# Annual Report on Nuclear Physics Activities 

July 1, 1982 - June 30, 1983

Editors:
H. Gemmeke, F. Käppeler, Ch. Weddigen

Institut für Kernphysik

# KERNFORSCHUNGSZENTRUM KARLSRUHE 

Institut für Kernphysik

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on

NUCLEAR PHYSICS ACTIVITIES

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This report surveys the activities in fundamental research from July 1, 1982 to June 30, 1983 at the three institutes of the KfK which are concerned with nuclear physics. The research program comprises laser spectroscopy, nuclear reactions with light ions, neutron physics, neutrino physics and physics at medium and higher energies.

ZUSAMMENFASSUNG

Der vorliegende Bericht gibt einen Überblick über die Arbeiten an den drei Kernphysikalischen Instituten des Kernforschungszentrums Karlsruhe im Zeitraum vom 1. Juli 1982 bis zum 30. Juni 1983. Das Forschungsprogramm umfaßt die Gebiete Laserspektroskopie, Kernreaktionen mit leichten Ionen, Neutronenphysik, Neutrino-Physik, sowie Mittel- und Hochenergiephysik.

## PREFACE

We have the pleasure to present the third report on nuclear physics activities at the Nuclear Research Center Karlsruhe.

In the past years, the collaboration between the three institutes active in nuclear physics, IK I, IK II and IAK II, has been intensified by joint planning, joint seminars and documented by a joint report. This tendency has now been formalized by a decision of our Supervisory Board to merge all groups active in nuclear physics into one 'Institute for Nuclear Physics' The former 'Institute for Applied Nuclear Physics II' is now called 'Institute for Nuclear Physics III'. In this way, the present report summarizes the work of the 'Institute for Nuclear Physics'.

Three groups of the Institute for Nuclear Physics I are engaged in work in various fields of nuclear physics and particle physics:

Fast Neutron Physics: Scattering experiments on very light nuclei are carried out using the polarized neutron beam of the Karlsruhe Cyclotron (POLKA). The main goal is to determine precise phase shifts from experiments with polarized neutrons on unpolarized and also on polarized protons. Moreover, the internal structure and dynamics of the nuclei up to the $A=5$ system are to be studied. Last year, the POLKA facility was used successfully in a number of experiments. The large-volume polarized proton target developed by the 'Experimental Methods' Group has meanwhile been put into operation successfully and has been in use for scattering experiments with the polarized neutron beam since May, 1983.

High Energy Physics: This Group runs the CELLO detector system at the $\mathrm{e}^{+} \mathrm{e}^{-}$ storage ring, PETRA, in Hamburg within the framework of an international collaboration. The detector serves for experiments to study $\mathrm{e}^{+} \mathrm{e}^{-}$collisions at the highest energies at present attainable. CELLO, with its modern liquid-argon calorimeter, lends itself particularly well to studies of the electromagnetic component in these reactions. This allows, e.g., precise studies of quantum electrodynamics, detailed studies of quark and gluon jets performed and the search for new quarks. The detector has meanwhile been upgraded and moved back into the beam since the summer of 1982. All its components are working satisfactorily
and are strongly involved in the search for a new quark during the continuous energy increases in PETRA.

Neutrino Physics: The newly founded Working Group is concerned with neutrino physics in the energy range between approximately 10 and 50 MeV , at the Spallation Neutron Source under construction at the RutherfordAppleton Laboratory (RAL) in England. This is a new field of work involving fundamental questions in the fields of elementary particle physics, nuclear physics and astrophysics.

The project was proposed by KfK. In the meantime, a bilateral agreement has been signed between $K f K$ and RAL. Several smaller working groups of the University of Oxford (Prof.Dr.N.E.Booth), Queen Mary College of London (Prof.Dr.J.A.Edgington), University of Erlangen (Prof.Dr.E.Finckh) and University of Tübingen (Prof.Dr.A.Faessler) have meanwhile joined the project. The detector system is being developed and built at KfK. It will be installed in a massive blockhouse of iron at the Rutherford SNS by the end of 1985.

The Institute for Nuclear Physics II is mostly working on medium energy physics at CERN and SIN.

One Group, continuing a long tradition in the field of exotic atoms at CERN, has concentrated its activities on the LEAR Project (Low-Energy Antiproton Ring). The project, the realization of which was backed very strongly by Karlsruhe, promises to offer unique possibilities for work with slow antiprotons. Spectra were obtained in a test run of 20 mi nutes which formerly took several months of measuring time. This also proved the successful conversion of the equipment to high counting rate capabilities. A Karlsruhe proposal makes use of the idea of the cyclotron trap proposed by Mr.Simons. This unit was tested with pions and myons at SIN, and the computed predictions were confirmed. It will now be brought to LEAR. Karlsruhe is involved personally and financially also in a technical upgrading of LEAR, namely the use of electron cooling。

The experiments at SIN focus on problems of pion interaction (scattering and absorption) with simple systems consisting of few nucleons. Theoretical assumptions, especially those about the existence of dibaryon states, are verified on the basis of additional information that can be obtained by using polarized targets.

Also the low-energy spectrometer designed by Kar1sruhe will mainly be used in studies of very simple systems at energies close to the pion threshold. The Coulomb nuclear interferences have been studied in preliminary experiments on carbon. The results of these experiments demonstrate that the proposed measuring technique allows to study pionhydrogen scattering in the region of Coulomb interference. This reaction is of great theoretical significance.

Research of the Institute for Nuclear Physics III takes place in the following fields:

Nuclear astrophysics: Capture cross sections of fast neutrons in the keV to MeV range are measured in order to understand in detail the build-up of heavy elements in stars. These measurements are supplemented by studies of nuclear spectroscopy on nuclei of specific importance to stellar neutron capture reactions.

Nuclear reactions: Alpha particle and ${ }^{6}$ Li beams from the Karlsruhe Isochronous Cyclotron are used for studying nuclear reactions at $26 \mathrm{MeV} /$ nucleon. The emphasis is at present on the break-up of ${ }^{6}$ Li in nuclear collision with the aim to study the momentum distribution of its constituents. A medium size magnetic spectrometer which will allow measurements down to very small reaction angles is being put into operation.

Laser spectroscopy: This technique is applied to sub-ng amounts of radioactive atoms in order to determine hyperfine structure and isotopic shifts of atomic transitions. The results yield information on nuclear moments and on the change of nuclear charge radii due to varying neutron number. Previous measurements on a long series of lead isotopes have been extended, and studies of tin isotopes have started.

Applied gamma-ray spectroscopy: Here instruments are developed to determine concentration and isotopic composition of fissile material. The instruments make use either of the intrinsic radioactivity or of X-ray absorption and fluorescence. Their main applications are in the safeguards of nuclear fuel and in process control during fabrication and reprocessing.

The Institute for Nuclear Physics III is also responsible for the operation of the Karlsruhe Isochronous Cyclotron. The cyclotron laboratory has been considerably extended during the last three years by the installation of a second cyclotron in an annex to the cyclotron building. This accelerator will go into operation in late 1983 (after a consderable delay caused by economic difficulties of the manufacturer) and will be operated on a commercial basis for producing isotopes for nuclear medicine and for irradiating machine parts for wear studies in industry.
chino
(A. Citron)
y. lent
(G. Schatz)

(B. Zeitnitz)

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1.1 NUCLEAR ASTROPHYSICS
1.1.1 NEUTRON CAPTURE IN s-WAVE RESONANCES OF ${ }^{56} \mathrm{Fe},{ }^{58} \mathrm{Ni}$, and ${ }^{60} \mathrm{Ni}$
K. Wisshak, F. Käppeler, G. Reffo ${ }^{+}$, F. Fabbri ${ }^{+}$, KfK-report 3516 (1983) and Nuc1. Sci. Eng. (submitted)

The exact determination of the capture widths of broad s-wave resonances in structural materials is an important problem because of two reasons: (i) By their large capture area, these resonances contribute significantly to the average capture cross section. (ii) In previous measurements their large ratio $\Gamma_{n} / \Gamma_{\gamma} \sim 10^{3}$ caused severe systematic uncertainties due to capture of resonance scattered neutrons. The sensitivity of the capture detectors to scattered neutrons limits the accuracy of present experiments at LINAC accelerators to $20-30 \%$. These uncertainties do not meet the requirements formulated for capture cross sections of structural materials (1).

In this work the neutron capture widths of s-wave resonances in ${ }^{56} \mathrm{Fe}(27.7 \mathrm{keV}),{ }^{58} \mathrm{Ni}(15.4 \mathrm{keV})$ and ${ }^{60} \mathrm{Ni}(12.5 \mathrm{keV})$ have been determined using a setup completely different from previous experiments. A pulsed 3-MV Van de Graaff accelerator and a kinematically collimated neutron beam, produced via the ${ }^{7} \mathrm{Li}(\mathrm{p}, \mathrm{n})$ reaction, was used in the experiments. Capture gamma-rays were observed by three Moxon-Rae detectors with graphite-, bismuth-graphite-, and bismuth-converter, respectively. The samples were positioned at a neutron flight path of only 9 cm and the distance from sample to detector is $\sim 16 \mathrm{~cm}$. Thus events due to capture of resonance scattered neutrons in the detectors or in surrounding materials are completely discriminated by their additional time of flight. The high neutron flux at the sample position allowed the use of very thin samples ( 0.15 mm 0.45 mm ), avoiding large multiple scattering corrections. The data obtained with the individual detectors were corrected for the efficiency of the respective converter materials. For that purpose, detailed theoretical calculations of the capture gamma-ray spectra of the measured isotopes and of gold, which was used as a standard, were performed. The final results are: $\Gamma_{\gamma}\left(27.7 \mathrm{keV},{ }^{56} \mathrm{Fe}\right)=1.06 \pm 0.05 \mathrm{eV}, \Gamma_{\gamma}\left(15.4 \mathrm{keV},{ }^{58} \mathrm{Ni}\right)=1.53 \pm 0.10 \mathrm{eV}$ and $\Gamma_{\gamma}\left(12.5 \mathrm{keV},{ }^{60} \mathrm{Ni}\right)=2.92 \pm 0.19 \mathrm{eV}$. The accuracy obtained with the present
experimental method represents an improvement of a factor 3-6 compared to previous experiments.

With these results Maxwellian average cross section were calculated for $\mathrm{kT}=30 \mathrm{keV}$. The following values were obtained for $\langle\sigma \mathrm{V}\rangle / \mathrm{v}_{\mathrm{T}}:(39.0 \pm$ $\left.2.5 \mathrm{mb},{ }^{58} \mathrm{Ni}\right),\left(31.4 \pm 1.6 \mathrm{mb},{ }^{60} \mathrm{Ni}\right)$ and $\left(13.9 \pm 1.1 \mathrm{mb},{ }^{56} \mathrm{Fe}\right)$. The investi= gated s-wave resonances contribute up to $40 \%$ to these average cross sections.

In a typical fast reactor spectrum these resonances cause about $10 \%$ of the total capture rate in structural materials and are therefore- together with the 1.15 keV resonance in ${ }^{56} \mathrm{Fe}$ - most important for reactor calculations.
(1) N. Dayday, Ed., "World Request List for Nuclear Data", INDC(SEC)-78/ URSF, International Atomic Energy Agency, Vienna (1981)

+ E.N.E.A. Bologna, Italy
1.1.2 NEUTRON CAPTURE IN s-WAVE RESONANCES OF ${ }^{64} \mathrm{Ni}$
K. Wisshak, F. Käppeler, R.L. Mack1in ${ }^{+}$, G. Reffo ${ }^{++}$, F. Fabbri ${ }^{++}$, KfK-report 3582 (1983) and Nuc1. Sci. Eng. (submitted)

The experimental method to determine the capture widths of $s$-wave resonances in structural materials with high accuracy as described in contribution 1.1 .1 has also been applied to ${ }^{64} \mathrm{Ni}$. This isotope has two s-wave resonances at 13.9 keV and 33.8 keV with very large ratios $\Gamma_{\mathrm{n}} / \Gamma_{\gamma}$ (2900 and 7700, respectively). These high values do not allow for an exact determination of the respective capture widths in LINAC experiments, and therefore this investigation was started in collaboration with the Oak Ridge National Laboratory. The results will serve to normalize the neutron sensitivity correction of LINAC measurements. In this way the accuracy of these experiments for resonances with large $\Gamma_{n} / \Gamma_{\gamma}$ ratios can be significantly improved in those energy regions which are not accessible to our method.

The final results $\left(\Gamma_{\gamma}(13.9 \mathrm{keV})=1.01 \pm 0.07 \mathrm{eV}\right.$ and $\Gamma_{\gamma}(33.8 \mathrm{keV})=$ $1.16 \pm 0.08 \mathrm{eV}$ ) are considerably smaller than the rough estimates obtained in previous work. The Maxwellian average neutron capture cross section for $k T=30 \mathrm{keV}$ is $\langle\sigma v\rangle / v_{T}=10.7 \pm 0.9 \mathrm{mb}$; it is determined to $63 \%$ by the two investigated resonances.

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$+{ }^{++}$E.N.E.A. Bologna, Italy.


### 1.1.3 THE CHEMICAL FIXATION OF Kr ISOTOPES IN ZEOLITE 5A-PREREQUISITE FOR THE DETERMINATION OF 25 keV NEUTRON CAPTURE CROSS SECTIONS WITH THE ACTIVATION METHOD

R. -D. Penzhorn ${ }^{+}$, G. Wa1ter, H. Beer, Zeitschrift für Naturforschung 38a (1983) 712

By chemical fixation of Kr in zeolite 5 A adequate samples can be obtained to determine the cross section of reactions such as ${ }^{84} \mathrm{Kr}(\mathrm{n}, \gamma){ }^{85} \mathrm{Kr}{ }^{\mathrm{m}}$ and ${ }^{86} \mathrm{Kr}(\mathrm{n}, \gamma){ }^{87} \mathrm{Kr}$. The employed zeolite loading was of the order 52-66 ( $\mathrm{cm}^{3}$ STP $\mathrm{Kr} / \mathrm{g}$ zeolite). The capture cross section of the reaction ${ }^{84} \mathrm{Kr}(\mathrm{n}, \gamma){ }^{85} \mathrm{Kr}^{\mathrm{m}}$ was determined at thermal and 25 keV neutron energy. The value obtained at 25 keV is of relevance to the stellar nucleosynthesis of heavy elements by the s-process.
$+\quad$ KfK, Institut für Radiochemie
1.1.4 NEUTRON CAPTURE CROSS SECTIONS AT 25 keV MEASURED BY ACTIVATION
G. Walter, H. Beer

The Maxwel1ian averaged neutron capture cross sections of ${ }^{71} \mathrm{Ga},{ }^{75} \mathrm{As}$, $79,81_{\mathrm{Br}}, 86_{\mathrm{Kr}}$ and $85,87_{\mathrm{Rb}}$ have been measured at $\mathrm{kT}=25 \mathrm{keV}$ by activation relative to the standard ${ }^{197} \mathrm{Au}$ 。

Neutrons were produced by the reaction ${ }^{7} \mathrm{Li}(\mathrm{p}, \mathrm{n}){ }^{7} \mathrm{Be} 27 \mathrm{keV}$ above the reaction threshold, the energy spectrum being almost Maxwellian (deviation $\sim 3 \%$ for the chosen experimental setup.

After activation the induced activities were counted in subsequent time intervals via a strong characteristic gamma-ray line with a $\gamma x$-Germanium detector ( $45 \mathrm{~cm}^{3}$, energy resolution 1.6 keV at 1.33 MeV ).

These cross sections form part of the data basis for a correct analysis of the s-process synthesis path through the elments between Germanium and Strontium. This mass region is important because there are two major sprocess branchings at ${ }^{79} \mathrm{Se}$ and ${ }^{85} \mathrm{Kr}$ which are expected to yield information on the s-process neutron density and temperature.

### 1.1.5 THE ABUNDANCE OF TECHNETIUM IN STARS

G. Schatz, Astron. Astrophys. 122 (1983) 327

It is known since more than 30 years that the unstable element technetium occurs in the atmosphere of some stars (1,2). All these are cool stars in a late stage of their evolution. Of the three processes responsible for the production of heavy elements in stars, the $\mathrm{r}^{-}, \mathrm{s}-$, and $\mathrm{p}-$ process, the latter is ruled out because it would lead to a large overabundance of molybdenum and, to a lesser extent, of samarium which is not observed. The r-process is associated with supernova explosions, and it is obvious from astronomical evidence that the technetium containing stars cannot be supernova remnants. It is therefore generally accepted that technetium in these stars is produced by the $s$-process somewhere in the stellar interior and mixed to the surface.

The only long-1ived $s$-process isotope of Tc is ${ }^{99} \mathrm{Tc}$ with a laboratory half-life of $2.1 \times 10^{5}$ a. Its fast neutron capture cross section has recently been measured (3). Using this result and the latest s-process systematics (4) it is then possible to calculate the abundance of ${ }^{99} \mathrm{Tc}$ or its ratio to the abundance of other elements. The result is

$$
N_{s}\left({ }^{99} \mathrm{Tc}\right) / \mathrm{N}_{\mathrm{s}}(\mathrm{Ba})=0.0162 \pm 0.0025
$$

The abundance ratio actually observed should be smaller than this figure for two reasons:
(i) Part of the barium now seen in the stellar atmosphere was incorporated into the star during formation whereas any technetium present at that time has since decayed.
(ii) The half-1ife of ${ }^{99} \mathrm{Tc}$ is considerably reduced at the high temperature of the stellar interior (5). This is due to possible allowed $B$-decays of low-lying exciting states whereas the ground-state $\beta$-decay is second forbidden. It can be estimated that the stellar half-1ife is 1 to 10 a at the most probable s-process temperature of $3.5 \times 10^{8} \mathrm{~K}$.

A quantitative determination of the stellar technetium abundance can therefore give clues to the time of mixing from the s-process zone to the stellar surface.
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### 1.1.6 NEUTRON CAPTURE CROSS SECTIONS OF THE STABLE XENON ISOTOPES AND THEIR APPLICATION IN STELLAR NUCLEOSYNTHESIS

H. Beer, F. Käppeler, G. Reffo ${ }^{+}$, G. Venturini ${ }^{+}$, Astrophysics and Space Science (in press)

The neutron capture cross sections of $124,132,134$ Xe have been measured by the activation technique at 25 keV neutron energy. These data were supplemented by calculated capture cross sections for $128,129,130,131_{\text {Xe }}$ via the statistical model based on carefully evaluated local parameter systematics. The complete set of capture cross sections obtained in this way is shown in Table I together with previous, more global calculations. The uncertainty of our theoretical results is estimated to $20 \%$.

Table I Comparison of the Xenon Cross Sections obtained in this work with Various Calculations at $\mathrm{kT}=30 \mathrm{keV}$.

| Xe- <br> Isotopes | $\sigma_{\mathrm{ny}}(\mathrm{mb})$ |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
|  | Harris (1) | $\begin{gathered} \text { Ho1mes } \\ \text { et a1. }(2) \\ \hline \end{gathered}$ | $\begin{aligned} & \text { ENDF/B-V } \\ & \text { (3) } \\ & \hline \end{aligned}$ | $\begin{array}{r} \text { Benzi } \\ \text { et } \quad \mathrm{al.}(4) \\ \hline \end{array}$ | present work calc. exp. |
| 128 | 510 | 232 | 206 | 239 | 249 |
| 129 | 1454 | 666 | 560 | 572 | 470 |
| 130 | 207 | 143 | 156 | 189 | 153 |
| 131 | 570 | 587 | 310 | 491 | 348 |
| 132 | 74.9 | 91 | 85 | 116 | $65 \quad 61+4$ |
| 134 | 37.5 | 47.5 | 41 | 60 | $33 \quad 29+2$ |



Fig. 1
Isotopic anomaly in chemically prepared residues of the Allende meteorite (5) as shown by the black dots. Open squares result from an interpretation in terms of an exotic component added to the solar system material. Open triangles indicate the effect of depleting solar system xenon in its s-process component.

With these cross sections we determined the solar xenon abundance through s-process systematics and decomposed the abundances of $129,131,132$ Xe into their $s^{-}$and $r$-process contributions. Furtheron, various isotopic anomalies in the xenon isotopes which were detected in several meteorites are discussed in detail. As an example, Fig. 1 shows that the anomaly found in the Allende meteorite (5) can be explained rather by an intermediate type $r$-process (or $n \beta$-process) than by the original interpretation by Clayton and Ward (6) as being due to a depletion of solar system xenon in its $\mathrm{s}^{-}$ process component.
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1.1.7 NEUTRON CAPTURE CROSS SECTIONS OF $142,143,144 \mathrm{Nd}$
G.J. Mathews ${ }^{+}$, F. Käppeler, K. Wisshak

The Nd isotopes are important for the s-process because of two reasons:
(i) $\quad 142 \mathrm{Nd}$ is an $s$-only nucleus with the magic neutron number 82 ; that means, it represents an important normalization point for the sprocess abundances.
(ii) The only significant isotopic anomalies in heavy elements (besides the noble gases) were found for Nd. Any interpretation has to check whether this anomaly might have to do with an enrichment or a depletion in smprocess material (1).

To improve the data basis for such an analysis, we have remeasured the capture cross sections of $142,143,144 \mathrm{Nd}$ between 6 and 250 keV neutron energy. The experimental setup consisted of two $C_{6} D_{6}$ detectors with offline pulse height weighting. Neutrons were produced at the Karlsruhe Van de Graaff accelerator via the ${ }^{7} \mathrm{Li}(\mathrm{p}, \mathrm{n})$-reaction. A flight path of 0.60 m and an overall time resolution of $\sim 1.2 \mathrm{~ns}$ allowed for a resolution of $\Delta \mathrm{E} / \mathrm{E}=$ 30 keV . Our results for the 30 keV Maxwellian averaged cross sections are shown in Table I together with previous data.

Table I $30-\mathrm{keV}$ Maxwellian Averaged Cross Sections (mb)

| Isotope | Ko 78 (2) | Mu 78 (3) | Na 79 (4) | this work |
| :--- | :---: | :---: | :---: | :---: |
| 142 Nd | $48 \pm 8$ | $57 \pm 7$ | $51 \pm 4$ |  |
| $143_{\mathrm{Nd}}$ | $265 \pm 50$ | $319 \pm 26$ | $258 \pm 11$ |  |
| 144 Nd | $65 \pm 10$ |  | $123 \pm 6$ |  |

We find good agreement with existing measurements for all isotopes but for ${ }^{144} \mathrm{Nd}$ where a severe discrepancy of more than a factor 2 shows up. The sprocess analysis is in progress.
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1.1.8 148,150 Sm: A TEST FOR s-PROCESS NUCLEOSYNTHESIS
R. R. Winters ${ }^{+}$, F. Käppeler, K. Wisshak, G. Reffo ${ }^{++}$, A. Mengoni ${ }^{++}$

The s-only isotopes 148,150 Sm allow to check the prediction of s-process nucleosynthesis, namely that the product $\sigma_{n \gamma} \cdot N_{s}$ ( $N_{s}$ being the resulting abundance) should be about equal for the isotopes involved in this process.

Similar tests using isotopes of Te and Ba verified the oN-equality within the respective cross section uncertainties of $10-15 \%$. However, for the example of Sm the situation was confused by a big discrepancy between previous measurements (see Table I).

We have remeasured the capture cross sections of $148,149,150{ }_{S m}$ in the energy range from 4 to 250 keV giving careful attention to reducing systematic uncertainties.

The absolute cross sections were determined with an uncertainty of $4.5 \%$ as is quoted in Table $I$, whereas for the cross section ratios for
which part of the uncertainties cancel out, a $3.3 \%$ uncertainty was found. This seems to be kind of a limit for present day techniques.

Table I Experimental values for the 30 keV Maxwellian Averaged Cross Section (mb)

| Isotope | Ma $67(1)$ | Ko 78 (2) | Mi 79 (3) | this work |
| :--- | :---: | :---: | :---: | :---: |
| ${ }^{148} \mathrm{Sm}$ | $257 \pm 50$ | $281 \pm 23$ |  | $269 \pm 12$ |
| 149 |  | $2490 \pm 200$ | $1489 \pm 65$ |  |
| $150_{\mathrm{Sm}}$ | $370 \pm 72$ | $690 \pm 51$ | $458 \pm 18$ |  |

Comparison of the results in Table I shows that good agreement is found only for ${ }^{148}$ Sm. For the other two isotopes severe descrepancies could be resolved by the present measurement.

For the astrophysically important cross section ratio $\sigma^{148} \mathrm{Sm} / \sigma^{150} \mathrm{Sm}$ we obtained
$\frac{(\sigma N)^{148} \mathrm{Sm}}{(\sigma N)^{150} \mathrm{Sm}}=0.90 \pm 0.03$.
This value being significantly different from unity implies a branching of the s-process capture path at ${ }^{147} \mathrm{Nd}$ and $147,148 \mathrm{Pm}$. In order to analyse these branchings, the capture cross sections of the unstable branching isotopes were calculated according to the Hauser-Feshbach formalism using a consistent set of carefully evaluated parameters.

With all these informations constraints for the shape and the normalization of the $\sigma N(A)$-curve as well as for the neutron density during the $s$ process were derived. The respective results are most affected by the uncertainty in the cross section ratio of ${ }^{148} \mathrm{Sm}$ and ${ }^{150} \mathrm{Sm}$. Further improvements in the measurement technique would therefore be important for this problem.
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# 1.1.9 ${ }^{178,179,180} \mathrm{Hf}$ AND ${ }^{180_{\mathrm{Ta}}(\mathrm{n}, \gamma)}$ CROSS SECTIONS AND THEIR CONTRIBUTION TO STELLAR NUCLEOSYNTHESIS 

H. Beer, R.L. Mack1in ${ }^{+}$, Phys. Rev. C 26 (1982) 1404

The neutron capture cross sections of $178,179,180$ Hf were measured in the energy range 2.6 keV to 2 MeV . The average capture cross sections were calculated and fitted in terms of strength functions. Resonance parameters for the observed resonances below 10 keV were determined by a shape analysis. Maxwellian averaged capture cross sections were computed for thermal energies $k T$ between 5 and 100 keV . The cross sections for $k T=30 \mathrm{keV}$ were used to determine the population probability of the $8^{-}$isomeric level in ${ }^{180}$ Hf by neutron capture as ( $1.24 \pm 0.06$ ) \% and the r-process abundance of ${ }^{180} \mathrm{Hf}$ as $0.0290\left(\mathrm{Si} \equiv 10^{6}\right)$.

These quantities served to investigate the origin of ${ }^{180} \mathrm{Ta}$ m by neutron capture nucleosynthesis. The detailed analysis showed that the s-process contributes only $\sim 10 \%$ to the buildup of the ${ }^{180} \mathrm{Ta}^{\mathrm{m}}$ solar abundance provided our estimated logft value of 6.4 for the ${ }^{180} \mathrm{Hf}^{\mathrm{m}} \rightarrow{ }^{180} \mathrm{Ta}^{\mathrm{m}}$ beta decay transition is correct. A final clarification must await a measurement of this fractional beta decay $f_{\beta^{-}}^{m}$. If it should turn out that $f_{\beta_{-}}^{m}$ is substantially greater than the present estimate one has to investigate the question of how the ${ }^{180} \mathrm{Ta}{ }^{\mathrm{m}}$ isomer could survive in the hot stellar photon bath under s-process temperature conditions because any excited level which could equilibrate the isomer with the 8.1 h ground state would rapidly destroy the result of the s-process synthesis. This would mean that ${ }^{180} \mathrm{Ta}^{\mathrm{m}}$ can be used as a stellar thermometer of the s-process.

There is still the possibiliby that post-r-process nucleosynthesis can produce a sizable amount of ${ }^{180} \mathrm{Ta}^{\mathrm{m}}$. To confirm this would mainly require a careful search for a branch in the ${ }^{180} \mathrm{Lu}$ decay feeding the ${ }^{180} \mathrm{Hf}^{\mathrm{m}}$ isomeric state. The study of these extended questions in the frame of $s-$ and $r$-process nucleosynthesis requires, however, a detailed knowledge of the ${ }^{180}$ Ta level scheme below 500 keV .

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### 1.1.10 THE SOLAR MERCURY ABUNDANCE

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For the determination of elemental solar abundances C1 type carbonaceous meteorites represent the primary source of information. This is true for at least the bulk of elements. Exceptions are the highly volatile noble gases but also mercury which shows an enormous variability in meteoritic samples (1).

In this situation it appears to be justified to refer to the concepts of nucleosynthesis and to try an interpolation via neighbouring nuclei.

In the present investigation the solar mercury abundance is estimated through s-process systematics. The analysis is based on the measurement of the neutron capture cross section of ${ }^{202} \mathrm{Hg}$ which is dominantly of s-process origin. The r-process contribution to this isotope was approximated by r-only ${ }^{204} \mathrm{Hg}$. With the measured cross section of $72 \pm 5 \mathrm{mb}$ we derived a solar mercury abundance of $0.24 \pm 0.04 / 10^{6}$ si using the oN-systematics of Käppeler et a1. (2).

A comparison of published solar abundances for mercury and the result of this work is given in Table I:

Table I Comparison of solar abundances for mercury

| Ref. | Solar mercury abundance |  |
| :--- | :---: | :---: |
| Suess and Urey | $(3)$ | 0.284 |
| Cameron | $(4)$ | 0.75 |
| Suess and Zeh | $(5)$ | 0.44 |
| Cameron | $(6)$ | 0.4 |
| Cameron | $(7)$ | 0.21 |
| Palme et a1. (8) (9) | 0.4 |  |
| Anders and Ebihara (9) | $0.52 \pm 0.26$ |  |
| this work | $0.24 \pm 0.04$ |  |

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### 1.1.11 DIRECT RADIATIVE CAPTURE: TEST OF THE LANE-LYNN MODEL AND DEVELOPMENT OF A METHODOLOGY FOR CALCULATIONS

A. Mengoni, KfK-report 3452 (1982)

The Lane-Lynn model for direct radiative capture (DRC) is extended to isotopes where no experimental information on the low-lying levels (spectroscopic factors etc.) is available. To this end we assumed that all the partial transitions to the final levels are lumped in the $s_{1 / 2}{ }^{+} p_{3 / 2}, s_{1 / 2}+p_{1 / 2}$ transitions only (where the p-wave final states are single particle states with complete strength). Within these assumptions the direct capture cross sections for calcium isotopes between $A=40$ and $A=48$ (Fig. 1) as well as for ${ }^{136}$ Xe were calculated. We showed that the uncertainties of these calculations are $\lesssim 20 \%$ if we compare the results to the calculations with the Lane-Lynn model for ${ }^{42,48} \mathrm{Ca}$ and ${ }^{136}$ Xe where the respective experimental informations are available. An additional uncertainty of $\simeq 10 \%$ has to be admitted in the calculations due to the uncertainty of the scattering radius which accounts for the effect of distant levels in the DRC model.


Fig. 1
Comparison of calculated DRC cross sections with experimental values from Ref. (1) for even-mass calcium isotopes at thermal energy. The calculated values $\sigma_{n, \gamma}^{D R C}$ are derived under the assumptions given in the text. The hatched error band of the calculated values corresponds to the uncertainties of the scattering radius $R^{\text {exp }}$.

The new approach allows for an evaluation of thermal cross sections and, more important, for the calculation of DRC cross sections of extremely neutron rich nuclei which are significant for nucleosynthesis of the heavy elements in the rapid neutron capture process ( $r$-process).
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### 1.1.12 NEUTRON CAPTURE RATES IN THE r-PROCESS: THE ROLE OF DIRECT RADIATIVE CAPTURE

G.J. Mathews ${ }^{+}$, A. Mengoni, F.-K. Thielemann ${ }^{++}$, W.A. Fowler ${ }^{+++}$, Ap. J. (in press)

The nuclear physics associated with neutron capture along the r-process path is reviewed, and it is noted that the usual statistical assumptions for the capture rates may not be valid. We have shown that direct radiative capture can introduce a significant, and perhaps more appropriate, contribution to the neutron capture rate for nuclei in the vicinity of the r-process path and should be included in r-process reaction networks.

