

Determination and correction of pointing errors in solar FTIR spectrometry

A. Reichert, P. Hausmann, and R. Sussmann





Mispointing determination: impact of mispointing below 10 km above 20 km 1/cos(z) approximation relative air mass change sza change 0.1 oer degree min. desired tracking range 0.0 60 from: Hase, 2000 apparent sza

- Inaccurate alignment of FTIR line of sight with the solar disc center causes errors in trace gas retrievals
- Errors are approximately proportional to mispointing in zenith direction
- These errors may exceed accuracy requirements

 → error diagnosis and correction necessary
- 2 | A. Reichert, P. Hausmann, R. Sussmann, KIT/IMK-IFU, Garmisch



Mispointing determination:



solar line shift measurements



- Neglecting differential rotation, measured shift is proportional to radial mispointing: $s[^{\circ}] = \frac{\Delta v/v \cdot 1^{\circ}}{3.9 \cdot 10^{-7} \cdot 60}$ (Gisi et al. 2011)
- **But:** No mispointing error correction directly possible from shift measurements since mispointing parallel to solar axis can not be determined
- 3 | A. Reichert, P. Hausmann, R. Sussmann, KIT/IMK-IFU, Garmisch



Mispointing determination:

basic idea of our method





- Multiple solar shift measurements at different solar axis orientations enable constraining both components of mispointing vector
- Each shift measurement constrains mispointing to lie on straight line, mispointing vector defined by intersection of lines with different axis orientations
- Underlying assumption: Mispointing does not vary significantly between measurements





orientation of solar rotation axis



- Orientation of solar rotation axis: apparent angle between zenith direction and solar axis
- Orientation has annual/daily cycle, sum of 3 contributions:
 1) zenith → Earth axis, 2) Earth axis → ecliptic axis,
 3) ecliptic axis → solar axis
- 5 | A. Reichert, P. Hausmann, R. Sussmann, KIT/IMK-IFU, Garmisch



transfer from sky to spectrometer coordinates



- Typical source of mispointing: non-ideal geometry of optical elements in the spectrometer
- → Approximately constant mispointing for spectrometer coordinates, not for sky coordinates
- Change of orientation of solar image due to tracker optics is calculated before mispointing calculation





differential solar rotation





- Angular velocity of solar rotation depends on solar latitude: differential rotation
- → solar shift measurement does not constrain mispointing on straight line parallel to axis but on line with constant ω(φ) (blue)
- For mispointing determination scheme: linear approximation (red)



Implementation: mispointing calculation 2 enith



- Combine multiple measurements in time bin + calculate mean mispointing for bin → reduced mispointing error
- Mean mispointing = weighted mean of intersection coordinates
- Intersection error: solar shift error from difference of shift fit in adjacent filters + difference of axis orientation for measurement pair
- 8 | A. Reichert, P. Hausmann, R. Sussmann, KIT/IMK-IFU, Garmisch





binning of measurement time series



- Tradeoff between constant mispointing within bin vs. mispointing uncertainty
- Selection of binsize for Zugspitze time series: minimum of XCH₄ diurnal variability
- 9 | A. Reichert, P. Hausmann, R. Sussmann, KIT/IMK-IFU, Garmisch



Results:

Zugspitze mispointing results



Karlsruhe Institute of Technology

← Mispointing in spectrometer coordinates, *x* and *y* component

← Mispointing in sky coordinates, zenith component





mispointing correction of retrieval results

- a posteriori method: calculate corrected airmass, multiply retrieved column by airmass_{org} / airmass_{corr}
- a priori method: repeat retrieval using corrected sza
- a priori method: 5 % bias in mispointing/ ~0.02 % bias in XCH₄ → fair approximation for column retrievals
- For profile retrievals: a priori correction preferable





Results:



corrected methane time series/trend





Summary and Conclusions

- Method to determine mispointing in solar FTIR measurements, enables correction of mispointing-induced errors in retrieval results
- Method relies on measurement of solar line shifts from multiple spectra at different orientations of the solar rotation axis
- Constraining zenith component of mispointing enables correction of retrieval results





Summary and Conclusions

- Mispointing correction for Zugspitze XCH₄ time series: effects of non-optimum tracking period are removed, trend consistency with nearby Garmisch site restored
- Benefit of presented method for refining other existing solar FTIR measurement time series
- More details: see Reichert et al., 2015, accepted for AMTD







Funding by KIT/IMK-IFU, the Bavarian State Ministry of the Environment and Consumer Protection as well as by the Deutsche Bundesstiftung Umwelt (DBU) is gratefully acknowledged. We thank Frank Hase for helpful discussions and comments.



