell^2)\right)^j}{\left(\lambda(1, x, x_\ell^2)\right)^{\frac{2(n-j)-1}{2}}},$$

where  $(a)_n = a(a+1)(a+2) \cdots (a+n-1)$  is the Pochhammer symbol,  $\lfloor j \rfloor$  is the floor function (rounding a real number to the nearest smaller integer number) and

$$\lambda(1, x, x_\ell^2) = x^2 - 2x(1 + x_\ell^2) + (1 - x_\ell^2)^2, \quad (3.10)$$

$$\lambda'(1, x, x_\ell^2) = \frac{d\lambda(1, x, x_\ell^2)}{dx} = 2x - 2(1 + x_\ell^2), \quad (3.11)$$

$$\lambda''(1, x, x_\ell^2) = \frac{d^2\lambda(1, x, x_\ell^2)}{dx^2} = 2. \quad (3.12)$$

We write

$$I_k(x_\ell^2, a) = \sum_{n=0}^{\infty} \frac{1}{n!} A_n(x_\ell^2) B_{n,k}(x_\ell^2, a) \quad (3.13)$$

with

$$A_n(x_\ell^2) \equiv \left. \frac{d^n}{dx^n} \sqrt{\lambda(1, x, x_\ell^2)} \right|_{x=0}$$

$$= \sum_{j=0}^{\lfloor \frac{n}{2} \rfloor} (-1)^j \left(-\frac{1}{2}\right)_{n-j} (2j-1)!! \binom{n}{2j} 2^{n-j} (1 + x_\ell^2)^{n-2j} (1 - x_\ell^2)^{1-2(n-j)}, \quad (3.14)$$

and

$$B_{n,k}(x_\ell^2, a) \equiv \int_0^a dx \left( (1 + x_\ell^2 - x)(1 + 2x + x_\ell^2) - 4x_\ell^2 \right) x^n \ln^k(x)$$

$$= -2I_{n+2,k}(a) + (1 + x_\ell^2)I_{n+1,k}(a) + (1 - x_\ell^2)^2 I_{n,k}(a). \quad (3.15)$$

The integrals in this expression are given by

$$I_{n,k}(a) \equiv \int_0^a dx x^n \ln^k x$$

$$= a^{n+1} \sum_{l=0}^k (-k)_l (n+1)^{-l-1} \ln^{k-l} a, \quad (3.16)$$

which completes the calculation of  $I_k$  in eq. (3.5).

Inserting eq. (3.4) into eq. (3.2) and using eq. (3.5) finally gives the full semi-hadronic width for decays of sterile neutrinos lighter than the  $D^+$  meson

$$\begin{aligned} \Gamma(N \rightarrow \ell X) = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \cdot 2(|V_{ud}|^2 + |V_{us}|^2) \quad (3.17) \\ \times \left[ I_0(x_\ell^2, (1-x_\ell)^2) c_{0,1} + a_{m_N} c_{1,1} I_0(x_\ell^2, (1-x_\ell)^2) \right. \\ + a_{m_N}^2 [c_{2,1} I_0(x_\ell^2, (1-x_\ell)^2) + 2c_{2,2} I_1(x_\ell^2, (1-x_\ell)^2)] \\ + a_{m_N}^3 [c_{3,1} I_0(x_\ell^2, (1-x_\ell)^2) + 2c_{3,2} I_1(x_\ell^2, (1-x_\ell)^2) \\ - (\pi^2 I_0(x_\ell^2, (1-x_\ell)^2) - 3I_2(x_\ell^2, (1-x_\ell)^2)) c_{3,3}] \\ + a_{m_N}^4 [c_{4,1} I_0(x_\ell^2, (1-x_\ell)^2) + 2c_{4,2} I_1(x_\ell^2, (1-x_\ell)^2) \\ - (\pi^2 I_0(x_\ell^2, (1-x_\ell)^2) - 3I_2(x_\ell^2, (1-x_\ell)^2)) c_{4,3} \\ \left. - (4\pi^2 I_1(x_\ell^2, (1-x_\ell)^2) - 4I_3(x_\ell^2, (1-x_\ell)^2)) c_{4,4} \right]. \end{aligned}$$

Logarithms involving the renormalization scale may be reconstructed by using the relation in appendix A. Sterile neutrinos heavier than the  $D^+$  meson also obey the relation above once the corresponding CKM element as well as threshold effects are included.

Additionally we also calculated the integral in eq. (3.5) for an arbitrary upper limit of the hadronic invariant mass, to permit the calculation of the integrated spectrum of the decay

$$\begin{aligned} \Gamma_{\text{cut}}(N \rightarrow \ell X) &= \int_0^{s_{\text{max}}} ds \frac{d\Gamma(N \rightarrow \ell X)}{ds}, \quad s_{\text{max}} \in [\mathcal{O}(1 \text{ GeV}), (m_N - m_\ell)^2], \\ &= \int_0^{x_{\text{max}}} dx \frac{d\Gamma(N \rightarrow \ell X)}{dx} \quad (3.18) \end{aligned}$$

where we defined  $x_{\text{max}} = s_{\text{max}}/m_N^2$ .  $\Gamma_{\text{cut}}(N \rightarrow \ell X)$  obeys eq. (3.17) with  $I_k = I_k(x_\ell^2, x_{\text{max}})$ . Up to  $k = 2$  we found a closed analytical form of the  $I_k(x_\ell^2, x_{\text{max}})$ . The expressions are found in appendix D, but the calculation from eq. (3.13) is easier even for  $k \leq 2$ .

## 4 Numerical analysis

### 4.1 Quality of the perturbation series

Perturbative QCD describes observables in an expansion of the strong coupling constant accurately if the relevant energy scale  $\mu$  is large enough such that  $\alpha_s$  is small. Any observable  $O_\mu$  at the scale  $\mu$  may be written as

$$O_\mu = \sum_n r_n \alpha_s(\mu)^n = r_0 + r_1 \alpha_s(\mu) + r_2 \alpha_s(\mu)^2 + \mathcal{O}(\alpha_s(\mu)^3). \quad (4.1)$$

However, at small scales ( $\mu \sim \mathcal{O}(1 \text{ GeV})$ ) the strong coupling constant becomes large and an expansion in  $\alpha_s$  is no longer guaranteed to converge. For the  $\tau$  lepton with a mass of  $m_\tau = 1.776 \text{ GeV}$  it was only possible to correctly describe the inclusive hadronic decay width through the inclusion of five-loop corrections. Only at the five-loop level does the perturbative

description become stable i.e. the  $\mathcal{O}(\alpha_s^4)$  corrections lead to a significantly reduced dependence on the renormalization scale [13]. In the case of a HSN with mass comparable to the  $\tau$  the kinematics is, up to the mass of the lepton attached to the HSN, the same. Setting  $m_\ell = 0$  leads to the same numerical coefficients as in  $\tau$  decay

$$\frac{\Gamma(N \rightarrow \ell X)}{|V_{ud}|^2 + |V_{us}|^2} = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \times \left[ 1 + a_{m_N} + 5.202a_{m_N}^2 + 26.366a_{m_N}^3 + 127.079a_{m_N}^4 \right], \quad (4.2)$$

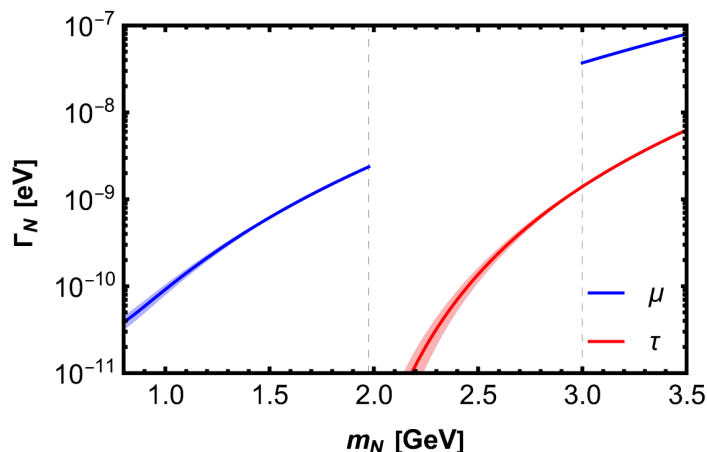
perfectly agreeing with ref. [13]. For a HSN mass around 3 GeV the inclusive semi-hadronic decay with  $\ell = \tau$  in the final state is kinematically accessible. The lepton mass ratio is then sizable,  $x_\tau = 0.592$ , and

$$\frac{\Gamma(N \rightarrow \tau X)}{|V_{ud}|^2 + |V_{us}|^2} = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} \times 0.071 \left[ 1 + a_{m_N} + 8.376a_{m_N}^2 + 74.605a_{m_N}^3 + 669.805a_{m_N}^4 \right], \quad (4.3)$$

largely reducing the decay rate and making the convergence of the perturbative series worse.

In figure 4 we show the full decay width. The widths for  $\ell = e$  and  $\ell = \mu$  are indistinguishable. For  $m_N \geq m_\mu + m_D$  the decay channel into charmed mesons opens up. For this reason we omit the region  $m_N \in [m_\mu + m_D, 3 \text{ GeV}]$  in figure 4 as here there are resonances of charmed mesons. They are dominated by non-perturbative effects making an accurate perturbative description in this region impossible. We show the  $\ell = \mu$  rate again beyond  $m_N \geq 3 \text{ GeV}$ . We have computed the error band by taking as uncertainty the difference between the minimal and maximal width with respect to the renormalization scale i.e.  $\sigma_\Gamma = (\max \Gamma_N(\mu) - \min \Gamma_N(\mu))/2$  for  $0.8 \text{ GeV} \leq \mu \leq 3.5 \text{ GeV}$ . Since we did not include charm mass corrections we expect an additional error of  $\mathcal{O}(m_D^2/q^2) = \mathcal{O}(40\%)$  for  $m_N = 3 \text{ GeV}$ , which is not shown. In figure 5 we plot the effect of each order in  $\alpha_s$  on the total decay width for  $\ell = \mu$ . For small HSN masses, close to the kinematic threshold of multi-hadron production, we see that including more corrections in  $\alpha_s$  does not improve the convergence. This is even more pronounced in figure 5(b) where the relative increase compared to the leading order is large for HSN masses in a window below  $\sim 1 \text{ GeV}$ . Beyond 1 GeV the perturbative behavior rapidly improves and higher order corrections become smaller with higher orders in  $\alpha_s$  in relation to the leading order. We specifically analyzed the scale dependence of the widths by expressing  $a_{m_N}$  in terms of logarithms of the renormalization scale  $\mu$  and the HSN mass as seen in appendix A. In figure 6 the scale dependence for different HSN masses is shown. The scale dependence is very strong close to the kinematic threshold. Increasing the mass of the HSN the scale dependence reduces. We see that HSN masses around 1.5 GeV, including the highest order corrections in  $\alpha_s$  available, lead to a flat scale dependence.

In the  $\ell = \tau$  final state lepton case the discussion is qualitatively similar to the  $\ell = e$  and  $\ell = \mu$  case. Looking only at the decay width in figure 7(a) the effect of poor convergence does not seem to be as pronounced, however, in absolute values it is comparable to the decay into a muon as seen in figure 7(b). As in the muon case, the quality of the perturbative series rapidly increases for larger HSN masses. In figure 8 the scale dependence for different



**Figure 4.** Hadronic decay widths  $\Gamma(N \rightarrow \ell X)$  for the tagging leptons  $\ell = \mu, \tau$  with a mixing angle of  $V_{N\ell} = 10^{-3}$ . The width for  $\ell = e$  looks like the  $\ell = \mu$  width. The numerics of the running coupling were calculated with the help of `RunDec` [34, 35]. We do not show the region  $m_N \in [m_\mu + m_D, 3 \text{ GeV}]$  where we expect large non-perturbative effects. The error band is taken as the difference between minimum and maximum of the width with respect to the renormalization scale  $\sigma_\Gamma = (\max \Gamma_N(\mu) - \min \Gamma_N(\mu))/2$  for  $0.8 \text{ GeV} \leq \mu \leq 3.5 \text{ GeV}$ . We do not include the uncertainty from our omission of quark masses, which can be sizable,  $\mathcal{O}(m_D^2/q^2) = \mathcal{O}(40\%)$  for  $m_N = 3 \text{ GeV}$ , beyond the charm production threshold.