As an example, the Cd isotopes with $120 \leq \mathrm{A} \leq 140$ have been investigated. The DRC calculations are shown in Fig. 1 in comparison with the statistical compound nucleus model. Once the neutron shell closure at $N=82$ is


Fig. 1
A comparison of Maxwellian averaged meutron capture cross sections at $\mathrm{kT}=100 \mathrm{keV}$ for Cd isotopes through the r-process path at $A=130$. DRC indicates the direct radiative capture component. HF1 and HF2 show two different approaches to the statistical Hauser-Feshbach mode1. It can be seen that beyond the closed neutron shell at $N=82$ the direct radiative capture component dominates the neutron capture rate.
reached, the level density decreases drastically, and therefore the compound nucleus capture cross section falls well below the DRC contribution which in turn remains practically constant. This holds at least for energies around 100 keV corresponding to typical r -process temperatures.

Since the DRC estimates presented here can be taken as realistic lower limits to the actual neutron capture rates, these cross sections can be utilized to infer an upper limit to the critical neutron density above which ( $n, \gamma$ ) equilibrium can probably be assumed in any r-process production calculation. This 1 imit we determine to be $n_{\text {crit }}<1.5 \times 10^{20} \mathrm{~cm}^{-3}$. In a future paper detailed systematics of the DRC reaction rates (including other multipolarities and angular momenta) will be summarized for application to r-process network calculations.

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### 1.2.1 THE NEUTRON CAPTURE CROSS SECTION OF ${ }^{243}$ Am IN THE ENERGY RANGE FROM 5 TO 250 keV

K. Wisshak, F. Käppeler, G. Rupp, KfK-report 3503 (1983) and Nuc1. Sci. Eng. (in press)

Neutron capture in ${ }^{243}$ Am contributes significantly to the production of ${ }^{244} \mathrm{Cm}$, which is of importance in fast reactor studies as a strong neutron emitter, and hence affects the total neutron production rate (1). Several requests for capture cross section measurements have been formulated ( $1,2,3,4$ ) asking for an accuracy of $\sim 10 \%$ in the kev range. The quoted applications are burn-up studies, the build-up of transplutonium elements, neutron shielding of transport casks for spent fuel elements, fuel reprocessing and storage as well as the long term radioactivity hazard.

Due to the very high radioactivity of the sample material no experimental data for the capture and subthreshold fission cross section of ${ }^{243} \mathrm{Am}$ are available from literature up to now. In the present experiment these cross sections were measured in the energy range from 5 to 250 keV using ${ }^{197} \mathrm{Au}$ and ${ }^{235} \mathrm{U}$ as the respective standards. Neutrons were produced via the ${ }^{7}$ Li $(p, n)$ and the $T(p, n)$ reaction with the Karlsruhe $3-M V$ pulsed Van de Graaff accelerator. Capture events were detected by two Moxon-Rae detectors with graphite and bismuth-graphite converters, respectively. Fission events were registered by a NE-213 1iquid scintillator with pulse-shape discriminator equipment. F1ight paths as short as $50-70$ mm were used to obtain an optimum signal-to-background ratio. After correction for the different efficiency of the individual converter materials the capture cross section could be determined with a total uncertainty of $3-6 \%$. The respective values for the fission cross section are $8-12 \%$. The final results for the capture cross section of ${ }^{243} \mathrm{Am}$ are shown in Fig. 1 together with predictions from model calculations.


Fig. 1 The neutron capture cross section of ${ }^{243} \mathrm{Am}$
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Nuc1. Sci. Eng., 63, (1977) 242.
1.2.2 BACKWARD ANGLE A MEASUREMENT FOR $\overrightarrow{\mathrm{n}} \mathrm{p}$ SCATTERING UP TO 50 MeV
F.P. Brady*, P. Do11, W. Heeringa, K. Hofmann, H.O. Klages,
J. Wilczynski

First measurements of $A$ were carried out at POLKA (1) by having the collimated neutron beam impinge on a target consisting of a vertical cylinder ( 76.2 mm high x 76.2 mm diameter) of NE 213 liquid scintillator at 6 m from the $\overrightarrow{\mathrm{n}}$ source. Scattered neutrons were detected at 2 m in 14 NE 213 1iquid scintillators (vertical cylinder 200 mm high x 140 mm diameter) at 7 angles covering the range from $16^{\circ}$ to $71^{\circ} 1 \mathrm{ab}$. Each event (a target-detector coincidence) was characterized by the source-target (or incident) time-of-flight (TOF), the target-detector TOF, the pulse heights in the target and detector scintillators, the pulse shape in the detector scintillator, the detector number, and the $\vec{n}$ spin direction. The first parameter, which was measured relative to the cyclotron RF, gave the incident neutron energy which was divided into 10 energy bins whose average energies were $17,19,22,25,27.5,30$, $33,36,40$ and 50 MeV .

Here we report on new measurements of $A_{y}$ for two backward $\vec{n}$ p CM angles. Using the set-updescribed elsewhere in this report, $A_{y}$ was measured by detecting and identifying recoil protons from a $\mathrm{CH}_{2}$ polyethylene target in $\triangle \mathrm{E}-\mathrm{E}$ telescopes at $\pm 14^{\circ}$ and $\pm 20^{\circ} \mathrm{lab}$. In this case the incident TOF was derived from the $\triangle E-E$ coincidence. The other parameters measured were the $\Delta E$ and $E$ pulse heights, telescope number, and $\overrightarrow{\mathrm{n}}$ spin direction. In each experiment all parameters were written on magnetic tape. After off-line analysis the data were corrected for finite geometry and multiple scattering effects. Then from events for each energy bin and each pair of left (L) and right (R) detectors and for neutron spin up $\uparrow$ and down $\downarrow$, the usual asymmetry $\varepsilon=(\sqrt{L \uparrow \cdot R \downarrow}-\sqrt{L \downarrow \cdot R \uparrow}) /$ $(\sqrt{L \uparrow \cdot R \downarrow}+\sqrt{L \psi \cdot R \uparrow})$ and hence $A_{y}=\varepsilon / P_{n}$ is calculated. The resulting values of $A_{y}$ (as \%) for two incident neutron energies are shown in Fig. 1. For comparison other data from earlier work are also shown. In general the POLKA results have much smaller uncertainties particularly at 30 MeV where the $\overrightarrow{\mathrm{n}}$ beam intensity is larger. The data at $22,25,27.5$ and 33 MeV also have small $\Delta A_{y}<0.5(\%)$. The largest $\Delta A_{y},<1.5(\%)$, are at 50 MeV . The uncertainties in the $151^{\circ}$ data from $p$ detection are about the size of the data points. (Similar data for $140^{\circ}$ are not yet analyzed.) This method gives high statistical accuracy and is free of the multiple scattering corrections required for neutron detection data at similar CM angles.

At 17 MeV our results confirm the very precise data from Duke (2).

For energies above 20 MeV the existing data base was considerably improved. At 50 MeV the results agree well with the data from UC Davis (3, 4). In Fig. 1 the curve through the data is from a phase shift analysis (5) which did not include the $151^{\circ}$ data from $p$ detection. A more recent analysis including the $151^{\circ}$ data has also been made (6), and the results show clearly the improved determination of some of the phase shifts, particularly the ${ }^{3} D_{3}$ phase.


Fig. 1: Ay data at 30 and 50 MeV . Eldred, Mutchler, Langsford, Garrett and Romero refer to refs. $7,8,9,3$, and 4 respectively.
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### 1.2.3 A PRELIMINARY PHASE SHIFT ANALYSIS FOR THE n-p SYSTEM UP TO 50 MeV

## H.O. Klages, J. Wilczynski

Although the two-nucleon-system has been studied extensively for many years, there remain severe problems in a unique description of the system. In phase shift analyses (1-3) of the experimental data below 100 MeV several parameters deviate from phase shifts predicted by theoretical models. For some phase shifts even different models come to inconsistent answers. Although there are new cross section measurements available (4) the parameters ${ }^{1} P_{1}$ and $\varepsilon_{1}$ remain still not well determined because of the insufficiency of the neu-tron-proton data base. Analysis shows that ${ }^{1} P_{1}$ and $\varepsilon_{1}$ are strongly correlated on the base of the existing neutron-proton data. Sensitivity calculations indicate that this correlation can be reduced by an accurate measurement of the angular distribution of the analyzing power $A y$ and a very precise determination of the differential cross section at backward angles.

Recently, at the Karlsruhe polarized neutron facility POLKA a measurement of the angular distribution of the analyzing power $A_{y y}$ in the energy range from 20 MeV to 50 MeV was conducted (5). The data of this experiment are still being analyzed. A measurement of the differential cross section of the neutron-proton scattering at backward angles in the energy range of interest is also in preparation at POLKA. So the experimental situation will soon be calling for a new analysis which should be able to determine the iso-spin-singletparameters ${ }^{1} \mathrm{P}_{1}$ and $\varepsilon_{1}$ in a more stringent way.

The phase shift analysis reported here was intialized by a new precise measurement of the neutron-proton analyzing power $A$ in the energy range from 17 MeV to 50 MeV (6). Calculations show that a precise knowledge of the analyzing power $A_{y}$ is sensitive to (among other phase shifts) all ${ }^{3} \mathrm{D}$ phase shifts. ${ }^{3} D_{3}$ phase shift is predicted differently by several theoretical models whereas $3^{3}{ }_{1}$ and ${ }^{3} D_{2}$ are in good agreement with values given by all models (7,8).

This single energy phase shift analysis was carried out at $25,30,40$ and 50 MeV by taking into account all available neutron-proton data. Phase shifts with $J \leq 3$ were fitted to the experiment data. Phase shifts contributing to higher $J$-values were fixed at values predicted by the Bonn OBEP (7). The energy dependence of all phase shifts was taken from the same OBE-model.

As expected, the addition of the new precise analyzing power data improved the knowledge of the ${ }^{3}$ D phase shifts. In Fig. $1 a-c$ the central-, tensorand spin-orbit phase parameters are presented which are linear combinations of the ${ }^{3} D_{J}$ phase shifts with $J=1,2,3$.

The phase shift analyses reported here could not determine the ${ }^{1} \mathrm{P}_{1}$ phase shift and the mixing parameter $\varepsilon_{1}$ which were discussed above, but investigate the behaviour of the ${ }^{3} D$ phase shifts. In the next step, as soon as the new analyzing power $A$ yy data are available, the uncertainties in the parameters ${ }^{1} P_{1}$ and $\varepsilon_{1}$ will be further reduced. The inclusion of proton data in a full analysis will lead to more reliable "experimental" phase shifts for the nucleon-nucleon system in this energy range.


Fig. 1a-c: $\Delta_{C}^{D}=+\frac{1}{15}\left(3 \cdot{ }^{3} D_{1}+5 \cdot{ }^{3} D_{2}+7 \cdot{ }^{3} D_{3}\right)$

$$
\begin{aligned}
& \Delta_{\mathrm{LS}}^{\mathrm{D}}=-\frac{1}{60}\left(9 \cdot{ }^{3} \mathrm{D}_{1}+5 \cdot{ }^{3} \mathrm{D}_{2}-14 \cdot \cdot^{3} \mathrm{D}_{3}\right) \\
& \Delta_{\mathrm{T}}^{\mathrm{D}}=-\frac{7}{120}\left(3 \cdot{ }^{3} \mathrm{D}_{1}-5 \cdot{ }^{3} \mathrm{D}_{2}+2 \cdot{ }^{3} \mathrm{D}_{3}\right)
\end{aligned}
$$

+ this work, o MAW X (Ref.11), Bonn OBE (Ref.7), a Paris Potential (Ref.8) , $\triangle$ A\&W (Ref.9) , BBS (Ref.10).
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### 1.2.4 SCATTERING OF POLARIZED NEUTRONS ON Ti AND TiH 2 SAMPLES

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Scattering experiments of polarized neutrons on a Ti and a $\mathrm{TiH}_{2}$ sample were performed, to investigate the possibility of extracting reliable asymmetries for $\vec{n}-p$ scattering. After a flight path of 5.2 m from the POLKA source the neutrons were scattered from the sample which was mounted between thick iron plates to simulate the surroundings of the polarized target. The scattered neutrons were detected by 5 pairs of symmetrically aranged detectors for which the distance to the scattering samples was optimized to achieve a kinematical separation between scattering from Ti and H . The following parameters where measured: The detector number, the total time-of-flight from the POLKA source via the sample to the detector, the proton recoil energy in the detector and the pulse shape in the detector. To subtract out carefully the amount of $n-p$ scattering for a given incident neutron energy to the left or right detector, time-of-flight spectra for all detectors were taken at zero degrees; the neutron $f 1 u x$ was monitored precisely by a pair of proton recoil telescopes before the scattering experiment. In addition background runs with no sample were performed. An independent check on the normalization is given by elastic scattering at $15^{\circ}$. Fig. 1 shows recoil energy spectra from scattering on $\mathrm{TiH}_{2}$ and after subtraction of scattering from Ti.


Fig. 1:
a) Pulse height spectrum for neutrons of 25 MeV scattered by $45^{\circ}$ on $\mathrm{TiH}_{2}$
b) Ti background spectrum subtracted

The countrate in the subtracted spectrum was evaluated as indicated by the arrows. A lower threshold had to be set to exclude scattering events which result from lower neutron energies from previous cyclotron bursts. Fig. 2 shows the asymmetries extracted for 9 energy intervalls from 19 MeV to 50 MeV incident neutron energy.


Fig. 2: Count rate asymmetries in the time-offlight spectra at $45^{\circ}$. Each data point corresponds to an energy bin in the range from 19 MeV to 50 MeV , calculated for the $n-p$ scattering

The upper part shows the asymmetries $\varepsilon$ for $\vec{n}-\mathrm{p}$ scattering, which agree fairly well with the $\varepsilon$ from a previous $\vec{n}-p$ experiment on a hydrocarbonate sample, the lower part shows $\varepsilon$ for $\overrightarrow{\mathrm{n}}-\mathrm{Ti}$ scattering in the same time-of-flight intervalls, which agree fairly well with optical model predictions. Similar data taken at $15^{\circ}, 50^{\circ}, 56^{\circ}$ and $63^{\circ}$ will prove the experimental methode to be reliable for measuring $A_{y y}$ with this polarized target and our continous energy polarized neutron beam.

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### 1.2.5 DETERMINATION OF THE PROTON POLARIZATION IN TiH 2 WITH POLARIZED NEUTRONS

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A brute-force polarized proton target has been developed for scattering experiments with fast polarized neutrons. The target material is titaniumhydride, which has the advantages of very high proton density and sufficient heat conductivity. During hydrogenation the titanium falls apart into a powder, which is pressed into copper cylinders afterwards. The sample is placed in the centre of a superconducting split pair magnet with a field strength of 8.2 T in this experiment, and it is cooled by a ${ }^{3} \mathrm{He}-{ }^{4} \mathrm{He}$ dilution refrigerator.

The temperature of the sample is measured with nuclear orientation thermometry using ${ }^{60}$ Co in a cobalt single crystal soldered onto the copper cylinder. In principle the polarization of the protons can be calculated from the magnetic field and the temperature. However, several insufficiently known effects, e.g. warming up by the beam and spin-lattice relaxation make it desirable to measure the polarization directly.

We have done this recently in a transmission experiment using polarized neutrons of 1.2 MeV . Up to about this energy the total cross section of protons for neutrons depends strongly and in a well-known way on their relative spin orientation. A measurement of the count rate difference $\varepsilon$ between parallel and anti-parallel spins is thus a direct and sensitive method to determine the proton polarization. The following relation holds for the quantities involved:

$$
\varepsilon=\frac{N^{\uparrow \uparrow}-N^{\uparrow \downarrow}}{N^{\uparrow \uparrow}+N^{\uparrow \downarrow}}=\tanh \left(-N L p_{p} p_{n} \frac{\sigma^{\uparrow \uparrow}-\sigma^{\uparrow \downarrow}}{2}\right)
$$

Here $N$ and $L$ are the proton density and the length of the target, $p_{p}$ and $p_{n}$ are the proton and neutron polarization and the sigma's are the total cross sections for both spin orientations.

Polarized neutrons were obtained by the reaction ${ }^{3} H(p, n){ }^{3}$ He at a reaction angle of $\Theta_{1 a b}=35^{\circ}$ with protons of 2.26 MeV . The spin direction of the collimated neutron beam was reversed by the field of a superconducting solenoid about every 15 min. The obtained proton polarization data are shown in Fig. 1 as crosses.


Fig. 1: The proton polarization obtained by the transmission experiment and the temperature found in the $\gamma$-anisotropy measurements as function of the cooling time.

After about 6 days of cooling a stable situation was attained. The result after this time is $p_{p}=60.4 \pm 0.7 \%$ when only the error in $\varepsilon$ is taken into account. The error increases to $2.5 \%$ when all other errors are included.

The circles in Fig. 1 are the temperatures obtained by the ${ }^{60}$ Co $\gamma$ measurements. The temperature found after 6 days was $10.7 \pm 0.2 \mathrm{mK}$, corresponding to a polarization of $p_{p}=65.4 \pm 0.9 \%$. With this target we have constructed the first brute-force polarized proton target useful for scattering experiments.

### 1.2.6 ANALYZING POWER OF THE ELASTIC $\vec{n}$-d SCATTERING FROM 14 MeV to 50 MeV

F.P. Brady ${ }^{*}$, P. Do11, E. Finckh ${ }^{* *}$, W. Heeringa, K. Hofmann, H.O. Klages, W. Nitz, J. Wilczynski

The three-nucleon systems have been studied extensively both in experiments and theoretical approaches during the last years (1,2,3). Most precision data were obtained for the $p-d$ system, whereas in the three-body calculations based on Faddeev-type equations the observables of the $n-d$ channel were studied. The Coulomb interaction was treated perturbatively in these calculations and the question remained whether the existing discrepancies were due to these approximations or to the details of the nucleon-nucleon input (4). Recently, differential cross section data for the elastic $n-d$ scattering in a wide range of energies were published (5) with an accuracy comparable to the $p-d$ results. For the analyzing power only at a few energies $(6,7,8)$ high precision data are available until now. Here we present the preliminary results of a measurement of the analyzing power $A_{y}$ for the elastic $\overrightarrow{\mathrm{n}}-\mathrm{d}$ scattering.

The experiment was performed using the continous energy polarized neutron beam at the Karlsruhe cyclotron. A deuterated liquid scintil1ator detector served as scattering sample. 20 neutron detectors were used in symmetric pairs at a distance of 1 m from the sample. Angular distributions were measured in the range from $39^{\circ} \mathrm{c} . \mathrm{m}$. to $153^{\circ} \mathrm{c} . \mathrm{m} . \mathrm{M}^{\circ}$ Multiparameter data acquisition was performed including pulse shape discrimination techniques for the neutron detectors and for the scattering sample. This enables us to distinguish clearly between signals from elastic scattering events, i.e. recoil deuterons, and from breakup protons in the scattering sample, as shown in Fig. 1.


Fig. 1: Two-dimensional view pulse height vs. pulse shape in the scattering sample. Shown are data for the neutron energy range from 36 to 50 MeV and for backward scattering angles.

The evaluation of analyzing power data for the FSI region of the breakup channel is described in contribution 1.2.6. The data for the elastic $\vec{n}-d$ scattering were analyzed in 12 energy bins from 14 MeV to 50 MeV . In the energy range above 20 MeV the accuracy of our results is such that the existing data base is considerably improved. In the next step our raw data have to be corrected for multiple scattering and for finite geometry.

This new set of precise data together with the recent results for the differential cross section will allow conclusive comparison with p-d observables and with the predictions of three-nucleon calculations. Fig. 2 shows our raw data at 30 MeV together with the $\overrightarrow{\mathrm{p}}$-d results of Johnston et al. (9) * and $\overrightarrow{\mathrm{n}}$-d data of Dobiasch et al. (7) $\Delta$.


Fig. 2: Comparison of $\overrightarrow{\mathrm{n}}-\mathrm{d}$ and $\overrightarrow{\mathrm{p}}-\mathrm{d}$ analyzing power data at 30 MeV (see text)
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### 1.2.7 ANALYZING POWER DISTRIBUTIONS FOR THE ${ }^{2} \mathrm{H}(\overrightarrow{\mathrm{n}}, \mathrm{n}) \mathrm{np}$ BREAKUP IN THE FSI RANGE

F.P. Brady ${ }^{*}$, P. Doll, E. Finckh ${ }^{* *}$, W. Heeringa, K. Hofmann, H.O. Klages W. Nitz, J. Wilczynski

In the study of the nucleon-nucleon potential it seems important to get more information about the off-shell behaviour of three nucleon scattering processes, because the off-shell matrix elements cannot be determined in two-body reactions. Furtheron the polarization effects in the region of the final-state-interaction (FSI) of the elastic nd scattering are very sensitive for the test of the influence of higher partial waves in three-body calculations (1). Theoretical considerations indicated for some phase space regions that one can get more information on the repulsive case in the $N-N$ interaction. This leads to better knowledge of the short-range nuclear forces (2). Experimental data for the breakup analyzing power exists only at 22.7 MeV (3), so there is still more need for precise data even at higher energies.

The experiment reported here was performed at the Karlsruhe cyclotron using the continuous energy neutron beam from POLKA (4). The experimental set-up is described in contribution 1.2 .5 to this report. The very good pulse shape properties of the deuterated liquid scintillator, which served as a scattering sample, enabled us to discriminate protons from the nd breakup channel and recoil deuterons from elastic scattering in the data analysis. Fig. 1 shows the separation of these processes during the off-1ine data analysis.

This distinction allows the determination of the analyzing power of the elastic nd scattering and the breakup channel simultaneously. The data were analyzed in 10 energy bins from 16 MeV to 40 MeV in an angular range from $30_{\mathrm{CM}}^{\circ}$ to $146.4^{\circ} \mathrm{CM}^{\circ}$. The position of the FSI region in the scattering matrix is calculated from the kinematical situation that the particles in the np pair have a relative kinetic energy smaller than 1 MeV . Fig. 2 shows the analyzing power distribution for the FSI events at 22.5 MeV , without corrections for background events, multiple scattering and finite geometry. The open circles in Fig. 2 show the breakup data of the $\vec{p} d$ experiment of ref. (3) and the dashed line represents the result of a calculation by Doleschall.



Fig. 1: Two-dimensional representation of neutron scattering events ( 25 MeV , $102_{1 a b}^{\circ}$ ) during the off-1ine data analysis. Left part: elastic nd scattering and double scattering events. Right part: $n-d$ breakup processes and inelastic carbon scattering with a proton in the exit channel


Fig. 2: Angular distribution of the analyzing power for the reaction $\mathrm{D}(\overrightarrow{\mathrm{N}}, \mathrm{n}) \mathrm{np}$ (FSI) at 22.5 MeV . Open circles and dashed line see text.
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1.2.8 MEASUREMENT OF THE ANALYZING POWER OF THE ELASTIC $\overrightarrow{\mathrm{n}}{ }^{3}{ }^{3}$ He SCATTERING UP TO 50 MeV
F.P. Brady ${ }^{*}$, P. Do11, B. Haesner ${ }^{* *}$, W. Heeringa, K. Hofmann, P. Jany, H.O. Klages, Chr. Maier, J. Wilczynski

In the framework of systematic investigations (1,2,3) of the four nucleon system using the $n+{ }^{3}$ He entrance channel, the analyzing power of the elastic $\overrightarrow{\mathrm{n}}^{-3}$ He scattering has been measured in the energy range from 16 to 50 MeV .

We made use of the continuous energy polarized neutron beam at the Karlsruhe cyclotron (4). A liquid ${ }^{3}{ }^{3}$ He scintillation detector (5) served as scattering sample. Neutrons were detected by 20 1iquid scintillation detectors at a distance of 1 m from the sample.

The analyzing power was measured at 14 angles from $38^{\circ} \mathrm{CM}$ to $158^{\circ} \mathrm{CM}$. Multiparameter data acquisition and off-1ine data analysis was performed. The data were split into 11 energy bins centered at $16,18,20,22,24,27$, 30, 33, 36,40 and 50 MeV , according to a recent measurement (3) of the differential cross section for the elastic $n-{ }^{3}$ He scattering up to 30 MeV .

Angular distributions for $A$ were determined with high statistical accuracy compared to previous data in the energy range above $15 \mathrm{MeV}(1,6,7)$. Whereas the data of Lisowski et al. (6) at 17.1 MeV are in good overall agreement with our results at 16 and 18 MeV , the angular distribution of Busse et a1. (7) at 22 MeV shows a different shape. By our new data at this energy the deviating results of Dobiasch et a1. (1) were confirmed.

For a final analysis in terms of phase shifts for the $\mathrm{n}^{-}{ }^{3}$ He system our data have to be corrected for multiple scattering and finite geometry effects. First calculations (8) show that these corrections are small due to the careful multiparameter off-line data analysis. Preliminary results are shown in Fig. 1 and 2 for $E_{n}=22 \mathrm{MeV}$ and $E_{n}=40 \mathrm{MeV}$, respectively.


Fig. 2: The angular distribution of the $\overrightarrow{\mathrm{n}}-{ }^{3} \mathrm{He}$ analyzing power at 40 MeV . The statistical accuracy of the results is indicated.

Fig. 1: Our results at 22 MeV (full circles) compared to the data of Busse et al. (7) (crosses) and to the work of Dobiasch et al. (1) (open circles and dotted line).

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### 1.2.9 MEASUREMENT OF THE SPIN CORRELATION PARAMETER A FOR $\overrightarrow{\mathrm{n}} \overrightarrow{\mathrm{p}} \overrightarrow{\mathrm{p}}$ SCATTERING IN THE ENERGY RANGE UP TO 50 MeV <br> R. Aures, A. Bischoff, F.P. Brady ${ }^{*}$, P. Doll, E. Finckh ${ }^{* *}$, W. Heeringa, K. Hofmann, H.O. Klages, J. Wilczynski, B. Zeitnitz

Recent nucleon-nucleon phase shift analyses (1-4) show that in the energy range below 100 MeV several isospin-singlet parameters are not well determined. The set of observables measured until now is by no means complete and the accuracy of the neutron-proton data in this energy range is not sufficient. Therefore, especially the ${ }^{1} P_{1}$ phase shift and the ${ }^{3} S_{1}-{ }^{3} D_{1}$ mixing parameter $\varepsilon_{1}$ from the phase shift analyses do not agree with model predictions ( 5,6 ). Sensitivity calculations reveal that certain observables depend strongly on variations of these parameters. The impact of ${ }^{1} P_{1}$ and $\varepsilon_{1}$ on the differential cross section and on the spin correlation parameters $A_{y y}$ and $A_{z z}$ is discussed briefly.

A variation of the ${ }^{1} P_{1}$ phase shift at 25 MeV by two standard deviations changes the differential cross section at far backward angles by about $2 \%$. The mixing parameter $\varepsilon_{1}$ does not effect this quantity but has a strong influence on $A_{y y}$ as can be seen in Fig. 1. The phase shift ${ }^{1} P_{1}$ is important for this observable at backward angles only. $A_{z z}$ however is very sensitive to variations of both $\varepsilon_{1}$ and ${ }^{1} P_{1}$. The predicted distributions for $A_{z z}$ depend strongly on the starting values for the calculations (7). A measurement of $A_{y y}$ or of $A_{z z}$ at forward and backward angles would fix $\varepsilon_{1}$ and ${ }^{1} P_{1}$ simultaneously. The few existing data for $A_{y y}(8,9)$ at backward angles are not sufficient for this purpose.

At the Karlsruhe polarized neutron facility, POLKA, an experiment was carried out to measure the analyzing power $A_{y y}$ for $\vec{n}-\vec{p}$ scattering. Data were taken in the energy range from 20 MeV to 50 MeV by detecting neutrons at c.m. angles of $\pm 30^{\circ}, \pm 90^{\circ}, \pm 100^{\circ}$ and $\pm 112^{\circ}$. The polarization of the continous energy neutron beam used in this experiment varies from 0.3 at 20 MeV to about 0.5 at 50 MeV neutron energy (10). In the scattering target a proton polarization of about $70 \%$ was achieved by "brute force polarization" of a $\mathrm{TiH}_{2}$ sample which was cooled to a temperature of 10 mK inside a strong magnetic field of $B=9 \mathrm{~T}$.

The scattering from hydrogen was separated by a difference measurement using a pure Ti scattering sample in the cryostat. In a test experiment it could be demonstrated that this method works neatly even with a continuous energy neutron beam. Using an unpolarized $\mathrm{TiH}_{2}$ target, we were able
to reproduce the analyzing power $A_{y}$ for the $\vec{n}-p$ scattering.



Fig. 1: Sensitivity of the spin correlation parameter $A_{y y}$ at 25 MeV to variations of the mixing parameter $\varepsilon_{1}$ and of the ${ }^{1}{ }^{\text {P }}{ }_{1}$ phase shift. Starting values are taken from the Paris potential (6).
a) $\varepsilon_{1}=1.6^{\circ}$ varied by $\pm 0.5^{\circ}$
b) ${ }^{1} \mathrm{P}_{1}=-7.0^{\circ}$ varied by $\pm 2.0^{\circ}$

The data of the $A_{y y}$ experiment are presently beeing analyzed. The estimated accuracy of the results will enable us to put sharper constraints on $\varepsilon_{1}$ (and probably ${ }^{1} \mathrm{P}_{1}$ ) in a new phase shift analysis including also the new Karlsruhe data (11) on $A_{y}$ in the energy range from 17 to 50 MeV .
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### 1.3.1 OPTICAL POTENTIALS AND ISOSCALAR TRANSITION RATES FROM 104 MeV

 ALPHA-PARTICLE SCATTERING BY THE N=28 ISOTONES ${ }^{48} \mathrm{Ca},{ }^{50} \mathrm{Ti}$ and ${ }^{52} \mathrm{Cr}$R. Pesl, H.J. Gils, H. Rebe1, E. Friedman ${ }^{+}$, J. Buschmàn, H. Klewe-Nebenius, S. Zagromski, KfK-report 3378 (1983) and Z. Physik A313 (1983) 111

Precisely measured differential cross sections for elastic and inelastic scattering of 104 MeV alpha-particles by ${ }^{48} \mathrm{Ca},{ }^{50} \mathrm{Ti}$ and ${ }^{52} \mathrm{Cr}$ are reported. The analyses aim primarily at the determination of strength, radial shapes and deformation of the scattering potentials, looking for isotonic differences of $N=28$ isotones. The mean square radii of the (real) potentials are discussed in terms of mean square radius differences of the matter distributions. The isoscalar transition rates derived by coupled channel analyses of the measured cross sections are compared with electromagnetic rates. In addition to the analyses on the basis of a slightly generalized extended optical model a semi-microscopic deformed folding model has been applied, using a density-dependent effective alpha-bound nucleon interaction. Though an excellent description of the data over the full angular range is obtained the resulting values of the deformation parameters appear to be not consistent with results from various different methods.
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1.3.2 ISOSCALAR OCTUPOLE TRANSITION RATES IN ${ }^{50} \mathrm{Ti},{ }^{52} \mathrm{Cr}$, and ${ }^{208} \mathrm{~Pb}$ FROM MODEL-INDEPENDENT ANALYSES OF $104 \mathrm{MeV} \alpha$-PARTICLE SCATTERING
V. Corcalciuc ${ }^{+}$, H. Rebel, R. Pes1, H.J. Gils,
J. Phys. G: Nuc1. Phys. 9 (1983) 177

Applying a recently proposed method for model-independent analyses experimental differential cross sections for $104 \mathrm{MeV} \alpha$-particle scattering have been analyzed. Reliable values of isoscalar ( $0^{+}-3_{1}^{-}$) octupole rates in ${ }^{50} \mathrm{Ti},{ }^{52} \mathrm{Cr}$ and ${ }^{208} \mathrm{~Pb}$ are presented and compared with electromagnetic rates.

