### Optimized and harmonized FTIR retrieval strategy for $CH_4$ and $N_2O$ columns and profiles

Sussmann, R., Forster, F., Borsdorff, T., De Mazière, M., Vigouroux, C., Blumenstock, T., Duchatelet, P., Hannigan, J., Hase, F., Jones, N., Klyft, J., Mahieu, E., Mellqvist, J., Notholt, J., Petersen, K., Strong, K., Taylor, J.

This talk presents the outcome of the Work Package "Optimized retrieval strategy for  $CH_4$  and  $N_2O$ " lead by IMK-IFU within the EC-HYMN project (update of HYMN deliverable document D4.4, dated Oct 2009, http://...)

The HYMN-applications are:

- revised historical time series of CH4 & N2O
- satellite validation
- model validation

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### 13 station participated to the CH4 & N2O harmonization effort

station	latitude	longitude	station altitude	number of columns in 2003/2004	tropopause height
Spitzbergen	78.92 °N	11.92 °E	20 m	113	8.95 km
Thule	76.53 °N	68.74 °W	225 m	177	8.51 km
Kiruna	67.84 °N	20.41 °E	419 m	338	9.62 km
Harestua	60.22 °N	10.75 °E	596 m	1234	10.20 km
Bremen	53.11 °N	8.85 °E	29 m	179	10.74 km
Zugspitze	47.42 °N	10.98 °E	2964 m	999	11.25 km
Garmisch	47.48 °N	11.06 °E	745 m	498	11.25 km
Jungfraujoch	46.55 °N	7.99 °E	3580 m	702	11.38 km
Toronto	43.66 °N	79.40 °W	174 m	185	13.25 km
Izaña	28.30 °N	16.48 °W	2367 m	207	14.44 km
Paramaribo	5.81 °N	55.21 °W	7 m	64	16.36 km
St-Denis	20.90 °S	55.48 °E	50 m	141	15.66 km
Wollongong	34.41 °S	150.88 °E	40 m	633	12.53 km

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### **Retrieval homogenization CH4: status April 2008**



harmonized treatment of interfering species?

identical spectroscopic line list for all partners ?

- Image: one common source of pT-input profiles ?
- Image: one consistent set of a priori profiles?
- In the set of regularization matrices and altitude grids?

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### Retrieval homogenization CH4: status May 2008

### one common micro-window set: <u>"UFTIR set"</u>



# CO2,HDO, CO2-NO2H2O, HDOHDOCO22NO2NO2

 $\Rightarrow$  one set of binput-files distributed

identical spectroscopic line list for all partners ?

- In the second second
- Image: one consistent set of a priori profiles?

In the set of regularization matrices and altitude grids?

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In the set of the s

identical spectroscopic line list for all:

- new CH4 lab measurements at Bremen in collab. with IMK-ASF
- Frank Hase retrieved a new set of line-parameters

⇒ one set of cfgl's distributed: HITRAN 04 incl. Hase update

common source of pT-input profiles

- Image: Image:
- In the set of regularization matrices and altitude grids?

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one common micro-window set: <u>"UFTIR set"</u>
 harmonized treatment of interfering species: <u>same binput-file</u>
 identical spectroscopy for all: <u>same cfgls (HIT04 & Hase update)</u>

Second control cont

one consistent set of a priori profiles?
 one consistent set of regularization matrices and altitude grids?

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### **Retrieval homogenization CH4: status June 2008**



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one common micro-window set: <u>"UFTIR set"</u>
 harmonized treatment of interfering species: <u>same binput-file</u>
 identical spectroscopy for all: <u>same cfgls (HIT04 & Hase update)</u>
 common source of pT-input profiles: <u>NCEP</u>
 one consistent set of a priori profiles: <u>Toon with Meier correction</u>

In the set of regularization matrices and altitude grids?

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CH4 regularization. Issue: for direct quantitative intecomparison the layering would have to be the same!



$$\mathbf{R} = \mathbf{S}_a^{-1} = \alpha \mathbf{L}^T \mathbf{L} \quad \in \mathfrak{R}^{n \times n}$$

with the Tikhonov regularization operator L and the regularization strength  $\alpha$ .