HSN masses is shown. The scale dependence reduces for larger masses. At a HSN mass of  $3 \text{ GeV}$  the  $\mathcal{O}(\alpha_s^4)$  corrections reduce the scale dependence significantly. For larger HSN masses already lower order results show a satisfactory small scale dependence.

## 4.2 Branching ratios of HSN

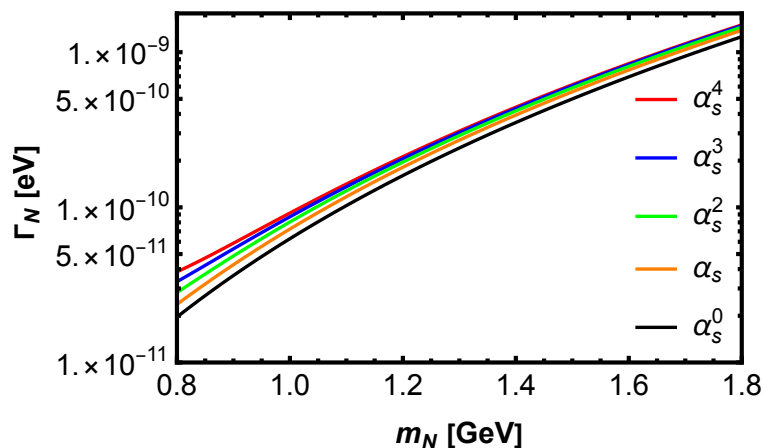
Without the neutral current contributions mediated by the  $Z$  boson one cannot calculate the total width, which is needed for branching ratio predictions. However ratios of branching ratios for the charged decays can be calculated

$$\kappa_\ell^P = \frac{Br(N \rightarrow \ell P)}{Br(N \rightarrow \ell X)} = \frac{\Gamma(N \rightarrow \ell P)}{\Gamma(N \rightarrow \ell X)}, \quad (4.4)$$

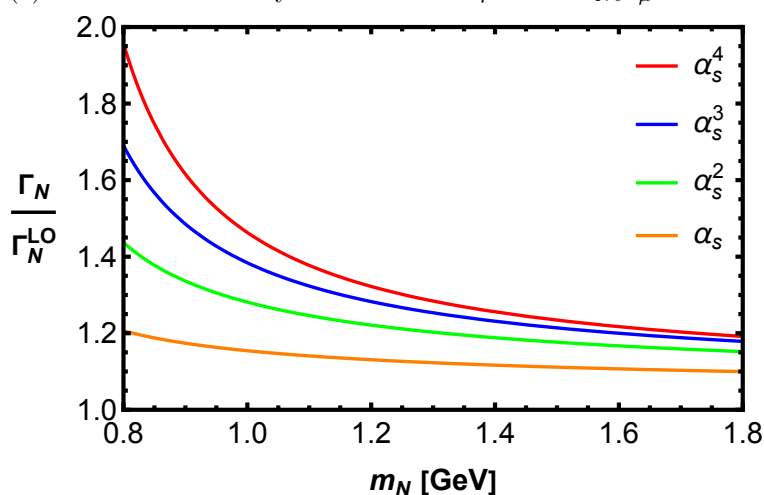
here  $P = \pi^+, K^+, D^+, \dots$  is some meson,  $\Gamma(N \rightarrow \ell P)$  is the width of a HSN into a meson

$$\Gamma(N \rightarrow \ell P) = \frac{G_F^2 m_N^3 |V_{N\ell}|^2}{16\pi} |V_{u'd'}|^2 f_P^2 \left[ (1 - x_\ell^2)^2 - x_P^2 (1 + x_\ell^2) \right] \sqrt{\lambda(1, x_\ell^2, x_P^2)}, \quad (4.5)$$

where  $f_P$  is the meson decay constant,  $|V_{u'd'}|$  is the relevant CKM matrix element, and  $x_i = m_i/m_N$ , and  $\Gamma(N \rightarrow \ell X)$  is the fully inclusive semi-hadronic decay rate as given in eq. (3.17). For our predictions of  $\kappa_\ell^P$  we assume  $m_N > 1.5 \text{ GeV}$  for electron and muon tagging and  $m_N > 3 \text{ GeV}$  for tau tagging. In figure 9 the  $\kappa$ -ratio for muon and tau tagging are shown. In both cases pions dominate the decay and only a few CKM-suppressed decays to kaons are predicted. For this reason we omit decays to  $D$  mesons as they are more suppressed than those into kaons due to the smaller phase space. In addition we also show the impact of the inclusion of  $D_s$  mesons.  $D_s$  mesons are produced copiously beyond the charm threshold as  $V_{cs} \approx 1$ . We exclude the region  $m_N \in [m_\mu + m_D, 3 \text{ GeV}]$  as here non-perturbative charm resonance effects



(a) Inclusive HSN decay width with  $\ell = \mu$ . Here  $V_{N\ell=\mu} = 10^{-3}$ .



(b)  $\Gamma_N$  at different orders of  $\alpha_s$ , normalized to the LO result.

**Figure 5.** Inclusive decay width  $\Gamma_N = \Gamma(N \rightarrow \mu X)$  as a function of  $m_N$ .

dominate the width. As long as the decay channels to charmed final states are absent, at least 12% of the charged-current decays of HSN are  $N \rightarrow \ell\pi$  for all considered values of  $m_N$ .

### 4.3 Constraints on $\theta_{N\tau}$ from the $\tau$ lifetime

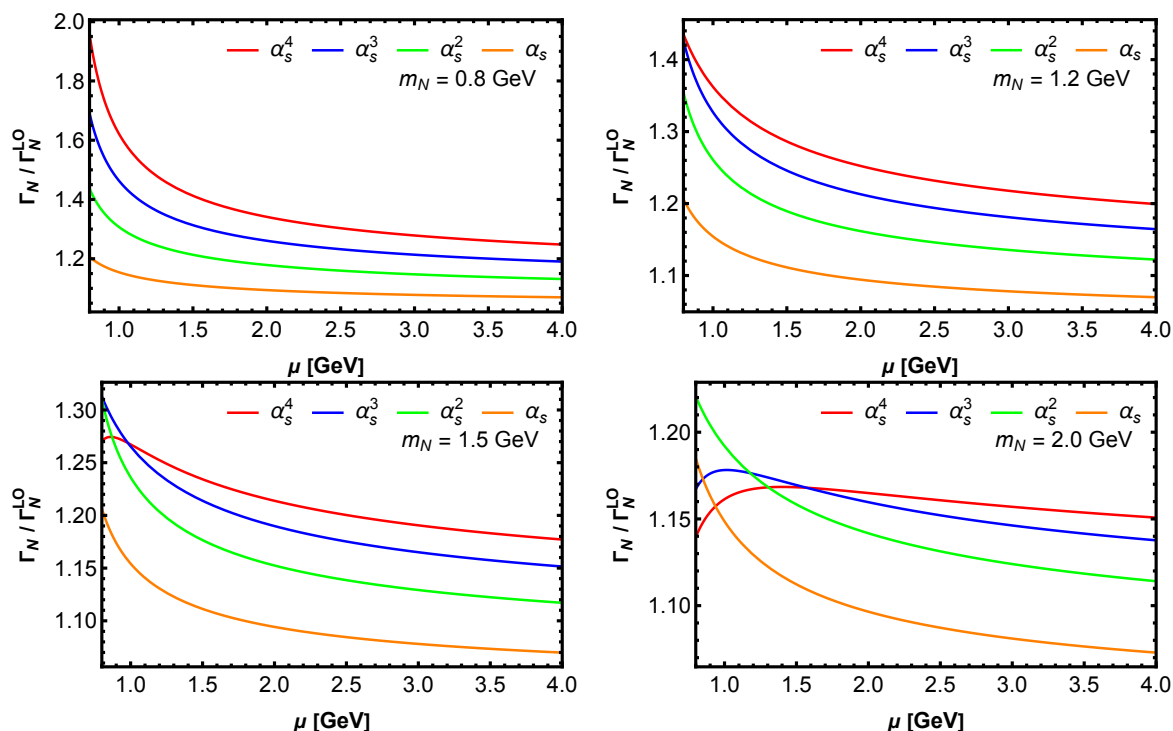
The current world average of the  $\tau$ -lifetime

$$\tau_\tau^{\text{exp.}} = (290.29 \pm 0.53) \text{ fs} \quad [40] \quad (4.6)$$

agrees with the SM theory prediction

$$\tau_\tau^{\text{SM}} = (288.59 \pm 2.31) \text{ fs} \quad [39, 41]. \quad (4.7)$$

Whenever a sterile neutrino interaction is present it comes with a factor of the mixing angle  $V_{N\tau} = \sin\theta_{N\tau}$ . At the same time SM neutrino interactions receive a contribution involving  $\cos\theta_{N\tau}$ . Mixing is possible between one or more generations of SM neutrinos. If there is mixing with  $\nu_e$  or  $\nu_\mu$  this has an impact on the Fermi constant. A complete analysis



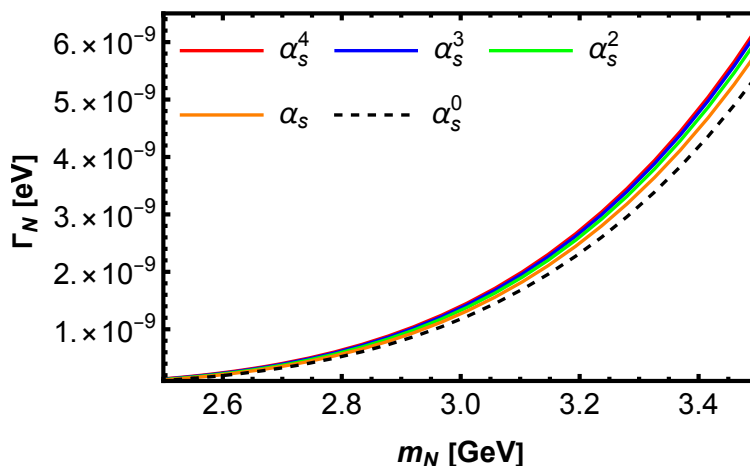
**Figure 6.** Scale dependence of the inclusive decay width  $\Gamma_N = \Gamma(N \rightarrow \mu X)$  for different HSN masses.

including also mixing with  $e$  or  $\mu$  requires a full electroweak fit (see ref. [42]) and is beyond the scope of this work. Hence we only look at mixing with  $\nu_\tau$ , and define  $\theta \equiv \theta_{N\tau}$ . Then a sterile neutrino will modify the total width, implying the following constraint on  $\theta$ :