Table I Transition rates for the $0^{+} \rightarrow 3_{1}^{+}$transition in ${ }^{50} \mathrm{Ti},{ }^{52} \mathrm{Cr}$ and ${ }^{208} \mathrm{~Pb}$ resulting from the analysis of radial moments of the real transition potentials in $\alpha$-particle scattering and from electromagnetic methods

| Quantity | ${ }^{50 \mathrm{Ti}}$ | ${ }^{52} \mathrm{Cr}$ | ${ }^{208} \mathrm{Cb}_{\mathrm{Pb}}$ |  |
| :--- | :---: | :---: | :--- | :--- |
| $\mathrm{G}_{3}$ (s.p.u. | $12.1 \pm 0.5$ | $11.1 \pm 0.3$ | $43.0 \pm 1,3$ | RMA of FB <br> potentia1s |
| $\mathrm{G}_{3}$ (s.p.u.) | 11.7 | 9.8 | 37.6 | RMA of WS <br> derivative |
| $\mathrm{G}_{3}$ (s.p.u.) | - | $6.1 \pm 0.3$ | 38.4 | electrom. $(1,2)$ <br> $\mathrm{G}_{3}$ (s.p.u.) |

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1.3.3 FUSION AND NONFUSION PHENOMENA IN THE ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ REACTION AT 156 MeV
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H. J. Gils, H. Rebe1, S. Zagromski, Nucl. Phys. (in press)

Reaction products from ${ }^{6}$ Li-induced reactions on ${ }^{40} \mathrm{Ca}$ at 156 MeV have been studied using the $d E \times E$ identification as well as the inclusice $\gamma$-ray method. The complete fusion cross-section has been found to be $\sigma_{f}=(67+20)$ mb. The $Z$ distribution of fusion evaporation residues is compared with statistical model predictions. The $Z$ spectrum of reaction products shows a maximum at $15 \lesssim Z \lesssim 20$, probably due to transfer and to incomplete fusion. It is suggested that the small value of the fusion cross-section is due to the strong competition of ${ }^{6}$ Li break-up processes.

[^0]
### 1.3.4 LOCAL DENSITY APPROXIMATION IN EFFECTIVE DENSITY-DEPENDENT $\alpha$ NINTERACTIONS.

## H. J. Gils

In folding models of the real optical potential for nucleon, 1ightand heavy-ion scattering the effective projectile-nucleon interactions used show up to be density-dependent. For finite range interactions a local density approximation (LDA) has necessarily to be introduced the form of which is not a priori given and may influence the properties of the folded optical potentials.

The effective $\alpha N$-interaction $V_{\alpha N}\left(\vec{r}_{N}, \vec{r}_{\alpha}, \rho_{m}\right)$ used in single folding models for $\alpha$-nucleus scattering is divided into a density-independent factor $f(|\vec{x}|)$ depending on the distance $|\vec{x}|=\left|\vec{r}_{N}-\vec{r}_{\alpha}\right|$ between the $\alpha$ particle and the interacting nucleon and a density-dependent factor $v_{D D}\left(\rho_{m}\right)$

$$
\begin{equation*}
V_{\alpha N}\left(\vec{r}_{N}, \stackrel{\rightharpoonup}{r}_{\alpha}, \rho_{m}\right)=f(|\overrightarrow{\mathrm{x}}|) \cdot \mathrm{v}_{\mathrm{DD}}\left(\rho_{\mathrm{m}}\right) \tag{1}
\end{equation*}
$$

where $\rho_{\mathrm{m}}$ is the target nuclear matter density. For the radial form factor $f(|\vec{x}|)$ a Gaussian plus Yukawa form is used for convenience with strength (V) and range (a) parameters $V_{G}, a_{G}, V_{Y}, a_{Y}$. The density dependence has previously been parametrized as (1)

$$
\begin{equation*}
v_{D D}\left(\rho_{m}\right)=1-\gamma \rho_{m}^{2 / 3} \tag{2}
\end{equation*}
$$

with the parameter $\gamma \approx 2 \mathrm{fm}^{2}$ determined from fits to experimental data. Subject of the local density approximation is the question which local density $\rho_{m}(\vec{r})$ (i.e. the local density at which space coordinate $\vec{r}$ ) has to be used in Eq. (2). For practical purposes most frequently the forms

$$
\begin{align*}
& \mathrm{v}_{\mathrm{DD}}^{\mathrm{N}}\left(\rho_{\mathrm{m}}\right)=1-\gamma \rho^{2 / 3}\left(\overrightarrow{\mathrm{r}}_{\mathrm{N}}\right)  \tag{3a}\\
& \mathrm{v}_{\mathrm{DD}}^{\mathrm{N} / 2}\left(\rho_{\mathrm{m}}\right)=1-\gamma \rho^{2 / 3}\left(\frac{\overrightarrow{\mathrm{r}}_{\mathrm{N}}+\overrightarrow{\mathrm{r}}_{\alpha}}{2}\right) \tag{3b}
\end{align*}
$$

have previously been used $(1,2)$ and were assumed to be equivalent due to the short range of $f(|\vec{x}|)$. If this is true, then also the form

$$
\begin{equation*}
v_{D D}^{\alpha}\left(\rho_{m}\right)=1-\gamma \rho^{2 / 3}\left(\vec{r}_{\alpha}\right) \tag{3c}
\end{equation*}
$$

should be equivalent to Eqs. (3a,b). In addition, we introduce a new form of LDA, namely

$$
\begin{equation*}
v_{D D}^{M}\left(\rho_{m}\right)=\left[1-m \gamma \rho^{2 / 3}\left(\vec{r}_{N}\right)\right]\left[1-(1-m) \gamma \rho_{m}^{2 / 3}\left(\vec{r}_{\alpha}\right)\right] \tag{4}
\end{equation*}
$$

which allows a continuous change between Eq. (3a) ( $m=1$ ) and Eq. (3c) ( $m=0$ ) and also approaches Eq. (3b) closely ( $m=0.5$ ).

The different forms of a LDA (Eqs. 3,4) have been compared in analyses of elastic $\alpha$ particle scattering cross sections of ${ }^{40} \mathrm{Ca}$ at $\mathrm{E}_{\alpha}=104 \mathrm{MeV}$ in order to find out whether these approaches are equivalent in describing the data and reproducing the phenomenological potentials as previously assumed (1,2). The nuclear density $\rho_{m}$ calculated by Brown et al. (3) has been used for this purpose and the parameters $V_{G}, a_{G}, V_{Y}, \gamma$ were varied together with the imaginary part (Woods-Saxon form) of the optical potential. The results compiled in Tables $I$ and II clearly indicate that the different forms are not equivalent.

The variation of the parameters over the "mixing" parameter $m$ of the form (Eq. 4) is shown in Fig. 1. For $m=0.2$ all parameters of $f(|\vec{x}|)$ are independent of $m$ and $\gamma$ confirming the factorization of $V_{\alpha N}$ into $f(|\vec{x}|)$ and $v_{D D}$ (Eq.1). With the "mixed" LDA and the parameters obtained for $m=0.2$ seven other $1 \mathrm{f}_{7 / 2}$ - shell nuclei could be fitted (4) yielding much better results than the previously used LDA (1).

Table Ia Parameters of the effective $\alpha N$-interaction obtained from analyses of ${ }^{40} \mathrm{Ca}(\alpha, \alpha)$ using different forms of the LDA

| LDA | $X^{2} / \mathrm{F}$ | $\mathrm{V}_{\mathrm{G}}$ <br> Eq. | $\mathrm{a}_{\mathrm{G}}$ <br> $(\mathrm{fm})$ | $\mathrm{V}_{\mathrm{Y}}$ <br> $(\mathrm{MeV})$ | $\mathrm{a}_{\mathrm{Y}}$ <br> $(\mathrm{fm})$ | $\gamma$ <br> $\left(\mathrm{fm}^{2}\right)$ |
| :--- | :--- | :---: | :---: | :---: | :---: | :---: |
| 3a | 3.3 | 50.8 | 1.797 | 56.9 | 1.0 | 2.136 |
| 3b | 2.9 | 21.4 | 1.926 | 162. | 0.83 | 1.800 |
| 3c | 2.6 | 19.1 | 2.020 | 102. | 0.86 | 1.382 |
| 4 | 2.6 | 20.6 | 2.009 | 111. | 0.85 | 1.645 |

Table IIb Volume integrals $J$ and radial moments of the folded real optical potentials and of the WS imaginary potential as compared to the phenomenological FB -potential

| LDA | $X^{2} / \mathrm{F}$ | $-J_{V} / 4 \mathrm{~A}$ <br> $\left(\mathrm{MeV} \mathrm{fm}^{3}\right)$ | $\left\langle\mathrm{r}^{2}\right\rangle_{V}^{1 / 2}$ <br> $(\mathrm{fm})$ | $\left\langle\mathrm{r}^{4}\right\rangle_{\mathrm{V}}^{1 / 4}$ <br> $(\mathrm{fm})$ | $\left\langle\mathrm{r}^{6}\right\rangle_{\mathrm{V}}^{1 / 6}$ <br> $(\mathrm{fm})$ | $-\mathrm{J}_{\mathrm{W}} / 4 \mathrm{~A}$ <br> $\left(\mathrm{MeV} \mathrm{fm}^{3}\right)$ | $\left\langle\mathrm{r}^{2}\right\rangle_{\mathrm{W}}^{1 / 2}$ |
| :--- | :--- | :--- | :---: | :---: | :---: | :---: | :---: |
| $(\mathrm{fm})$ |  |  |  |  |  |  |  |

Fig. 1


Variation of the parameters of the effective $\alpha N$-interaction (Eqs.3,4) over the "mixing" parameter $m$ and corresponding values of $\chi^{2} / F$.
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### 1.3.5 $1 \mathrm{f}_{7 / 2}$ ISOTONE AND ISOTOPE DIFFERENCES OF NUCLEAR MATTER DENSITIES

H. J. Gils, H. Rebel, and E. Friedman ${ }^{+}$

The differences of the nuclear matter densities of $1 \mathrm{f}_{7 / 2}$ isotopes $40,42,43,44,48 \mathrm{Ca}$ and isocones ${ }^{50} \mathrm{Ti},{ }^{51} \mathrm{~V},{ }^{52} \mathrm{Cr}$ including two odd mass nuclei have been investigated by folding model analyses of elastic $\alpha$ particle scattering cross sections. The single-folding model analyses use a densitydependent effective $\alpha N$-interaction with an improved local density approximation (1). The nuclear matter densities $\rho_{m}$ were parametrized as the sum of an initial distribution function $\rho_{o}(r)$ (two parameter Fermi form) and a FourierBessel series (FB) (2)

$$
\begin{equation*}
\rho_{m}(r)=\rho_{o}(r)+\sum_{i=1}^{N} \beta_{n} j_{o}\left(\frac{n \pi r}{R_{c}}\right) \tag{1}
\end{equation*}
$$

The parameters $\beta_{n}$ were varied together with the imaginary part of the optical potential in order to fit the experimental cross sections. The values of $x^{2}$ per degree of freedom obtained in the fits were as small as in potential analyses using a "model-independent" FB potential. The resulting folded optical potentials were in excellent agreement with the $F B$ potentials (3) confirming the validity of the procedures used. As an example, Fig. 1 shows the nuclei matter density of 50 Ti extracted from the analyses including the error band (shaded area) obtained from the error correlation matrix (4).


> Fig. $1 \quad$ Nuclear matter density of 50 Ti with error band (shaded area) from FB-folding model analysis.

Differences of the nuclear matter densities as compared to the closed she 11 nuclei 40,48 Ca are shown in Fig. 2 for the isotope chain (upper row) and the isotone chain (lower row). The differences show a systematic behaviour of the even nuclei which is similar for isotopes and isotones. The additional nucleons seem not only to be built into the $1 \mathrm{f}_{7 / 2}$ shell but also into ra-
dial regions corresponding to neighbouring shells. The odd nuclei ${ }^{43} \mathrm{Ca},{ }^{51} \mathrm{~V}$ were found to have rms radii larger than those of their even neighbours (c.f. Table I), a fact which is also obvious from the shape differences (Fig. 2) (which peak at larger radii).


Fig. 2 Differences of nuclear matter densities of Calcium isotopes and ( $\mathrm{N}=28$ ) isotones with error band (shaded area)

Table I Root-mean-square radii of nuclear matter densities as obtained from elastic $\alpha$ particle scattering
Target $\quad{ }^{40} \mathrm{Ca} \quad{ }^{42} \mathrm{Ca} \quad 4^{44} \mathrm{Ca} \quad 4^{48} \mathrm{Ca} \quad{ }^{50} \mathrm{Ti} \quad{ }^{52} \mathrm{Cr}$
$\left\langle\mathrm{r}^{2}\right\rangle_{\mathrm{m}}^{1 / 2}(\mathrm{fm}) \quad 3.38 \pm .02 \quad 3.43 \pm .02 \quad 3.45 \pm .02 \quad 3.51 \pm .03 \quad 3.55 \pm .02 \quad 3.54 \pm .02$
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H. Klewe-Nebenius, B. Neumann, R. Pes1, G. Bechtold, KfK 2838 (1980)
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### 1.3.6 ISOSCALAR TRANSITION RATES AND SINGLE FOLDING MODEL WITH DENSITY-DEPENDENT FORCES

D.K. Srivastava ${ }^{+}$, H. Rebel

Isoscalar transition rates in nuclei are most interesting nuclear structure quantities as they are rather directly related to the nuclear transition densities. However, a great deal of confusion has prevailed over the last years about the comparability of these values obtained by electromagnetic procedures and the result derived from inelastic scattering of hadrons, in particular of $\alpha$-particles thus clouding the immense value of such studies. Having recognized the unsuitability of the Bernstein procedure (1) implicit folding model analyses $(2,3)$ on the basis of Satch1er's theorem have been applied in a number of studies to account for the disparity between the values obtained by electromagnetic methods and $\alpha$-particle scattering. Though the implicit folding model analysis proves to be by far more successful than the crude Bernstein procedure, a strong 1-dependent disparity remained largely unaccounted for. We have found that a proper accounting of the densitydependence of the $\alpha$-particle-bound nucleon interaction removes this difficulty and establishes the unique value of medium energy $\alpha$-particle scattering for the study of isoscalar transition rates.

In this semimicroscopic method the real part of the deformed optical potential is given by

$$
\begin{equation*}
\mathrm{U}_{\mathrm{R}}\left(\overrightarrow{\mathrm{r}}_{\alpha}\right)=\int \rho(\overrightarrow{\mathrm{r}})\left(1-\gamma \rho^{2 / 3}\left(\frac{\overrightarrow{\mathrm{r}}_{\alpha}+\overrightarrow{\mathrm{r}}}{2}\right)\right) v_{o} e^{-\left|\overrightarrow{\mathrm{r}}-\overrightarrow{\mathrm{r}}_{\alpha}\right| / \mathrm{b}^{2}} \mathrm{dr}^{3} \tag{1}
\end{equation*}
$$

where the values of $\alpha, v_{o}$ and $b$ are standardized by the fit to elastic $\alpha$ particle $-{ }^{40} \mathrm{Ca}$ scattering data. The target density $\rho$ is deformed by any reasonable procedure. Concentrating on its multipole moments $q_{1}^{\rho}$, defined by

$$
\begin{equation*}
\mathrm{q}_{1}^{\mathrm{f}}=\int \mathrm{f}(\overrightarrow{\mathrm{r}}) \mathrm{r}^{1} \mathrm{Y}_{10}(\hat{\mathrm{r}}) \mathrm{d} \overrightarrow{\mathrm{r}} \tag{2}
\end{equation*}
$$

the dependence on the deformation prescription is largely eliminated in these integral quantities. With the assumption that $\rho\left(\frac{\vec{r}+\vec{r}_{\alpha}}{2}\right) \approx \rho(\vec{r})$ within the range of the force, it can be shown that

$$
\begin{equation*}
\frac{\mathrm{q}_{1}^{\mathrm{u}}}{\mathrm{~J}}=\frac{\mathrm{q}_{1}^{\rho}}{\mathrm{A}} \cdot \mathrm{C}_{1} \tag{3}
\end{equation*}
$$

$J_{u}$ being the volume integral of the real optical potential and $C_{1}$ an 1 - and A-dependent correction factor as shown in Fig. 1.


Fig. 1 Correction of the normalized multipole moments of a deformed Fermi distribution due to the density dependence

For $\gamma=$ o or for a homogeneous distribution $\rho$ with vanishing surface diffuseness $C_{1}=1$. In realistic cases (say a Fermi distribution) $C_{1} \neq 1$ and, when included, it resolves the discrepancies previously found (3).
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1.3.7 LIMITATION OF THE $\alpha+{ }^{40}$ Ca FUSION CROSS SECTION
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J. Buschmanr, S. Zagromski, H.J. Gils, K. Grotowski ${ }^{++++}$, H. Rebel

In context of experimental studies of break-up-fusion reactions induced by $156 \mathrm{MeV}{ }^{6} \mathrm{Li}$ (1), we investigated separately the fusion of $\alpha-\mathrm{par}{ }^{-}$ ticles with ${ }^{40} \mathrm{Ca}$. In the interactions of the strongly clusterized ${ }^{6} \mathrm{Li}$ with the target some kind of incomplete fusion reactions are expected to con-
tribute dominantly (2). The $\alpha+{ }^{40} \mathrm{Ca}$ fusion cross sections have been measured at $E_{\alpha}=60,80$ and 104 MeV (LAB) by observing and analysing the inclusive $\gamma$-ray spectra emitted from the reaction products (3). The $\alpha$-particle beam of the Karlsruhe Isochronous Cyclotron was focused onto a $43 \mathrm{mg} / \mathrm{cm}^{2} \mathrm{na}-$ tural Ca target, and the beam current was measured by means of a well shielded Faraday cup located 6 m behind the reaction chamber (and using additional focussing after the target). Inclusive $\gamma$-ray spectra were registered by a Ge(Li) detector with $15 \%$ detection efficiency for the $1332 \mathrm{keV}{ }^{60}$ Co 1 ine. The $\gamma$-spectra containing about 260 lines each were analyzed with the computer code SAMPO and were used for identifying the reaction products with masses $A=9-43$. As statistical model calculations support, the main part of the reaction cross section (represented by the observed mass spectra with $\mathrm{A}<41$ ) can be attributed to fusion. The experimental results which are shown in Fig. 1 together with results for lower energies (4) and with the total reaction cross section derived from optical model analysis of elastic scattering $\left(\sigma_{R}{ }^{O M}\right.$ ) (5), provide information about that region of the fusion excitation function where the fusion probability decreases, limited by the competition of other reaction channels.

Due to the very close masses of the target and the compound nucleus $A=44$ it is difficult to identify the evaporation part well distinguished from target-like products originating from transfer processes followed by $\gamma-$ deexcitation. However, we can present upper limits for fusion cross sections by summing up the evaporation region in the mass spectrum.

Recently Matsuse, Arima and Lee (6) introduced a new concept of a critical distance of two colliding heavy ions where they lose their identity and start fusing. Using the (ground-state) nuclear ms radii

$$
\begin{equation*}
\left\langle r^{2}\right\rangle_{A}=\frac{3}{5}\left(1.12 \cdot A^{1 / 3}\right)^{2}\left[1+3.84 A^{-2 / 3}\right] \tag{1}
\end{equation*}
$$

the critical distance $\left\langle d^{2}>1 / 2\right.$ between two colliding nuclei $\left(A_{1}, A_{2}, A=\right.$ $A_{1}+A_{2}$ ) is defined by

$$
\begin{equation*}
\left.A<r^{2}\right\rangle_{A}=A\left\langle r^{2}\right\rangle_{A_{1}}+A_{2}\left\langle r^{2}\right\rangle A_{2}+\frac{A_{1} A_{2}}{A}\left\langle d^{2}\right\rangle \tag{2}
\end{equation*}
$$

Assuming the effective potential close to $\left\langle d^{2}\right\rangle 1 / 2$ being essentially a harmonic oscillator potential (modified by the Coulomb potential), the effective barrier height $E_{F E(J)}$ of the fused system and for the spin $J$ at the critical distance can be calculated. Adopting the proposed value (6) of the oscillator constant, the calculated cross section ( $\sigma_{\text {fulc }}^{c a l}$ ) and the fusion entry line $E_{F E}$ are displayed in Figs. 1 and 2, respectively.


Fig. 1 Fusion excitation function for the $\alpha+{ }^{40} \mathrm{Ca}$ system

For lower energies (see Fig. 2) the fusion barrier line EB appears to be a limit. It is clear that for higher energies none of the yrast lines, $E_{y}$ nor $E_{s y}(7)$, of the compound nucleus ${ }^{44}$ Ti*, does limit fusion cross sections. The result is in agreement with results obtained for the $\alpha+{ }^{233} 3_{U}$ system (6). The cross sections predicted by the model of Matsuse et al. for the fusion of $\alpha+{ }^{40} \mathrm{Ca}$ are much lower than our measured data. If these values are adopted as upper limits for complete fusion, the cross section difference has to be ascribed to multiparticle transfer reactions.

The critical distance calculated for the $\alpha+{ }^{40} \mathrm{Ca}$ system $\left\langle\mathrm{d}^{2}\right\rangle{ }^{1 / 2}=$ $r_{c r}=0.76$ fm agrees well with previously reported results for ${ }^{44} \mathrm{Ti} *$ revealing that the fusion cross section is limited in this case by compound nucleus properties (8).

Fig. 2
Fusion cut-off lines in the E*-J diagram of the compound nucleus ${ }^{44} \mathrm{Ti} *$ formed via $\alpha+{ }^{40} \mathrm{Ca}$. The fusion entry line $E_{F E}$, and the fusion barrier $E_{B}$ are calculated with parameters taken
from Ref. 6, $V_{c}$ is the Coulomb barrier, $V_{o}$ the effective potential at $r=<d 2>^{1 / 2}$. The yrast 1ines are calculated with the rigid-body moment of inertia, $E_{S Y}=$ $\mathrm{E}_{\mathrm{Y}}+\Delta \mathrm{Q}$ with $\Delta \mathrm{Q}=$ $(10 \pm 2.5) \mathrm{MeV}$
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### 1.3.8 INVESTIGATION OF THE ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ INTERACTION IN A COINCIDENCE EXPERIMENT

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It was recently found, that for the ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ reaction at 156 MeV (LAB) the complete fusion exhausts on 1 y a rather small part of the total reaction cross section (1), while ${ }^{6}$ Li break-up channels provide about $35 \%$ (2). Our present experiments study some particular contributions of incomplete fusion channels to the ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ total reaction cross section. For that purpose the $\gamma$-ray spectra emitted from heavy residual nuclei were measured in coincidence with those fragments which moved with the beam velocity of the ${ }^{6}$ Li projectiles. Detection angles were $\theta_{L}=9^{\circ}$ and $12^{\circ}$. Applying the standard procedure similar to that of Wilczynski et al. (3) we determined the cross sections for the $\left({ }^{6} \mathrm{Li}, \mathrm{p} \gamma\right),\left({ }^{6} \mathrm{Li}, \mathrm{d} \gamma\right),\left({ }^{6} \mathrm{Li}, \alpha \gamma\right)$ channels:

$$
\begin{array}{ccc}
+14 & 30+12 & 78 \\
-9 & 7 & 142 \\
-71
\end{array}
$$

We claim that the above reaction channels belong to a class of binary reactions called incomplete fusion. This statement is based on a proper selection of the experimental situation, first of all on the coincidence condition with beam velocity fragments of the projectile. The binary character of the investigated reactions is corroborated by a reasonable agreement with predictions of the Wilczyński "sum rule" model (3).

Incomplete fusion channels together with complete fusion exhaust a relatively small part of the total reaction cross section. The balance of all known reactions initiated by the ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ collision at 156 MeV indicates a missing cross section of the order of few hundred millibarns. The prediction of the "sum rule" model suggests that at least part of this cross section may be attributed to (deep-inelastic) pick-up reactions.
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1.3.9 INCLUSIVE $\gamma$-RAY SPECTRA EROM ${ }^{208} \mathrm{~Pb}$ BOMBARDED BY $73-156 \mathrm{MeV}{ }^{6}$ Li IONS
M. Albinska ${ }^{+}$, H. Klewe-Nebenius ${ }^{++}$, H. Rebel, H.J. Gils, J. Buschmann, S. Zagromski, J. Albinski ${ }^{+++}$

Projectile break-up in the field of the target nucleus followed by a fusion process of one of the fragments has been shown (1) to be a dominant mode in ${ }^{6} \mathrm{Li}$ induced nuclear reactions in the energy range above $15 \mathrm{MeV} / \mathrm{amu}$. In order to study these "internal break-up" processes in detail for the ${ }^{6} \mathrm{Li}+{ }^{208} \mathrm{~Pb}$ system, we started coincidence measurements between 1 ight charged ejectiles and $\gamma$-rays produced in the collisions. In a pre-experiment of the coincidence studies we investigated the inclusive $\gamma$-ray spectra after bombarding a $46 \mathrm{mg} / \mathrm{cm}^{2}$ thick ${ }^{208} \mathrm{~Pb}$ target by $73,84,93,107,113,125,130$ and 156 MeV (LAB) ${ }^{6}$ Li ions. The Ge(Li) detector ( $15 \%$ efficiency for the $1332 \mathrm{keV}{ }^{60}$ Co 1ine) was placed at $90^{\circ}$ to the beam axis in order to avoid Doppler broadening. Prompt, delayed and various background spectra were measured. The prompt spectra comprise about 250 1ines each. Up to now the spectra measured at 156 MeV were analysed using the computer code SAMPO. About $85 \%$ of the observed lines could be identified. As a by-product the $\gamma$-background arising from activation of the target chamber and detector shielding material has been carefully investigated, mainly originating from interactions of secondary neutrons with A1 and Fe.

From the $\gamma$-spectra the production cross sections for target residues (resulting from neutron-evaporation- $\gamma$-deexcitation) can be determined. In


Fig. 1 Mass spectrum of the reaction products when bombarding ${ }^{208} \mathrm{~Pb}$ by 156 MeV ${ }^{6}$ Li ions (cross section uncertainty in order of $15 \%$ ).
this way we may obtain global information about complete and incomplete fusion channels. Fission products, if existing, should be visible in the mass spectrum for $A \approx 100-110$. Within the accuracy of about $15 \%$, inherent to the in-beam $\gamma$-method we may neglect the compound nucleus deexcitation way by fission.

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(1) B. Neumann, J. Buschmann, H. Klewe-Nebenius, H. Rebe1 and H.J. Gils
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1.3.10 INTERNAL BREAK-UP CONTRIBUTIONS IN THE ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ INTERACTION AT 156 MeV
M. Albińska ${ }^{+}$, J. Albiński ${ }^{++}$, R. Planeta ${ }^{+++}$, H. Klewe-Nebenius ${ }^{++++}$, H. Rebel

Recent experimental investigations of nuclear reactions induced by $156 \mathrm{MeV}{ }^{6} \mathrm{Li}$ ions bombarding ${ }^{40} \mathrm{Ca}$ have demonstrated that the complete fusion comprises only $3 \%$ of the total reaction cross section while the reaction cross section measured by observing the $\gamma$-rays from the reaction products is considerably larger (1). The missing part can be ascribed to incomplete fusion. By further measurements (2) one of the expected incomplete fusion channe1s, the fusion of $\alpha$-particles at the beam-velocity corresponding to $156 \mathrm{MeV}{ }^{6}$ Li projectiles has been investigated. The available experimental information is compiled in Fig. 1 displaying the $Z$ spectra of the reaction products. These were derived by observation of the $\gamma$-rays after bombarding ${ }^{40} \mathrm{Ca}$ by $156 \mathrm{MeV}{ }^{6} \mathrm{Li}$ and $104 \mathrm{MeV} \alpha$-particles, respectively ( ${ }^{40} \mathrm{Ca}\left({ }^{6} \mathrm{Li}\right.$, x $\gamma$ ) and $\left.{ }^{40} \mathrm{Ca}(\alpha, x \gamma)\right)$ or by observing the recoiling target residues with large momentum transfer ( ${ }^{40} \mathrm{Ca}\left({ }^{6} \mathrm{Li}\right.$, recoils)) . We recognize that ( $\alpha+{ }^{40} \mathrm{Ca}$ ) fusion processes are a considerable contribution to that what has been observed in the ${ }^{6}$ Li induced reactions. The complementary contribution from ( $d+{ }^{40}$ Ca) fusion has not been studied experimentally. Statistical model calculations (using the LANCELOT code) predict the corresponding evaporation region slight1y shifted to larger A values as compared to the ( $\alpha+{ }^{40} \mathrm{Ca}$ ) case. It can be expected that the $\mathrm{d}+{ }^{40}$ Ca reaction will complement the cross section found in the region with $Z>16$. The $Z$ spectrum for $Z \geq 12$ results from different
process (not considered here) and is certainly affected by 0 and $C$ contaminations of the ${ }^{40} \mathrm{Ca}$ target.


Fig. 1 Experimental charge spectra for ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}(156 \mathrm{MeV})$ and $\alpha+{ }^{40} \mathrm{Ca}$ ( 104 MeV (LAB)) reactions. The inset indicates the angular momentum localization for complete and incomplete fusion processes in ${ }^{6} \mathrm{Li}+$ ${ }^{40} \mathrm{Ca}$ collisions at 156 MeV .

The critical angular momentum for the ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ system derived from the fusion cross section within the sharp cut-off approximation is equal to $1_{c r}^{\mathrm{Li}+\mathrm{Ca}}=8 \mathrm{~h}$ (1). In the same way the $1_{\mathrm{cr}}^{\alpha+\mathrm{Ca}}=16 \mathrm{~h}$ value was calculated. Transforming $1_{c r}^{\alpha+C a}$ to the $\mathrm{Li}+\mathrm{Ca}$ channel we get the angular momentum window $\Delta 1=8-24 h$ for the $(\alpha+C a)$ incomplete fusion reaction. Assuming that the remaining cross section in the evaporation region of the ${ }^{6} \mathrm{Li}+{ }^{40} \mathrm{Ca}$ interaction is due to fusion of deuterons, the corresponding angular momentum window would be $\Delta I^{\prime} \equiv 18-27 \mathrm{~h}$, overlapping with the window for $\alpha$-fusion. As the $1_{\text {max }}^{\mathrm{Li}+\mathrm{Ca}}=45 \mathrm{~h}$ (at 156 MeV ) the angular momentum window $\Delta 1^{\prime \prime}=27-45 \mathrm{~h}$ is left for other reaction mechanisms. It should be noted that the incomplete fusion cross sections may be somewhat overestimated due to inclusion of the cross sections for $Z=20$ and 21 which may originate also from direct reaction channels.

We note that the angular momentum window calculated within the Wilczynski-mode1 (3) and using experimental half-way radii of the trans-
ferred light particle do not agree with the above results. This is not unexpected for a light projectile like ${ }^{6}$ Li。
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### 1.3.11 DETERMINATION OF THE MULTIPOLARITY OF PROMPT ELECTROMAGNETIC TRANSITIONS FROM ANGULAR DISTRIBUTION OF CONVERSION ELECTRONS

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A spectrometer to measure the angular distribution of conversion eletrons in charged particle induced reactions is described. The spectrometer consists of a magnetic filter built with small permanent magnets and a $S i(L i)$ detector. The parameters of the instrument are discussed and attenuation factors due to the finite solid angles are evaluated. Spectra of conversion electrons following the reaction ${ }^{169} \mathrm{Tm}(\mathrm{d}, \mathrm{xn})$ and ${ }^{182} \mathrm{Os}(\mathrm{d}, \mathrm{xn})$ at an incident deuteron energy of $\mathrm{E}_{\mathrm{d}}=52 \mathrm{MeV}$ taken at different angles are shown.

The measured angular distributions of the conversion electrons is used to assign multipolarities E1, $\mathrm{M} 1, \mathrm{E} 2$ and ( $\mathrm{M} 1+\mathrm{E} 2$ ) to prompt $\alpha$-transitions.
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### 1.3.12 NEUTRON DECAY OF THE GIANT QUADRUPOLE RESONANCE IN ${ }^{90} \mathrm{Zr}$

W. Eyrich ${ }^{+}$, K. Fuchs ${ }^{+}$, A. Hofmann ${ }^{+}$, B. Müh1dorfer ${ }^{+}$, H. Rebel, U. Scheib ${ }^{+}$

The study of isoscalar giant resonances in heavy nuclei has been continued. In the meantime we completed the measurements and analyses of our ( $\alpha, \alpha^{\prime} n$ ) experiment on ${ }^{90} \mathrm{Zr}$ investigating excitation and decay of resonant strength in the region of the giant quadrupole resonance (GQR). One aim of the experiment is to extract model-independent branching ratios of the $n$-decay into the individual low lying states of ${ }^{89} \mathrm{Zr}$. Similar to ${ }^{207} \mathrm{~Pb}$ (1) the spins of these states act as a "filter" which allows a separation of strengths with multipolarities $1 \geq 4$ from E2 and EO strengths. This can be seen more quantitatively from statistical model calculations which are shown in Fig. 1.