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### What is Tikhonov Regularization: you already know Tikhonov L<sub>0</sub>!

$$\mathbf{R} = \alpha \mathbf{L}_0^T \mathbf{L}_0 = \alpha \times \begin{pmatrix} 1 & 0 & 0 & \cdots & 0 \\ 0 & 1 & \ddots & \ddots & \vdots \\ 0 & \ddots & 1 & \ddots & 0 \\ \vdots & \ddots & \ddots & 1 & 0 \\ 0 & \cdots & 0 & 0 & 1 \end{pmatrix} \in \Re^{n \times n}$$
  
In this case  $\mathbf{R} = \mathbf{S}_a$ 

i.e., <u>Tikhonov  $L_0$  is mathematically identical to a diagonal</u>, <u>altitude constant  $S_a$ </u>

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### What is Tikhonov $L_1$ : it constrains the derivative of the profile with resp. to alt.

$$\mathbf{R} = \alpha \mathbf{L}_{1}^{T} \mathbf{L}_{1} = \alpha \times \begin{pmatrix} 1 & -1 & 0 & \cdots & 0 \\ -1 & 2 & \ddots & \ddots & 1 \\ 0 & \ddots & \ddots & \ddots & 0 \\ \vdots & \ddots & \ddots & 2 & -1 \\ 0 & \cdots & 0 & -1 & 1 \end{pmatrix} \in \Re^{n \times n}$$
(1)  
regularization strength  $\alpha$ 

with regularization strength  $\alpha$ .

Case  $\alpha \to \infty$  any change in profile shape totally is forbidden, any altitude constant change fully allowed:  $dofs \to 1$ .

Case  $\alpha \rightarrow 0$  is a totally unconstrained profile retrieval with  $dofs \rightarrow n =$  number of model layers (oscillations)

## L<sub>1</sub> constrains only the profile shape and any altitude constant change is fully allowed

(2)

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Possible profile units:

- absolute VMR (default in PROFFIT)
- per centage VMR (default in SFIT 2)
- log VMR (option in PROFFIT and, since recently, option in SFIT 2)
- number density
- partial column

Examples for L<sub>1</sub>:

•

- regularization in units of absolute VMR favors shifting of VMR profiles to higher/lower VMR's
- regularization in units of %-VMR favors scaling of VMR profiles

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IMK-IFU pledger: Tikhonov L<sub>1</sub> on the %-VMR-scale is probably the most robust standard procedure very well applicable to all species

Arguments

• VMR profile scaling (L<sub>1</sub> with  $\alpha \rightarrow \infty$ , *dofs*=1) is one of the best-tested retrieval approaches (SFIT 1, WFM-DOAS, ...)

• VMR profile scaling is more realistic than VMR profile shifting (i.e.,  $L_1$  on absolute VMR scale) (avoids neg. VMR's)

• starting from high  $\alpha$ ,  $\alpha$  can be empirically reduced (*dofs* increased) to allow for some additional flexibility in the profile to account for true profile variations <u>and/or cloud impact on the spectra</u>  $\Rightarrow$  to get even more precise columns than by profile scaling

• whatever the dofs ( $\alpha$ ) is, there is *per definitionem* never any under-estimation of true profile-scaling-type variability using L<sub>1</sub>

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### Tikhonov $L_1$ : how find optimum regularization strenght $\alpha$ ?



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### $L_0$ versus $L_1$ : diurnal variation as a function of $\alpha$ (dofs) - $L_1$ more robust



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### **Tikhonov vs UFTIR: main facts**

	Tikhonov	UFTIR		
Mean DOFS	2.73	3.03		
Residuals	No significant difference			
CH <sub>4</sub> Tot Col	No significant difference			
CH <sub>4</sub> VMR	No oscillations	Still oscillates		
CH₄ Part Col	~10% bias			

Values obtained for one year of observations @ Jungfraujoch (~300 spectra) Consistent set of regularization matrices and altitude grids: status Oct 2008

Had to decide at Garmisch HYMN meeting in Oct 2009 how to proceed:

a) all use same Tikhonov-Regularization ("%-VMR-L1") & äquidistant altitude grid (IMK-IFU)

versus

b) freedom in regularization matrices & grids ("UFTIR strategy", BIRA)

<u>Decided for HYMN (CH4 & N2O):</u> obligatory use of Tikhonov ("%-VMR-L<sub>1</sub>") but freedom in regularization strength and altitude grid

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So, finally, a word about altitude grids has become necessary...: Status Nov 2008

Tikhonov matrices shown in this talk were for altitude-constant grid only.