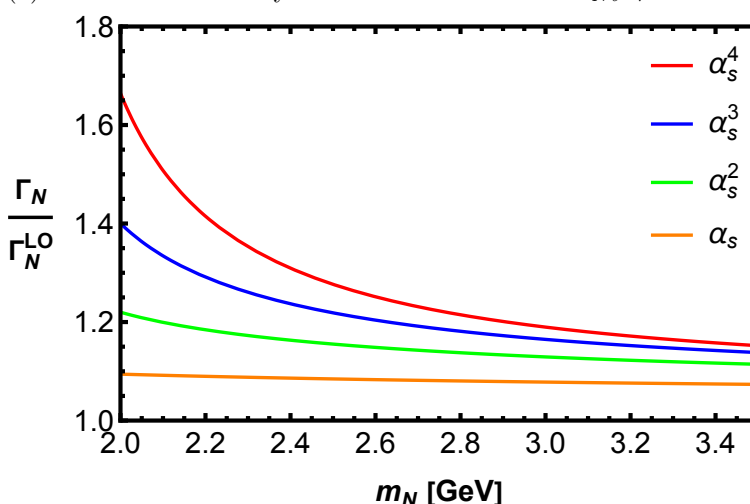
$$\tau_\tau^{\text{exp.}} = \frac{1 - \mathcal{B}_\tau^s}{\cos^2 \theta (1 - \mathcal{B}_\tau^s) \Gamma_\tau^{\text{SM}} + \sin^2 \theta \Gamma^N}, \quad (4.8)$$

where  $\mathcal{B}_\tau^s = 0.0292$  [39] is the branching fraction of  $\tau$  decays into strange final states,  $(1 - \mathcal{B}_\tau^s) \Gamma_\tau^{\text{SM}}$  is the SM width involving leptonic and non-strange hadronic decays, and  $\Gamma^N$  is the non-strange sterile neutrino contribution, i.e. the  $\tau \rightarrow NX$  decay rate with  $X$  being a non-strange hadronic final state.  $(1 - \mathcal{B}_\tau^s) / \tau_\tau^{\text{exp.}}$  is the experimentally determined total decay rate into non-strange particles; our approach follows ref. [39].  $\Gamma_\tau^{\text{SM}}$  is calculated as  $\Gamma_\tau^{\text{SM}} = 1 / \tau_\tau^{\text{SM}}$  and  $\Gamma^N$  follows from eq. (3.17) by replacing  $V_{N\ell} \rightarrow 1$ ,  $V_{us} \rightarrow 0$ ,  $m_N \rightarrow m_\tau$ , and  $I_k(x_\ell^2, (1 - x_\ell)^2) \rightarrow I_k(m_N^2/m_\tau^2, (1 - m_N/m_\tau)^2)$ . For the strong coupling constant we use  $\alpha_s^{n_f=3}(m_\tau) = 0.316$  which we obtain by using the weighted average  $\alpha_s^{n_f=5}(M_Z) = 0.1182$  of all measurements listed in the PDG [39], except the one using  $\tau$  decay, and running  $\alpha_s(M_Z)$  down to  $m_\tau$ . In figure 10 we show the renormalization scale dependence of the  $\tau$  decay width. At  $\mathcal{O}(\alpha_s^4)$  the scale dependence remains sufficiently flat up to HSN masses around  $m_N = 600$  MeV. For larger HSN masses the perturbative description becomes poorer.

Sterile neutrinos solely produced through mixing with  $\nu_\tau$  will always decrease the  $\tau$  width and thus prolong the predicted lifetime, possibly bringing it into tension with the experimental value. This is so, because the decay rate to  $\nu_\tau$  is decreased by a factor of  $\cos^2 \theta_{N\tau}$  while the additional decay channel with  $N$ , proportional to  $\sin^2 \theta_{N\tau}$  involves phase-space suppression. Neglecting electroweak corrections in  $\Gamma^N$  we use eq. (4.8) to determine the parameter space



(a) Inclusive HSN decay width with  $\ell = \tau$ . Here  $V_{N\ell=\tau} = 10^{-3}$ .



(b)  $\Gamma_N$  at different orders of  $\alpha_s$ , normalized to the LO result.

**Figure 7.** Inclusive decay width  $\Gamma_N = \Gamma(N \rightarrow \tau X)$ .

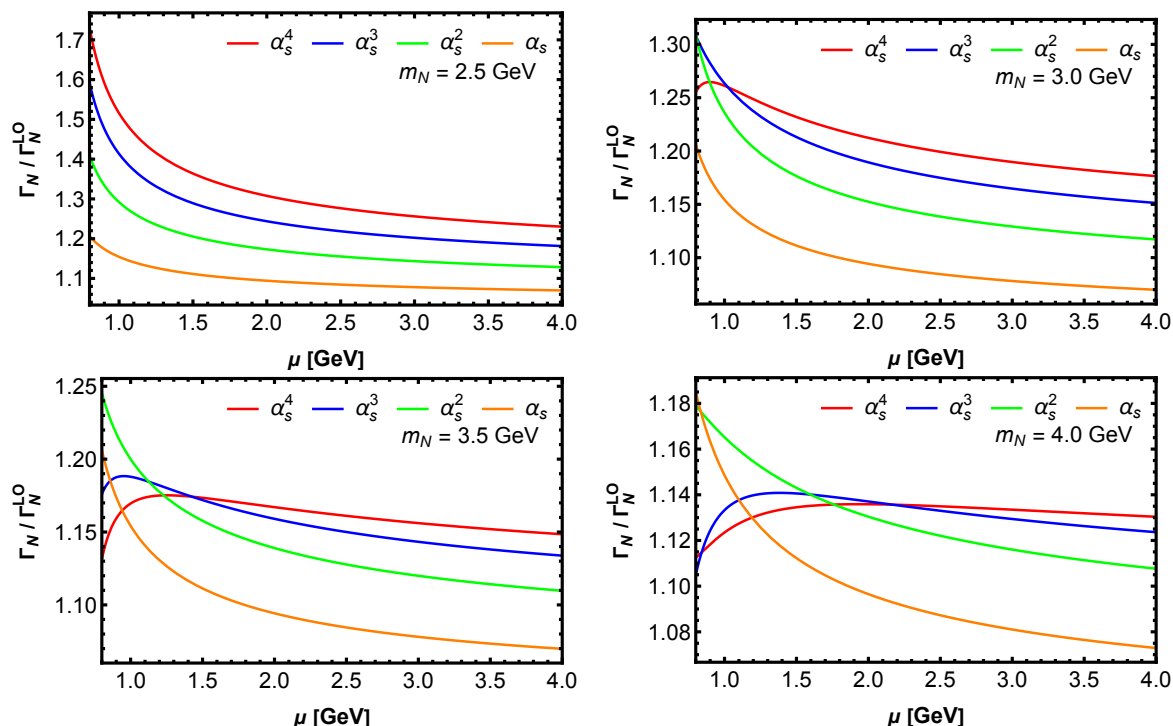
for  $\theta$  and  $m_N$  in agreement with the measured  $\tau$  lifetime, see figure 11, which shows the  $1\sigma$  region of the allowed mixing angle and HSN mass, up to  $m_N \leq 600$  MeV. We further show the allowed region when the experimental value in eq. (4.6) is varied in the  $3\sigma$  range, while we do not triple the theoretical uncertainty. For  $m_N > 600$  MeV we continue our  $1\sigma$  and  $3\sigma$  lines for illustration, bearing in mind that the quality of the perturbative calculation diminishes significantly for  $m_N \gtrsim 600$  MeV, while the perturbation series for the inclusive hadronic  $\tau$  decay width converges well for  $m_N \lesssim 600$  MeV.

Eq. (4.8) holds for any value of  $m_N$ , in particular, it also holds for  $m_N > m_\tau$ . In this case  $\Gamma_N = 0$  and eq. (4.8) becomes

$$\sin^2 \theta = 1 - \frac{\tau_\tau^{\text{th.}}}{\tau_\tau^{\text{exp.}}} = 0.00586 \pm 0.00816, \quad (4.9)$$

which leads to

$$|\sin \theta| \leq 11.8 \cdot 10^{-2} \quad \text{at } 1\sigma. \quad (4.10)$$



**Figure 8.** Scale dependence of the inclusive decay width  $\Gamma_N = \Gamma(N \rightarrow \tau X)$  for different HSN masses.

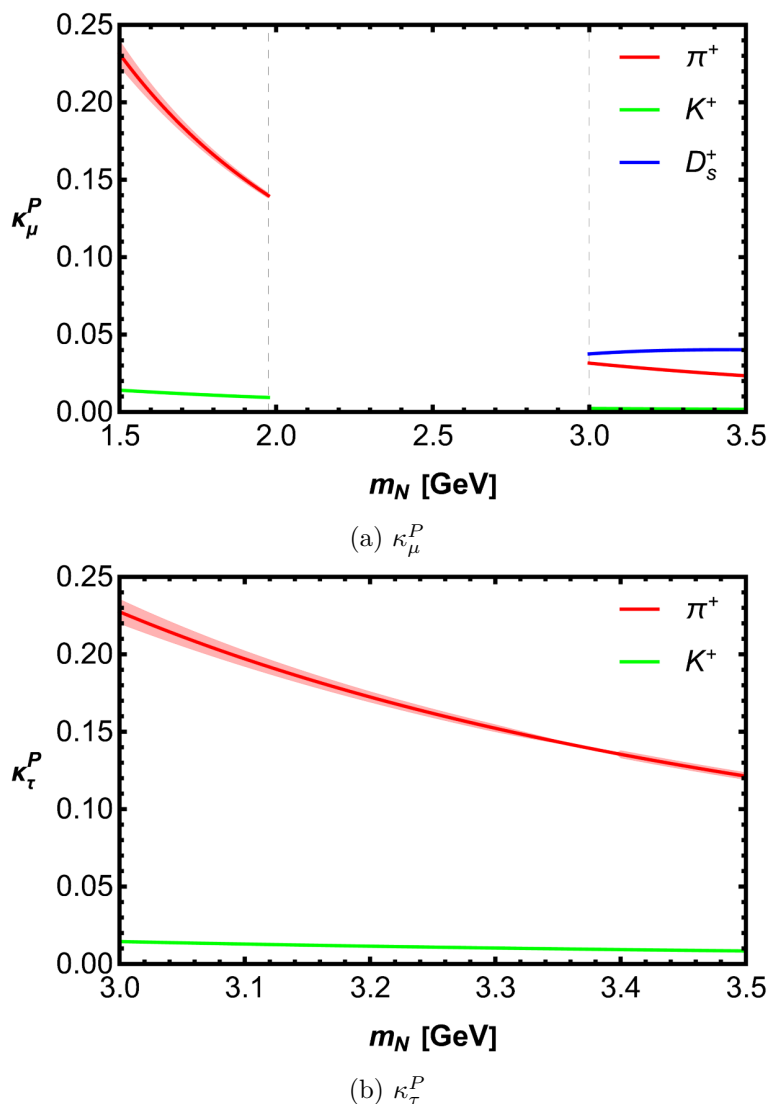
#### 4.4 Contributions on $\theta_{N\tau}$ for $m_N > m_\tau$ from $\tau$ branching fractions

HSNs heavier than  $\tau$  affect  $\tau_\tau$  in eq. (4.8) indirectly through the cosine of the mixing angle. To study this scenario we look at the decays  $\tau^- \rightarrow \pi^- \nu_\tau$ ,  $\tau^- \rightarrow K^- \nu_\tau$ , and  $\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell$  [20, 43]. This will allow us to bypass the theoretical uncertainty of the total hadronic width feeding into  $\tau_\tau$  in eq. (4.7). The decay widths of interest read

$$\begin{aligned} \Gamma(\tau \rightarrow P \nu_\tau) &= \cos^2 \theta \Gamma_{\text{SM}}(\tau \rightarrow P \nu_\tau) \\ &= \frac{G_F^2 m_\tau^3 \cos^2 \theta}{16\pi} |V_{uq}|^2 f_P^2 \left(1 - \frac{m_P^2}{m_\tau^2}\right)^2 (1 + \delta_{\text{RC}}^{(P)}), \end{aligned} \quad (4.11)$$

$$\begin{aligned} \Gamma(\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell) &= \cos^2 \theta \Gamma_{\text{SM}}(\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell) = \frac{G_F^2 m_\tau^5 \cos^2 \theta}{192\pi^3} \\ &\times \left[ (1 - 8x_\ell^2 + 8x_\ell^6 - x_\ell^8 - 12x_\ell^4 \ln(x_\ell^2)) + \frac{\alpha(m_\tau)}{\pi} H_1 + \frac{\alpha^2(m_\tau)}{\pi^2} H_2 \right], \end{aligned} \quad (4.12)$$

where  $x_\ell = m_\ell/m_\tau$ ,  $P$  is either  $K^-$  or  $\pi^-$ ,  $f_P$  is the meson decay constant, and  $V_{uq}$  is the CKM matrix element. We have added the  $\cos \theta$  factor to account for the effect of HSN mixing. Here the  $\delta_{\text{RC}}^{(P)}$  are the radiative QED corrections governed by  $\alpha(m_\tau)$ . We use the values  $\delta_{\text{RC}}^{(\pi)} = 0.0194 \pm 0.0061$  and  $\delta_{\text{RC}}^{(K)} = 0.0204 \pm 0.0062$  presented in ref. [44] which are based on