Fig. 1
Relative branching ratios for the decay of EO, E2, E4, and E6 strenghts from the GQR region in ${ }^{90}$ Zr into the 5 lowest states in ${ }^{89} \mathrm{Zr}$ (calculated from the statistical model).

Above excitation energies of $E_{x} \approx 13.5 \mathrm{MeV} \mathrm{E} 0$ and E 2 strengths predominantly decay into the $1 / 2^{-}$and $3 / 2^{-}$"low spin" states whereas $E 4$ and E6 strength mainly feed the $9 / 2^{+}$"high spin" state. A representative experimental decay spectrum of the $G Q R$ region of 90 Zr is displayed in Fig. 2 a . The strong population of the $9 / 2^{+}$ground state indicates the existence of strengths with multipolarities $1 \geq 4$ in the observed energy region. In Fig. $2 b$ the strength distribution in ${ }^{90} \mathrm{Zr}$ decaying into the $9 / 2^{+}$and the $1 / 2^{-}$. states is shown, respectively. The strength decaying into the $1 / 2^{-}$and $3 / 2^{-\quad}$ state with a
predominant multipolarity $1=2$ is concentrated strongly around the known maximum ( $E_{x} \approx 14.5 \mathrm{MeV}$ ) of the $G Q R$ in ${ }^{90} \mathrm{Zr}$ whereas the strength decaying into the $9 / 2^{+}$state shows a relatively unstructured distribution. Fine structures as seen in the case of ${ }^{208} \mathrm{~Pb}$ (1) do not appear in these spectra.


Fig. 2a
Spectrum of the $n$-decay of the GQR region in ${ }^{90} \mathrm{Zr}$ into the individual states of $8^{89}$ Z.

(1)
H. Steuer et al., Phys. Lett. 47 (1981) 1702

Physikalisches Institut der Universität Erlangen-Nürnberg

### 1.3.13 DIRECT DECAY COMPONENT OF THE ISOSCALAR GIANT MONOPOLE RESONANCE IN ${ }^{90} \mathrm{Zr}$

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A powerful tool to get detailed information about decay properties of isoscalar giant resonances is the coincident observation of the decay particles after excitation through inelastic hadron scattering. In the preceding annual report we reported first results of an $\alpha^{\prime} n$ coincidence experiment to study the neutron decay of the isoscalar giant monopole resonance (GMR) in ${ }^{90} \mathrm{Zr}$. In the meantime this experiment and its analysis have been completed.

In order to obtain a model independent determination of a possible direct decay component we measured the neutron decay spectra of the GMR region for a maximum and a minimum of the excited resonant strength at 8 positions. In the case of a pure statistical decay this spectra should exhibit an evaporation shape. An experimental excess of fast neutrons corresponding to the decay into the low lying hole states of the residual nucleus gives an estimate of the direct decay component.

In Fig. 1 the spectra of the decay neutrons from the center of the GMR in ${ }^{90} \mathrm{Zr}$ are shown for a maximum (lower part) and a minimum (upper part)


Fig. 1 Neutron decay spectra of the reaction ${ }^{90} \mathrm{Zr}$ ( $\alpha, \alpha^{\prime}$ ) for an excitation maximum (lower part and minimum (upper part) of the GMR
of the excited resonant strength summed over the eight measured positions of the n-detectors together with the prediction for a pure statistical decay. The spectrum corresponding to the maximum of the GMR shows a significant deviation from the statistical curve at the high energy tail whereas the spectrum of the minimum is in good agreement with the evaporation curve. The quantitative analysis leads to a direct decay component of the resonant strength in the GMR region of $15 \pm 3 \%$. This value is very similar to our results on ${ }^{208} \mathrm{~Pb}$ and seems to be typical for the GMR in heavy nuclei.

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### 1.3.14 STUDY OF THE LOW ENERGY OCTUPOLE GIANT RESONANCE IN ${ }^{90} \mathrm{Zr}$ VIA ( $\alpha, \alpha^{\gamma} \gamma$ ) COINCIDENCE EXPERIMENTS

W. Eyrich ${ }^{+}$, K. Fuchs ${ }^{+}$, A. Hofmann ${ }^{+}$, H. Rebel, U. Scheib ${ }^{+}$, B.Müh1dorfer ${ }^{+}$

Within the last 10 years the existence of various new giant multipole resonances has been successively confirmed by electron and hadron scattering experiments. One of these excitation modes is the low energy octupole resonance (LEOR) which up to now has been studied in only few experiments. Due to its relatively low excitation energy ( $\mathrm{E}_{\mathrm{x}} \approx 30 \cdot \mathrm{~A}^{-1 / 3} \mathrm{MeV}$ in medium nuclei) the LEOR decays predominatnly by gamma emission.

Using the $104 \mathrm{MeV} \alpha$-beam of the Karlsruhe cyclotron we started an ( $\alpha, \alpha^{\prime} \gamma$ ) coincidence experiment to study the LEOR in ${ }^{90} \mathrm{Zr}$. In Fig. 1 a single $\alpha$-scattering spectrum is shown for an excitation maximum of the resonant strength in the LEOR region. From the measured angular distribution a predominant multipolarity of $1=3$ could be confirmed for the excitation region between 6 and 9 MeV . To get information about the microscopic structure of this strength branching ratios of the decay have to be deduced. In the present measurement gamma-rays with energies up to 3 MeV are detected in coincidence with the scattered $\alpha$-particles for 8 positions of the $\gamma$-detectors. In the $\gamma$-spectrum coincident to the LEOR region which is shown for one position of the $\gamma$-detectors in Fig. 2 the photo-peaks of various subsequent $\gamma$-transitions between low lying states in ${ }^{90} \mathrm{Zr}$ can be seen. An enhanced transition to these states, especially to the $2^{+}$state at $E_{x}=2.18 \mathrm{MeV}$ has to be associated with a coupling of the LEOR to nuclear surface oscillations. To investigate these structure aspects more quantitatively the spectra of all measured $\gamma$-positions have to be taken into account and side feeding cor-
rections have to be considered. This work is still in progress.



Fig. 1
${ }^{90} \mathrm{Zr}$ ( $\alpha, \alpha^{\prime}$ ) singles spectrum at an excitation maximum of the low energy octupole resonance (LEOR)

Fig. 2
$\gamma$-spectrum coincident to the region of the LEOR in ${ }^{90} \mathrm{Zr}$ from an ( $\alpha, \alpha^{\prime} \gamma$ ) measurement

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1.3.15 STUDY OF ISOSCALAR ELECTRIC GIANT RESONANCES BY $156 \mathrm{MeV}^{6}$ LiSCATTERING AND ( ${ }^{6} \mathrm{Li},{ }^{6} \mathrm{Li}^{\prime} \mathrm{n}$ )-COINCIDENCE EXPERIMENTS
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B. Mühldorfer, H. Schlösser

In our ( $\alpha, \alpha^{\prime} n$ ) coincidence experiments (1) investigating excitation and decay of the isoscalar electric giant resonances (GR) in ${ }^{208} \mathrm{~Pb}$ and ${ }^{90} \mathrm{Zr}$ it turned out that the values of the extracted percentages of the sum rules
and the deduced direct decay components of the different GR's depend to some extent on the ratio of the resonant strength to the underlying background and on the decay mode of the background. To get more information about these dependencies we started ( ${ }^{6} \mathrm{Li},{ }^{6} \mathrm{Li}$ ) -scattering experiments and ( ${ }^{6} \mathrm{Li},{ }^{6}{ }_{\mathrm{Li}}{ }^{\prime} \mathrm{n}$ )coincidence experiments on ${ }^{90} \mathrm{Zr}$ using the $156 \mathrm{MeV}{ }^{6} \mathrm{Li}$ beam of the Karlsruhe cyclotron. We measured ( ${ }^{6} \mathrm{Li}^{6},{ }^{6} \mathrm{Li}{ }^{\prime}$ )angular distributions on ${ }^{208} \mathrm{~Pb},{ }^{90} \mathrm{Zr}$ and ${ }^{24} \mathrm{Mg}$ using two telescopes consisting of surface barrier detectors. In Fig. 1 (upper part) an $\alpha$ and a ${ }^{6}$ Li scattering spectrum on ${ }^{90} \mathrm{Zr}$ are shown for an excitation maximum of the E2 and E0 giant resonances. Both spectra obvious1y differ essentially in their background. The sma11 background at higher excitation energies in the ${ }^{6}$ Li spectrum is mainly due to the lowlying break-up threshold of ${ }^{6}$ Li. The different background structure facilitates the study of the influence of the background on the extracted properties of the giant resonances. In the lower part of the figure a coincidence spectrum between ${ }^{6}{ }^{6}$ Li detector and a $n$-detector from the ( ${ }^{6}{ }_{\text {Li }},{ }^{6}{ }_{\text {Li }}{ }^{\prime} n$ ) experiment is displayed. Besides the peak of the prompt $\gamma$-quanta one can recognize the neutron peak of interest on a low background of accidental coincidences. The set-up of


Fig. 1 (top): $\alpha$ - and ${ }^{6}$ Li-scattering spectra from ${ }^{90} \mathrm{Zr}$ for scattering angles corresponding to a maximum excitation of the GQR. (bottom): Coincidence spectrum between a $6_{\text {Li-detector }}$ and a $n$-detector.
this experiment is similar to that of the previous ( $\alpha, \alpha^{\prime} n$ ) measurements (1). It proved to be suitable also for ( ${ }^{6} \mathrm{Li},{ }^{6} \mathrm{Li}{ }^{\prime} \mathrm{n}$ ) -coincidence experiments, which seem to be feasible without further difficulties.
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1.3.16 $\beta^{-\gamma}$ ANGULAR CORRELATIONS IN THE DECAY OF ${ }^{28} \mathrm{P}$
U. Scheib ${ }^{+}$, W. Eyrich ${ }^{+}$, H. Forke1 ${ }^{+}$, A. Hofmann ${ }^{+}$, D. Steuer ${ }^{+}$

One way to look for the possible existence of second class currents in nuclear $B$-decay is the measurement of the asymmetry coefficient $\varepsilon$ in $\beta-\gamma$ angular correlations of a $\beta$-decay followed by $\gamma$-radiation. For allowed transitions the correlation function $W$ is of the form:

$$
W\left(E_{\beta}, \theta\right)=1+\varepsilon\left(E_{\beta}\right)\left(3 / 2 \cos ^{2} \theta-1 / 2\right)
$$

For statements about the existence of second class currents, the measurements should be performed in isospin triplets. If for both B-transitions in those triplets the energy dependence of $\varepsilon$ is measured and the results are combined, the contributions of first class currents cancel and those from second class current are summed up. This is only true, however, if isospin impurity effects caused by the Coulomb interaction can be neglected. In reality, isospin impurity effects enter to some extent into the theoretical calculations for the different transitions. Therefore, only small A isospin triplet systems (where the theoretical wavefunctions are fairly well known and the impurity effects are small) are suited for those investigations. Promising candidates are the $A=20, A=24$ and $A=28$ systems. Since in the $A=20$ system the asymmetry coefficients of the $B-\gamma$ angular correlation have already been measured (1), we started to investigate the $A=24$ and $\mathrm{A}=28$ systems. The measurement of a new system can be expected to be more successful than remeasurements of the same system even with smaller error bars, because the model dependence of the theoretical calculations yields uncertainties which are of the same order of magnitude as the experimental ones.

The measurements of the $B^{-}$decay of the nuclei ${ }^{24} \mathrm{Na}$ and ${ }^{28} \mathrm{~A} 1$ have been done at the Erlangen tandem accelerator (2). The investigation of the cor-
responding $\mathrm{B}^{+}$-decays of ${ }^{24} \mathrm{Al}$ and ${ }^{28} \mathrm{P}$ are in progress at the Karlsruhe isochronous cyclotron. For experimental reasons we began with ${ }^{28}$ P.
${ }^{28} \mathrm{P}$ is produced via the ${ }^{28} \mathrm{Si}(\mathrm{p}, \mathrm{n})$ reaction with an incident proton energy of 18 MeV . Because of the short half-1ife of ${ }^{28} \mathrm{P}$ of only 268 ms we constructed a fast target handling system. In Fig. 1 the schematic detectortarget arrangement is shown. The target is moved pneumatically every second between the beam and the measuring position. During the measuring time the the beam is stopped inside the cyclotron in order to reduce background radiation. The detectors, two $3^{\prime \prime} \times 3^{\prime \prime} \operatorname{NaI}(T 1)$-crystals for the $\gamma$-rays and two plastic scintillator telescopes for the positrons, are situated symmetrically around the "beam off" position of the target.

Fig. 1 Schematic view of the
experimental setup.


With this setup we measured simultaneously two counting rates for an angle of 90 degrees between the positron and the coincident $\gamma$-ray and two counting rates for an angle of 180 degrees. The coincident events were stored event by event on magnetic tape (an event is defined by the energy of the positron and the number of the respective counter, the energy and detector number of the coincident $\gamma$-ray and the time difference between the detection of these two events (TPC-spectrum)). This enabled us to set energy and time windows afterwards thus avoiding possible systematic errors caused e.g. by gain shifts of the detectors.

The experiments and the data evaluation are not yet finished but from the results obtained so far from our experiments in Karlsruhe and Erlangen
we expect that we will be able to give new limits for the contribution of second class currents to the charged hadron current of the weak interaction.
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H. Forke1, Diplomarbeit Erlangen 1983

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### 1.4 NUCLEAR THEORY

### 1.4.1 CALCULATION OF THE MOMENTUM DISTRIBUTION OF THE ALPHA-DEUTERON MOTION IN ${ }^{6}$ Li

## R. Beck, F. Dickmann

Experimental data ( 1,2 ) on the momentum distriubtion of the $\alpha-d$ motion in the nucleus ${ }^{6}$ Li provide a valuable testing ground for theories (3) which aim at a study of clustering effects in this nucleus. In this contribution we report on a calculation of the $\alpha-$ motion in the ground and first excited state of ${ }^{6}$ Li in the framework of the Generator Coordinate Method (GCM) using as trial function a superposition of antisymmetrized ( $A$ ) ( $\alpha-\mathrm{d}$ ) and ( ${ }^{5} \mathrm{He}-\mathrm{p}$ ) two -cluster wave functions (4)
where

$$
\Phi_{\alpha-\mathrm{d}}\left(\underline{S}_{1}\right)=A \Phi_{\alpha}\left(S_{\alpha}\right) \quad \Phi_{\mathrm{d}}\left(\mathcal{S}_{\mathrm{d}}\right), \Phi_{5_{\mathrm{He}-\mathrm{p}}}\left({\underline{S_{2}}}\right)=A \Phi_{5_{\mathrm{He}}}\left(S_{5_{\mathrm{He}}}\right) \Phi_{\mathrm{p}}\left(\mathcal{S}_{\mathrm{p}}\right)
$$

The functions $\Phi_{\alpha}, \Phi_{d}, \Phi_{5}{ }^{\text {He }}$ and $\Phi_{p}$ are Slater determinants of harmonic os-
 leons are distributed in the lowest orbits. The operator $P$ (JM) projects onto the eigenspace of angular momentum (5). The generator coordinates ${\underset{\sim}{\sim}}_{1}$ and ${\underset{\sim}{2}}_{2}$
 corresponding to the clusterizations ( $\alpha-\mathrm{d}$ ) and ( ${ }^{5} \mathrm{He}-\mathrm{p}$ ), are determined by solving the Schrödinger equation with a phenomenological central two-body interaction (4) in the space of the functions of Eq. (1).

If we expand the overlap amplitude of the ${ }^{6}$ Li ground and first excited state with a state where the $\alpha$ particle and the deuteron are moving relative to each other with momentum $\underset{\sim}{k}$ in partial waves,

$$
\begin{equation*}
\left\langle\psi_{\alpha-d}(k, 1 \nu)\right| \psi_{6}{ }_{L i}{ }^{(L M)\rangle}=\underset{L M_{L}}{\Sigma} \phi_{L L}(k)\left(L M_{L} ; L M \mid 1 \nu\right) \sqrt{4 \pi} Y_{L M_{L}}^{*}(\hat{k}) i^{L} \tag{2}
\end{equation*}
$$

we obtain the momentum distribution

$$
\rho_{L}(\mathrm{k})=\sum_{\mathrm{L}} \quad \phi_{L \mathrm{~L}}^{2} \quad(\mathrm{k})
$$

The wave function $\Phi_{\alpha-d}(k, 1 \nu)$ where $(1 \nu)$ stand for the spin quantum numbers of the deuteron is written in the GC-representation

$$
\psi_{\alpha-d}(\underset{\sim}{k})=\int d \underset{\sim}{s} f(\underset{\sim}{s}, \underset{\sim}{k}) \Phi_{\alpha-d}(\underset{\sim}{s}), \quad f(\underset{\sim}{s}, k)=\text { const } \cdot \exp (i \underbrace{}_{\Omega}{ }^{k}+3 K^{2} /(8 R))
$$

where $B$ is the common width parameter of the single particle functions. The matrix elements in Eq. (2) as well as those of the Hamiltonian are evaluated employing the algebraic programming system REDUCE 2 (6).

The result for the ground state is shown in Fig. 1. The calculated momentum distribution seems to have a somewhat smaller width than that of the data. A reanalysis (7) of these data, however, shows that the experimental width should be reduced by $\sim 60 \%$.

The momentum distribution for the first excited state is shown in Fig. 2. The value $k=0.5 \mathrm{fm}^{-1}$ for $i$ ts maximum corresponds to an excitation energy of 3.9 MeV . This is in reasonable agreement with the caluclated excitation energy of 3.3 MeV .


Fig. $1{ }^{6} \mathrm{Li} \rightarrow \alpha+d$ momentum distribution for the ground state. The data are deduced from ${ }^{6} \mathrm{Li}(\mathrm{p}, \mathrm{pd}) \alpha$ experiments at 590 MeV (solid circles, Ref.1) and 155 MeV (open circles, Ref.2). The theoretical curve shows a diffraction minimum at $k=0.71\left(\mathrm{fm}^{-1}\right)$.


Fig. 2 Same as Fig. 1 for the first excited state.
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W. Dollhopf, C.F. Perdrisat, P. Kitching, W.C. Olsen, Phys. Lett. 58 B (1975) 425
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### 1.4.2 ELASTIC AND INELASTIC CHARGE FORM FACTORS OF ${ }^{6}$ Li: APPLICATION OF ALGEBRAIC PROGRAMMING TO CALCULATIONS IN THE GENERATOR COORDINATE METHOD (GCM)

R. Beck, F. Dickmann

Applications of the GCM in the framework of the nuclear cluster model are often hampered by the fact that the calculation of the many-body matrix elements gets complicated due to the requirements of the Pauli principle and of angular momentum projection. In this study (1), we show that both problems can be handled conveniently using the algebraic programming system REDUCE 2 (2).

As an example, we choose the model of ${ }^{6}$ Li introduced by Krivec and Mihailovic (3). The model space is described in contribution 1.4 .1 to this report. REDUCE 2 is applied for one- and two-body matrix elements as well as for many-body matrix elements using various cofactor expansions.

During these algebraic calculations we want to keep track of the analytic dependence of the matrix elements on certain quantum numbers of individual nucleons, e.g. the magnetic quantum number of a particle in a p-orbit. This information is needed when the many-body functions are projected onto the eigenspace of the operators of angular momentum and parity (4). This requirement is met by REDUCE 2 through the possibility of introducing certain operators depending on symbolic variables.

The interface between algebraic and numerical processing of expressions is a rather critical part of the whole calculation which needs some careful analytical work prior to the application of REDUCE 2. It is shown that the algebraic expressions obtained by REDUCE 2 may be recast into a form suitable as input to a FORTRAN program.

Results for the elastic and inelastic charge form factor of ${ }^{6}$ Li are shown in Figs. 1 and 2 using two different effective two-body interactions. The rms radius and $B(E 2)$ deduced from the limit values of the form factors as q tends to zero are 1 isted in Table I .

Table I rms radius and $B(E 2)$ for ${ }^{6}$ Li. V2 stands for force 2 of Ref.5, while $B$ stands for the Brink force of Ref. 6

|  |  | V 2 | B | experimental |
| :--- | :--- | :---: | :---: | :---: |
| $\sqrt{\left\langle\mathrm{r}^{2}\right\rangle}$ | $(\mathrm{fm})$ | 2.65 | 2.78 | 2.54 |
| $\mathrm{~B}(\mathrm{E} 2)$ | $\left(\mathrm{fm}^{4}\right)$ | 12.9 | 19.5 | 25.1 |



Fig. 1 The elastic charge form factor of ${ }^{6}$ Li. The full 1 ine corresponds to the force V2, the broken line to the force B. References for the experimental data may be found in Ref. 6 .


Fig. 2 Square of the inelastic charge form factor of ${ }^{6}$ Li. The various symbols are the same as in Fig. 1
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## 2. <br> LASER SPECTROSCOPY

2.1 HIGH RESOLUTION MEASUREMENTS OF ISOTOPE SHIFTS AND HYPERFINE STRUCTURE IN STABLE AND RADIOACTIVE LEAD ISOTOPES
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G. Meise1, H. Rebel, G. Schatz, B.A. Brown ${ }^{+}$,
J. Phys.G: Nuc1. Phys. 9 (1983) 443

We present measurements of isotope shifts and hyperfine structure in the lead resonance line for a total of 15 isotopes. The experimental accuracy is of order 4 MHz . Using independent measurements of the nuclear parameter $\lambda$ for the stable isotopes we have derived $\lambda$ for all measured isotopes. The derived $\lambda$ values are compared with various theoretical predictions for the lead nuclei. We also give values for the nuclear magnetic dipole and electric quadrupole moments deduced from our measurements.

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2.2 ISOTOPE SHIFTS AND HYPERFINE STRUCTURE SPLITTINGS OF THE $\left(6 \mathrm{p}^{2}{ }^{3} \mathrm{P}_{\mathrm{o}}-6 \mathrm{p} 7 \mathrm{~s}{ }^{3} \mathrm{P}_{1}{ }^{\circ}\right) \mathrm{PbI}$ RESONANCE LINE IN $196,197,197 \mathrm{~m}, 214 \mathrm{~Pb}$
M. Anselment, S. Göring, A. Hanser, J. Hoeffgen, G. Meise1, H. Rebel, G. Schatz, W. Faube1 ${ }^{+}$

The high-resolution measurements of isotope shifts and hyperfine structure in radioactive lead isotopes (1) have been extended using essentially the same laserspectroscopic methods and the experimental arrangement as described in Ref. 1 . The $196,197,197 \mathrm{~m}_{\mathrm{Pb}}$ nuclides were produced by $\alpha$-particle irradiation of enriched Hg and the samples were prepared with the aid of an electromagnetic mass separator. ${ }^{214} \mathrm{~Pb}(\mathrm{Ra} \mathrm{B})$ was collected as daughter product of a readily emanating ${ }^{226}$ Ra sample. The isotope shifts relative to ${ }^{208} \mathrm{~Pb}$ and other quantities are given in Table I.

With the calibration used in Ref. 1 we have extracted the values of the nuclear parameter $\lambda$ and the variation $\delta\left\langle r^{2}\right\rangle_{A-208}$ of the nuclear charge radii. A11 information on $\delta\left\langle\mathrm{r}^{2}\right\rangle_{\mathrm{A}-208}$ presently available, is displayed in Fig. 1 and compared with theoretical results. We note the conspicuous trend of the odd-even staggering of the charge radii.

Table I Results of the laserspectroscopic measurements

| A | Isotope shift <br> $(\mathrm{MHz})$ | Nuclear parameter <br> $\lambda\left(\mathrm{fm}^{2}\right)$ | A-factor <br> $(\mathrm{MHz})$ | B-factor <br> $(\mathrm{MHz})$ |  |
| :--- | :---: | ---: | :--- | ---: | :--- |
| 196 | $-11441(30)$ | $-0.572(16)$ | - | - |  |
| 197 | $-11402(19)$ | $-0.566(9)$ | -5324 | $(11)$ | 9 |
| 197 m | $-10827(17)$ | $-0.516(2)$ | $-1262.7(3.0)$ | -54 | $(39)$ |
| 214 | $+11500(99)$ | $+0.545(36)$ | - | - |  |



Fig. 1 Mean square charge radii of Pb isotopes as compared with theoretical predictions from the droplet model and Hartree-Fock calculations.
(1) R.C. Thompson, M. Anselment, K. Bekk, S. Göring, A. Hanser, G. Meise1, H. Rebel, G. Schatz, B.A. Brown, J. Phys. G: Nuc1. Phys. 9 (1983) 493
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### 2.3 PRECISE rf TUNING FOR Cw DYE LASERS USING OPTICAL SIDEBANDS

M. Anselment, S. Chongkum, H. Hoeffgen, G. Meisel,

The current measurements of hyperfine structure and isotope shifts of Sn and Pb (1) require to determine optical frequency differences of order $10^{7}$ to $10^{9} \mathrm{~Hz}$ at a level of $5 \times 10^{14} \mathrm{~Hz}$ to a high precision. Therefore the laser frequency must be stabilized to reduce frequency jitter. This is achieved by locking it to a stabilized optical reference resonator, resulting in a fixed frequency laser system. Among the techniques to restore tunability, only $r f$ methods are considered here since only these are precise enough for our purposes. One of them is the heterodyne method (2). Since this approach requires two complete dye lasers and allows only for limited accuracy the heterodyne setup was replaced by a sideband system (3). It has higher precision at increased dependability, and, since it requires only one dye laser, costs are reduced.

The sideband method uses an amplitude modulation of the laser beam with some radio frequency $\nu_{r f}$. The spectrum of the modulated light contains the unmodulated ("carrier") frequency $\nu_{d y e}$ as well as sideband frequencies ( $\nu_{d y e}+$ $\nu_{r f}$ ) and ( $\nu_{d y e}-\nu_{r f}$ ) (Fig.1). The side frequencies can be used in a variety of ways to achieve error free laser tuning (3).


Fig. 1 Frequency spectrum of the amplitude modulated laser light. The left sideband frequency $\left(\nu_{1 s}\right)$ and the right sideband frequency ( $\nu_{r s}$ ) are displaced by $\nu_{r f}$ off the carrier ( $\nu_{d y e}$ ). The dashed sinusoidal curve shows the time dependence of the left sideband if an additional frequency modulation (fm) $\Delta v$ on $v_{r f}$ is applied. The small horizontal arrows at $v_{1 s}$ and $\nu_{r s}$ indicate that the fm induced displacement variations are opposite in phase for the two sidebands.

In the present setup one of the sidebands is locked to the reference resonator in the same way as in the fixed frequency system (Fig. 2). If the modulating frequency is altered, the laser frequency stabilization influences the laser frequency in such a way that the selected sideband frequency remains constant since it is locked to the fixed reference etalon; in other words, a fixed side frequency with changing modulation frequency $\nu_{r f}$ leads to a laser frequency change by exactly the same amount. Thus error free rftuning of the dye laser frequency is achieved. The radio frequency $\nu_{r f}$ is generated under program control by a frequency synthesizer; to speed up laser lock, the processor simultaneously generates a feed-forward voltage which anticipates most of the expected feedback control signa1.

In our system a continuous scan width of 1 GHz is possible which is limited by the rf synthesizer and the driving rf power amplifier available. This range can be doubled by using both sidebands in turn.

Since altogether three optical frequencies are present in the modulated beam, it is necessary to ensure that the proper sideband is brought in


Fig. 2 Experimental setup of the laser sideband tuning system. The small graphs at the top show the frequency spectrum of the laser light before (right) and after (1eft) passing through the amplitude modulating unit.
coincidence with a transmission peak of the reference etalon and is then selected to lock the laser. For this purpose a new feature was introduced, namely a frequency modulation of $\nu_{r f}$ with speed $\nu_{f m}$ and width $\Delta \nu$. Thus only the sidebands but not the carrier are frequency modulated. Locking a sideband frequency to a peak of the reference etalon is accomplished by phase sensitive detection at $\nu_{f m}$ with a lock-in amplifier. Both sidebands can be distinguished from each other by their relative modulation phase which is $180^{\circ}$ (Fig.1). Thus one sideband is identified by the choice of lock-in-phase. Switching from one sideband to the other is then simply performed by inverting the phase-sensitive detector response.

Since frequency doubled laser light (UV) is used in the present experiments the rf controlled tuning width is effectively doubled, resulting in a total tuning range of 4 GHz .
(1) Preceding contributions 2.1 and 2.2 of this report.
(2) R.C. Thompson, M. Anse1ment, K. Bekk, S. Göring, A. Hanser, G. Meise1,
H. Rebel, G. Schatz, B.A. Brown, J. Phys. G: Nuc1. Phys. 9 (1983) 443
(3) B. Burghard, W. Jitschin, G. Meise1, App1. Phys. 20 (1979) 141

### 2.4 LASERSPECTROSCOPY OF Sn ATOMS

M. Anse1ment, A. Hanser, J. Hoeffgen, S. Göring, G. Meisel, H. Rebel,
G. Schatz

From the point of view of nuclear structure models currently under discussion there is a considerable interest in precise information about charge and matter distributions in nuclei around the magic proton number $Z=50$. We started laserspectroscopic measurements of isotope shifts and the hyperfine structure of the ( $5 \mathrm{~s}^{2} 5 \mathrm{p}^{23} \mathrm{P}_{\mathrm{o}}-5 \mathrm{~s}^{2} 5 \mathrm{p} 6 \mathrm{~s}^{3} \mathrm{P}_{1}, \lambda=286 \mathrm{~nm}$ ) transition in stable and radioactive Sn atoms. The results presently available are measured with an experimental method somewhat simplified as compared to the previously used heterodyning procedure (1). The frequency doubled light of a ring dye laser excited the resonance fluorescence in an atomic beam of Sn isotopes. The frequency of the dye laser is swept over the whole frequency range of interest by its internal scan mode. Synchronously with the laser a multichannel analyzer (in multiscaling mode) is scanned, recording the photon count rate (measured by a photomultiplier looking perpendicular to the atomic beam) versus laser frequency. Fig. 1 displays a spectrum obtained by this procedure which allows for a quick survey over the horizons of the


Fig. 1 Isotope shifts and hyperfine splittings of the stable Sn isotopes resonance lines. However, it is limited in accuracy due to a slight nonlinearity of the internal laser scan. The linearity of the measured spectra can be improved by linearisation via the transmission maxima (equally spaced in the frequency scale) of a high finesse cavity (spectral range ca. 300 MHz ). Table I gives the presently available results improved in this way. It includes the radioactive nuclide ${ }^{113} \mathrm{Sn}$ and compares with results from the Oxford group (2).

Table I Measured isotope shifts and values for the $A$ and $B$ factors of the odd isotopes

| Sn-Is otope | $\begin{gathered} \text { Positions [MHz] } \\ \text { this work } \end{gathered}$ | Isotope s this work | $\begin{array}{r} \text { hift }[\mathrm{MHz}] \\ \text { Oxford } \end{array}$ | A-Factor this work | $\begin{aligned} & {[\mathrm{MHz}]} \\ & \text { Oxford } \end{aligned}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 112 | -1219.3(50) | -1219.3(50) | -1214.4 (5) |  |  |
| $113 \mathrm{~F}=1 / 2$ | $3122.2(70)$ | $-1085.9(52)$ | - | -4208.1(66) | - |
| $\mathrm{F}=3 / 2$ | -3189.9(70) |  |  |  |  |
| 114 | - 897.4(50) | - 897.4(50) | $-894.9(13)$ |  |  |
| $115 \mathrm{~F}=1 / 2$ | 3594.2(50) | - 803.5(37) | -801.7(11) | $-4397.8(47)$ | -4395.4 (21) |
| $\mathrm{F}=3 / 2$ | -3003.2(50) |  |  |  |  |
| 116 | - 576.7(50) | - 576.7(50) | $-574.4(6)$ |  |  |
| $117 \mathrm{~F}=1 / 2$ | 4315.9(7.0) | - $471.5(41)$ | $-467.6(7)$ | -4787.4(57) | -4790.7(17) |
| $\mathrm{F}=3 / 2$ | -2865.3(50) |  |  |  |  |
| 118 | - 266-9(50) | - 266.9(50) | $-269.2(7)$ |  |  |
| $119 \mathrm{~F}=1 / 2$ | 4835.6(50) | - 176.8(37) | - $175.3(11)$ | $-5012.5(47)$ | $-5014.8(19)$ |
| $\mathrm{F}=3 / 2$ | -2683.1(50) |  |  |  |  |
| 120 | 0.0 | 0.0 | 0.0(10) |  |  |
| 122 | 238.4(50) | 238.4(50) | 234.1(7) |  |  |
| 124 | 439.7(50) | 439.7(50) | 441.6(7) |  |  |

Future measurements will use the much more precise "sideband" method, including further radioactive Sn nuclides.
(1) R.C. Thompson, M. Anselment, K. Bekk, S. Göring, A. Hanser, G. Meisel, H. Rebel, G. Schatz, B.A. Brown, J. Phys. G: Nuc1. Phys. 9 (1983) 443
(2) D. Stacey, private communication

### 2.5 A WAVEMETER FOR RAPID AND PRECISE WAVELENGTH MEASUREMENTS OF VISIble Laser light

A. Steiger, G. Meisel

To measure isotope shifts and hyperfine splittings in atoms the energy differences of interest have been determined as small changes of optical laser frequencies (1). In some projected experiments, however, the differences are so wide that the methods used so far are not applicable. A solution of the problem is to measure the full optical frequencies of each transition and to obtain the sp1ittings as small differences of them. Such biased measurements require a high level of accuracy, namely $10^{-7}$ or better. The appropriate method is optical interferometry. Among the various interferometric instruments ("wavemeters") a modified version of the Fabry-Perot wavemeter of Fischer et a1. (2) was chosen. Its concept, first described by Byer et a1. (3), is to use a set of Fabry-Perot etalons and a grating monochromator to determine the wavelength. A fast periodic readout of the wavemeter is achieved by monitoring the interferometer patterns by 1 inear photodiode arrays; their gignals are digitized and fed into an on1ine computer which calculates the actual wavelength value within less than a second.