"Freedom" in layering means:

Image: we had to recalculate the Tikhonov matrices for each group individually



### In the second second

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So, finally, a word about altitude grids has become necessary...: Status Nov 2008

♦ it is possible with sufficient accuracy:

Zugspitze 2004 columns differ only by 0.02 % between an equidistant and an exponential layering – after recalculation of the L1 matrix



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2481.3000	2482.6000	4	0.00	0	0	IP
N20	0 9500					
CH4	0 9500					
CO2	0 9500					
H2O	0 9500					
2526.4000	2528.2000	3	0.00	0	0	IP
N20	0 9500					
C02	0 9500					
CH4	0 9500					
2537.8500	2538.8000	3	0.00	0	0	IP
N20	0 9500					
HDO	0 9500					
CH4	0 9500					
2540.1000	2540.6000	1	0.00	0	0	IP
N20	0 9500					

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### **Retrieval homogenization N2O: prior**



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#### Description of strategy:

Sussmann et al., D4.4 document on HYMN retrieval strategy: http://... type "HYMN" to google

Sussmann, R., F. Forster, T. Borsdorff, B. Dils, M. De Mazière, C. Vigouroux, T. Blumenstock, M. Buchwitz, J.P. Burrows, P. Duchatelet, C. Frankenberg, J. Hannigan, F. Hase, N. Jones, J.Klyft, E. Mahieu, J. Mellqvist, J. Notholt, K. Petersen, O. Schneising, K. Strong, J. Taylor: A novel Tikhonovbased approach for harmonized high-accuracy retrieval of methane columns and profiles from NDACC FTIR network measurements. Application to global validation of ENVISAT/SCIAMACHY biases, Geophysical Research Abstracts, Vol. 11, EGU2009-7869-2, 2009, http://meetingorganizer.copernicus.org/EGU2009/EGU2009-7869-2.pdf, EGU General Assembly 2009 (Talk).

#### Why we did that effort:

Sussmann., R., Forster, F., Borsdorff, T., Dils, B., De Mazière, M., Vigouroux, C., Blumenstock, T., Buchwitz, M., Burrows, J.P., Duchatelet, P., Frankenberg, C., Hannigan, J., Hase, F., Jones, N., Klyft, J., Mahieu, E., Mellqvist, J., Notholt, J., Petersen, K., Schneising, O., Strong, K., Taylor, J.: Satellite validation of column-averaged methane on global scale: ground-based data from 13 FTIR stations versus last generation ENVISAT/SCHIAMACHY retrievals, ACP or AMT in preparation 2009.

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120	N2O regularization. Issue: for direct quantitative intercomparison the layering would have to be the same Appendix viewgra					
120	Izana/Kiruna: Tikhonov L0+L1 (->which relation?, tuned versus what); layer-stee independent. dofs $\approx 3.5$	eps increasing w	ith altitude, pe	rcentage variability	altitude	
100	Image: Bremen       4 km off diag, constant layering, dofs?         ISSJ       no off diag, based on works performed by Arndt Meier, layer-s         Reunion       5 km off diag, layer-steps increasing with altitude, dofs ≈3.2	teps increasing v	with altitude, d	ofs ≈ 3.65		
80	<ul> <li>Harestua 4 km off diag, layer-steps increasing with altitude, dofs?</li> <li>Zugspitze: Tikhonov first derivative, percentage variability altitude independent profile oszillations, dofs ≈ 3</li> </ul>	ent, exponential (	66-layering, op	timized diurnal var	. and	
60						
40						
20	<pre></pre>					
0		1.00 nisch-Par	1.20 tenkirche	1.40 • <b>n</b>	1.60	
T	Rali per centage VMR		<sup>r</sup> CH <sub>4</sub> and $N_2$ O			