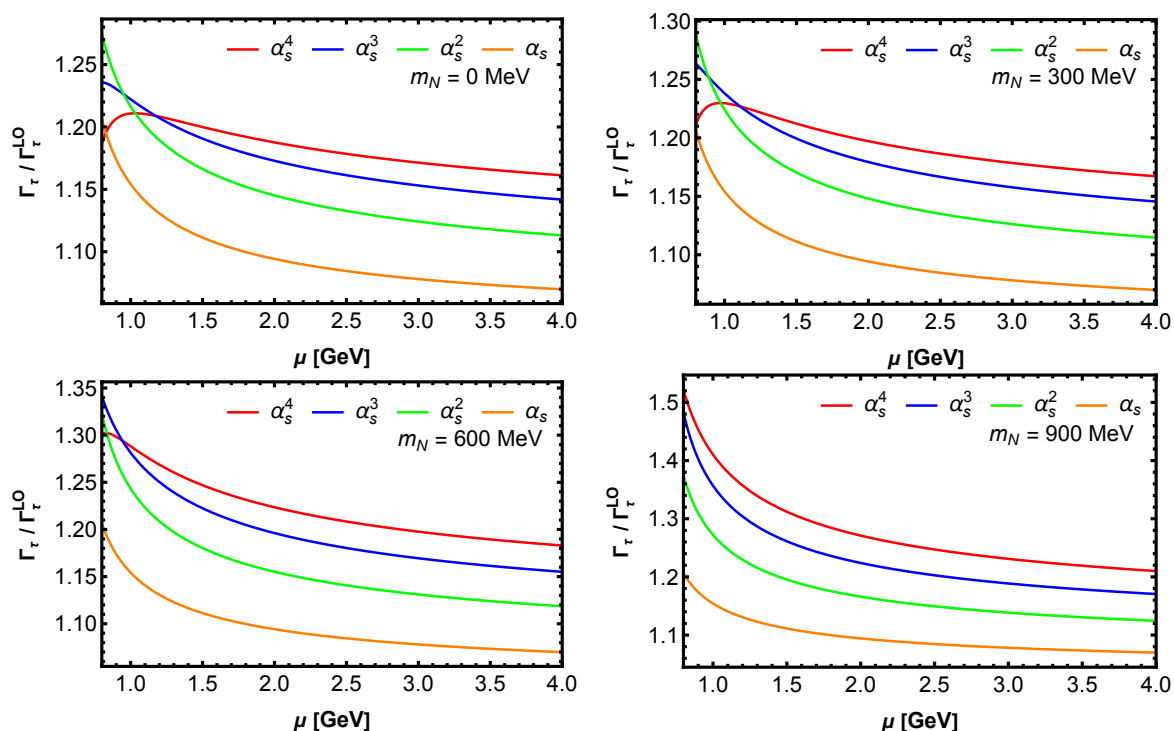
**Figure 9.**  $N \rightarrow \ell P$  branching fractions (normalized to  $Br(N \rightarrow \ell N)$ ) for  $P = \pi^+, K^+, D_s^+$ , see eq. (4.4). We use  $f_D = (212.0 \pm 0.7)$  MeV,  $f_{D_s} = (249.9 \pm 0.5)$  MeV [36–38], and  $|V_{cd}| = 0.221 \pm 0.004$ , and  $|V_{cs}| = 0.975 \pm 0.006$  [39]. The error bands are the same as in figure 4.

refs. [45–48]. The  $H_i$  are the radiative QED corrections for leptonic decay, reading [49–53]

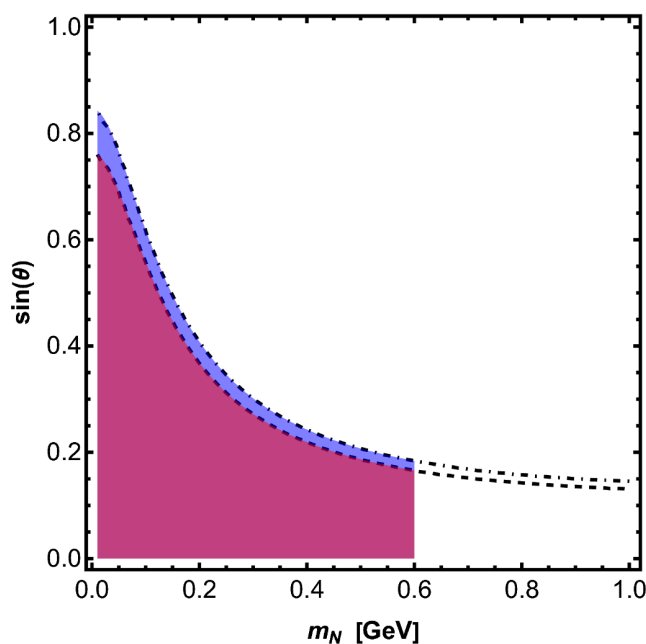
$$H_1 = \left( \frac{25}{8} - \frac{\pi^2}{2} \right) - (34 + 24 \ln x_\ell) x_\ell^2 + 16\pi^2 x_\ell^3 + \mathcal{O}(x_\ell^4) \quad (4.13)$$

$$H_2 = \frac{156815}{5184} - \frac{518}{81} \pi^2 - \frac{895}{36} \zeta_3 + \frac{67}{720} \pi^4 + \frac{53}{6} \pi^2 \ln 2 - (0.042 \pm 0.002) - \frac{5}{4} \pi^2 x_\ell + \mathcal{O}(x_\ell^2) \quad (4.14)$$

We use the measured  $\tau$  lifetime  $\tau_\tau = (290.29 \pm 0.53)$  fs [40], the branching ratios, decay constants, and parameters listed in table 1. We have determined the pion decay constant from the ratio  $f_{K^+}/f_{\pi^+}$  and  $f_{K^+}$ .



**Figure 10.** Scale dependence of the inclusive  $\tau$  decay width with a final state HSN for different HSN masses  $m_N$ .



**Figure 11.** Allowed parameter space deduced from the measured  $\tau$  lifetime in eq. (4.6) via eq. (4.8) at the  $1\sigma$  (purple) and  $3\sigma$  (purple and blue) levels. For  $m_N \leq 600$  MeV we trust the perturbative calculation. The dashed lines for  $m_N > 600$  MeV indicate the perturbative result in a region where the description is no longer reliable.

$\text{Br}^{\text{exp.}}(\tau \rightarrow \nu_\tau e \bar{\nu}_e)$	$0.1785 \pm 0.0004$	HFLAV [40]
$\text{Br}^{\text{exp.}}(\tau \rightarrow \nu_\tau \mu \bar{\nu}_\mu)$	$0.17366 \pm 0.00036$	HFLAV [40]
$\text{Br}^{\text{exp.}}(\tau \rightarrow \pi \nu_\tau)$	$0.1082 \pm 0.0005$	HFLAV [40]
$\text{Br}^{\text{exp.}}(\tau \rightarrow K \nu_\tau)$	$(0.697 \pm 0.010) \cdot 10^{-2}$	HFLAV [40]
$f_{K^+}/f_{\pi^+}$	$1.1934 \pm 0.0019$	FLAG [36–38, 54–56]
$f_{K^+}$	$(155.7 \pm 0.3) \text{ MeV}$	FLAG [37, 38, 55–57]
$f_{\pi^+}$	$(130.5 \pm 0.3) \text{ MeV}$	
$ V_{ud} $	$0.97367 \pm 0.00032$	PDG [39]
$ V_{us} $	$0.22431 \pm 0.00085$	PDG [39]
$1/\alpha(m_\tau)$	$133.50 \pm 0.02$	[58, 59]

**Table 1.** Table of branching ratios, decay constants, and CKM elements used.

This leads to the following mixing angles allowed by the experimental bounds for the leptonic decays

$$\left( \frac{\text{Br}^{\text{exp.}}(\tau \rightarrow \nu_\tau e \bar{\nu}_e)}{\tau_\tau^{\text{exp.}}} \cdot \frac{1}{\Gamma_{\text{SM}}(\tau \rightarrow \nu_\tau e \bar{\nu}_e)} \right)^{1/2} = \cos \theta = 1.00201 \pm 0.00145 \quad (4.15)$$

$$\left( \frac{\text{Br}^{\text{exp.}}(\tau \rightarrow \nu_\tau \mu \bar{\nu}_\mu)}{\tau_\tau^{\text{exp.}}} \cdot \frac{1}{\Gamma_{\text{SM}}(\tau \rightarrow \nu_\tau \mu \bar{\nu}_\mu)} \right)^{1/2} = \cos \theta = 1.00205 \pm 0.00139 \quad (4.16)$$

and for the semi-hadronic decays

$$\left( \frac{\text{Br}^{\text{exp.}}(\tau \rightarrow \pi \nu_\tau)}{\tau_\tau^{\text{exp.}}} \cdot \frac{1}{\Gamma_{\text{SM}}(\tau \rightarrow \pi \nu_\tau)} \right)^{1/2} = \cos \theta = 0.99717 \pm 0.00462 \quad (4.17)$$

$$\left( \frac{\text{Br}^{\text{exp.}}(\tau \rightarrow K \nu_\tau)}{\tau_\tau^{\text{exp.}}} \cdot \frac{1}{\Gamma_{\text{SM}}(\tau \rightarrow K \nu_\tau)} \right)^{1/2} = \cos \theta = 0.99091 \pm 0.00884 \quad (4.18)$$

here we roughly estimated the errors by assuming they are all uncorrelated and distributed Gaussian. The leptonic decays all prefer unphysical values with  $\cos \theta > 1$  at the  $1\sigma$  level, the semi-hadronic decays, however, permit non-zero mixing angles at the  $1\sigma$  level. For the decay  $\tau \rightarrow \pi \nu_\tau$  we find the mixing angle at  $1\sigma$

$$\tau \rightarrow \pi \nu_\tau : \quad |\sin \theta| \leq 12.2 \cdot 10^{-2} \quad \text{at } 1\sigma \quad (4.19)$$

and for  $\tau \rightarrow K \nu_\tau$  the mixing angle

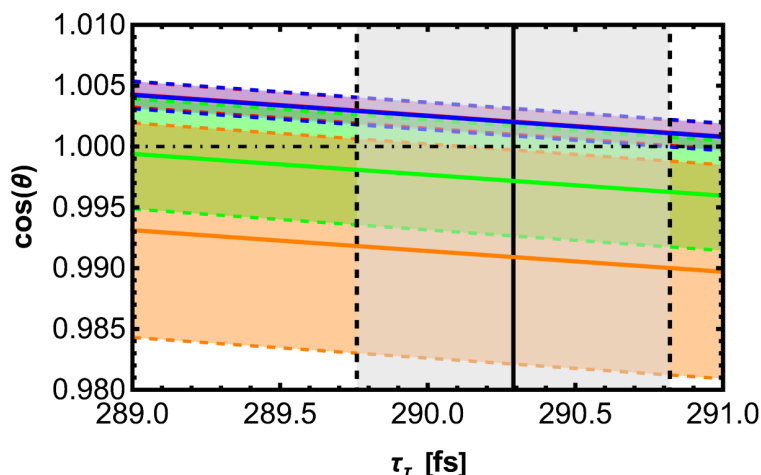
$$\tau \rightarrow K \nu_\tau : \quad |\sin \theta| = \left( 13.5_{-11.3}^{+5.4} \right) \cdot 10^{-2} \quad \text{at } 1\sigma. \quad (4.20)$$

In figure 12 we show the  $1\sigma$  band of the cosines of the mixing angle as a function of the  $\tau$  lifetime. Combining the semi-hadronic and leptonic decays yields  $\cos \theta = 1.00168 \pm 0.00098$ , while combining only the two semi-hadronic decays leads to  $\cos \theta = 0.99582 \pm 0.00410$  corresponding to

$$\tau \rightarrow P \nu_\tau, \quad P = \pi, K \text{ combined} : \quad |\sin \theta| = \left( 9.1_{-7.8}^{+3.7} \right) \cdot 10^{-2} \quad \text{at } 1\sigma. \quad (4.21)$$

Taking the weighted average of the leptonic decays leads to  $\cos \theta = 1.00203 \pm 0.00100$ . These decays permit  $\cos \theta \geq 0.99901$  at the  $3\sigma$  level leading to:

$$\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell \quad \ell = e, \mu \text{ combined} : \quad |\sin \theta| \leq 4.4 \cdot 10^{-2} \quad \text{at } 3\sigma. \quad (4.22)$$



**Figure 12.** Cosine of the mixing angle vs. the  $\tau$  lifetime. The grey band is the  $1\sigma$  band of the world average of the  $\tau$  lifetime. The solid lines are the central values and the dashed lines delimit the  $1\sigma$  band. The bands from  $\tau \rightarrow \nu_\tau e \bar{\nu}_e$  (blue) and  $\tau \rightarrow \nu_\tau \mu \bar{\nu}_\mu$  (red) overlap. The constraints from both  $\tau \rightarrow \pi \nu_\tau$  (green) and  $\tau \rightarrow K \nu_\tau$  (orange) fully comply with the measured  $\tau$  lifetime. The area under the dot-dashed line at  $\cos \theta = 1$  delimits the physically allowed region.