The optical setup is shown in Fig.1. The incoming cw laser 1ight of about 1 mW power is split into several beams. A small part is imaged onto the entrance slit of the grating monochromator (GCA McPherson Instr. EU 700). The image of the entrance slit is monitored by the linear photodiode array Do (Fig.1); the array of type RL1024 G by Reticon contains 1024 photodiodes which cover over their length of 25 mm a spectral range of about 50 nm at a fixed grating position. The diode-to-diode spacing of $25 \mu_{\mathrm{m}}$ allows one to determine the wavelength to within about 100 gHz . Etalons No. 1 and 2 are solid fused silica etalons of 0.3 mm and 5 mm thickness; No. 3 has 100 mm airgap with a Zerodur spacer. All etalon surfaces are coated with a broadband dielectric coating of $95 \%$ reflectivity from 400 to 700 nm . A


Fig. 1 Schematic diagram of the wavemeter. DO to D3 are photodiode arrays.
system of cylindrical lenses in front of each etalon is designed to illuminate 3 to 5 interference rings. The interference patterns are projected onto the diode arrays D1 to D3. Fig. 2 shows a sample fringe pattern as recorded with an oscilloscope. These patterns are periodic with the laser frequency, the period is the free spectral range (FSR) $\delta \nu$ of the etalon. Therefore, an etalon pattern can fix the frequency only within its FSR. The monochromator result, however, allows one to select the correct period for etalon No. 1. In the same way, each etalon result serves to eliminate the ambiguity for the next thicker one. Thus, the precision to which the FSR of the third etalon is known essentially limits the accuracy of the final wavelength/frequency value.

The status of the instrument is as follows:

1. The interface electronics for reading the diode arrays and digitizing their output has been built and successfully tested.
2. The monochromator has been modified so that data taking by a diode array is possible.
3. Etalons No. 1 and 2 have been tested with their respective optics and diode arrays.
4. Special etalon No. 3 is being assembled.
5. The online computer (Nova 2 ) will be ready in its final configuration in the near future.
6. The optical setup is being constructed.


Fig. 2 Signals as recorded by a diode array when etalon No. 1 is illuminated with a He-Ne laser at 632.8 nm . The experimental line width is about three diodes.
(1) H. Rebe1, G. Schatz, Laser Induced Fluorescence Spectroscopy in Atomic Beams of Radioactive Nuclides, in Lasers in Nuclear Physics, ed. C.E. Bemis and H.K. Carter (Harward Acad. Publ., Harward 1982)
(2) A. Fischer, R. Kullmer, W. Demptöder, Optic Communications, 39 (1981) 277
(3) R.L. Byer, J. Paul, M.D. Duncan, A Wavelength Meter, Laser Spectroscopy III (Springer, Heidelberg 1977)
2.6 QUADRUPOLE MOMENTS OF ODD Pb ISOTOPES FROM HYPERFINE STRUCTURE OF THE $\left(6 \mathrm{p}^{2}{ }^{3} \mathrm{P}_{\mathrm{o}}-6 \mathrm{p} 7 \mathrm{~s}{ }^{3} \mathrm{P}_{1}{ }^{\circ}\right)$ ATOMIC TRANSITION IN PbI
J. Dembczynski, H. Rebe1

Recently (1), the hyperfine structure (hfs) of the ( $6 \mathrm{p}^{2}{ }^{2} \mathrm{P}_{\mathrm{o}}-$ $6 \mathrm{p} 7 \mathrm{~s}{ }^{3} \mathrm{P}_{1}{ }^{\mathrm{o}}$ ) transition in PbI has been precisely measured for a series of stable and radioactive Pb isotopes. While the $A$ factor can be calibrated in terms of the muclear magnetic dipole moments by use of the value known for 207

Pb from an independent measurement, the interpretation of the B factors in terms of nuclear quadrupole moments requires a detailed consideration of the atomic structure involved. In the framework of an analysis (2) of fineand hyperfine structure splittings and $g$-factors of the relevant atomic con-
figurations, the quantities necessary for a determination of the quadrupole moment from the $B$-factor could be semi-empirically deduced. The theoretical intermediate coupling expression of the $B-f a c t o r$ of a $6 p 7 s{ }^{3} P_{1}$ level is given by

$$
\begin{equation*}
B\left(6 p 7 s{ }^{3} P_{1}\right)=0.021311 b_{6 p}^{02}+0.004175 b_{6 p}^{11} \tag{1}
\end{equation*}
$$

where the negligibly small contribution from the $6 d$ electron due to mixing with the 6 p 6 d configuration is omitted. The effective hfs one-electron parameters $b_{6 p}^{k k}(\mathrm{MHz})$ are related (3) to the quadrupole moment $Q(b)$ and to $\left\langle\mathrm{r}^{-3}\right\rangle[$ a.u.] by

$$
\begin{equation*}
\left.b_{\mathrm{nl}}^{\mathrm{kk}}=234.974<\mathrm{r}^{-3}\right\rangle_{\mathrm{nl}}^{\mathrm{kk}} \cdot \mathrm{Q} \tag{2}
\end{equation*}
$$

( $k k=02,11$ for the $p$-electron).
Additionally, we know from relativistic atomic structure calculations (3)

$$
\begin{equation*}
\mathrm{b}_{6 \mathrm{p}}^{11} / \mathrm{b}_{6 \mathrm{p}}^{02}=\left\langle\mathrm{r}^{-3}\right\rangle{ }_{6 \mathrm{p}}^{11} /\left\langle\mathrm{r}^{-3}\right\rangle \frac{\mathrm{o}}{6 \mathrm{p}},=0.680 \tag{3a}
\end{equation*}
$$

and

$$
\begin{equation*}
\left\langle r^{-3}\right\rangle \frac{02}{6 p} /\left\langle r^{-3}\right\rangle \frac{01}{6 p}=0.8262 \tag{3b}
\end{equation*}
$$

Using the semi-empirical result

$$
\begin{equation*}
\left\langle r^{-3}\right\rangle \frac{01}{6 p}=20.3(1.2) \quad[\text { a.u. }] \tag{3c}
\end{equation*}
$$

derived from the observed magnetic-dipole hfs (2) we find

$$
\begin{equation*}
\left.\mathrm{Q}\left({ }^{\mathrm{A}} \mathrm{~Pb}\right)[\mathrm{b}]=-0.00868(52) \mathrm{B}\left({ }^{\mathrm{A}} \mathrm{~Pb}\right) ; 6 \mathrm{p} 7 \mathrm{~s}^{3} \mathrm{P}_{1}\right) \tag{4}
\end{equation*}
$$

With the B-factors given in Ref.1, we have the values in Table I which are about 40 \% larger than those quoted in Ref. (1).

Table I Nuclear quadrupole moments of Pb nuclei

| A | 199 | 201 | 202 m | 203 | 205 | 209 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $\mathrm{Q}[\mathrm{b}]+0.8(8)$ | $-0.09(90)+.58(15)$ | $+.10(10)$ | $+.23(6)$ | $-.27(16)$ |  |  |

(1) R.C. Thompson, M. Anselment, K. Bekk, S. Göring, A. Hanser, G. Meise1, H. Rebe1, G. Schatz, B.A. Brown: J. Phys. G: Nuc1. Phys. 9, (1983) 443
(2) J. Dembczynski, H. Rebe1, to be published
(3) J. Lindgren, A. Rosén, Case Stud. Atom. Phys. 4 (1974) 250
2.7 PRELIMINARY FINE- AND HYPERFINE STRUCTURE ANALYSIS OF THE ${ }^{8}$ S ( 7 p 7 s ) CONFIGURATION IN AmI

## J. Dembczynski

There is a number of experimental data (1) on the ( $\left.5 \mathrm{f}^{7}\right)^{8} \mathrm{~S}(7 \mathrm{p} 7 \mathrm{~g})$ configuration in the Am atom which provide a basis for constructing a level scheme (Fig.1).


Fig. 1
Preliminary energy scheme for AmI

In the space of the ( 7 p 7 s ) configuration the eigenvectors have been calculated (Table I) and some observed transitions could be identified as

Table I Energy levels and eigenvector composition of the $\left(5 E^{7}\right)^{8} S(7 p 7 s)^{2 S_{1}+1} P$; $2 S^{+1} P_{J}$ configuration in AmI (preliminary)

| Term | J | Energy $\left(\mathrm{cm}^{-1}\right)$ | Composition in \% |
| :---: | :---: | :---: | :---: |
| $\left(5 f^{7}\right)^{8} \mathrm{~S}(7 \mathrm{p} 7 \mathrm{~s})^{3} \mathrm{P} ;{ }^{10} \mathrm{P}$ | $7 / 2$ | 15608.15 | $91\left({ }^{3} \mathrm{P}\right)^{10} \mathrm{P}+8\left({ }^{3} \mathrm{P}\right)^{8} \mathrm{P}$ |
|  | 9/2 | 16511.69 | $83\left({ }^{3} \mathrm{P}\right)^{10} \mathrm{P}+15\left({ }^{3} \mathrm{P}\right)^{8} \mathrm{P}$ |
|  | 11/2 | 18300. ${ }^{\text {t }}$ | $100\left({ }^{3} \mathrm{P}\right)^{10} \mathrm{P}$ |
| $(7 \mathrm{p} 7 \mathrm{~s})^{3} \mathrm{P} ;{ }^{8} \mathrm{P}$ | 5/2 | 18429.09 | $84\left({ }^{3} \mathrm{P}\right){ }^{8} \mathrm{P}+39\left({ }^{3} \mathrm{P}\right){ }^{6} \mathrm{P}$ |
|  | $9 / 2$ | 19625. ${ }^{\text {t }}$ | $82\left({ }^{3} \mathrm{P}\right)^{8} \mathrm{P}+16\left({ }^{3} \mathrm{P}\right)^{10} \mathrm{P}$ |
| $(7 \mathrm{p} 7 \mathrm{~s})^{1} \mathrm{P} ; 8^{8}$ | 5/2 | 28056. ${ }^{\text {t }}$ | $\left.99\left({ }^{1} \mathrm{P}\right)^{8} \mathrm{P}+1{ }^{3} \mathrm{P}\right)^{6}{ }^{6}$ |
|  | 7/2 | 28009.78 | $97\left({ }^{1} \mathrm{P}\right)^{8} \mathrm{P}+2\left({ }^{3} \mathrm{P}\right)^{6} \mathrm{P}$ |
|  | 9/2 | 28480.35 | $96\left({ }^{1} \mathrm{P}\right)^{8} \mathrm{P}+3\left(^{3} \mathrm{P}\right)^{8} \mathrm{P}$ |
| $(7 \mathrm{p} 7 \mathrm{~s})^{3} \mathrm{P} ;{ }^{6} \mathrm{P}$ | 3/2 | 21644. ${ }^{\text {t }}$ | $100\left({ }^{3} \mathrm{P}\right){ }^{6} \mathrm{P}$ |
|  | 5/2 | 21239.91 | $84\left({ }^{3} \mathrm{P}\right){ }^{6} \mathrm{P}+15\left({ }^{3} \mathrm{P}\right)^{8} \mathrm{P}$ |
|  | 7/2 | 20654. ${ }^{\text {t }}$ | $59\left({ }^{3} \mathrm{P}\right){ }^{6} \mathrm{P}+37\left(^{3} \mathrm{P}\right)$ |

$t$ : theoretically predicted.
ground state transitions. Comparing the experimental levels with the calculated ones, some perturbation, most likely from the 7 p 6 d configuration, is observed, which will influence the hyperfine structure splitting and isotope shifts in these transitions. This result requires a more complex study of the $\left(5 f^{7}\right)^{8} S(7 p 7 s+7 p 6 d)$ configuration, if correct values for nuclear quadrupole moments will be extracted from hyperfine structure splitting measurements. In order to enable such studies highly precise hfs data of at least six ground state transitions in one isotope are necessary.
M. Fred, F.S. Tomkins, J. Opt. Soc. Am. 47 (1957) 1076

## 2. 8 A NEW PULSED AND TUNEABLE VACUUM-ULTRA VIOLET AND X-RAY LINE SOURCE

H. Poth and A. Wolf, Phys. Lett. 94A (1983)135

It is proposed to use the radiation emitted in electron capture by ions occuring in electron cooling of stored stable particle beams as a tunable photon source for vacuum-ultraviolet (VUV) and X-ray spectroscopy. By choosing a suitable ion beam and storage energy, photon energies from VUV to the X-ray region can be obtained. The photon line width can be made smaller than 0.2 eV and its energy can be tuned in steps smaller than $10^{-4}$. The photon source can be continuous or pulsed with a pulse duration of 100 ns and less and repetition rates of 1 MHz and more. With present electron cooling systems and storage-ring configurations, about $10^{8}-10^{12}$ photons can be produced into a $4 \pi$ solid angle. The corresponding intensities for observation in a solid angle of 1 msr range from $10^{4}$ in the VUV to more than $10^{8}$ in the X-ray region.
3. NEUTRINO PHYSICS
3.1 NEUTRINO PHYSICS AT THE RUTHERFORD SNS
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E. Finckh, W. Kretschmer, K. Stauber

University of Erlangen
N.E. Booth

Oxford University
J.A. Edgington

Queen Mary College, London
T.A. Gabriel, R.A. Lillie

Oak Ridge National Laboratory

The Karlsruhe neutrino project at the Rutherford SNS has now officially been approved by a signed agreement between the Kernforschungszentrum Karlsruhe and the Science and Engineering Research Council (SERC) represented by the Rutherford Appleton Laboratories (RAL).

A collaboration has been established between KfK and University of Karlsruhe, University of Erlangen, Oxford University, Queen Mary College London, Oak Ridge National Laboratory, and University of Tübingen.

The physics programme, properties of the $V$-source etc. have been described in detail in ref. 1.

### 3.1.1 THE NEUTRINO FACILITY AT THE RUTHERFORD SNS

Instead of using an underground cave below the SNS target station for the experimental neutrino area as described in the original proposal the neutrino detectors will now be installed in a neutrino block house at one side of the target station (see Fig.1). The distance to the V-source as well as the dimensions of the shieldings were matched approximately to the same conditions as for the cave solution. The design work on the block house plinth, an extra neutrino crane and an extension of the existing experimtal hall has been finished and construction work will start during October 1983.


NEUTRINO BLOCKHOUSE

## 

The order for about 6000 tonnes of iron for shielding purposes will be given in october 1983 too. The $v$-block house will consist of continuously cast iron slabs vertically installed into a mounting frame and staggered to minimize the free path for background radiation. A movable door of about 600 tonnes will allow access to the block house. Moving of the detector will be provided by an air pad support structure. The erection of the $V$-block house will take place during the second half of 1984 with support from RAL.

Fig, 1: Neutrino facility at the Rutherford SNS

The detector is supposed to provide an energy resolution in the order of $\sigma \approx 10 \% / \sqrt{E(M e V)}$, a time resolution of about 1 nsec and a spatial resolution of about 15 cm . This together with the time structure of the neutrinos from the pulsed spallation source should provide an unambiguous signature for the various events of interest.

Detailed considerations on various scintillator materials as well as experimental investigations have been carried out to find the most appropriate scintillator. Although more difficult to handle price/performance considerations finally led to the decision to use mineral oil based liquid scintillator with a light output of about $50 \%$ anthrazene. To achieve a maximum fiducial volume the detector will consist of one large tank with $60,000 \ell$ of 1 iquid scintillator. It will be segmented into 512 optically isolated bars of $18 \times 18 \mathrm{~cm}^{2}$ cross section and 350 cm length, viewed by photo tubes at each end (Fig. 2a). The optical segmentation consists of thin walls of tightly sealed double lucite sheets. The thin air gap in between provides totally internal reflection which guides the scintillation light to phototubes at the end of each bar (Fig. 2b). Gd-painted paper between the lucite double layer allows to detect neutrons which after thermalization produce $8 \mathrm{MeV} \gamma$-rays via the $\mathrm{Gd}(\mathrm{n}, \gamma)$ reaction. The time difference between the two


Fig. 2a: 50 ton liquid scintillation detector

1. 512 liquid scintillator modules
2. passive detector shield
3. active antishield
4. neutrino blockhouse
phototubes provides a spatial resolution of about 15 cm along the bar axis.
The whole detector is surrounded by an 18 cm 1 layer of iron and an active antishield consisting of scintillator slabs. This arrangement not only detects cosmic particles entering the detector but also promises effective suppression of neutral background radiation from outside which otherwise cannot be identified unambiguously. As this background mainly originates from cosmic bremsstrahlung in the last 20 cm layer of the shielding around the detector this may be identified by a detected charged particle entering this layer through the active anticounter.

## SCINTILLATOR INVESTIGATIONS

Different liquid and solid scintillators have been investigated concerning light output ard light attenuation. The light output was measured by the compton edge of a $\gamma$-spectrum of ${ }^{60}$ Co taking NE 110 (Nuclear Enterprise) for reference and is listed in Table 1.

TABLE 1
Light output for various solid and liquid scintillators in \% Anthrazene with reference to NE 110

a) Nuclear Enterprises, Edingburgh, Scottland
b) KSH Plastics, Geel, Belgium
c) Kyowa Gas, Tokyo, Japan
d) Röhm, Darmstadt

- The light attenuation normally given by the attenuation length (1/e fall off) is strongly dependent on geometry as the number of reflections and thus the reflectivity enters into the light output of long scintillator rods. Fig. 3 shows the attenuation curves for two acrylic based scintillators GS2037 (RÖHM) with different dimensions. The solid lines are calculations with an analytical method taking into account the various orders of reflection before reaching the phototube. The dashed curve is a MC-simulation (s. ch. 3.2.3). With a given reflectivity of the walls and light absorption in the scintilla-


Fig. 3:
Position dependent light output for acrylic scintillator bars (GS2037); solid line: analytical order of reflection calculation, dashed line: Monte Carlo simulation.


Fig. 4:
Monte Carlo calculation for position dependent light output a) liquid scintillator (NE235H). B) solid scintillator (GS2037).
tor both methods allow to calculate position dependent light output curves for any geometry.

Fig. 4 shows the MC calculations for a mineral oil based scintillator NE 235H (Nuc1ear Enterprise) and the GS2037 scintillator (RÖHM) both for realistic dimensions. No significant difference in the performance is to be seen.

The decision to use a mineral oil based scintillator was finally based on cost arguments. In addition GS2037 contains $15 \%$ oxy $^{-}$ gen which causes additional background reactions and decreases the number of ${ }^{12} \mathrm{C}$ and ${ }^{1} \mathrm{H}$ atoms. All other suitable polystyrene or polyvinyltoluene based solid scintillators like NE110 (Nuclear Enterprise), BC412 (BICRON) or SCSN2O (KYOWA) could not be taken into account as they are much too expensive.

## MONTE CARLO CALCULATIONS

The energy resolution mainly determined by the number of photoelectrons in the phototubes has been investigated with detailed Monte Carlo calculations and cross checked with experimental results from prototype measurements. The crucial parameters which besides the actual geometry enter into these calculations are the specific light output, the emission and absorption spectra of the scintillator, reflectivity of the scintillator walls, fractional scintillator area covered by the phototubes, photo ef-
ficiency spectrum and photoelectron collection efficiency. A critical point is the reflecitivity which has to be as good as possible as it enters exponentially into the 1 ight collection with an average of about 30 reflections before reaching the phototube. Due to the fabrication procedure of the totally internal reflecting sheets there are always some small blind areas along the corners of each bar which has to be taken into account. In addition it was necessary to consider the spectral response of emission and absorption of the scintillator light as well as the spectral photo cathode efficiency to understand the results from prototype measurements. Fig. 5 shows the MC results of the light out-


Fig. 5:
Position dependent light output for liquid scintillator (NE235H) $(15 \times 15 \times 350) \mathrm{cm}^{3}$ viewed by a $2^{\prime \prime}$ phototube; solid line: Monte Carlo calculation. put along a 350 cm long liquid scintillator (NE 235H) with totally reflecting walls with cross sections $15 \times 15 \mathrm{~cm}^{2}$ viewed by a $2^{\prime \prime}$ phototube. The deviation from a pure exponential decrease of 1ight output with distance from the phototube is not a geometry effect but stems from wavelength dependent light absorption which has been measured in a separate experimental setup. Agreement of the MC calculations with experimental data from triggered cosmics measured with the prototype detector is fairly good so that the properties of such a scintillator module are now quite well understood.

## PHOTOTUBES

To achieve optimum energy resolution due to photoelectron statistics one would like to cover as much of the scintillator surface as possible with photosensitive material. For the face area of a scintillator bar of $18 \times 18 \mathrm{~cm}^{2}$ cross section this would mean to use a phototube with a large area photocathode. Unfortunately the collection efficiency of a $5^{\prime \prime}$ phototube is only half as good as that of a $2^{\prime \prime}$ or $3^{\prime \prime}$ tube. Thus we decided to look at each face of a scintillator bar with two $3^{\prime \prime}$ phototubes with common HV supply and signal output. A complicated and expensive light guide could only provi-
de a light collection improvement by a factor of 1.2 and is therefore omitted. For a dynamic range of 500 for the detector signal a gain of $10^{6}$ and a linearity up to 100 mA is required for the phototubes. For a time resolution of about 1 nsec the rise time of the phototube should not be worse than 3 nsec. Various tubes of different manufactorors are under test now.

## OPTICAL SEGMENTATION

For optimum energy resolution any dead volume which could absorb the small amount of energy available has to be avoided. Therefore the optical boundaries of the 512 scintillator bars have to be as thin as possible. On the other hand they have to provide optimum reflectivity. 1 mm lucite sheets and even thinner plastic foils have been considered to use for totally internal reflecting double layers. Besides the lucite sheets only a Hostafan foil showed satisfactory reflectivity. The sealing at the edges of the double layer has to be minimized for optical reasons. On the other hand it is very delicate to handle as any small even invisable crack in the sealing causes the reflectors to be filled with liquid thus being demaged. Reliable sealing techniques as well as to find out easy and cheap fabrication procedures are under investigation.

ELECTRONICS AND DATA HANDLING

Design of the electronics for the scintillation detector has been worked out and first modules are being built and tested. Due to the characteristic signature of true events within a fairly high background the detector as a whole has to provide:
i precise measurement of the total energy absorbed
ii absolute event time with respect to the proton spill
iii relative timing of the individual bars contributing to an event
iv ability to detect a time coincidence within one bar with minimum delay and complete energy and time information (delayed coincidence signature).

Fig. 6 shows a block diagram of the electronics. High voltage for the phototubes is provided by a computer controlled HV system for 1024 individu al channels (LeCroy 1440). For each of the 512 scintillator modules the analogue signal of the phototubes (energy), the time difference between the two ends of the module ( $\mathbb{A}$ ), and the relative time with respect to the first fired bar ( t ) is stored within an event time of 70 nsec . Digitizing is ini-
tialized by a software controlled trigger processor after 300 nsec into fast $A D C$ 's with buffer memories. A multihit TDC initialized as well by the trigger processor measures the absolute event time. A second event following a first one even in the same scintillator module may be detected after a minimum deadtime of $3 \mu \mathrm{sec}$. Readout of the ADC - and TDC-buffer registers via "intelligent" CAMAC control takes place after 16 msec i.e. 4 msec before the next proton spill of the SNS. The "valid event" decision in the trigger processor is based on a "sum energy ( $\Sigma E$ )/multiplicity (M)" matrix mask set by RAM's. In addition a neutrino induced event has to show up in the first 9 usec due to the neutrino time spectra. Having detected such an event the trigger processor is switched to a "watch dog" mode for the second part of a delayed coincidence event and will reject any other event which concerning the $\sum E / M$-matrix do not promise to be a valid one. For sing le events the trigger processor applies reasonable scale down suppression factors. The purpose is to measure the background explicitely but to minimize the background related deadtime.

Detector control will be provided by a light pulser system with fiber optics fed by a $U V N_{2}-1$ aser. Logic control will be done by a software controlled event pattern pulser system.

Data acquisition and experiment control will be performed via CAMAC by a LSI $11 / 23$ microcomputer. A DMA interface connects the Q-Bus of the LSI $11 / 23$ with the UNIBUS of the VAX $11 / 750$. The VAX will be used for one line data handing and programme executing. 2 MByte memory and $2 \times 80$ MByte mass storage allows to store large data arrays and programs. The system is already operating and programme development has started.


Fig. 6:
Blockdiagram of the electronics for one out of 512 scintillator modules.

### 3.1.3 PROTOTYPE RESULTS

Two prototype liquid scintillator detectors have been built using totally internal inflecting sheets for optical segmentation. One is a 350 m long module with $15 \times 15 \mathrm{~cm}^{2}$ cross section viewed by a phototube at each end. The other is a $1000 \ell$ detector with 9 modules of $17 \times 17 \mathrm{~cm}^{2}$ cross section and 175 cm length each (Fig. 7). This detector is triggered by large area multiwire proportional chambers on top and bottom of the scintillator tank.


Fig. 7: 1000 \& 1iquid scintillator detector (NE235H)

1. scintillator tank with 9 optically isolated modules
2. large area proportional counters (cosmic trigger)

With this detector cosmic runs are planned to study in detail the exact position dependence of energy, time and spatial resolution as well as the correlation between neighbouring modules.

These informations will serve as a basis for the actual "detector Monte Carlo" programme. The prototype II detector which has gone under operation in August 83 will also be used to test the electronics and data hand1 ing system.

Measurements with prototype I have been performed with $2^{\prime \prime}$ phototubes. The position dependent light output for triggered cosmics has already been shown in Fig. 5.

Fig. 8 shows the pulse height spectrum for minimum ionizing particles


Fig. $8:$
Pulse height spectrum of m.i. particles passing a $(15 \times 15) \mathrm{cm}^{2}$ liquid scintillator module at 175 cm from a $2^{\prime \prime}$ phototube. Solid line: $X^{2}$ fit with Landau energy loss distribution and Gaussian resolution curve with $\sigma=8 \%$ 。
with energy loss of $\Delta E=25 \mathrm{MeV}$ seen by a $2^{\prime \prime}$ phototube at a distance of 175 cm from the position of energy absorption. The solid line represents a $X^{2}-f i t$ where the appropriate Landau distribution has been folded with a Gaussian resolution curve of $18 \%$ FWHM. Extrapolation to the 50 ton detector with two $3^{\prime \prime}$ phototubes at each end of a rscintillator module approaches roughly the design goal of $\sigma=10 \% / \sqrt{E(\mathrm{MeV})}$. The time resolution between the two phototubes at each end which has been corrected for amplitude walk was 1.5 nsec FWHM. Fig. 9 shows the measured time difference versus the position


Fig. 9 :
Time difference between phototubes at a 350 cm long scintillator module versus position of energy absorption.
Solid line: Theoretjcal distribution from Monte Carlo simulation.
of energy absorption together with the results for the average light path from a Monte Carlo calculation.

Further experiments with prototype $I$ are going on to test improvements of the optical segmentaticn as well as different phototubes.

### 3.1.4 THE LIQUID ARGON TEST DETECTOR

A 50 ton 1 iquid argon time projection chamber is planned to measure the differential cross section of the neutrino electron scattering.

In a small test device described in the annual report of 1982 signals from $\alpha$ particles of an ${ }^{241}$ Am source could be observed. In a drift chamber configuration with a grid in front of the anode the electron spectrum of ${ }^{207}$ Bi conversion source has been measured. Further developments to increase the drift gap up to $>10 \mathrm{~cm}$ and to test different electrodeconfigurations for time projection techniques are going on.
(1) R. Maschuw, B. Zeitnitz, Neutrino Physics at the Pulsed Spallation Neutron Source SNS, KfK 3362 (Jun. 1982) and Addendum (Nov. 1982)

### 3.2 INELASTIC SCATTERING OF NEUTRINOS FROM ARGON

S. Furui, Z. Phys. A308 (1982) 339

Institute for Theoretical Physics of the University of Tübingen

In the planned neutrino experiments with a liquid argon target only the electrons ${ }^{1)}$ are detected in the final state. Therefore, it's necessary to know the total and differential cross section for the reaction

$$
\nu_{\mathrm{e}}+{ }^{40} \mathrm{Ar} \rightarrow{ }^{40} \mathrm{~K}+\mathrm{e}^{-}
$$

We used the statistical model ${ }^{2)}$ to calculate total cross sections for allowed and first forbidden transitions and found that the quenching of the Gamow-Teller giant resonance can't be neglected. In the case of the Fermi transition, the results of the statistical model and the microscopic model of Donnelly and Peccei ${ }^{3)}$ are almost identical. We calculated the differential cross sections for allowed transitions in the microscopic model and normalize the intensity to get the identical total cross sections in both models. Because of the Lorentz transformations, the differential cross section of neutrino-electron scattering has a pronounced foreward peak in the lab system, in which respect it differs from the neutrino-nucleus reaction.
(1) R. Maschuw and B. Zeitnitz, Neutrinophysics at the Pulsed Spallation Neutron Source SNS, KfK 3364 (1982)
(2) N. Itoh, Y. Kohyama and Y. Fujii, Nuc1. Phys. A287 (1974) 501
(3) I.W. Donnelly and R.D. Peccei, Phys. Rep. 50C (1979) 1

## 3.3 <br> STUDY OF AN HIGH ACCURACY HELICITY MEASUREMENT OF THE ELECTRON NEUTRINO

A. Draut, P. Blüm, L.M. Simons

The theory of weak interaction is based on the $V-A$ hypothesis. Although the latter is generally accepted, some theorists have proposed a $V+A$ part, whose amplitude is believed to be less than $10 \%$.

A method to reduce this upper limit would be the measurement of the neutrino helicity

$$
H=\langle\vec{\sigma} \cdot \vec{p}\rangle / \sigma \cdot p
$$

with an accuarcy better than $1 \%$. The values of the helicity of the $\nu_{e}{ }^{-}$, measured up to now, have an accuracy of about $20 \%$ (1).

Therefore we have studied the feasibility of an experiment using the ${ }_{4}^{7} \mathrm{Be}$ decay:

$$
\begin{aligned}
&{ }_{4}^{7} \mathrm{Be}+\mathrm{e}^{-} \rightarrow{ }_{3}^{7} \mathrm{Li} \mathrm{~B}^{\text {弚 }}+\nu \\
& \rightarrow{ }_{3}^{7} \mathrm{Li}+\gamma
\end{aligned}
$$

Because $e^{-}$capture is mainly a two body decay (competing processes are of the order of $10^{-4}$ ) the neutrino's momentum is transferred to the ${ }^{7} \mathrm{Li}^{*}$ and it is polarized according to the spin of the neutrino. After the decay of Li* ${ }^{\text {* }}$ the $\gamma$ carries the information about the momentum as a Doppler shift of its energy and about the spin as a partial circular polarization.