#### 4.5 Detecting HSN for $m_N < m_\tau$ from spectra

HSNs light enough to be produced in  $\tau$  decays are difficult to detect in branching ratios, because

$$\Gamma(\tau \rightarrow \nu_\tau Y) + \Gamma(\tau \rightarrow NY) = \Gamma_{\text{SM}}(\tau \rightarrow \nu_\tau Y) + \mathcal{O}\left(\frac{m_N^2}{m_\tau^2} \cdot \sin^2 \theta\right) \quad (4.23)$$

for  $m_N < m_\tau$  and where  $Y$  is any final state. To this end one better employs precise measurements of the charged lepton energy spectrum

$$\frac{d\Gamma(\tau \rightarrow \nu \ell \bar{\nu}_\ell)}{dE_\ell} = \frac{G_F^2 m_\tau^4 V_\nu^2}{2\pi^3} F(x_E, x_\ell, x_\nu), \quad (4.24)$$

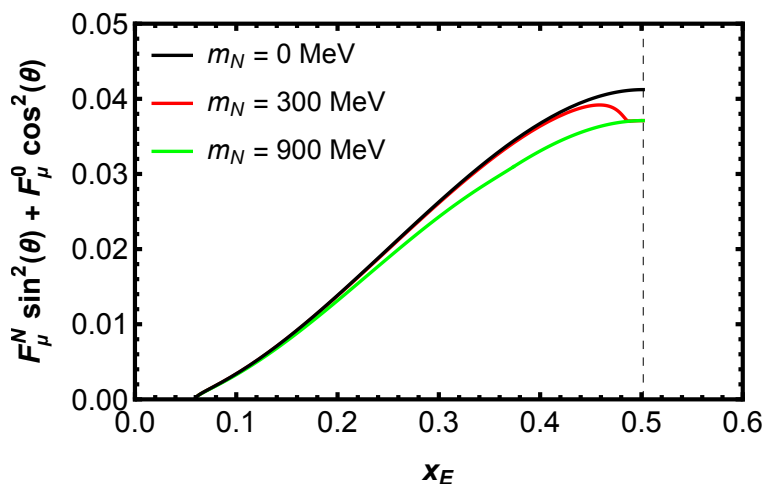
here  $\nu = \nu_\tau$ ,  $N$  is either a SM or a sterile neutrino,  $x_E = E_\ell/m_\tau$ ,  $x_i = m_i/m_\tau$ ,  $V_\nu = \cos \theta$  for SM and  $V_\nu = \sin \theta$  for sterile neutrinos, and

$$F(x_E, x_\ell, x_\nu) = \frac{\sqrt{x_E^2 - x_\ell^2} (1 - 2x_E + x_\ell^2 - x_\nu^2)^2}{6(1 - 2x_E + x_\ell^2)^3} \quad (4.25)$$

$$\times \left[ 8x_E^3 - 2x_E^2(5 + 5x_\ell^2 + x_\nu^2) \right. \\ \left. + x_E(3 + 10x_\ell^2 + 3x_\ell^4 + 3x_\nu^2(1 + x_\ell^2)) - 2x_\ell^2(1 + x_\ell^2 + 2x_\nu^2) \right].$$

Since the experiment cannot distinguish between sterile and SM neutrino the measured lepton energy is

$$\begin{aligned} \frac{d\Gamma_{\text{lep}}^{\text{exp.}}}{dE_\ell} &= \frac{d\Gamma(\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell)}{dE_\ell} + \frac{d\Gamma(\tau \rightarrow N \ell \bar{\nu}_\ell)}{dE_\ell} \quad (4.26) \\ &= \frac{G_F^2 m_\tau^4 V_\nu^2}{2\pi^3} \left[ F(x_E, x_\ell, x_N) \sin^2 \theta + F(x_E, x_\ell, 0) \cos^2 \theta \right]. \end{aligned}$$



**Figure 13.** Here we show  $F_\mu^N \sin^2 \theta + F_\mu^0 \cos^2 \theta = F(E_\mu/m_\tau, x_\mu, x_N) \sin^2 \theta + F(E_\mu/m_\tau, x_\mu, 0) \cos^2 \theta$  for  $\ell = \mu$ , the black line corresponds to the SM case  $m_N = 0$ , the mixing angle assumed here corresponds to  $\sin \theta = 0.32$ .

The hadronic spectrum is derived analogously,

$$\frac{d\Gamma(\tau \rightarrow \nu X)}{ds} = N_c \frac{G_F^2 m_\tau^3 V_\nu^2 \cdot 2(|V_{ud}|^2 + |V_{us}|^2)}{192\pi^3} S(m_\tau, m_Z) G(x, x_\nu), \quad (4.27)$$

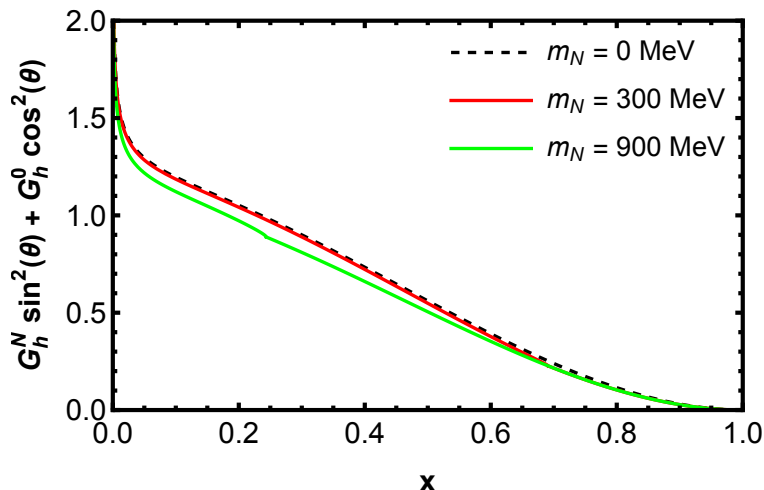
where  $S(m_\tau, m_Z) = 1.01907$  [59, 60] accounts for electroweak corrections,  $x = s/m_\tau^2$ ,  $x_\nu = m_\nu/m_\tau$ , and

$$G(x, x_\nu) = \left[ \left(1 + x_\nu^2 - x\right) \left(1 + 2x + x_\nu^2\right) - 4x_\nu^2 \right] \times \sqrt{\lambda(1, x, x_\nu^2)} \cdot 12\pi \text{Im} \Pi_V^{1+0}(m_\tau^2 x). \quad (4.28)$$

As for the leptonic decay the measured quantity is the sum of SM and sterile neutrino contributions

$$\begin{aligned} \frac{d\Gamma_{\text{had}}^{\text{exp.}}}{ds} &= \frac{d\Gamma(\tau \rightarrow \nu_\tau X)}{ds} + \frac{d\Gamma(\tau \rightarrow NX)}{ds} \\ &= N_c \frac{G_F^2 m_\tau^3 V_\nu^2 \cdot 2(|V_{ud}|^2 + |V_{us}|^2)}{192\pi^3} S(m_\tau, m_Z) \\ &\quad \times \left[ G(x, x_N) \sin^2 \theta + G(x, 0) \cos^2 \theta \right]. \end{aligned} \quad (4.29)$$

In figures 13 and 14 we show the spectrum for the leptonic decay with  $\ell = \mu$  in the final state and the hadronic spectrum, respectively, in both cases with  $\sin \theta = 0.32$ . A HSN with a mass of  $m_N = 300$  MeV leads to a significantly modified spectrum. The perturbative calculation of the hadronic spectrum can be compared to data for large  $s \gtrsim 1$  GeV. In addition, one can compare integrated spectra from  $s = 0$  to a chosen  $s_{\text{max}} \gtrsim 1$  GeV, exploiting quark-hadron duality as described in appendix B.



**Figure 14.** The plot shows  $G_h^N \sin^2 \theta + G_h^0 \cos^2 \theta = G(s/m_\tau^2, x_N) \sin^2 \theta + G(s/m_\tau^2, 0) \cos^2 \theta$ , the mixing angle is chosen to give  $\sin \theta = 0.32$ .

#### 4.6 Comparisons

It is interesting to compare the allowed values of the mixing angle with other experimental exclusions. Both ATLAS [61, 62] and CMS [63–65] have published upper and lower limits on the mixing angle. ATLAS performed an analysis for the HSN exclusively mixing with the  $\tau$  neutrino and is less competitive than CMS. The CMS experiment specifically looked at  $\nu_\tau$ - $N$  mixing and is sensitive to HSN masses  $1 \text{ GeV} \leq m_N \lesssim 2.2 \text{ GeV}$  and  $m_N \geq 3 \text{ GeV}$ .

Below  $m_N < 3 \text{ GeV}$  there are exclusion limits for  $N$ - $\nu_\tau$  mixing from DELPHI [66], ArgoNeuT [67], BABAR [68], and Belle [69]. In the mass range  $0.4 \text{ GeV} \leq m_N \leq 1.3 \text{ GeV}$  ArgoNeuT and BABAR provide the strongest bounds. Beyond  $m_N \geq 1.3 \text{ GeV}$  DELPHI and CMS provide the strongest bounds, with the DELPHI bound being strongest for  $m_N \gtrsim 1.9 \text{ GeV}$ . Furthermore, there are recasts of data from the CHARM [70] and WA66 experiments [71]. While originally these experiments were not understood to be sensitive to  $N$ - $\nu_\tau$  mixing for HSN with masses  $m_N \approx 1 \text{ GeV}$ , the recasts provide the tightest constraints up to  $m_N = 1.6 \text{ GeV}$ . Only the DELPHI experiment is sensitive to masses in the range  $2.2 \text{ GeV} \lesssim m_N < 3 \text{ GeV}$ .

For  $m_N < m_\tau$  we find the allowed range  $|V_{N\tau}| = |\sin \theta| \lesssim 0.2$  from the  $\tau$  lifetime for  $m_N \approx 600 \text{ MeV}$ . In this mass regime our bound from  $\tau$  decay is not competitive, the strongest bound stems from the BABAR analysis, which finds  $|\sin \theta| \lesssim 2.49 \cdot 10^{-2}$  [68], and the CHARM and WA66 recasts,  $|\sin \theta| \lesssim 1.73 \cdot 10^{-3}$  [70, 71], which, however, are more model-dependent as explained below.

For  $m_N > m_\tau$  we find the mixing angle in eq. (4.10) from the  $\tau$  lifetime. The semi-hadronic  $\tau$  decays exclude mixing angles for HSN masses  $m_N > m_\tau - m_P$  where  $P = \pi, K$  and lead to eq. (4.19) for  $\tau \rightarrow \pi\nu_\tau$ , and eq. (4.20) for  $\tau \rightarrow K\nu_\tau$ . In combination, the semi-hadronic decays lead to eq. (4.21), valid for  $m_N > m_\tau - m_P$ . Kobach and Dobbs (2015) [43] have updated the bound of ref. [20]. They used leptonic  $\tau$  decays to constrain the mixing angle for  $m_N > m_\tau$  and derived the bound  $|\sin \theta^{\text{KD}}| \lesssim 7.07 \cdot 10^{-2}$  at the  $\sim 2\sigma$  level. This is weaker than our bound obtained from leptonic  $\tau$  decays given in eq. (4.22) at the  $3\sigma$  level. While our bounds obtained from semi-hadronic  $\tau$  decays (eq. (4.21)) and the  $\tau$  lifetime (eq. (4.10)) are

weaker than the constraint from leptonic decays, they are less model-dependent as explained below. It is interesting to note that  $\theta = 0$  is ruled out by  $1\sigma$  in eq. (4.21). Bearing in mind that our error estimates do not consider correlations, these numbers may well indicate that better data may lead to evidence for  $\theta \neq 0$  in the foreseeable future.

The experiment most sensitive in the mass range  $1.9 \text{ GeV} \lesssim m_N \leq 3 \text{ GeV}$  is the DELPHI experiment, determining  $\sin^2 \theta^{\text{DELPHI}} \lesssim 6 \cdot 10^{-5} \Rightarrow |\sin \theta^{\text{DELPHI}}| \lesssim 7.75 \cdot 10^{-3}$  for  $m_N \approx 2 \text{ GeV}$ ; the bound becomes weaker for smaller HSN masses [66]. The DELPHI experiment studied  $Z$  decays and was therefore sensitive to HSN and SM neutrinos of all flavors. However, they searched for the HSN through charged-current and neutral-current decays of HSN and derived the bound  $|\sin \theta^{\text{DELPHI}}| \lesssim 1.55 \cdot 10^{-2}$  from the non-observation of HSN decays into  $\tau$  leptons. The DELPHI bound is tighter than ours, but is also more model-dependent: while both our and the DELPHI study assume that  $N$  only interacts with SM particles through  $N$ - $\nu_\tau$  mixing, we do not assume anything on other interactions of  $N$ . A heavy  $N$  could be a mediator to a dark sector and possibly decay into lighter sterile neutrinos and/or dark bosons. In such scenarios the branching ratios of the decay modes studied by DELPHI will decrease and the bounds on  $|\theta|$  will become weaker and the same remark applies to the recasts of CHARM and WA66 data. For the same reason we also put less emphasis on the bounds on  $|\theta|$  in eq. (4.22) found from leptonic decays. Unlike decays into final states with a single neutrino, leptonic  $\tau$  decays can be contaminated by decays into final states containing one or more massive dark bosons. Since the central values of the measured leptonic decay rates are higher than the SM predictions, we study this scenario in the next subsection.

Recently it was pointed out that indeed the LHC might already be capable of probing mixing angles for mixing of sterile neutrinos with  $\tau$  neutrinos far more efficiently [72]. A reinterpretation of the existing LHC data would be highly interesting.

#### 4.7 BSM scenarios with other light invisible particles

Leptonic  $\tau$  decays  $\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell$  are interesting as there are two neutrinos produced in the final state: what is studied is  $\tau \rightarrow \ell + E_{\text{miss}}$  and the squared missing mass is much larger than zero, so that there can be contributions from  $\Gamma(\tau \rightarrow \ell X_{\text{dark}})$  or  $\Gamma(\tau \rightarrow \ell X_{\text{dark}} X_{\text{dark}})$  with massive dark particles  $X_{\text{dark}}$ . A natural candidate for an invisible boson emitted in a  $\tau$  decay is a majoron coupling to  $\nu_\tau$  [73–75].

We write  $\Gamma_{\text{NP}}$  for the dark contributions i.e.  $\Gamma(\tau \rightarrow \ell + E_{\text{miss}}) = \Gamma(\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell) + \Gamma_{\text{NP}}$ . This constrains the new physics contribution to

$$\Gamma_{\text{NP}} = \frac{\text{Br}^{\text{exp.}}(\tau \rightarrow \nu_\tau e \bar{\nu}_e)}{\text{Br}^{\text{exp.}}(\tau \rightarrow \pi \nu_\tau)} \Gamma_{\text{SM}}(\tau \rightarrow \pi \nu_\tau) - \Gamma_{\text{SM}}(\tau \rightarrow \nu_\tau e \bar{\nu}_e) = (3.92 \pm 3.81) \mu\text{eV} \quad (4.30)$$

$$\Gamma_{\text{NP}} = \frac{\text{Br}^{\text{exp.}}(\tau \rightarrow \nu_\tau \mu \bar{\nu}_\mu)}{\text{Br}^{\text{exp.}}(\tau \rightarrow \pi \nu_\tau)} \Gamma_{\text{SM}}(\tau \rightarrow \pi \nu_\tau) - \Gamma_{\text{SM}}(\tau \rightarrow \nu_\tau \mu \bar{\nu}_\mu) = (3.85 \pm 3.69) \mu\text{eV}, \quad (4.31)$$

where we set  $\cos \theta = 1$  because of the tight constraints on  $|\sin \theta|$  found above.

## 5 Conclusions

In this paper, we have calculated the inclusive semi-hadronic charged-current decay width of HSNs up to  $\mathcal{O}(\alpha_s^4)$  in the strong coupling constant by using the known results of the

electroweak gauge boson correlator which has been calculated up to the five-loop level. Decay rates into massless leptons can be trivially obtained from SM  $\tau$  decay rates, while  $N \rightarrow \tau + \text{hadrons}$  decays involve new phase-space integrals. We derive analytic expressions in terms polylogarithms (up to  $\mathcal{O}(\alpha_s^3)$ ) and a series representation (valid for all calculated orders) for these rates. We then calculated the perturbative series and analyzed its stability in order to determine the sterile neutrino mass range for which the perturbative calculation is valid and robust theory predictions are possible. Reliable predictions for the semi-hadronic width are found for  $m_N \geq 1.5 \text{ GeV}$  for  $\ell = e, \mu$  in the final state. For the final state with  $\ell = \tau$  we find stability for  $m_N \geq 3 \text{ GeV}$ . Our result can also be used for  $\tau \rightarrow N + \text{hadrons}$  in the mass range  $m_N \lesssim 600 \text{ MeV}$ . In  $\tau$  decays the perturbative series is only stable after including the  $\mathcal{O}(\alpha_s^4)$  corrections.

We have further studied the impact of an  $N$ - $\nu_\tau$  mixing angle  $\theta$  on the  $\tau$  lifetime. If  $N$  is light enough to be produced in  $\tau$  decay, the SM decay amplitude is modified by a factor of  $\cos \theta$  and there is an additional  $\tau \rightarrow N + W^*$  mode with amplitude proportional to  $\sin \theta$ . We find the parameter space of  $m_N$  and  $\theta$  constrained by the  $\tau$  lifetime for HSN masses  $m_N \lesssim 600 \text{ MeV}$ . Furthermore, we present expressions for the lepton energy spectrum in  $\tau \rightarrow \ell \nu_\ell N$  and the hadron invariant mass spectrum in  $\tau \rightarrow N + \text{hadrons}$ , which may be used to detect HSN effects in the future.

HSNs heavier than the  $\tau$  lepton may still leave imprints on observables through the cosine of the mixing angle. The lifetime analysis leads to  $|\sin \theta| \leq 11.8 \cdot 10^{-2}$  at  $1\sigma$  in this case. For the four decays  $\tau \rightarrow \nu_\tau \ell \bar{\nu}_\ell$  with  $\ell = e, \mu$  and  $\tau \rightarrow P \nu_\tau$  with  $P = \pi^-, K^-$  we have compared the measured widths with the theoretical predictions. Only the data on the semi-hadronic decays permit a non-zero mixing angle and we derive the constraints  $|\sin \theta| \leq 12.2 \cdot 10^{-2}$  at  $1\sigma$  from  $\tau \rightarrow \pi \nu_\tau$  and  $|\sin \theta| = \left(13.5_{-11.3}^{+5.4}\right) \cdot 10^{-2}$  at  $1\sigma$  from  $\tau \rightarrow K \nu_\tau$ , respectively. Combining both semi-hadronic decays leads to  $|\sin \theta| = \left(9.1_{-7.8}^{+3.7}\right) \cdot 10^{-2}$  at  $1\sigma$ . Our results are complementary to those from CMS, which constrain  $\theta$  for HSN masses above  $3 \text{ GeV}$ , and less model dependent than those of DELPHI below  $3 \text{ GeV}$ . In another application of our calculation we have derived relative branching fractions between certain exclusive decays and the inclusive semi-hadronic decay and determined the fraction of  $N \rightarrow \pi^+ \ell^-$  decays of all  $N \rightarrow \ell^- + \text{hadrons}$  decays.

Finally we have derived a bound on the  $\tau$  decay width into final states containing one or more new invisible particles, for example majorons. These decays contaminate  $\tau \rightarrow \ell \bar{\nu}_\ell \nu_\tau$ , whose decay rate is indeed a bit larger than the SM prediction. This finding calls for more precise experimental studies of the leptonic  $\tau$  decays.

## Acknowledgments

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## A QCD $\beta$ -function

We use the  $\beta$ -function with the following normalization

$$\frac{da_\mu}{d \ln \mu^2} = \beta(a_\mu) = -a_\mu^2 \left[ \beta_0 + \beta_1 a_\mu + \beta_2 a_\mu^2 + \dots \right], \quad (\text{A.1})$$

where  $a_\mu = a(\mu) = \alpha_s(\mu)/\pi$  and with the coefficients  $\beta_i$  [76, 77]:

$$\beta_0 = \frac{1}{4} \left( 11 - \frac{2}{3} n_f \right), \quad (\text{A.2})$$

$$\beta_1 = \frac{1}{16} \left( 102 - \frac{38}{3} n_f \right), \quad (\text{A.3})$$

$$\beta_2 = \frac{1}{64} \left( \frac{2857}{2} - \frac{5033}{18} n_f + \frac{325}{54} n_f^2 \right). \quad (\text{A.4})$$

Eq. (A.1) leads to

$$\begin{aligned} a(\mu) = a(\mu_0) &- a(\mu_0)^2 \beta_0 \ell_{\mu\mu_0} + a(\mu_0)^3 \left[ -\beta_1 \ell_{\mu\mu_0} + \beta_0^2 \ell_{\mu\mu_0}^2 \right] \\ &+ a(\mu_0)^4 \left[ -\beta_2 \ell_{\mu\mu_0} + \frac{5}{2} \beta_0 \beta_1 \ell_{\mu\mu_0}^2 - \beta_0^3 \ell_{\mu\mu_0}^3 \right] + \mathcal{O}(a(\mu_0)^5), \end{aligned} \quad (\text{A.5})$$

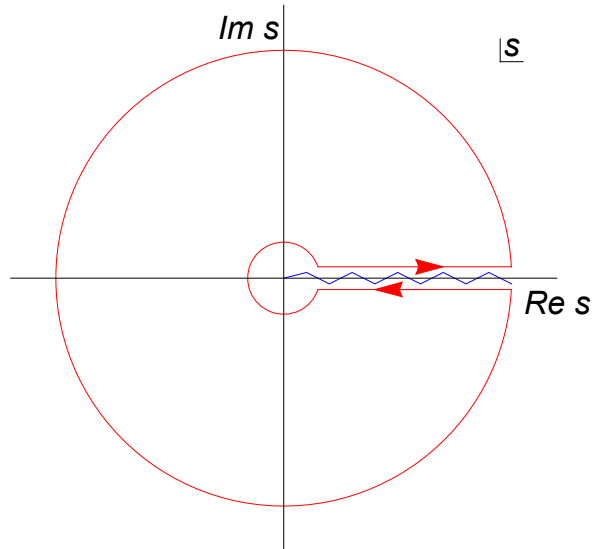
where  $\ell_{\mu\mu_0} = \ln \left( \frac{\mu^2}{\mu_0^2} \right)$ .