The Be decay was analysed in order to extract the information about the momentum by means of nuclear resonance fluorescence (NRF) and of the circular polarization by scattering resonantly at a polarized Li target.

In order to achieve the NRF we need a high temperature of the source to increase the recoil velocity of $\mathrm{Li}^{+}$.

Due to the long mean lifetime $\left(10^{-13} \mathrm{~s}\right)$ of $\mathrm{Li}{ }^{*}$, large slowing down effects occur in solid and liquid sources. However we found good conditions in a gaseous source at low pressure. Because of the high temperature, the NRF cross section rises several orders of magnitude above the background.

In particular we investigated the use of the very volatile $\mathrm{BeCl}_{2}$ in an Ar atmosphere within a closed quartz tube. The high temperatures can be reached by electrodeless high frequency excitation.

Detailed calculations gave a11 parameters needed to obtain the neutrino's helicity from counting the $\gamma$ rays scattered resonantly from the Li target, polarized alternatiyely in or opposite to the $\gamma$ direction.

Thus the desired accuracy can be reached by measuring only several days with a source of an activity of 1 Ci.
(1) Goldhaber, Grodzins, and Sunyar, Phys. Rev. $109(1958) 1015$

## 4. INTERMEDIATE ENERGY PHYSICS <br> 4.1 LEPTONIC ATOMS

### 4.1.1 PERTURBED ANGULAR CORRELATIONS IN MUONIC ATOMS

## L.M. Simons

The idea to use perturbed angular correlations to determine the ratio of the line width $\Gamma$ and the hyperfine splitting energy $\Delta E_{F, F^{\prime}}$ of two hyperfine levels characterized by the quantum numbers $F$ and $F^{\prime}$ in muonic atoms was discussed several years ago by the author and was independently discovered by Batkin (2).

One makes use of the fact, that the time integrated attenuation coefficient $G_{k k}$ for the hyperfine interaction deviates from the hard core value in a rather simple way (3). Thus the effect to be measured is the deviation of the intensities in an angular correlation from the predicted fully disturbed values which is proportional to the interference term

$$
\sum_{F \neq F^{\prime}} \frac{(2 F+1)\left(2 F^{\prime}+1\right)}{2 J+1}\left\{\begin{array}{lll}
F F^{\prime} & \lambda \\
I ~ I ~ J
\end{array}\right\}^{2} \frac{1}{1+\left(\frac{\Delta E_{F F^{\prime}}}{\Gamma}\right)^{2}}
$$

where $I=$ quantum number of the nuclear $\operatorname{spin}, J=$ total quantum number of the muonic state without hyperfine interaction, and $\lambda=$ degree of orientation.

Usually $\mathrm{J}=3 / 2$ (muonic $2 \mathrm{p}_{3 / 2}$ state) should be taken for an angular correlation $3 \mathrm{~d} \rightarrow 2 \mathrm{p}_{3 / 2} \rightarrow 1_{\mathrm{s}}$ where only intensities are observed $(\lambda=2)$. Ideally the $3 \mathrm{~d} \rightarrow 2 \mathrm{p}_{1 / 2} \rightarrow 1_{\mathrm{s}}$ transition chain may serve as a calibration transition because it will be isotropic and can be used to disentangle systematic errors.

A modern coincidence arrangement will use planar Ge-detectors measuring the $3 \mathrm{~d} \rightarrow 2 \mathrm{p}$ transitions in coincidence with the $2 \mathrm{p} \rightarrow \mathrm{I}_{\mathrm{s}}$ transitions detected by CsF detectors. In order to keep accidental coincidences small a muon channel in a momentum mode with negligible micro time structure has to be used. The method is useable at first sight only for the measurement of magnetic moments in the cases where quadrupole moments are small. Usually a "normally" large quadrupole moment will lead to a complete splitting with $\Delta \mathrm{E}_{\mathrm{FF}}{ }^{\prime} \gg \Gamma$ 。

Fortunately enough for all nuclei with $I=3 / 2$ the quadrupole splitting of the $F=1, F^{\prime}=3$ levels is zero in first order approximation. Therefore muonic atoms with these nuclei are ideally suited to measure magnetic moments (Bohr-Weisskopf effect) or the influence of higher order terms of the quadrupole splitting.
(1) L.M. Simons, The depolarization of negative muons in muonic atoms, Proceedings of the IKO study week on physics with low energy beams of pions and muons, I.K.O. Ansterdam, (1978)
(2) I.S. Batkin, Sov. J. Nuc1. Phys. 31 (1980)3
(3) R.M. Steffen and K. Alder, Chapter 13 in The electromagnetic interaction in nuclear spectroscopy, (North Holland, Amsterdam, 1973)

### 4.1.2 THE QUADRUPOLE MOMENT OF ${ }^{97}$ Mo

W. Kunold, W. Oesterle, M. Schneider, L.M. Simons, R. Abela ${ }^{\dagger}$

In order to increase the accuracy of an earlier experiment (1) of our group a measurement of the $2 p_{3 / 2}$ hyperfine splitting of ${ }^{97}$ Mo was performed at the $\mu \mathrm{El-chann} 1$ at the Swiss Institute for Nuclear Research. Both the $3 \mathrm{~d} \rightarrow 2 \mathrm{p}_{3 / 2}$ and the $2 \mathrm{~s} \rightarrow 2 \mathrm{p}_{3 / 2}$ transitions have been observed with two high resolution intrinsic planar Ge-detectors. Special care was taken to avoid background events and neutron induced events. For the analysis muonic energies and relative intensities were computed by the programs MUON and RURP. One excited nuclear state had to be taken into account and no neutron induced background was found during the evaluation process in contrast to Ref. 1. The results are:

$$
\begin{array}{ll}
2 \mathrm{~s}_{1 / 2} \rightarrow 2 \mathrm{P}_{3 / 2} \text { transition: } & \mathrm{Q}=.218 \pm 008 \text { barn } \\
3 \mathrm{~d}_{5 / 2} \rightarrow 2 \mathrm{p}_{3 / 2} \text { transition: } & \mathrm{Q}=.221 \pm 012 \text { barn }
\end{array}
$$

The weigthed average is $Q=.219 \pm 007$ barn.
(1) P. Heitlinger, Präzisionsmessung des Kernquadrupolmoments von 42 Mo 97 mit Hilfe der Methode myonischer Atome, Dissertation Universität Karlsruhe (1980)
$+\quad$ Institut für Physik, Universität Basel

### 4.1.3 LASER-INDUCED ELECTRON-ION CAPTURE AND ANTIHYDROGEN FORMATION

R. Neumann ${ }^{+\quad}$, H. Poth, A. Winnacker ${ }^{+}$, A. Wolf, (1)

The influence of intense light on the recombination of electrons and ions during electron cooling has been considered in this work. It appeared that, under realistic conditions at LEAR, an enhancement of the proton-electron recombination rate by a factor of $10^{2}$ can be expected during a laser pulse sent along the common trajectory of electrons and protons. This enhancement is due to induced radiative electron capture which can be understood considering a proton within the electron gas as an atomic system in which only continuum states are occupied and which, hence, is strongly inverted.

As opposed to the spontaneous recombination already observed in earlier electron cooling experiments, the initial kinetic energies of the electrons involved, and the final levels of the atoms formed, are well defined in induced recombination by the energy of the photons sent in. The induced recombination rate, as a function of the laser frequency, significantly depends on the shape of the electron velocity distribution; therefore, its measurement appears very useful for the diagnostics of electron cooling.

Further, induced recombination might find an application in producing a pulsed beam of antihydrogen atoms in well defined atomic states provided a high quality positron beam would be available in a set-up similar to that used in electron cooling. An estimation, under realistic conditions at LEAR, considering positron sources available, reveals an antihydrogen intensity of a few atoms per second.

An experiment investigating electron-proton recombination with the LEAR electron cooler, in collaboration with a laser physics group from Heidelberg, is being prepared. The dye laser system to be constructed for this purpose sha11, in addition, be used for a Thomson scattering experiment with the LEAR electron cooler, before its installation at LEAR. As has been discussed in a separate report (2), the spectral analysis of photons scattered by the electron beam would allow to determine the spacecharge related energy profile of the beam at a count rate of $\sim 0.4$ photons/sec.
(1) dto., CERN-EP/83-66, submitted to Z. Phys.
(2) A. Wolf, unpublished report
$+\quad$ Physikalisches Institut der Universität Heidelberg

### 4.2 HADRONIC ATOMS

### 4.2.1 ANTIPROTONIC X-RAYS AT LEAR

L. Adiels ${ }^{+}$, G. Backenstoß ${ }^{++}$, I. Bergström ${ }^{+}$, P. Blüm, S. Charalambous ${ }^{++++}$, S. Dedoussis ${ }^{++++}$, K. Fransson ${ }^{+}$, A. Hancock, J. Hauth, H. Koch, T. Köhler, A. Kreiss1, A. Nilsson ${ }^{+}$, P. Pavlopoulos ${ }^{++}$, H. Poth, U. Raich, D. Rohmann, J. Repond ${ }^{++}$, M. Suffert ${ }^{+++}$, L. Tauscher ${ }^{++}$, D. Tröster ${ }^{++}$, K. Zioutas ${ }^{+++}$.

The aim of the experiment PS 176 at LEAR is the measurement of the strong interaction between antiprotons and nuclei at kinetic energies near zero. Furthermore, it is intended to investigate the $\gamma$-spectra of the residual nuclei in order to get information on the break up of nuclei after $\overline{\mathrm{p}}$-absorption. The measurement of high energy $\gamma$-rays in the energy region up to several hundred MeV aims at the search for deeply bound p -nuclear systems of considerably long life times.

The experiment is realized in a collaboration with Basel, Stockholm, Strasbourg and Thessaloniki, and in cooperation with Munich. During the past year the set up - as already presented in the last Annual Report (4.2.1) was built in the Karlsruhe workshops, tested and transported to CERN. The target ladder was added from Munich, and the telescope counters and moderators were supplied by Basel. An electronic scheme for simultaneous runs with six and more Si/Ge - diodes or NaI - detectors for promt and delayed coincidences was worked out. The data analysis system using a fast $\mu$-processor at the front end and our PDP 11/34 at the rear end was finally tested. In the spring 1983, the electronic hut was installed and the analog and digital electronics for the detector system was set up. A series of careful testo $n^{\circ}$ the diodes was performed resulting in an excellent timing behaviour for all energies of interest. In parallel the on- and off-line programs for the data acquisition system and the off-line evaluation were written.

For the first test run at LEAR with protons the complete system was set up. The few hours with protons were used to optimize the beam focus, to measure the range curve and to fix the relative timing between the counter telescope and the $X$-ray detectors.

In the first days of August antiprotons of $300 \mathrm{MeV} / \mathrm{c}$ momentum were available during a short period of about two hours in total. The system was tuned so well that not on1y range curves and beam dimensions could be deter-
mined but even first data on antiprotonic ${ }^{16} 0$-X-rays could be measured. Six diodes were in operation covering an energy range between 3 keV and 6 MeV and a time range from $0-16 \mu \mathrm{~s}$. In addition high energy $\gamma$-rays were measured with a $15 \times 10^{\prime \prime}$ NaI provided by Strasbourg, and a neutron spectrum was obtained using a special counter delivered from Stockholm. The first spectrum of antiprotonic atoms seen at LEAR is shown in Fig. 1. The peak to background ratio was very good because of the excellent on-line timing conditions. The com-

$$
\overline{\mathrm{p}}^{16} \mathrm{O}
$$



Fig. 1: $\overline{\mathrm{p}}^{-16} 0$ spectrum measurde at LEAR. About 20 minutes measuring time, $5000 \overline{\mathrm{p}} / \mathrm{s}$ in average.
parison with the former data and the fact that the $\overline{\mathrm{p}}$-rate delivered by LEAR was by more than a factor ten lower than the design value gives confidence for a successful and interesting measuring program of $\overline{\mathrm{p}}$-X-rays in the future.

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++ Institute for Physics, University of Basle, Switzerland
+++ Centre de Recherches Nucléaires and Université Louis Pasteur, Strasbourg,
    France
++++ Department of Nuclear Physics, University of Thessaloniki, Greece
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### 4.2.2 FEASIBILITY STUDIES FOR HIGH PRECISION MEASUREMENTS OF THE MASS AND THE MAGNETIC MOMENT OF THE ANTIPROTON

G. Büche, H. Koch, A. Kreiss1

The LEAR facility at CERN will. provide for high fluxes of antiprotons in the near future. This opens the way for a lot of experiments which can be tried for the first time or can be continued on a higher level of quality standards. To this second category can be attributed the intention for higher accuracies in the values of the mass and the magnetic moment of the antiproton. We have started to find criteria how to extend these accuracies in connection with the measurement of x -rays from antiprotonic atoms. On one hand, the discussion includes the energy resolution of the experimental equipment as well as the absolute and relative calibrations of photon energy, respectively, as far as the measurement is concerned. On the other hand, the evaluation of data with respect to the mass or the magnetic moment comprises a careful analysis of a whole series of electromagnetic corrections to the Dirac atomic energies as well as the influence of the strong interaction of antiprotons to the low lying levels. Until mid of 1983 several computer programs were made available at Karlsruhe which are suitable for calculations of the energies of atomic levels by integration of the quantum mechanical equations and of the intensities of X -ray components from a cascade model. A systematic study of antiprotonic $X$-ray transitions with respect to ( $Z, A$ ) of the nucleus and to the size of the various corrections is in preparation in order to minimize the expected uncertainties in a determination of $m-$ and $\mu \bar{p}$.

### 4.3.1 MEASUREMENTS OF $i T_{11}$ in $\pi-d_{\text {pol }}$ SCATTERING BETWEEN 140 AND 325 MeV

G.R. Smith, E.L. Mathie, E.T. Boschitz, M. Meyer ${ }^{+}$, F. Vog1er ${ }^{+}$
M. Daum ${ }^{++}$, S. Mango ${ }^{++}$, J.A. Konter ${ }^{++}$

Due to the importance of measurements of the vector analysing power $\mathrm{iT}_{11}$ in $\pi-\mathrm{d}$ scattering for the investigation of dibaryon resonances we have drastically enhanced the data taking rate by employing a different measuring technique and an improved polarized deuteron target.

The experiment was performed at the $\pi M 3$ area of the Swiss Institute for Nuclear Research. The polarized deuterium target was the same as for the earlier measurements of $\mathrm{iT}_{11}(1,2)$ except for the following improvements: The teflon cell containing the deuterated butanol target was replaced by a thin ( $100 \mu \mathrm{~m}$ ) brass cell in order to reduce the background and energy loss of the outgoing (recoil) deuterons. The corresponding deterioration of the observed deuteron NMR signal was compensated for by employing a FET in an impedance matching circuit contained physically in the cryostat (at a magnetic field of 2.5 T and a temperature of $0.5^{\circ} \mathrm{K}$ ). As usual the magnitude of the target polarization was determined from comparisons of the dynamic polarization NMR signal with the thermal equilibrium (natural polarization) NMR signal.

The particle detection system was quite different from that employed previously. In order to enhance the data taking rate we introduced a multiple detector time of flight "spectrometer". Pions were detected simultaneously at 6 angles in coincidence with their corresponding recoil deuterons. Scintillation counter telescopes were used for both pions and deuterons. The deuteron counters were followed by absorbers (whose thickness was adjusted for each setting to discriminate between recoil deuterons and the higher energy protons of background reactions) and a veto counter. In order to obtain optimum timing resolution, the pion and deuteron TOF scintillators were each viewed at both ends by phototubes. The pion solid angles in each arm was 30 msr , and the angular acceptance was $\Delta \theta_{\pi}=5^{\circ}$. Pions could be detected in the angular range $30^{\circ}$ to $140^{\circ}$ in $2.5^{\circ}$ steps, with the deuteron arms being adjusted correspondingly. In this new $\pi-$ d coincidence set-up, the following conditions were imposed on a


Fig. 1: The vector analysing power $\mathrm{iT}_{11}$ as a function of energy and c.m. angle.
$\pi-d$ signature: pion energy loss, deuteron energy loss, deuteron range, and the time of flight difference (TOF) between signals in the pion arm and the deuteron arm.

The results are presented in Fig. 1. The data taken ear1ier with the pion spectrometer are well confirmed except for the datum at 130 degrees at 256 MeV . As one can see, there is a clear trend of the data from 140 to 325 MeV . For angles larger than 100 degrees $\mathrm{i}_{11}$ changes little with
energy. Around 70 degrees however, a pronounced dip develops at the higher energies. This behaviour is not accounted for by predictions from Faddeev type calculations $(4,5)$ or by Glauber respectively impulse approximation calculations.
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(3) N. Giraud, Y. Avishai, C. Fayard, G.H. Lamot, Phys. Rev. C19 (1979) 465
(4) A.S. Rinat, E. Hamme1, Y. Starkand, A.W. Thomas, Nuc1. Phys. A329 (1979)285

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### 4.3.2 SEARCH FOR RAPID ANGULAR AND ENERGY DEPENDENCE OF $\mathrm{d} \sigma / \mathrm{d} \Omega$ AND $\mathrm{iT}_{11}$ IN LARGE ANGLE $\pi-\mathrm{d}$ SCATTERING

E.L. Mathie, G.R. Smith, E.T. Boschitz, M. Meyer ${ }^{+}$, F. Vogler ${ }^{+}$, M. Daum $^{++}$, S. Mango ${ }^{++}$, J.A. Konter ${ }^{++}$

Recent measurements of the tensor polarisation $t_{20}$ below the (3,3)resonance have produced puzzling results. Two sets of conflicting data exist at a pion laboratory energy near 142 MeV and $\Theta_{\mathrm{cm}}>120^{\circ}$. One set shows a tensor polarisation which varies relatively slowly with angle and is always more negative than -0.3 in this angular region. The other set displays a smooth relatively flat distribution at $T_{\pi}=117 \mathrm{MeV}$. With increasing energy a double hump distribution develops, becomes largest at $T_{\pi}=134 \mathrm{MeV}$ and reduces to a flat distribution at $T_{\pi}=151 \mathrm{MeV}$. Most of these data have positive values.

If the latter data set is correct a strong angular and energy dependence may also show up in the observables $\frac{d \sigma}{d \Omega}$ and $i T_{11}$. Therefore, we have made accurate measurements of the angular and energy dependence of $i_{11}$ and the differential cross section between $T_{\pi}=117$ and 15.1 . MeV. The main goal was to provide a broader, more restrictive data base for realistic phase shift analyses. It was further hoped that the energy and angular behaviour of $i T_{1]}$ in this region would help to indirectly confirm or rule out one of the conflicting $t_{20}$ data sets. The experimental results are presented in Fig. 1 and Fig. 2.

Neither the differential cross-section nor the vector analysing


Fig. 1: Differential cross sections for $\pi-d$ elastic scattering


Fig. 2: Vector Analysing power $\mathrm{iT}_{11}$ for $\pi-\mathrm{d}$ elastic scattering
power show within the measured accuracy rapid changes as a function of angle and energy in the backward hemisphere.

Obviously, the question arises whether or not the smooth dependence of $i T_{11}$ and $d \sigma / d \Omega$ together with the oscillating nature of the SIN $t_{20}$ angular distribution can be reproduced by any set of reasonable phase shifts. The only published attempt of doing so (1) found numerous solutions and predicted oscillatory behaviour in $\mathrm{iT}_{11}$ which is not observed at 134 MeV . It remains to be seen if the extensive $i T_{11}$ and $d \sigma / d \Omega$ presented here will help further reduce the number of solutions.
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### 4.3.3 POSITIVE PION ABSORPTION ON DEUTERIUM AT ENERGIES ABOVE THE 3,3 RESONANCE

E.L. Mathie, G.R. Smith, E. Boschitz, J. Hoftiezer, M. Meyer ${ }^{+}$

The experiment was performed at the $\pi M 3$ pion channel at the Swiss Institute for Nuclear Research (SIN). A coincidence technique requiring the detection of the incident pion and both outgoing protons was used. The incident pions were detected with a plastic scintillation detector and one of the three elements of a hodoscope detector. A coincidence requirement with the cyclotron rf effectively eliminated the muon and electron contamination in the beam. The forward going (higher energy) proton was was detected by two scintillation detectors (F1 and F2). The F1 detector simply constrained the particle direction. The F2 detector both defined the solid angle, and through the use of two phototubes, constant fraction discriminators, and mean timing, enabled an excellent time of flight measurement for the forward going proton. The backward going proton was detected by two similar detectors (R1 and R2), the latter also arranged to give an optimum timing signal. The time of flight difference was then calculated for the forward and recoil particles in the F2 and R2 detectors respectively. This difference was found to have a 450 psec FWHM, and separation of the $\pi d \rightarrow 2 p$ (free absorption with the quasifree absorption reaction part of the same peak), $\pi p$ quasielastic, and $\pi d$ quasielastic was good. In all cases the measurements were made with a $\mathrm{CD}_{2}$ target and with a $\mathrm{CH}_{2}$ target, to enable an explicit background subtraction.

The number of $\pi d, 2 p$ events was determined for each case from the histogram of the time of flight difference for the outgoing protons,

Fig. 1: Center of mass system differential cross sections for $\pi d \rightarrow p p$. The error bars shown are typical of all the data and the curves are the empirical fit to the data with a series of even order Legendre polynominals

$$
\frac{d \sigma}{d \Omega}=\frac{1}{4 \pi} \sum_{i=0}^{3} a_{2 i} P_{2 i}\left(\cos \Theta^{*}\right)
$$




Fig. 2: The $a_{o}$ fit parameter as a function of incident pion energy with data from other recent experiments; the empirical fits of Spuller and Richard-Serre; and various theoretical calculations. The calculations of Chai and Niskanen clearly do not describe the data. The most successful theory (Bankleider) does not adequately describe the data in the peak region.


Fig. 3: The ratios $a_{2 i} / a_{0}$ as a function of incident pion energy. The meaning of the curves is the same as in Fig. 5. The failures of the most successful theory (Blankleider) are reflections of the underestimate of $a_{o}$ in the energy region $80-180 \mathrm{MeV}$.
suitably corrected for quasifree events from the carbon in the $\mathrm{CD}_{2}$ target (based on the background measurements). Taking into account various corrections due to detector efficiency, Coulomb multiple scattering, pion decay etc. the data presented in Fig. 1 were obtained. In Figs. 2
and 3 the fit parameter $a_{0}$ and the ratio $a_{2} / a_{0}$ are shown for our data in comparison to earlier data and theoretical predictions.
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4.3.4 ABSORPTION OF $\pi^{+}$AND $\pi^{-}$IN ${ }^{3} \mathrm{He} \mathrm{AT} 120 \mathrm{MeV}$
S. Cierjacks, S. Ljungfelt, U. Mankin, G. Schmidt, H. Ullrich, G. Backenstoss ${ }^{+}$, M. Izycki ${ }^{+}$, M. Steinacker ${ }^{+}$, P. Weber ${ }^{+}$, H.J. Weyer ${ }^{+}$, M. Furic ${ }^{++}$, T. Petkovic ${ }^{++}$

Our understanding of pion absorption in nuclei is mainly based on the knowledge of the absorption process in deuterium. Kinematically complete experiments on ${ }^{3}$ He give access to new observations: 1) Absorption on a $T=1$ nucleon pair; 2) The influence of a third nucleon.
Experimentally the two points can be well separated. Quasi-free absorption process (QFA) and those with final state interaction (FSI) take place under different kinematical conditions. Also, events where three nucleons share comparable amounts of energy show a special kinematical signature.

A surprisingly large value for the $Q F A$ ratio $R=\left(\pi^{-} n p \rightarrow n n\right) /\left(\pi^{-} p p \rightarrow n p\right)$ has been discovered for the first time in our previous experiment with stopped $\pi^{-}$(1). The representation of the data in the form of a Dalitz plot has also revealed pronounced FSI effects. Both observations urged a continuation of the measurements with pions in flight.

The measurements have been performed on the $\pi E 1$-channel at SIN. Incoming pions with $220 \mathrm{MeV} / \mathrm{c}$ momentum have been identified by a conventional counter telescope. A 1iquid ${ }^{3}$ He target has been used. Emitted neutrons, protons and deuterons have been detected in a large area ( $2 \mathrm{~m} \times 0.5 \mathrm{~m}$ ) position sensitive TOF counters capable of particle identification. Charged particles have also been detected in a large total absorption hodoscope consisting of a $4 \times 3$ array of plastic scintillators ( $17 \times 1 \times 30$ $\mathrm{cm}^{3}$ each) in connection with two MWP chambers for track identification.

The following results have been obtained:
a) As for stopped ions, a strong enhancement of the events appears on the perimeter of the Dalitz plot. Since the measurement did not yet cover the complete kinematically allowed region, the existence of more
events in the middle of the plot cannot strictly be excluded.
b) The number of events in the FSI-region is smaller than in our experiment with stopped pions and also smaller than in the preliminary TRIUMF-data at 65 MeV (2).
c) An important results is the value of $R^{\prime}=\left(\pi^{+} n p \rightarrow p p\right) /\left(\pi^{-} p p \rightarrow n p\right)=$ $32 \pm 11$. This quantity corresponds completely to R.Ets observed value is much higher than 6.5 as expected from isospin considerations alone.

First theoretical calculations for QFA using the isobar model have been published recently by Lee et al. (37, Toki et al. (5), and Silbar et a1. (5). All authors give explanations for the $T=0$ enhancement over the $T=1$ absorption but seem to overestimate the effect, at least at lower energies. A summary of the experimental and theoretical results on R is given in Fig. 1.


Fig. 1: The dependence of $R$ on pion energy. Experimental points and theoretical curves are shown.

In conclusion we can say that we confirmed the dominant role of QFA in ${ }^{3}$ He also for pion absorption in flight. The observed isospin dependence, however, is still not well understood.
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4.3.5 MEASUREMENT OF THE ${ }^{12} \mathrm{C}\left(\pi, \pi^{1} \gamma\right){ }^{12} \mathrm{C}\left(2^{+}, \mathrm{T}=0 ; 4.44 \mathrm{MeV}\right)$

ANGULAR CORRELATIONS: A TEST OF THE ISOBAR-HOLE MODEL
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It has been shown that the effects of the nuclear medium on pionnucleon interactions in nuclei cannot be neglected in the analysis of elastic and inelastic pion scattering from nuclei. At pion energies near the ( 3,3 )-resonance the pion-nucleus interaction is dominated by the formation, propagation, and decay of intermediate $\Delta$-hole states.

Recently, it was pointed out that the double-differential cross sections for ( $\pi, \pi^{\prime} p$ ) quasi-elastic scattering from ${ }^{16} 0$ (to ${ }^{15} \mathrm{~N}$ states in the continuum) are fitted much better with the $\Delta$-hole model than in the standard closure approximation to the first order theory (static-model). Here, we report on the first measurement, and the analysis of ( $\pi, \pi \gamma$ ) angular correlation functions for the transition to a discrete state in ${ }^{12} \mathrm{C}\left(2^{+}, 4.44 \mathrm{MeV}\right)$. The double-differential cross section $\mathrm{d}^{2} \sigma / \mathrm{d} \Omega_{\pi} \mathrm{d} \Omega_{\gamma}$ for ( $\pi, \pi \gamma$ ) contains information on the magnetic substate populations (which cannot be obtained from $\left(\pi, \pi^{\prime}\right)$ ) and is therefore expected for provide more detailed information about the reaction mechanism than a single differential cross-section.

The experimental results are shown in Fig. 1. The errors in $\mathrm{d}^{2} \sigma / \mathrm{d} \Omega_{\pi} \mathrm{d} \Omega_{\gamma}(\sim \pm 27 \%$ were derived from the uncertainties in background subtraction, dead time losses, $\gamma$-detector efficiency and statistics. The horizontal "error bars" simply indicate the angular width of the $\gamma$-detector


Fig. 1: Double differential cross sections at $\phi_{\pi}=40^{\circ}$ (upper part) and $\phi_{\pi}=20^{\circ}$ (lower part) for the ${ }^{12} \mathrm{C}\left(\pi^{+}, \pi^{+1} \gamma_{4.44}\right)$ reaction at $T_{\pi}=162 \mathrm{MeV}$. Data points: this experiment. Solid 1ines: full $\Delta$-hole mode1 predictions (including L•S term) ; dotted lines: $\Delta$-hole model predictions without * Anshed-dotted lines: static mode1 1 ; dashed lines: sta
$\left(\Delta \phi_{\gamma}= \pm 5^{\circ}\right)$ as determined by its aperture. The finite opening of the pion detector $\left(\Delta \phi_{\pi} \simeq \pm 5^{\circ}\right)$ also resulted in an averaging of the double differential cross section over a range of $\phi_{\pi}$. The theoretical results at $\phi_{\pi}=20^{\circ}$ needed to be averaged over this range of angles since $d^{2} \sigma / d \Omega \pi d \Omega \gamma$ changes very rapidly as a function of $\phi_{\pi}$ near $20^{\circ}$. At $40^{\circ} d^{2} \sigma / d \Omega_{\pi} d \Omega_{\gamma}$ is nearly constant across the size of the pion detector so that non-averaged and averaged predictions are almost the same.

Angular correlation functions were calculated with the $\Delta$-hole code of Lenz et al.. The results from these calculations are shown in Fig. 1. We have also done calculations with the code ARPIN of T.-S.H. Lee (dashed lines, labelled static model (2) in which a different transition density was used.

As one can see from Fig. 1 the angular correlation at $\theta_{\pi}=20^{\circ}$ is more sensitive to the different theoretical predictions than at $40^{\circ}$. The data at $\Theta_{\pi}=20^{\circ}$ clearly favour the predictionsifrom the $\Delta$-nucleus dynamics. More accurate data can certainly be obtained detecting the pions in the SUSI spectrometer (which has not been done in this first preliminary. experiment). They will allow a more definite statement about the need of
a $L \cdot S$ term in the $\Delta$-hole-model.
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### 4.4.1 THE DIFFERENTIAL CROSS SECTION OF THE REACTION PP $\rightarrow$ TId FOR PROTON ENERGIES BETWEEN 500 AND 600 MeV

J. Hoftiezer, Ch. Weddigen, P. Chatelain ${ }^{+}$, B. Favier ${ }^{+}$, F. Foroughi ${ }^{+}$, C. Nussbaum ${ }^{+}$, J. Piffaretti ${ }^{+}$, S. Jaccard ${ }^{++}$, P. Walden ${ }^{++{ }^{+++}}$, Nuc1. Phys. A402 (1983)429

The reaction $p p \rightarrow \pi^{+} d$ is of basic interest from several points of view. Its understanding is essential for the interpretation of pion interactions with nuclear matter as regards absorption and production. As to two particle reaction it is amenable to the investigation of detailed balance, time reversal and isospin symmetry by comparing with results from the inverse reaction and with $n p \rightarrow \pi^{0} d$, respectively. Inelastic channels in pp scattering, such as the $p p \rightarrow \pi d$ reaction may have considerable feedback to the elastic channel, in which peaked structures in spin correlated cross sections gave rise to speculations on the possible existence of dibaryon resonances. These 'resonances' are highly inelastic. Therefore - if real - they should show up as resonance structures in the production amplitudes describing in $p p \rightarrow \pi d$ reaction, in addition to the well-known $N \Delta$ resonant like production mechanism near 600 MeV .

An amplitude analysis calls for precise data. Therefore we measured the differential cross section $\sigma(\theta)$ of the $p p \rightarrow \pi d$ reaction for seven energies between 516 and 582 MeV with an overall absolute precision better than two percent. Pions and deuterons were detected in coincidence with scintillation detectors, using time-of-flight techniques for event identification. Special techniques for absolute energy calibration, normalization and measurement of relative distributions were used. Detailed corrections have been applied to the data for multiple scattering, pion decay, detection of decay muons, and for finite solid angle acceptance. The data have been fit with Legendre polynomials to obtain the expansion coefficients $a_{2 i}^{00}$. In Fig. 1 our results (solid symbols) are compared with those obtained by Boswell et al. (1) for the inverse reaction and with theoretical predictions. Blankleider et al. (2) reproduce $a_{o}^{o o}$ and $a_{2}^{o o}$ reasonably well, whereas for $a_{4}^{o 0}$ Niskanen (3) obtained better results.

Our corrected results show only small differences from preliminary data reported earlier (4). The effect of the applied corrections is that we can now





$$
\begin{gathered}
p p-\pi d \\
\sigma(\theta)=\frac{1}{4 \pi}\left[a_{2 i}^{\circ o} P_{2 i}(\cos \theta)\right.
\end{gathered}
$$

Fig. 1:Comparison of our results (full symbo1s) for Legendre polynomial expansion coefficients with recent data from the reverse reaction (Ref. 1, open symbols) and with theoretical predictions of Niskanen (3) (solid lines) and Blankleider (2) (broken lines). For data with arrows no uncertainties are given in Ref. 1.
exclude an $a_{6}^{o o}$ contribution within the quoted precision of $\pm 50 \mu \mathrm{~b}$ in the energy range between $E_{p}=516$ and 582 MeV . Thus the conclusions in Ref. 4 remain valid.
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### 4.4.2 THE ANALYSING POWER A Yo OF THE REACTION pp $\rightarrow \pi \mathrm{d}_{\mathrm{H}}$ FOR PROTON ENERGIES BETWEEN 500 AND 600 MeV

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The interpretation of inelastic pp-scattering at medium energies is still an unresolved problem. From a theoretical point of view a unified treatment taking the most important elastic and inelastic channels into account and including two and three particle unitarity promised to clarify the situation. However, for the channel pp $\rightarrow \pi d$ significant differences remain between theory and primary observables. This is a stimulating situation for experimentalists. By providing precise data on polarization observables and on the differential cross section it should be possible to determine the most important production amplitudes in a direct manner. Accordingly, our group (NESIKA) first performed precision measurements (2) on the differential cross section $\sigma(\theta)$ for the $p p \rightarrow \pi d$ reaction. Here, we report on our second series of experiments namely the measurement of the analyzing power $A_{y o}$ for pion production via $p p \rightarrow \pi d$ in the same energy range. The availability of precise data on $\sigma(\theta)$ at identical energies is a unique advantage for the interpretation of our $A_{\text {yo }}$ results.

The experiment was performed at the PM1 beam line of SIN. The apparatus and the data handling system were similar to those used for our $\sigma(\theta)$ measurements and are described in more detail in Ref. (2). Our results for A yo are plotted in Fig. 1 as a function of the $C M$ pion production angle $\theta$. The indicated errors of $A_{y o}$ rarely exceed one per cent.

The product

$$
\begin{equation*}
\sigma(\theta) A_{y o}=\frac{1}{4 \pi} \sum_{j} b_{j}^{y o} P_{j}^{l}(\cos \theta) \tag{1}
\end{equation*}
$$

can be expanded using associate Legendre polynomials $\mathrm{P}_{\mathrm{j}}^{1}$. The expansion coefficents $b_{j}^{y o}$ can be expressed as bilinear forms of production amplitudes (3). The curves shown in Fig. 1 are five-parameter fits to our results.

The expansion coefficients $b_{j}^{y o}$ are plotted as a function of the proton energy in Fig. 2 and compared to theoretical predictions. These are the very first (4) based on a coupled channel formalism for the $N \Delta$ intermediate state and recent results (5) of a "unified" model calculation. This treats the elastic channels ( $p p$ and $\pi d$ ) together with the transition channel ( $p p \leftrightarrow \pi d$ ) taking two- and three-particle unitarities into account. It can be seen that

Fig. 1: Results for the analysing power $A_{y o}$ as a function of the CM pion production angle $\theta$. The curves are Legendre polynomial fits according to Eq.(1).

Fig. 2: Comparison of our results for Legendre polynomial expansion coefficients $b_{i}^{\text {yo }}$ (Eq. (1)) with theoretical predictions of Niskanen (4) ( $\quad, ~ N i$ )
and of Blankleider (5) ( - , Bi).
Quantities concerning $b_{5}^{\text {yo }}$ are
magnified by 10.



the large coefficients ( $j$ odd) describing interferences between initial-state triplet and the dominant singlet amplitudes are reasonably well reproduced, whereas the triplet-triplet interferences ( $j$ even) show discrepancies in sign and magnitude.
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4.4.3 THE SPIN CORRELATION COEFFICIENT $A_{x z}$ OF THE REACTION $p p \rightarrow \pi d$ FOR PROTON ENERGIES BETWEEN 500 AND 600 MeV
J. Hoftiezer, G.S. Mutchler ${ }^{+++}$, Ch. Weddigen, J.A. Konter ${ }^{+}$, S. Mango ${ }^{+}$, A. Berdoz ${ }^{++}$, B. Favier ${ }^{++}$, F. Foroughi ${ }^{++}$(1)

In the framework of the NESIKA collaboration a series of experiments were performed at SIN on the reaction $p p \rightarrow \pi d$ at proton energies between 500 and 600 MeV to provide precise data for a possible partial wave analysis of this reaction in the energy range from 500 to 600 MeV . In the first experiment (2) the differential cross-section $\sigma(\theta)$ was measured at seven energies. The result was that a three-term Legendre polynomial expansion was sufficient to describe $\sigma(\theta)$ in our energy range. In the second experiment (3) the analysing power $A_{y o}(\theta)$ was measured at four energies. An associate Legendre polynomial fit to the product $\sigma(\theta) A_{\text {yo }}$ yielded significant expansion coefficients $b_{l}^{\text {yo }}$ up to the fifth order.

Here, we describe an experiment on the spin correlation coefficient $A_{X Z}(\theta)$ which was performed at 516,542 , and 582 MeV . $A_{X z}$ is complementary to Ayo in the sense that the complex quantity

$$
\begin{equation*}
\sigma(\theta)\left(A_{x z}+i A_{y o}\right)=\frac{1}{4 \pi} \sum_{k}{\underset{\sim}{f}} P_{k}^{1}(\cos \theta) \tag{1}
\end{equation*}
$$

can be expanded using associate Legendre polynomials $P_{k}^{1}$, where the complex expansion coefficients $\underset{\sim}{b}{ }_{k}=b_{k}^{x z}+i b_{k}^{y o}$ can be expressed as bilinear forms of reaction amplitudes. A kinematical limit is given by

$$
A_{x z}^{2}+A_{y o}^{2} \leq 1
$$

Our measurements were performed at the PMI beam of SIN. Since the available beam time with the polarized target was short, all of our $A_{x z}$ results had to be obtained within a total of twelve days of target operation.

Fig. 1 shows the experimental set-up. The polarized target is viewed


Fig. 1:Experimental set-up for the measurement of the spin correlation coefficient $A_{x z}$ of the $p p \rightarrow \pi d$ reaction (not to scale). The target polarization $P_{T}$ is inverted by changing the HF-frequency, while the beam polarization $P_{B}$ remained constant. $A, B$, and $C$ are scintillation detectors. The $\mathrm{CH}_{2}$ target upstream of the target is viewed by two scintillation telescopes (not shown) and serves for the relative normalization of different runs. MWIC $=$ multiwire proportional ionization chamber.
by two time-of-flight telescopes, the detectors B and A producing start and stop signals, respectively. The copper Cu absorbers isotachic protons which do not reach the coincident detectors $C$. The $C M$ production angel $\theta$ of the undetected pions is deduced from the lab. angle $\delta$ and at the Jacobian cusp ( $\delta_{\max }=12^{\circ}$ ) from the deuteron TOF and the energy deposite in the $C$ detectors. A specially designed dummy target with equivalent amounts of carbon and ${ }^{3} \mathrm{He}$ was used to measure background deuterons.

Fig. 2:Angular distributions of the spin correlation coefficient $A_{x z}$ with a five-parameter fit according to Eq. (1). The scales for the 582 and 516 MeV data are given on the left', while the scale for the 542 MeV data is given on the right. Broken curves are taken from recent predictions of partial wave analy ses (4) with a dominant $D$ (dashdotted) and dominant $S$ (dashed) production partial wave at 500,550 , and 600 MeV , respectively. The pa-rity-forbidden coefficient $A_{y z}$ at 582 MeV and $97^{\circ}$ was measured as a trial. $\theta=$ pion $C M$ production angle.


In Fig. 2 our results are compared with predictions of a recent phaseshift analysis (4). Our data do not definitively favour the $S$ - or the $D^{-}$dominant solution. However, it can be expected that our results will help eliminate ambiguities in such analyses, which are a powerful tool to search for possible hidden anomalies of inelastic pp-scattering at medium energies.
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### 4.4.4 PRECISION MEASUREMENT OF THE ANALYZING POWER A yo FOR PROTON - PROTON ELASTIC SCATTERING AT 582 MeV

A. Berdoz ${ }^{+}$, B. Favier ${ }^{+}$, F. Foroughi ${ }^{+}$, Ch Weddigen

The analyzing power is a frequently used input observable for $N-N$ phase shift analysis (PSA) the results of which serve for different purposes such as comparison with results of interaction models, calculation of finalstate interactions, or search for anomalies in the energy dependence of special partial waves ('dibaryons'). Existing data (1) near 580 MeV had a mean precision of only $1.9 \%$ and discrete differences up to $6.4 \%$ between PSA predictions and experiment occurred.

After the measurement of the absolute differential cross section (2) at a CM angel $\theta=90^{\circ}$ and of relative angular distributions (3) for the $p-p$ elastic scattering we measured the analyzing power $A_{y o}$ at $582 \pm 1 \mathrm{MeV}$ under 13 angels between $\Theta=20$ and $80^{\circ}$. The experiment was performed in June ' 83 at the PM1 beam line of SIN. The experimental technique was essentially the same as for the $A_{y o}$ measurement (4) for inelastic pp scattering.

Fig. 1 shows preliminary experimental results for $A_{y o}$ together with a PSA prediction for 578 MeV taken from Ref. 1. The mean statistical error

Fig. 1: Preliminary results for the analyzing power $A_{\text {yo }}$ of elastic $p-p$ scattering as a function of CM scattering angel $\Theta$. The statistical errors ( $\because .24 \%$ ) are smaller than the dots. The PSA prediction for 578 MeV is taken from Ref. 1.

(.24\%) is smaller than the experimental points. The angular distribution of Ayo is found to be smooth without discrete deviations from the PSA prediction.
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### 4.5.1 NEUTRON AND CHARGED-PARTICLE PRODUCTION CROSS SECTIONS <br> FOR 590 MeV PROTONS ON THIN METAL TARGETS WITH $12 \leq \mathrm{A} \leq 238$

S. Cierjacks, Y. Hino, F. Raupp, L. Buth ${ }^{+}$

Differential production cross sections for neutrons and charged particles from 590 MeV proton bombardment of various thin metal targets have been measured at the SIN cyclotron. Analyses of the experimental results have presently provided absolute neutron production cross sections for $\mathrm{C}, \mathrm{Al}, \mathrm{Fe}, \mathrm{Nb}, \mathrm{In}, \mathrm{Ta}, \mathrm{Pb}$ and U at laboratory angles of $23^{\circ}, 30^{\circ}, 90^{\circ}$ and $150^{\circ}$. The analysis of the charged particle data (which is in a final stage and whose completion is expected by the end of 1983) will provide, in addition, differential production cross sections for secondary protons, deuterons, tritons and pions ( $\pi^{+}+\pi^{-}$) for the same target materials and laboratory angles of $23^{\circ}, 45^{\circ}, 90^{\circ}, 135^{\circ}$ and $157^{\circ}$. The systematic investigations of this type were primarily performed for testing the accuracy of cross sections predicted by modern intranuclear cascadeevaporation models and used in several high-energy nucleon-meson transport codes, e.g. the HETC Monte Carlo Code (1). While measured integral neutron yields from thick heavy metal targets are presently well described by these models, similar comparisions of differential yields and production cross sections have revealed significant discrepancies, particularly in cascade energy region.

Differential cross section measurements were performed by the time-of-flight technique employing the microstructure pulses from the SIN cyclotron ( $\sim 200 \mathrm{ps}$ pulse width at a pulse repetition frequency of 16.84 MHz ). The principal detector was an NE 213 liquid scintillator used for its pulse shape discrimination (PSD) properties. Particle identification in terms of neutrons and charged particles was accomplished by a thin plastic scintillator placed in fromt of the main detector. While neutron events were accumulated by operating both detectors in anticoindence, charged-particle data were collected in coincidence measurements.

A typical result obtained from these experiments is shown in Fig. 1. This figure displays the differential neutron production cross sections
for various elemental target materials at a laboratory angle of $30^{\circ}$. All invidual curves exhibit the well-known two-component shape resulting from evaporation reactions ( $\mathrm{E}_{\mathrm{n}} \lesssim 15 \mathrm{MeV}$ ) and cascade processes ( $\mathrm{E}_{\mathrm{n}} \gtrsim 15 \mathrm{MeV}$ ), respectively.

Fig. 1: Measured differential cross sections for neutrons from 590 MeV proton bombardment of various thin metal targets


Fig. 2: Comparison of measured and calculated cross sections for tantalum. It can be seen that HETC-calculations provide reasonable predictions in the evaporation range ( $\mathrm{E}_{\mathrm{n}} \check{1}_{15 \mathrm{MeV} \text { ), }}$ but underpredict the data increasingly with energy in the cascade region (E $\mathrm{E}_{\mathrm{n}} \gtrsim_{15 \mathrm{MeV} \text { ). }}$


The experimentally determined cross sections are presently being used for extensive tests of model predictions provided by the High Energy

Nuclon-Meson Transport Code (HETC). The available data revealed that the neutron production cross sections in the evaporation region can sufficiently be predicted by this code. In the cascade region, however, large discrepancies continue to persist. A recent comparison of measured and calculated production cross sections is shown in Fig. 2. The results given for neutron production in tantalum illustrate the large existing discrepancies in the cascade region. At $E_{n} \simeq 100 \mathrm{MeV}$ the calculated data are almost by an order of magnitude smaller than the measured ones.
(1) HETC, ORNL-4744, Oak Ridge Nat. Laboratory, 1972
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4.5.2 PROTON - ANTIPROTON ANNIHILATION AT REST INTO $\pi^{0} \omega, \pi^{0} \eta$, $\pi^{0} \gamma, \pi^{0} \pi^{0}$, AND $\pi^{0} \eta^{\prime}$
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In an experiment on the inclusive $\gamma$-spectrum after $\mathrm{p} \overline{\mathrm{p}}$-annihilation at rest, we were able to extract branching ratios for so far unobserved neutral twobody annihilation channels like $\pi^{0} \omega, \pi^{0} \eta$, and the channel $\pi^{0} \gamma$. A new value for the $\pi^{0} \pi^{0}$ and an upper limit for the $\pi^{0} \eta^{p}$ channel were deduced.

Owing to the fact that in a twobody reaction the particles are monoenergetic, the spectrum of the photons from the electromagnetic decay of a meson exhibits an uniform shape between the limits

$$
\begin{equation*}
E_{\max / \min }=\frac{1}{2} E(1 \pm \beta) \tag{1}
\end{equation*}
$$

where $E$ is the total energy of the decaying particle and $\beta$ is its velocity. The edges of these distributions are sharp if the particles involved have narrow widths, and thus show up in the inclusive $\gamma$-spectrum from $\overline{p p}-$ annihilation at rest. The lower the background and the better the resolution, the higher are the chances of observing these edges. Thus at the high-energy end of the spectrum such edges could be observed. In order to analyse these con-


Fig. 1: Experimental spectrum after subtraction of the continuous background. The solid line shows the final fit. The contributing distributions are indicated.

Table 1:Branching ratios of twobody annihilation channels in $\bar{p} \bar{p}-a n n i h i l a t i o n$ at rest
Channe1 Yield Error (stat.) Error (syst.)

| $\pi^{0} \eta^{\prime}$ | 0.30 | $\pm 0.53$ | +0.58 |
| :--- | :--- | :--- | :--- |
| $\pi^{0} \omega$ | 2.38 | $\pm 0.27$ | +0.32 |
|  |  | $\pm 0.07$ | -0.52 |
| $\pi^{0} \eta$ | 0.82 | $\pm 0.25$ | $\pm 0.06$ |
| $\pi^{0} X$ | 1.19 | $\pm 0.004$ | $\pm 0.44$ |
| $\pi^{0} \gamma$ | 0.015 | $\pm 0.03$ | $\pm 0.003$ |
| $\pi^{0} \pi^{0}$ | 0.06 |  | $\pm 0.01$ |

tributions we subtracted a proper continuous background, which comes from annihilations into more than two particles. We defined a shape function by means of a Monte Carlo simulation (2) and subtracted the background after appropriate normalization. Fig. 1 shows the resulting experimental spectrum, which was analysed. The single distributions used to describe the spectrum and the final fit are shown as well. The results for the branching ratios are summarized in Table 1.
(1) dto., submitted to Nuclear Physics B
(2) Th. Köhler Diplomarbeit, Universität Karlsruhe (1980), unpublished

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The main research activity of the High Energy Physics group of the IK I has been the exploitation of the large spectrometer facility CELLO at PETRA. Since June 1982 G. Flügge is spokesman of the CELLO collaboration. The CELLO detector was built by a collaboration of the three German and three french institutions listed above. Following a major upgrading in 1982 the detector was moved back into the beam in August 1982. After a running in phase, during which various difficulties caused by the new components and the increased background had to be solved, the detector is successfully taking data since the beginning of 1983.

In autumn 1982 five groups of the former PLUTO collaboration, the DESY groups F14 and F33 and the Universities of Glasgow, Hamburg, and Te1 Aviv joined the CELLO collaboration. Together with these groups a major up-
grading of the CELLO detector was proposed and accepted by the PRC and Directorate of DESY. In particular, a new inner detector is under construction.

The main activities of the Karlsruhe group were to
a) improve and maintain the LAr electronics and cryogenics and participate to the planing for the new central detector
b) process and analyse the CELLO data using the KfK computer. Physics analysis concentrated on hadronic final states and test of QCD (in particular total cross sections, event shape, inclusive electrons and electroweak interference effects in $\tau$ pair production.

### 5.1 HARDWARE ACTIVITIES

### 5.1.1 CELLO OPERATION

Experience with the upgraded detector was obtained in a first run in 1982. A high background showed up in the experiment consisting, mainly of synchrotron radiation due to the high beam energies. Many attempts were undertaken to cope with this background. Among others, a new beam pipe was built and additional triggers were installed using the information from the LAr calorimeters. These triggers are not sensitive to the synchrotron radiation as the track triggers from the inner chambers are. On the other hand, the calorimeter trigger suffers from pick-up noise in the sensitive amplifier system ( $10^{4}$ channels). Special efforts was undertaken to reduce all sort of noise. As a result, a trigger reacting on a single photon in only one modul of the calorimeter could be installed which has a threshold of approximately 2.5 GeV .

During the running period 1983 the calorimeter, CELLO's trigger- and cryogenic system operated without any problems. For the future perspectives of the CELLO detector, the group participates in the design and construction of a new inner detector (1), the "Stereo Wire Chamber". The institute contributes to the mechanical and electronic construction.
(1) DESY-CELLO-83-01, Proposal to Upgrade the CELLO Detector.

### 5.1.2 DEVELOPMENT IN THE INSTITUTE

For the calorimeter a new trigger system has been built using flash

ADC's and RAM's. With this unit a fast decision is taken already on the trigger level. Using the shower propagation within the calorimeter, the trigger discriminates between pattern of real shower and background like cosmic ray shower or pick-up signals.

For future applications we investigate the possibility to use the scintillation light in liquid argon, mainly for fast trigger proposes. The emitted light is in the far ultraviolett region at 125 nm , which makes it difficult to guide the light in large detector systems. Among others, it could be demonstrated that the light can be shifted to 175 nm by a small amount of Xenon (see Fig. 1). This shift circumvents the problem, since in this wavelength region aluminium is a reasonably good mirror.

In another line of development we try to replace the cryogenic liquid argon in the ionisation chamber by a medium which is liquid at room temperature. Good results with tetramethylsilane (TMS) could be obtained in a small test chamber representing a typical submodul (tower) of a calorimeter. The pulse height corresponds to $30 \%$ of that in liquid argon.


Fig. $1:$
Scintillation Light from Liquid Argon excited by Am ${ }^{241}$ $\alpha ' s$ shifted with 460 ppm added Xenon to $1750 \AA$ and reflected at the inner side of an aluminium tube.

### 5.2 TEST OF QCD AND HADRONIC FINAL STATES

### 5.2.1 THE INFLUENCE OF FRAGMENTATION MODELS ON THE DETERMINATION OF THE STRONG COUPLING CONSTANT IN $\mathrm{e}^{+} \mathrm{e}^{-}$ANNIHILATION INTO HADRONS $*$

Hadronic events obtained with the CELLO detector at PETRA were compared with first order QCD predictions using two different models for the fragmentation of quarks and gluons, the Hoyer model and the Lund model. Both models are in reasonable agreement with the data, although they do not completely reproduce the details of many distributions. Several methods have been applied to determine the strong coupling constant $\alpha_{s}$. Although within one model the value $\alpha_{s}$ varies by $20 \%$ among the different methods, the values determined using the Lund model are larger by $30 \%$ or more (depending on the method used) than the values determined with the Hoyer model. Our results using the Hoyer model are in agreement with previous results based on this approach.

TABLE 1 Values of $\alpha_{s}$ obtained at $\sqrt{s}=34 \mathrm{GeV}$ with the Lund Mode1 (LM) and the Hoyer Model (HM). (First order in QCD).

The error in the determination of $\alpha_{s}$ is statistical only (including statistical Monte Carlo error).

| METHOD | LUND MODEL | HOYER MODEL | $\frac{\alpha_{S}(\mathrm{LM})}{\alpha_{S}(\mathrm{HM})}$ |
| :--- | :--- | :--- | :--- |
| $\mathrm{S} \geq 0.25 \mathrm{~A} \leq 0.1$ | $0.280 \pm 0.045$ | $0.190 \pm 0.030$ | 1.47 |
| $0 \geq 0.20$ | $0.260 \pm 0.040$ | $0.190 \pm 0.020$ | 1.37 |
| $0 \geq 0.30$ | $0.255 \pm 0.050$ | $0.200 \pm 0.035$ | 1.28 |
| $\#$ of 3 c1usters | $0.235 \pm 0.025$ | $0.145 \pm 0.020$ | 1.62 |
| C1uster Thrust | $0.235 \pm 0.025$ | $0.155 \pm 0.015$ | 1.52 |
| E.W.A.C $*$ | $0.250 \pm 0.040$ | $0.150 \pm 0.020$ | 1.67 |

* Energy Weighted Angular Correlation


### 5.2.2 INCLUSIVE PRODUCTION OF ELECTRONS AND MUONS IN MULTIHADRONIC EVENTS AT PETRA *

The $e^{+} e^{-}$annihilation into hadrons at PETRA energies is well described by assuming the pair production of $u, d, s, c$, and $b$ quarks and gluon bremsstrahlung. Since weak decays of the hadrons containing light quarks ( $u, d, s$ ) have guided the understanding of electroweak interactions, it is of great interest to explore the weak decays of hadrons containing heavy quarks ( $c, b$ ). Assuming standard mixing angles, in the Kobayashi-Maskawa model (1) bottom is expected to decay predominantly into charm, and charm to strangeness.

The leptons resulting from the semileptonic decay of bottom hadrons are expected to dominate the region of transverse momentum larger than $1.0 \mathrm{GeV} / \mathrm{c}$ with respect to the jet axis. This feature combined with a more spherical event shape is used in the selection of hadronic $b \bar{b}(g)$ events.

The search for inclusive leptons in hadronic events provides therefore an indirect study of heavy quark production. From the inclusive lepton yield we obtain the fraction of "prompt" leptons, which result only from semileptonic $b-$ or $c$-decays, including the cascade decay $b \rightarrow c \rightarrow I V X$ by subtracting all trivial lepton sources as $\pi / K$-decay, $\gamma$-conversion etc. The prompt lepton signals yield then with suitable additional cuts a measurement of the semileptonic branching ratios of heavy quarks.

A selection of events containing mostly $\mathrm{b} \overline{\mathrm{b}}(\mathrm{g})$ allows a measurement of the charge angular asymmetry predicted in the Weinberg-Salam model (2).

The final state hadrons result from the fragmentation of the original quarks. This fragmentation process is poorly known, and models with several free parameters are used to describe it. The measurement of the rate and spectra of inclusive leptons could thus help to understand better the fragmentation functions of heavy quarks. The model dependence of the determination of the semileptonic branching ratios has carefully been studied by comparing the data with the Hoyer et al. (3) and Lund (4) MC event generators.

The data were taken with the CELLO detector (5) at PETRA at average c.m. energies of $W=14,22$, and 34 GeV . The corresponding integrated luminosities are $1.1 \mathrm{pb}^{-1}, 2.5 \mathrm{pb}^{-1}$, and $7.9 \mathrm{pb}^{-1}$.

To identify electrons all tracks measured in the inner detector have been extrapolated to the liquid argon calorimeter. The study of MC events

[^1]yields the electron efficiency in multihadronic events of typically $60 \%$ and the pion misidentification probability is about $3 \times 10^{-3}$. For the generation of multihadronic annihilation events we employed the Hoyer Model (3) using Feynman-Field fragmentation (6). The various prompt $b-$ or $c$-decays and background contributions to the elctron candidates with $\mathrm{P}_{\perp}>0.5 \mathrm{GeV}$ at $W=14 \mathrm{GeV}$ and with $\mathrm{P}_{\perp}>1 \mathrm{GeV} / \mathrm{c}$ at $\mathrm{W}=22$ and 34 GeV are shown in Table 1 .

| Source of electrons | $W=14 \mathrm{GeV}$ | $\mathrm{W}=22 \mathrm{GeV}$ | $\mathrm{W}=34 \mathrm{GeV}$ |
| :--- | :---: | :---: | :---: |
| Light quarks (u,d,s) |  |  |  |
| Misidentified $\pi^{\prime} \mathrm{s}$ | $5.9 \%$ | $5.5 \%$ | $7.6 \%$ |
| Converted photons, $\pi / \mathrm{K}$ decay | $11.8 \%$ | $2.8 \%$ | $11.8 \%$ |
| IC | $0.0 \%$ | $0.2 \%$ | $1.4 \%$ |
| DIS | $1.3 \%$ | $2.8 \%$ | $3.5 \%$ |
| Heavy quarks (c,b) | $4.2 \%$ | $4.2 \%$ | $8.0 \%$ |
| Misidentified $\pi^{\prime} s$ | $3.8 \%$ | $4.4 \%$ | $10.9 \%$ |
| Converted photons, $\pi / \mathrm{K}$ decay | $73.0 \%$ | $80.1 \%$ | $56.8 \%$ |
| Prompt electrons (a11) | $36.6 \%$ | $60.0 \%$ | $34.2 \%$ |
| b $\rightarrow$ enly | 6 | 8 | 20 |



Fig. 1

Fig. 1:
Distribution of the transverse momentum $p_{\perp}$ for electrons compared to MC siumlation
a) A11 electron candidates.

- MC expected yield
- prompt electrons as expected from MC events
b) Prompt electrons.
- prompt electrons as expected from MC events
- prompt electrons from the semileptonic decay $b \rightarrow e v X$ as expected
from the MC

Fig. 1 shows the comparison between the data and the MC prediction for the $p_{\perp}-s p e c t r u m$ of the electrons at the highest $c . m$. energy. The sample of electron candidates is displayed in Fig. la together with the total MC expected yield (solid line) and the MC expected prompt lepton yield (dashed line). A good agreement between data and $M C$ is observed. The $p_{\perp}$-distribution of prompt electrons obtained after background subtraction is shown in Fig. lb and compared to the MC expected prompt electron yield (solid line). The dashed line gives the contribution of the semileptonic b-decay only.

Inclusive muon candidates are selected using the information provided by the muon detector. To identify muons, space points reconstructed in the muon chambers are associated with the charged particles recorded by the central detector and extrapolated to the chambers. The contributions to the muon yield for muons with $p>1.6 \mathrm{GeV} / \mathrm{c}$ and $\mathrm{p}_{\perp}>0.5 \mathrm{GeV} / \mathrm{c}$ at 14 GeV and $\mathrm{p}_{\perp}>1.0 \mathrm{GeV} / \mathrm{c}$ at $\mathrm{W}=22 \mathrm{GeV}$ and $W=34 \mathrm{GeV}$ is given in Table 2 .

TABLE 2

| Source of muons | $W=14 \cdot \mathrm{GeV}$ | $\mathrm{W}=22 \cdot \mathrm{GeV}$ | $\mathrm{W}=34 \mathrm{GeV}$ |
| :--- | :---: | :---: | :---: |
| $\pi / \mathrm{K}$-decay | $10 \%$ | $9 \%$ | $9 \%$ |
| hadron - punch through | $18 \%$ | $22 \%$ | $22 \%$ |
| Random associations | $3 \%$ | $5 \%$ | $14 \%$ |
| Prompt muons (a11) | $69 \%$ | $64 \%$ | $55 \%$ |
|  | $\mathrm{~b} \rightarrow \mu$ only | $47 \%$ | $40 \%$ |
| Number of muon candidates | 25 | 14 | 22 |

Fig. 2a shows the transverse momentum of the muon candidates with respect to the thrust axis for the data at $W=34 \mathrm{GeV}$ and MC genereated event events (solid line). The dashed line shows the fraction of prompt muons expected from MC events. The corresponding distribution for the prompt muons as obtained after background subtraction is shown in Fig. 2 b and compared to the MC expected yield (solid line). The dashed line gives the corresponding fraction of muons expected from the semileptonic b-decay.

The influence of fragmentation models to semileptonic branching ratios is studied with the Hoyer MC by changing the decay matrix element of heavy quarks to $V-A$, by introducing the mass dependent QCD matrix element and by variing the strong coupling constant $\alpha_{S}$. Secondly fragmentation parameters as $\sigma_{q}$, the quark ratios $u: d: s$ and the $p_{\perp}$ distribution of the primordi-


Fig. 2a

Fig. 2: Distribution of the transverse momentum $\mathrm{P}_{\perp}$ for muons compared to MC simulation
a) All muon candidates

- MC expected yie1d
-- prompt muons as expected from MC events.
b) Prompt muons.
- prompt muons as expected from MC events
-- prompt muons from the semileptonic decay $b \rightarrow \mu \nu X$ as expected from the MC.
al $1^{\text {st }}$ rank meson were changed. Thirdly the differences between the Hoyer and Lund generators were investigated. All together the overall systematic error is og the order of $18 \%$ for the $c$-decay and $5 \%$ for the $b-$ decay. The variations of the not well known fragmentation of heavy quarks leads to an additional systematic error of about $15 \%$ of the semileptonic c-decay and $5 \%$ of the b -decay.

We derive from the prompt leptons the branching ratios of the semileptonic decay of the $c-$ and $b$-quark by comparing the $p_{\perp}$ spectra of the data with the MC prediction of fixed branching ratios.

As discussed previously the leptons from b-decays are expected to have high transverse momenta. To enrich the b-sample we therefore take events with $p_{\perp}>0.5 \mathrm{GeV} / \mathrm{c}$ at $\mathrm{W}=14$ and with $\mathrm{p}_{\perp}>1.0 \mathrm{GeV} / \mathrm{c}$ at $\mathrm{W}=22$ and $W=34 \mathrm{GeV}$. Table 3 shows the derived branching ratios:

TABLE 3

| Branching ratios |  |  |  |
| :---: | :---: | :---: | :---: |
| W <br> $(\mathrm{GeV})$ | $\mathrm{b} \rightarrow \mathrm{evX}$ | $\mathrm{b} \rightarrow \mu \nu \mathrm{X}$ | $\mathrm{c} \rightarrow \mu \nu \mathrm{X}$ |
| 14 | $4.0 \pm 11.0 \pm 4.3 \%$ | $26.0 \pm 8.2 \pm 3.9 \%$ | $20.5 \pm 7.2 \pm 4.9 \%$ |
| 22 | $21.3 \pm 10.5 \pm 2.1 \%$ | $3.8 \pm 6.2 \pm 3.7 \%$ | $15.0 \pm 4.9 \pm 3.4 \%$ |
| 34 | $15.5 \pm 9.0 \pm 2.8 \%$ | $6.3 \pm 4.8 \pm 3.5 \%$ | $8.1 \pm 4.3 \pm 3.9 \%$ |
| Average | $14.1 \pm 5.8 \pm 3.0 \%$ | $8.8 \pm 3.4 \pm 3.5 \%$ | $12.3 \pm 2.9 \pm 3.9 \%$ |

The systematic errors include those discussed in the lepton identification and background subtraction and those due to the fragmentation and methods specific to the MC event generation. They were added in quadrature. We find an average branching ratio of $14.1 \pm 5.8$ (stat.) $\pm 3.0$ (syst.) \% for $b \rightarrow e \nu X$ and of $8.8 \pm 3.4 \%$ (stat.) $\pm 3.5$ (syst.) \% for $b \rightarrow \mu \nu X$, in good agreement with previous results (7).