## B Contour integration

In the framework of perturbative calculations it is possible to calculate scattering or decay processes into quarks. However, in an experiment only hadrons are detectable and not quarks. This poses a problem as hadrons at different masses would appear as resonances in e.g. cross-sections. These resonances however are not accounted for by perturbative calculations using quarks. This conundrum was resolved by Poggio, Quinn and Weinberg [78]; they showed that a perturbative calculation smeared over a suitable energy range correctly averages over the resonances. This approach was then refined by showing that it is possible to transform the integration over the suitable energy range into an integration around a suitable complex contour [79, 80]. This was used for  $\tau$  decays [23, 81], the integral in eq. (3.2) can be rewritten in terms of a contour integration around a keyhole contour (see figure 15). Thus the decay width is proportional to the discontinuity over the branch cut along the positive real axis. In this way the integral may be rewritten in the following way (setting the lepton mass to zero for brevity)

$$\begin{aligned} \Gamma(N \rightarrow X\ell) &\sim 12\pi \int_0^{m_N^2} ds \left( 1 - \frac{s}{m_N^2} \right)^2 \left[ \left( 1 + 2\frac{s}{m_N^2} \right) \text{Im}\Pi^{(1+0)}(s) - 2\frac{s}{m_N^2} \text{Im}\Pi^{(0)}(s) \right] \\ &= 6\pi i \oint_{s=m_N^2} ds \left( 1 - \frac{s}{m_N^2} \right)^2 \left[ \left( 1 + 2\frac{s}{m_N^2} \right) \Pi^{(1+0)}(s) - 2\frac{s}{m_N^2} \Pi^{(0)}(s) \right], \end{aligned}$$

the integral only depends on  $s$  at the value of the mass of the decaying particle and so low energy effects are already accounted for in the decay width. Indeed in this way the inclusive decay width correctly averages over all possible resonances induced by the mesons.



**Figure 15.** Contour integration permitting to replace the integral over the resonance region by the integral over the large circle, which is far away from resonances. Thus, while the differential spectrum is not perturbatively calculable in that region, the integrated spectrum is found correctly in this way.

### C Easy-to-use formulae for phase-space integrals

For easier use we have expanded eqs. (3.7), (3.8) as

$$I_1^{\text{approx., T}} = \frac{2}{525}x_\ell^{14} + \frac{1}{60}x_\ell^{12} + \frac{2}{15}x_\ell^{10} - \frac{43}{24}x_\ell^8 + \frac{34}{3}x_\ell^6 + \left(\frac{15}{2} - 2\pi^2\right)x_\ell^4 + \frac{10}{3}x_\ell^2 - \frac{19}{24} + (x_\ell^4 - 8x_\ell^2 - 6)x_\ell^4 \log(x_\ell) + \mathcal{O}(x_\ell^{16}) \quad (\text{C.1})$$

$$I_2^{\text{approx., T}} = -\frac{787}{110250}x_\ell^{14} + \frac{1}{1800}x_\ell^{12} + \frac{217}{450}x_\ell^{10} + \frac{1}{6}(2\pi^2 - 5)x_\ell^8 + \frac{2}{3}(37 - 4\pi^2)x_\ell^6 - 2(-12\zeta(3) + 3 + \pi^2)x_\ell^4 - \frac{56}{9}x_\ell^2 + \frac{265}{144} - \frac{(8x_\ell^8 + 35x_\ell^6 + 280x_\ell^4 - 2100x_\ell^2 + 16800)x_\ell^6 \log(x_\ell)}{1050} + \mathcal{O}(x_\ell^{16}). \quad (\text{C.2})$$

Additionally we have also fitted eq. (3.7), (3.8) to polynomials (see figure 16)

$$I_1^{\text{approx., P}} = 1.27683x_\ell^9 - 1.11042x_\ell^8 - 10.3142x_\ell^7 + 21.4421x_\ell^6 - 9.05503x_\ell^5 - 6.19752x_\ell^4 + 1.5042x_\ell^3 + 3.24256x_\ell^2 + 0.00309017x_\ell - 0.7917 \quad (\text{C.3})$$

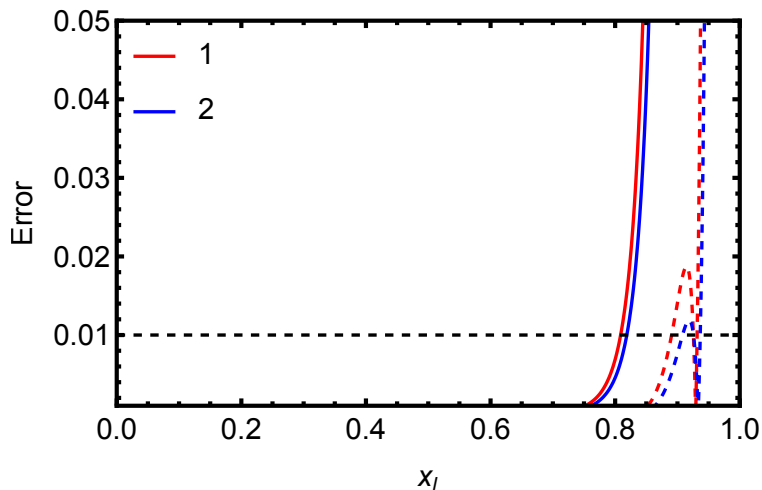
$$I_2^{\text{approx., P}} = -4.83878x_\ell^9 + 29.1877x_\ell^8 - 53.4361x_\ell^7 + 34.6139x_\ell^6 - 7.13307x_\ell^5 + 6.63816x_\ell^4 - 0.720101x_\ell^3 - 6.14877x_\ell^2 - 0.00323628x_\ell + 1.84032 \quad (\text{C.4})$$

### D Integrated spectrum

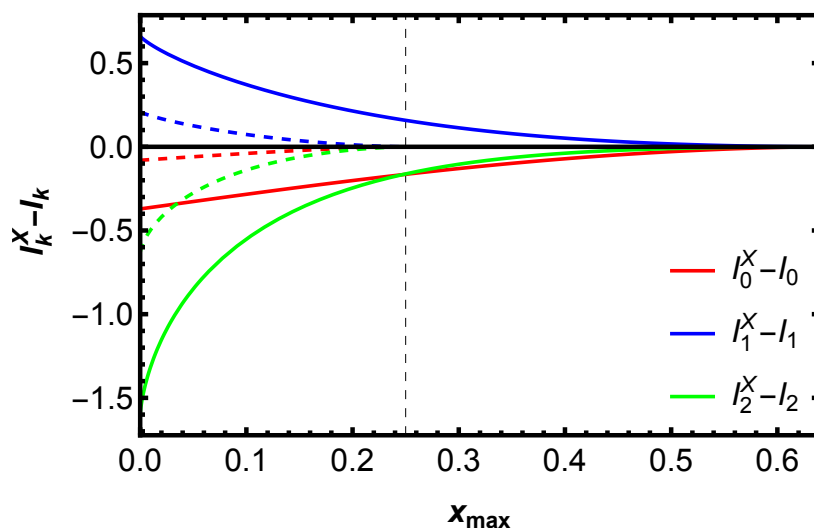
Here we present the analytical solutions of the integral

$$I_k(x_\ell^2, x_{\max}) = \int_0^{x_{\max}} dx \left( (1 + x_\ell^2 - x)(1 + 2x + x_\ell^2) - 4x_\ell^2 \right) \sqrt{\lambda(1, x, x_\ell^2)} \ln^k(x), \quad (\text{D.1})$$

for  $k < 3$ . In figure 17 the  $I_k(x_\ell^2, x_{\max}) - I_k(x_\ell^2, (1 - x_\ell)^2)$  are shown.



**Figure 16.** The absolute error  $|1 - I_i^{\text{approx., T/P}}/I_i|$  of the approximations in eqs. (C.1)–(C.4). The solid red and blue curves correspond to the Taylor expansion and the dashed blue and red curves are the polynomial fit. The dashed black line indicates the 1% level. For  $x_\ell \lesssim 0.8$  both approximations have uncertainties  $|1 - I_i^{\text{approx., T/P}}/I_i| < 1\%$ .



**Figure 17.** Here  $I_k^X - I_k$  is plotted for two different values of  $x_\ell = 0.2$  (closed line) and  $x_\ell = 0.5$  (dashed line). The vertically dashed line lies at the end of the spectrum of the dashed curves i.e.  $(1 - 0.5)^2 = 0.25$ . At the end of the spectrum the  $I_i^X - I_i = 0$ .

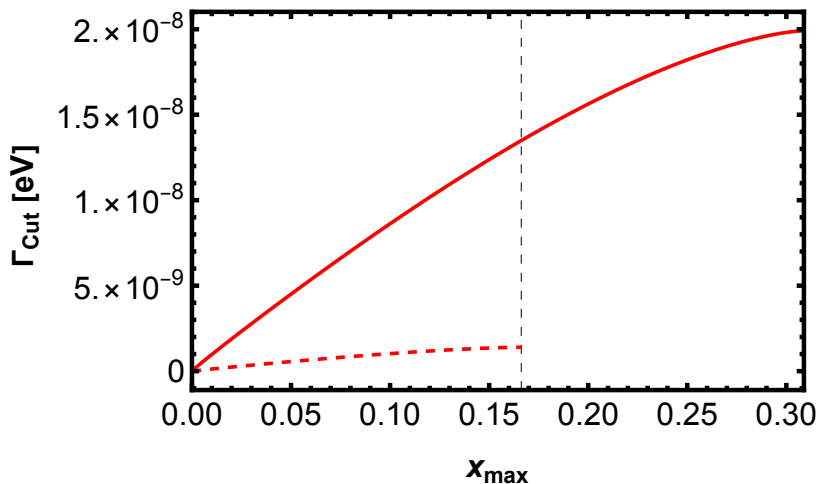
The  $I_k(x_\ell^2, x_{\text{max}}) - I_k(x_\ell^2, (1 - x_\ell)^2)$ , which we denote as  $I_k^X - I_k$  and use  $x_{\text{max}} = \hat{x}$  as a shorthand notation, take the following form

$$\begin{aligned}
 I_0^X - I_0 = \frac{1}{2} \left( - \left( \hat{x}^2 + 7 \right) x_\ell^2 - \left( \hat{x} + 7 \right) x_\ell^4 + x_\ell^6 \right. \\
 \left. + \left( \hat{x} - 1 \right)^2 \left( \hat{x} + 1 \right) \right) \left( \hat{x} - \left( x_\ell + 1 \right)^2 \right) X_r - 24 x_\ell^4 \text{ArcTanh}(X_r)
 \end{aligned}
 \tag{D.2}$$

$$\begin{aligned}
I_1^X - I_1 &= 12x_\ell^4 \left( \Delta\text{Li}_2 \left( -\frac{X_-}{2x_\ell} \right) - \Delta\text{Li}_2 \left( \frac{X_-}{2} \right) + \Delta\text{Li}_2 \left( \frac{1-X_r}{2} \right) \right) \\
&\quad + \frac{1}{24} X_r \left( \hat{x} - (x_\ell + 1)^2 \right) \\
&\quad \times \left( -3\hat{x}^3 + 12 \left( -(\hat{x}^2 + 7)x_\ell^2 - (\hat{x} + 7)x_\ell^4 + x_\ell^6 + (\hat{x} - 1)^2(\hat{x} + 1) \right) \ln(\hat{x}) \right. \\
&\quad \quad \left. + 5(\hat{x}^2 + \hat{x}) + (\hat{x}(5\hat{x} + 4) + 85)x_\ell^2 + 5(\hat{x} + 17)x_\ell^4 - 19x_\ell^6 - 19 \right) \\
&\quad + \left( -8x_\ell^2 - 6x_\ell^4 \left( 4\ln(\hat{x}) + 2\ln \left( 1 - X_r^2 \right) - 4\ln(2) + 2 \right) \right. \\
&\quad \quad \left. - 8x_\ell^6 + x_\ell^8 + 1 \right) \text{ArcTanh}(X_r) \\
&\quad - \left( 8x_\ell^2 + 24x_\ell^4 \ln(x_\ell) - 8x_\ell^6 + x_\ell^8 - 1 \right) \text{ArcTanh} \left( -X_r \frac{1+x_\ell}{1-x_\ell} \right)
\end{aligned} \tag{D.3}$$

$$\begin{aligned}
I_2^X - I_2 &= X_r \left[ \left( (\hat{x}^2 + 7)x_\ell^3 - (\hat{x} - 1)(\hat{x}^2 + 3)x_\ell^2 + (\hat{x} + 3)x_\ell^6 + (\hat{x} + 7)x_\ell^5 \right. \right. \\
&\quad \left. \left. + (7 - 3\hat{x})x_\ell^4 - (\hat{x} - 1)^2(\hat{x} + 1)x_\ell + \frac{1}{2}(\hat{x} - 1)^3(\hat{x} + 1) - \frac{x_\ell^8}{2} - x_\ell^7 \right) \ln^2(\hat{x}) \right. \\
&\quad \left. + \left( \frac{1}{12}(-3\hat{x}^4 + 8\hat{x}^3 - 24\hat{x} + 19) + \left( -2\hat{x} - \frac{11}{2} \right) x_\ell^6 - \frac{5}{6}(\hat{x} + 17)x_\ell^5 \right. \right. \\
&\quad \left. \left. + \frac{1}{6}(38\hat{x} - 85)x_\ell^4 + \frac{1}{6}(-\hat{x}(5\hat{x} + 4) - 85)x_\ell^3 \right. \right. \\
&\quad \left. \left. + \frac{1}{6}(\hat{x}(\hat{x}(4\hat{x} - 3) + 38) - 33)x_\ell^2 + \frac{1}{6}(\hat{x}(\hat{x}(3\hat{x} - 5) - 5) + 19)x_\ell \right. \right. \\
&\quad \left. \left. + \frac{19x_\ell^8}{12} + \frac{19x_\ell^7}{6} \right) \ln(\hat{x}) + \frac{1}{144} \left( 9\hat{x}^4 - 32\hat{x}^3 + 288\hat{x} - 265 \right) \right. \\
&\quad \left. + \left( 2\hat{x} + \frac{113}{24} \right) x_\ell^6 + \frac{23}{72}(\hat{x} + 41)x_\ell^5 + \frac{1}{72}(943 - 446\hat{x})x_\ell^4 \right. \\
&\quad \left. + \frac{1}{72}(\hat{x}(23\hat{x} + 28) + 943)x_\ell^3 + \frac{1}{72}(\hat{x}((9 - 16\hat{x})\hat{x} - 446) + 339)x_\ell^2 \right. \\
&\quad \left. + \frac{1}{72}(\hat{x}((23 - 9\hat{x})\hat{x} + 23) - 265)x_\ell - \frac{265x_\ell^8}{144} - \frac{265x_\ell^7}{72} \right] \\
&\quad + x_\ell^4 \left[ -16\mathcal{P}_{(0,7,5,0)} \text{ArcTanh}^2(X_r) \right. \\
&\quad \left. + 24\ln \left( 1 - X_r^2 \right) \left( \text{Li}_2 \left( -\frac{X_-x_\ell}{X_+} \right) - \text{Li}_2 \left( -\frac{X_-}{X_+x_\ell} \right) \right) \right. \\
&\quad \left. + 24\text{Li}_3 \left( -\frac{X_-}{X_+x_\ell} \right) - 24\text{Li}_3 \left( -\frac{X_-x_\ell}{X_+} \right) + 48\text{Li}_3 \left( -\frac{X_-}{X_r - 1} \right) \right. \\
&\quad \left. + 48\text{Li}_3 \left( \frac{X_r + 1}{(X_r - 1)x_\ell} + 1 \right) + 24\text{Li}_3 \left( -\frac{(X_r + 1)x_\ell}{X_r - 1} \right) \right. \\
&\quad \left. + 24\text{Li}_3 \left( \frac{x_\ell - X_r x_\ell}{X_r + 1} \right) + 30\text{ArcTanh}(X_r) - 48\zeta(3) \right] \\
&\quad + \frac{1}{6}x_\ell^8 \mathcal{P}_{(0,19,0,-3)} + x_\ell^6 \mathcal{P}_{(0,-\frac{46}{3},-2,4)} + x_\ell^2 \mathcal{P}_{(0,2,\frac{46}{3},-4)}
\end{aligned} \tag{D.4}$$

$$\begin{aligned}
 & + \ln(x_\ell) \left[ x_\ell^8 \mathcal{P}_{(0,1,3,0)} - 8x_\ell^6 \mathcal{P}_{(0,1,3,0)} - 12x_\ell^4 \mathcal{P}_{(\pi^2,0,2,-1)} \right. \\
 & + \ln(\hat{x}) \left( -12x_\ell^4 \left( 4\text{ArcTanh}(X_r) + 1 \right) + 2x_\ell^8 - 16x_\ell^6 \right) \\
 & - 12x_\ell^4 \ln^2(\hat{x}) + 48x_\ell^4 \ln(X_-) \text{ArcTanh}(X_r) \\
 & + \ln(X_+) \left( -24x_\ell^4 \left( 2\ln \left( 1 - X_r^2 \right) + 2\text{ArcTanh}(X_r) - 1 \right) - 4x_\ell^8 + 32x_\ell^6 \right) \\
 & + 48x_\ell^4 \ln^2(X_+) + 96x_\ell^4 \text{ArcTanh}(X_r) \ln(1 - x_\ell) \\
 & \left. - 16x_\ell^2 \text{ArcTanh}(X_r) + 2\text{ArcTanh}(X_r) \right] \\
 & + \ln^2(x_\ell) \left( 12x_\ell^4 (2\ln(X_+) - \mathcal{P}_{(0,2,0,0)}) + \frac{p(x_\ell)}{2} \right) - 12x_\ell^4 \ln^3(x_\ell) \\
 & + \ln(\hat{x}) \left[ x_\ell^8 \mathcal{P}_{(\frac{19}{12}, -2, 0, 0)} + x_\ell^6 \mathcal{P}_{(-\frac{26}{3}, 16, 0, 0)} + x_\ell^2 \mathcal{P}_{(\frac{26}{3}, 0, -16, 0)} + \mathcal{P}_{(-\frac{19}{12}, 0, 2, 0)} \right. \\
 & - 24x_\ell^4 \left( \text{Li}_2 \left( -\frac{X_-}{X_+ x_\ell} \right) - \text{Li}_2 \left( -\frac{X_- x_\ell}{X_+} \right) \right. \\
 & \left. + 2\text{ArcTanh}(X_r)^2 + \text{ArcTanh}(X_r) \right) - 2p(x_\ell) (\ln(X_+) - \ln(x_\ell + 1)) \left. \right] \\
 & + \frac{1}{2} \ln^2(\hat{x}) \left( p(x_\ell) - 48x_\ell^4 \text{ArcTanh}(X_r) \right) + 2\ln^2(X_+) p(x_\ell) \\
 & + \ln(X_+) \left[ \left( x_\ell^8 - 1 \right) \mathcal{P}_{(-\frac{19}{6}, 2, 2, 0)} - \left( x_\ell^4 - 1 \right) x_\ell^2 \mathcal{P}_{(-\frac{52}{3}, 16, 16, 0)} \right. \\
 & + 48x_\ell^4 \left( \text{Li}_2 \left( -\frac{X_-}{X_+ x_\ell} \right) - \text{Li}_2 \left( -\frac{X_- x_\ell}{X_+} \right) \right) \\
 & \left. - 4p(x_\ell) \ln(x_\ell + 1) + 144x_\ell^4 \text{ArcTanh}^2(X_r) \right] \\
 & + 48x_\ell^4 \text{ArcTanh}^2(X_r) \ln(X_-) \\
 & + \ln(1 - x_\ell) \left[ 8x_\ell^4 \left( 6\Delta \text{Li}_2 \left( \frac{X_-}{1 - X_r} \right) + 6\Delta \text{Li}_2 \left( -\frac{X_-}{2x_\ell} \right) \right. \right. \\
 & - 6\Delta \text{Li}_2 \left( -\frac{1}{2} (X_r - 1)(x_\ell + 1) \right) \\
 & - 12\text{Li}_2 \left( \frac{X_r + 1}{2} \right) + 6\ln(X_-) \left( \ln(X_r + 1) + \ln \left( \frac{x_\ell + 1}{2} \right) \right) \\
 & - 6\ln(X_+) \left( \ln \left( \frac{1 - X_r}{2} \right) + \ln(x_\ell + 1) \right) + 6\ln(2) \ln \left( 1 - X_r^2 \right) \\
 & \left. - 6\ln^2(X_r + 1) + 12\text{ArcTanh}^2(X_r) + \pi^2 - 6\ln^2(2) \right) \\
 & \left. + 2p(x_\ell) \left( \ln(\hat{x}) - 2\ln(X_+) + \ln \left( 1 - X_r^2 \right) \right) \right] \\
 & + \frac{1}{6} \left( \mathcal{P}_{(0,0,-19,3)} + p(x_\ell) \left( 6\text{Li}_2 \left( -\frac{X_-}{X_+ x_\ell} \right) + 6\text{Li}_2 \left( -\frac{X_- x_\ell}{X_+} \right) \right) \right)
 \end{aligned}$$



**Figure 18.** Integrated spectrum  $\Gamma_N^X(x_{\max})$ . The width is plotted with a mixing angle of  $V_{N\ell} = 10^{-3}$  and the lepton is a tau  $\ell = \tau$ . In the full curve the mass of the HSN is  $m_N = 4$  GeV whereas in the dashed curve it is  $m_N = 3$  GeV; the dashed vertical line marks the endpoint of the integrated spectrum for the dashed red curve.

$$\begin{aligned}
 &+ 12 \ln \left( (1 - X_r^2) \ln(x_\ell + 1) + \pi^2 \right) \\
 &+ \text{Li}_2 \left( \frac{X_-}{1 - X_r} \right) q_-(x_\ell) - \text{Li}_2 \left( \frac{X_+}{1 + X_r} \right) q_+(x_\ell)
 \end{aligned}$$

where the  $I_i$  are given in eq. (3.6)–(3.8) and  $X_r = \sqrt{(\hat{x} - (1 - x_\ell)^2)/(\hat{x} - (1 + x_\ell)^2)}$ . We defined the following abbreviations

$$X_\pm = (1 - x_\ell) \pm X_r(1 + x_\ell), \tag{D.5}$$

$$\mathcal{P}_{a,b,c,d} = a + b \ln(1 - X_r) + c \ln(1 + X_r) + d \ln^2(1 - X_r^2) \tag{D.6}$$

$$\Delta \text{Li}_n(f(X_r)) = \text{Li}_n(f(X_r)) - \text{Li}_n(f(-X_r)) \tag{D.7}$$

$$p(x_\ell) = 1 - 8x_\ell^2 + 8x_\ell^6 - x_\ell^8 \tag{D.8}$$

$$q_\pm(x_\ell) = 1 - 8x_\ell^2 - 12x_\ell^4 - 8x_\ell^6 + x_\ell^8 - 24x_\ell^4 \ln(\hat{x}) \pm 24x_\ell^4 \ln \left( \frac{\hat{x}(1 - X_r^2)}{X_r^2} \right) \tag{D.9}$$

where  $f(X_r)$  is an arbitrary function depending on  $X_r$ .

Using eq. (D.1) we plot the integrated decay width  $\Gamma_{\text{cut}}$  defined as

$$\begin{aligned}
 \Gamma_{\text{cut}}(N \rightarrow \ell X) = N_c \frac{G_F^2 m_N^5 |V_{N\ell}|^2}{192\pi^3} & \tag{D.10} \\
 & \times 2(|V_{ud}|^2 + |V_{us}|^2) \left[ I_0 c_{0,1} + a_{m_N} c_{1,1} I_0 \right. \\
 & + a_{m_N}^2 (c_{2,1} I_0 + 2c_{2,2} I_1) \\
 & + a_{m_N}^3 (c_{3,1} I_0 + 2c_{3,2} I_1 - (\pi^2 I_0 - 3I_2) c_{3,3}) \\
 & \left. + a_{m_N}^4 (c_{4,1} I_0^Q + 2c_{4,2} I_1 - (\pi^2 I_0 - 3I_2) c_{4,3} - (4\pi^2 I_1 - 4I_3) c_{4,4}) \right]
 \end{aligned}$$

where  $I_k = I_k(x_\ell^2, x_{\max})$ , see figure 18.

**Data Availability Statement.** This article has no associated data or the data will not be deposited.

**Code Availability Statement.** This article has no associated code or the code will not be deposited.

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