To extract the branching ratios for the semileptonic decay of $c$ quarks we use cuts in the thrust distribution of $T>0.88,0.88$, and 0.94 at the three c.m. energies ( $W=14,22$, and 34 GeV ). Due to large contributions from background processes to the elctron signal we are only able to determine the muonic branching ratio. Column 3 in Table 3 contains the $c \rightarrow \mu \nu X$ branching ratios with an average of $12.3 \pm 2.9$ (stat.) $\pm 3.9$ (syst.) \% again in agreement with previous measurements (8).

Using the sample of enriched $b \bar{b}-$-events, we determined the asymmetry due to the electro-weak interference. We use the thrust axis as an approximate measurement of the quark direction of flight and the sign of the lepton to tag the charge of the parent quark. After background subtraction we fit the angular distribution of the asymmetry. For b $\bar{b}-p a i r ~ p r o d u c t i o n ~ a n ~ a s y m m e-~$ try of $-27.3 \%$ is predicted. Due to the various background contributions we expect a reduced asymmetry of $-5.8 \%$ for electrons and $-8.0 \%$ for muons at $W=34 \mathrm{GeV}$ with an uncertainty of $2.2 \%$ for electrons and $3.2 \%$ for muons respectively. This error on the predicted value includes all systematic uncertainties discussed above: lepton identification, model dependence and branching ratios. Even though we cannot really test in a definitive way these predictions with our limited statistics, our measurements of $-38 \% \pm 21 \%$ for electrons and of $43 \% \pm 31 \%$ for muons are compatible with the expected va1ues.
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### 5.3 TWO-PHOTON PHYSICS

### 5.3.1 LEPTON PAIR PRODUCTION IN DEEP INELASTIC e- $\gamma$ SCATTERING *

At PETRA we have measured the process $e+\gamma \rightarrow e+\ell^{+}+\ell^{-}$, where $\ell$ is either an electron or muon, at an average $Q^{2}$ of $9.5 \mathrm{GeV}^{2} / \mathrm{c}^{2}$. The total number of collected events is 240 . We find that our data agree with QED predictions based either on exact Feynman graph calculations or on the photon structure function expressions.

### 5.3.2 EXPERTMENTAL STUDY OF THE HADRONIC PHOTON STRUCTURE FUNCTION **

We have measured at PETRA the process $\mathrm{e} \gamma \rightarrow \mathrm{e}+$ hadrons at an average $Q^{2}$ value of $9 \mathrm{GeV}^{2} / \mathrm{c}^{2}$. The total number of observed events attributed to this process is 215. Our data are compared to calculations based on the estimation of the photon structure function $F_{2}$ in the quark parton model and in QCD.

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### 5.4 QED PROCESSES AND ELECTRO-WEAK INTERFERENCE

5.4.1 INVESTIGATION OF TWO PHOTON FINAL STATES IN $e^{+} e^{-}$ANNIHILATION AT $(\sqrt{\mathrm{s}})=34.2 \mathrm{GeV} *$

Two photon final states in $\mathrm{e}^{+} \mathrm{e}^{-}$annihilation have been analyzed at CM energies around 34 GeV . Good agreement with QED is observed. Lower 1imits for the QED cutoff parameters of $\Lambda_{+}>59 \mathrm{GeV}$ and $\Lambda_{-}>44 \mathrm{GeV}$ are determined. A search for two photons with missing energy yields an upper limit for the production of neutral particles which decay into a photon and a non-interacting particle. Constraints on the mass and the coupling strength of supersymmetry photinos are discussed.


### 5.4.2 MEASUREMENT OF THE $\tau$ LIFETIME **

Using 3 prong-decays of $\tau$-leptons from the reaction $e^{+} e^{-} \rightarrow \tau^{+} \tau^{-}$at 34.2 GeV and 22 GeV the $\tau 1 i f e t i m e$ has been determined by a maximum $1 \mathrm{ikeli}-$

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* Phys.Lett. 123B (1983) 127
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** Nuc1.Phys. B211 (1983) 369
hood fit to the decay length distribution. The result

$$
\tau_{\tau}=4.7 \begin{aligned}
& +3.9 \\
& -2.9
\end{aligned} \times 10^{-13} \mathrm{~s}
$$

agrees well with the lifetime of $(2.8 \pm 0.25) \times 10^{-13}$ s expected from $\mu-\tau$ universality and the known branching ratio for $\tau \rightarrow e v \bar{\nu}$.

### 5.4.3 $\tau$ BRANCHING RATIOS AND POLARIZATION LIMITS IN $e^{+} e^{-}$INTERACTIONS AT $\sqrt{\mathrm{s}}=34 \mathrm{GeV} *$

Considerable knowledge on the $\tau$ lepton has been obtained in low energy $e^{+} e^{-}$interactions both at SPEAR and DORIS (1). At center of mass energies far above the production threshold as provided by PEP/PETRA however, the signature of $\tau^{+} \tau^{-}$events with respect to background is much cleaner, offering the possibility to extracts a virtually background-free sample of $\tau$ pairs without restriction to specific decay topologies. Our trigger logic incorporated a two-charged particle trigger with a momentum cutoff at $250 \mathrm{MeV} / \mathrm{c}$ and various calorimetric triggers. The results are based on 526 $\tau^{*} \tau^{-}$events corresponding to an integrated luminosity of $11.2\left(\mathrm{pb}^{-1}\right)$ at a center of mass energy of 34 GeV . All $\tau^{+} \tau^{-}$decay topologies are recorded by the detector. We report here on measurements of the decay channels $\tau \rightarrow \rho \nu$, $\pi \nu, e v \nu, \mu \nu \nu$.
$\tau \rightarrow \rho \nu . \rho$ candidates were selected from 1 prong decays with one or two additional photons in the same hemisphere. Due to the large momenta of the decay particles, the two photons from the $\pi^{0}$ merge into one shower in about $60 \%$ of the decays. Thus in addition to the 1 charged +2 photon final state also the 1 charged +1 photon topology was considered. A minimal opening angle of $6^{\circ}$ between the charged particle and any photon is required to reduce background due to shower fluctuations from hadronic interactions of the charged particle in the calorimeter.

Whenever two photons were accepted (about $40 \%$ the $\rho \rightarrow \pi \pi^{0}$ decays) a kinematical fit to the $\pi^{0}$ mass was performed in order to improve the $\rho$ mass resolution. The spectrum of the effective $\gamma \gamma$ mass prior to the fit is shown in Fig. la together with the prediction from our Monte-Carlo calculation. The effective mass for the charged particle (assuming a $\pi$ mass) and the photon (or the 2 photons after the kinematic fit) is shown in Fig. 1 b . A clear $\rho$ signal is observed. The following background contributions have been con-

[^2]

Fig. 1: a) Distribution of the invariant two photon mass for $\rho$ candidates where 2 showers were reconstructed (histogram) and the MC prediction (curve).
b) ( $\pi \pi^{0}$ ) invariant mass spectrum (histogram), the MC prediction is superposed (curve). Also shown in the background contribution (shaded area).
sidered (see Fig. 1b) :

- $\tau$ decaying to $\left(\pi \pi^{0}\right) \nu+\geq 1 \pi^{0}$ which are not detected, e.g. from $\tau \rightarrow \mathrm{A}_{1} \nu$.
- $\tau \rightarrow K^{*} \nu$. Since we do not distinguish experimentally $K^{\prime} s$ from $\pi^{\prime} s$, $K^{*}(890)$ decays also feed into our $\rho$ sample.

Background due to other processes, such as electrons, muons, or pions together with a radiative photon yielding an invariant mass within the $\rho$ band, is found to be negligible.
$\tau \rightarrow \mu \nu \nu$. Muons were identified in the calorimeter as minimum ionizing particles and by their ability to penetrate the iron filter. A minimal particle momentum of $2 \mathrm{GeV} / \mathrm{c}$ and an associated hit in the muon chambers were required. The pion punch through has been determined by Monte-Carlo simula-
tion.
$\tau \rightarrow e V V, \pi \nu$. Electrons and pions were separated with the help of their characteristic shower pattern in the calorimeter. Both energy-momentum matching and shower development in depth were employed. A considerable background from $\tau \rightarrow \rho \nu \rightarrow \pi \pi^{0} \nu$, due to undetected photons from $\pi^{0}$ decay, is expected. In order to reject $\rho^{\prime} s$ in our pion sample we consider the showers within a cone of $45^{\circ}$ half opening angle around the charged particle. The pion efficiency and the remaining $\rho$ background have been determined by Mon-te-Carlo techniques.

Branching ratios have been obtained by normalizing to the corrected total number of $\tau$ events, i.e. to the measured total cross section for $\mathrm{e}^{+} \mathrm{e}^{-} \rightarrow \tau^{+} \tau^{-}$at $\sqrt{s}=34 \mathrm{GeV}$. We obtain the following values:
$B R(\tau \rightarrow \rho \nu)=.228 \pm .025 \pm .021$
$\mathrm{BR}(\tau \rightarrow \pi \nu)=.099 \pm .017 \pm .013$
$B R(\tau \rightarrow e \nu \nu)=.183 \pm .024 \pm .019$
$B R(\tau \rightarrow \pi \nu \nu)=.176 \pm .026 \pm .021$
where the first error is statistical and the second one is systematic. Table 1 summarizes the four $\tau$ decay samples, selection efficiencies, and background contamination with their systematic errors.

TABLE 1 Summary of the four decay samples and their resp. efficiencies and background contributions

|  | $\rho$ | $\pi$ | $e$ | $\mu$ |
| :---: | :---: | :---: | :---: | :---: |
| No. of events | 101 | 34 | 60 | 47 |
| Efficiency | $.45 \pm$ | .04 | $.48 \pm$ | .04 |
| Contamination <br> from | 0 | $1.4 \pm \pm 1.4$ | .04 | $.70 \pm$ |

Our results for $\tau \rightarrow \pi \nu$, evV, $\mu \nu \nu$ are in good agreement with determinations at lower energies (1,2). The branching ratio for $\tau \rightarrow \rho \nu$ has somewhat improved in the overall error as compared to previous publications (3) based
on larger $\tau^{+} \tau^{-}$samples. We attribute this fact to our large selection efficiency for $\rho^{\prime} s$ and to the low background in the $\tau{ }^{+} \tau{ }^{-}$sample (see Table 1 ).

Polarization. In the case of polarized $\tau$ 's the angular decay asymmetry in the $\tau$ rest frame (4) leads to a characteristic distortion of the laboratory momentum spectra of the decay particles. The spectra of $\rho, \pi, e$ and $\mu$ were determined by repeating the above described background subtraction and efficiency correction bin by bin in the respective raw momentum distributions. In Fig. 2 the corrected laboratory momentum spectra are displayed.


Fig. 2: Corrected laboratory momentum spectra for $\rho, \pi$, $e$, and $\mu$, integrated over the scattering angle. The errors shown include both the statistical (dominant) and the systematic uncertainty. Also shown is the expectation for no polarization (full lines) and for $\pm 100 \%$ polarization (dashed lines), assuming a standard V-A charged current decay for the $\tau$.

For the two body decays $\tau \rightarrow \rho \nu, \pi \nu$ an unpolarized $\tau$ will produce flat laboratory momentum spectra, whereas a non-zero polarization results in a nonzero slope (see Fig. $2 a, b$ ). For the leptonic decays $T \rightarrow e \nu \nu, \mu \nu \nu$ the effects is less pronounced but still useful in deriving polarization limits (see Fig. 2c,d).

For the subsequent analysis of electroweak contributions, the laboratory momentum spectra have been determined in the forward (FW) and backward (BW) hemisphere separately, leading to a measurement of the average polarization in each hemisphere.

A fit combining the eight 1 aboratory momentum spectra (FW and BW) derived for the four decay channels yields a polarization asymmetry defined as $\mathrm{A}(\mathrm{P})=\left(\left\langle\mathrm{P}_{\mathrm{FW}}\right\rangle-\left\langle\mathrm{P}_{\mathrm{BW}}\right\rangle\right) / 2$ of

$$
A(P)=(+1 \pm 22) \%
$$

corresponding to the $95 \%$ C.L. 1imit of $A(P)<44 \%$ 。
In the standard $V$, $A$ ansatz for the weak neutral current (5) the final state polarization for the $\tau^{-}$is a function of the laboratory scattering angle:

$$
\begin{equation*}
P(\theta)=-g(s) \cdot\left(v_{e} a^{a}+v_{\tau} e \frac{2 \cos \theta}{1+\cos ^{2} \theta}\right. \tag{1}
\end{equation*}
$$

with $\left.g(s)=\left(G_{F} / \sqrt{2} 4 \pi \alpha\right)\right) \cdot\left(s_{Z}^{2} /\left(s-M_{Z}^{2}\right)\right) \cdot G_{F}$ is the Fermi coupling constant, $\alpha$ the fine structure constant, $s$ the square of the center of mass energy and $M_{2}$ the mass of the weak neutral boson assumed to be $89 \mathrm{GeV} / \mathrm{c}^{2}$. The first term, independent of $\cos \theta$, contains the product of $v_{e}$ and $a_{\tau}$ which are known from ve scattering, Bhabha scattering, and the $\tau^{+} \tau^{-}$charge asymmetry. The term propoertional to $\cos \theta$ contains $a_{e}$, measured by ve scattering in combination with the Bhabha results, and the unknown quantity $v_{\tau}$ 。

According to equ. (1) our result translates into a vector coupling constant of the $\tau$ to the neutral weak current of $\nu_{\tau}=-0.1 \pm 2.8$. The error is dominated by statistics*. Comparing to the value for $v_{e}=-0.04 \pm 0.06$ (6), our result is certainly consistent with the assumption of universality of the weak leptonic couplings. Due to the larger error on $v_{\tau}$, however, no stronger conclusions can be drawn. Still, our measurement constitutes the first attempt to determine $v_{\tau}$.

[^3]Assuming universality ( $v_{e}=v_{\tau}=v, a_{e}=a_{\tau}=a$ ) one can try to test a general property of the weak Lagrangian (6) which supposedly factorizes into two currents coupled to a single neutral weak boson. With the universality condition, $P(\theta)$ assumes the form

$$
\begin{equation*}
P(\theta)=-g(s) 4 h_{V A} \frac{(1+\cos \theta)^{2}}{1+\cos ^{2} \theta} \tag{2}
\end{equation*}
$$

Factorization (single $z^{0}$ ) then implies $4 h_{V A}=v \cdot a$. Fits incorporating the universality constraint yield $4 \mathrm{~h}_{\mathrm{VA}}=+1.0 \pm 1.4$, which can be compared to the $95 \%$ C.L. 1imit $v \cdot a<0.14$ from ve and $e^{+} e^{-}$data (7).
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6. DEVELOPEMENT AND APPLICATIONS
6.1 DETECTORS

### 6.1.1 CALCULATED EFFICIENCY OF A $4 \pi$ DETECTOR OF BGO OR BaF 2 FOR MONOENER-

 getic gamma-RAys and gamma cascades following neutron captureK. Wisshak, F. Käppe1er, G. Schatz, KfK-report 3580 (1983) and Nuc1. Instr. Meth. (submitted)

The applicability of a spherical shell of $B G O$ or $\mathrm{BaF}_{2}$ as a $4 \pi$ detector for high precision measurements of neutron capture cross sections was investigated. First the efficiency of both scintillator materials for monoenergetic gamma-rays was calculated in the energy range from 0.5 to 10 MeV . Configurations with different thicknesses and inner radii were considered. Second, neutron capture cascades were calculated for several isotopes with widely different capture gamma-ray spectra according to the statistical model. Both informations together allowed to determine the efficiency of an actual detector for neutron capture events in dependence of the threshold energy. A thickness of 10 cm BGO or $17.5 \mathrm{~cm} \mathrm{BaF}_{2}$ proved to be sufficient to registrate more than $95 \%$ of all capture events above a threshold energy of 3 MeV . This reduces the systematic uncertainty due to the detector effi-


Fig. 1
Calculated efficiency of a spherical she11 of BGO $\left(\mathrm{R}_{\mathrm{i}}=10 \mathrm{~cm}, \mathrm{R}_{\mathrm{o}}=20 \mathrm{~cm}\right)$ for gamma-ray cascades following neutron capture in ${ }^{56} \mathrm{Fe},{ }^{197} \mathrm{Au}$ and ${ }^{241} \mathrm{Am}$.
ciency in an absolute cross section measurement to less than $1 \%$ and in a relative measurement using a gold standard to less than $0.5 \%$.

### 6.1.2 SENSITIVITY OF A $4 \pi$ DETECTOR OF BGO OR BaF 2 TO SCATTERED NEUTRONS

K. Wisshak, F. Käppeler

The most important source of background, using a $B G O$ or $\mathrm{BaF}_{2}$ detector in neutron capture cross section measurements, is due to capture of sample scattered neutrons in the scintillator material. This was investigated quantitatively by means of a Monte Carlo simulation. Each interaction in the detector was randomly assigned to capture or scattering according to the respective cross sections. Scattering by up to three isotopes is treated separately in order to consider different angular distributions and energy losses. The free path length between two interactions is randomly selected from an exponential distribution. The code follows the neutron in time and energy until it leaves the detector or is captured.

The calculations were performed under the following assumptions:
1.) The primary neutron spectrum is produced using a Van de Graaff accelerator and the ${ }^{7} \mathrm{Li}(\mathrm{p}, \mathrm{n})$ reaction
2.) The maximum neutron energy is 70 keV corresponding to a proton energy 10 keV above the reaction threshold.
3.) The flight path is 50 cm in case of a $\mathrm{BaF}_{2}$ detector and 1 m for the BGO detector. As the time resolution of $\mathrm{BaF}_{2}$ is about a factor of ten better compared to BGO, the energy resolution is still better for $\mathrm{BaF}_{2}$.
4.) The scattering cross section of the hypothetical sample material is a factor of ten larger than the capture cross section. This is a typical value for the samples of interest to nuclear astrophysics. Fig. 1 shows the result for two configurations with compatible gammaray efficiency. The solid line is the time-of-flight (TOF) distribution of the primary capture events, while the dashed line is the background due to capture of scattered neutrons. These curves represent a conservative limit for the expected signal-to-background ratio. In an actual measurement there are various possibilities to reduce the background. For instance, the shape of the cross section will be determined by selecting pulse height channels around the binding energy of the measured isotope. If this binding energy is
different from that of ${ }^{135} \mathrm{Ba}(9.1 \mathrm{MeV})$ a much better signal to background ratio is obtained. The absolute normalization of the cross section is performed in the energy range from 70 to 30 keV . This region is undisturbed by scattering events and thus all pulse height channels above a certain threshold energy can be used in the evaluation.

From the calculations the following conclusions can be drawn:
1.) For the two configurations shown in Fig. 1 the neutron sensitivity is about the same. In the energy range from 10 to $80 \mathrm{keV} 90-95 \%$ of the scattered neutrons leave the crystal, while $5-10 \%$ are captured.
2.) The capture events are strongly delayed in time. Only $\sim 1 \%$ the neutrons are captured in the first 100 nsec after leaving the sample.
3.) A much better signal to background ratio is obtained using the $\mathrm{BaF}_{2}$ configuration. This is mainly due to the shorter flight path which causes the same number of primary capture events to be registered in a much shorter TOF interval.
4.) For the $\mathrm{BaF}_{2}$ configuration there is an energy interval from 70 to 30 keV which is undisturbed by capture of scattered neutrons.

In summary, we find that $\mathrm{BaF}_{2}$ is superior to $B G O$ as far as neutron sensitivity is concerned, especially if one keeps in mind that the better time resolution allows for a shorter flight path (resulting in a fourfold neutron flux at the sample position) and for a better energy resolution at the same time.


Fig. 1 Calculated TOF spectrum of capture events (solid line) and background due to the capture of scattered neutrons (dashed line) for a $4 \pi \mathrm{detec}{ }^{-}$ tor of BGO and $\mathrm{BaF}_{2}$.

### 6.1.3 PEAK-TO-BACKGROUND IMPROVEMENT OF SEMICONDUCTOR X-RAY DETECTORS AT 10 keV

D. Gotta, M. Schneider, L.M. Simons

To study the strong interaction effects in antiprotonic atoms formed with hydrogen isotopes with the 'cyclotron trap' (1), the energies (about 10 keV ) and the small intensities of the $\mathrm{n} \rightarrow 1$ and $\mathrm{n} \rightarrow 2$ transitions (of the order of $10^{-2}-10^{-3} \rightarrow$ ?) should be measured with high accuracy. Especially important therefore is a good peak-to-background ratio of the detectors to be used.

The most important sources of background in the interesting energy region (about 10 keV ) are: i) Compton-scattered electrons with an energy around 10 keV or depositing an equivalent energy before leaving the crystal, ii) Compton ganmas of higher energies produced in regions of poor charge collection, iii) backscattered electrons or gammas from the vicinity of the detector producing events in the low energy region, iv) spurious pulses from poor charge collection zones.

The largest background contributing factor (i) can be reduced by the following method: The interaction of a low energy $X$-ray takes place in the first ' $\mu \mathrm{m}$ ' of the crystal (mean penetration depth of a 10 keV X-ray in Ge: $\sim 50 \mu \mathrm{~m}$ ), while the gammas interact uniformly over the whole detector (thickness several mm). This leads to different collection times of the deposited charge and therefore to different pulse-shapes. The gamma events have a much shorter risetime than the pulses from X -ray interactions. This makes pulse-shape investigations to improve the peak-to-background ratio feasible.

In the first tests we used an intrinsic planar Ge-detektor (IGP 87) of 7 mm thickness and an area of 80 mm . With different sources (e.g. ${ }^{207}$ Bi with X-rays and gammas or ${ }^{60}$ Co only with gammas) and different combinations of integration and differentation times two-dimensional spectra (energy vs. pulse-shape) are obtained. In Fig. 1 such a spectrum from ${ }^{207}$ Bi is shown, where two pulse-shape groups with different risetimes are shown. In the second group (slower risetime) the $\mathrm{Pb} \mathrm{L} X$-rays ( 10,551 and 12.614 keV ) and in the first group the events from Compton-scattered gammas are visible.

Taking a pulse-shape window from the second group a drastic reduction of the low energy background is achievable in the energy spectrum. Below 10 keV the reduction attaines $70 \%$ compared to the ungated spectrum, while the reduction of the $\mathrm{Pb} \mathrm{L} X$-ray is on $1 \mathrm{y} 5 \%$.

Fig. 1: Two-dimensional spectrum from ${ }^{207} B i$ (vertical: energy, horizontal:
risetime) with two pulse-shape groups


Fig. 2: Cut through the guard-ring reject detector crystal with inner (i) and outer (0)
detector. The dashed line indicates the electronically reduced inner detector volume

(al1 measure in mm)
To reduce the background contributions from ii), iii) and iv) a guardring reject detector has to be used. Its consists of an inner and outer detector operated in anticoincidence (Fig. 2). With such a detector similar pulse-shape tests are made. It turned out that with the pulse shape technique the inner detector delivers the same results in background reduction as the inner detector in anticoincidence with the outer one. This is understood through the electronic reduction of the volume of the inner detector through the veto of the outer one. In Fig. 2 the resulting inner volume is outlined, while in Fig. 3 the strong suppression of the low energy background through the guard-ring reject configuration is shown with the ${ }^{207}$ Bi source.

These excellent reduction strongly supports the validity of our assumptions made for the pulse-shape technique, because both methods are equally sensitive to the penetration depth.
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P. Blüm, D. Gotta, R. Guigas, H. Koch, W. Kunold, M. Schneider, L.M. Simons, CERN-proposal PS 175/1980

Fig. 3: Spectra of the ${ }^{207} \mathrm{Bi}$ source. a) Inner detector only (the additional peak results from the reset-pulse of the outer detector through crossmtalk). b) Inner detector with anticoincidence of the outer one.



### 6.1.4 AN X-ray ABSORPTTON FLUORESCENCE DETECTOR

D. Gotta, M, Schneider L.M. Simons

One of the planned experiments with the so called 'cyclotron trap' (1) aims to study the muonic $2 \mathrm{~s} \rightarrow 1 \mathrm{~s}$ transitions in gases, especially in Neon. This Ml X-ray transition should show a remarkable forward/backward asymmetry as an effect of parity violating neutral currents.

The problem in observing the M1 $2 \mathrm{~s} \rightarrow 1 \mathrm{~s} \mathrm{X}$-ray line in muonic Neon (207.79 keV) is its proximity to the $2 \mathrm{p} \rightarrow 1 \mathrm{~s}$ transitions (about 207.4 keV ) and the extremely low yield of about $10^{-4}$. To separate the $2 \mathrm{~s} \rightarrow 1 \mathrm{~s}$ from the $2 \mathrm{p} \rightarrow$ ls transitions we considered a triggering scheme via a coincidence measurement with the $3 \rightarrow 2$ transitions. It turned out that there exists a possibility to separate the $3 \mathrm{~d} \rightarrow 2 \mathrm{p}(39.0 \mathrm{keV})$ and the $3 \mathrm{p} \rightarrow 2 \mathrm{~s}(38.6 \mathrm{keV})$ lines with an absorption K -edge. Thus with the coincidence between $3 \mathrm{p} \rightarrow 2 \mathrm{~s}$ and $2 \mathrm{~s} \rightarrow 1 \mathrm{~s}$ one could reach a good separation and also a significant background reduction. For muonic Neon a $400 \mu \mathrm{~m} \mathrm{La}$ foil with a K-edge at 38.925 keV can separate the $3 \mathrm{~d} \rightarrow 2 \mathrm{p}$ and the $3 \mathrm{p} \rightarrow 2 \mathrm{~s}$ transitions. While the $3 p \rightarrow 2 s$ line should be attenuated by a factor of about 2.5 , the $3 \mathrm{~d} \rightarrow 2 \mathrm{p}$ line should be drastically reduced by a factor of about 200.

Fig. 1: Schematic view of the 'absorption filter' with different foils, a Mo grid and the $\mathrm{BaF}_{2}$-detector


In Fig. 1 a possible configuation of such an 'absorption filter' is shown. The Zr foil in front of the La foil is used to attenuate the muonic $\mathrm{Ne}(\mathrm{n} \rightarrow 3$ )-transitions, while the $S m$ foil attenuated the ( $\mathrm{n} \geq 5 \rightarrow 2$ )-1ines. In the La foil fluorescence X-rays are produced isotropically. These X-rays should be attenuated in the following Mo grid of about 2 mm length, with $1 \times 1 \mathrm{~mm}$ spaces and a wire-thickness of $100 \mu \mathrm{~m}$.

For the Ne X-rays coming from a much greater distance the grid has a much higher transparency than the fluorescence $X$-rays, which should be strongly attenuated. Finally the munic Ne $X$-rays should reach a $\mathrm{BaF}_{2}{ }^{-}$ detector (with a K-edge at 37.441 keV ). Thus the Lak X -rays will not be detected.

To demonstrate the feasibility of this method we used an arrangement which simulates this behaviour. We used a Ce source with X-rays at 33.034, $33.44,37.8$, and 38.7 keV . As the absorber material Iodine was chosen, because its $K$-edge ( 33.170 keV ) 1ies exactly between the $K_{\alpha 1}$ and $K_{\alpha 2}$ energies. In the tests an Iodine solution in ethanol was taken and in Fig. 2 the effect of the K-edge absorption is demonstrated. The $K_{\alpha 2}-1$ ine is attenuated by a factor of 3.15 , while the $K_{\alpha-}$ - line is suppressed by a factor 2400 . The effective Iodine thickness is about $300 \mu \mathrm{~m}$. The reduction of the Iodine fluorescence $X$-rays (about $6.0 \%$ ) . through a Mo grid with a thickness of 2 mm is shown in Fig. 3.
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Fig. 2: La-K X-rays without (a) and with (b) Iodine absorption material

$\leftarrow$ Fig. 3: Reduction of the I fluorescence X-rays with a Mo grid.
(a) without, (b) with grid.

### 6.1.5 PERFORMANCE OF A HIGH-RESOLUTION MODULAR NaI-DETECTOR

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M. Suffert ${ }^{+++}$, K. Zioutas ${ }^{++++}$

A NaJ-detector with a big volume (1/12 of the total solid angle) and modular structure was used for the measurement of high energy $\gamma$-quanta, in order to search for Baryonium states (1) and rare neutral annihilation channels (2) in the $\bar{p} p-s y s t e m$. Its properties and performance were described in a recent publication (3).

It turned out that the energy resolution as well as the separation of particles of different kinds was superior as compared to a conventional $12 \times 10^{\prime \prime}$ monocrystalline NaI (4). In order to achieve a good energy resolution in the energy range between 20 and 1000 MeV a careful intercalibration of the 54 modules had to be done. At low energies, where the electromagnetic shower does not escape from one single module, the intercalibration was done with radioactive sources giving a zero approximation calibration for all energies by linear extrapolation. For the final intercalibration the reaction $\pi^{-}$stop $p \rightarrow n \gamma$ was used yielding a 129 MeV monoenergetic $\gamma$-1ine. Events were selected, where more than $60 \%$ of the full energy was concentrated in one module. The addition of the low energy parts of the shower in the surrounding modules (calibrated with sources) must give a total energy of 129 MeV and served thus as a first order calibration of the central module. It turned out that only three iterations of this kind were necessary in order to get the excellent energy resolution of $\Delta \mathrm{E} / \mathrm{E}=3.3 \% \times(\mathrm{E} / \mathrm{GeV})^{-1 / 4}$ ( $\Delta E / E=5.5 \%$ at 129 MeV ), a value which was never reached by the big brother of our detector, the Stanford crystal ball.

The running times for the search for Baryonium states lasted for seve ral months in a row, and a careful stabilization of the detector and the electronics was necessary. This was achieved by a light pulser system consisting of a Xe-flash-tube the light of which was randomized and sent to all 54 photomultipliers. The light bulb itself was stabilized via the light output of a $2 \times 2^{\prime \prime}$ NaI which was illuminated by a radioactive source. All elements of the system could be optimized such that a stability of better than $1 \%$ over a period of several months was reached.

In order to get a backgroundless $\gamma$-spectrum, particularly at energies below 100 MeV , the topology of events of different kinds was studied, and the results were used for a good particle separation. Finally, all charged particles and multiple hits could be eliminated, and it was possible, to get a low energy $\gamma$-spectrum free of neutron-background, which facilitated the identification of small structures in this energy domain.
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### 6.1.6 A LARGE-AREA POSITION-SENSITIVE TTME-OF-FLIGHT COUNTER FOR NEUTRONS AND CHARGED PARTICLES

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A large position-sensitive time-of-flight counter for subnanosecond timing has been designed for pion absorption studies in few-nucleon systems. The new counter has been extensively tested and successfully used in a recent experiment on $\pi^{+} / \pi^{\dot{-}}$ absorption in ${ }^{3} \mathrm{He}$ at 170,220 and $270 \mathrm{MeV} / \mathrm{c}$ performed at the SIN cyclotron. Kinematically complete experiments of this type require the coincidence measurement of either a neutron and a charged particle or of two charged particles, respectively, depending on the charge of the incident pion. Therefore, the time-of-f1ight counter was designed to serve both purposes, i.e. to detect high-energy neutrons as well as energetic charged particles.

The new counter consists primarily of thirty NE 102A plastic scintillator rods, each 200 cm long, 10 cm high, and 5 cm thick. Each rod is viewed from both ends by a Valvo XP 2230 photomultiplier. This provides the event position in the rod from time differences in the two timing signals, and individual flight times of the projectiles (from the target to the detector) by the corresponding mean times. The thirty rods are arranged in a $10 \times 3$ rod matrix to give a total counter size of 200 x $132 \times 15 \mathrm{~cm}^{3}$ with an effective counter area of $200 \times 100 \mathrm{~cm}^{2}$ facing the incident particle beam. The total set of rods is grouped into five vertically separated modules consisting of six rods each with two layers in height and three layers in depth. While the optically separated rods are packed as closely as possible in each module, the different modules


Fig. 1: Reverse side view of the new position-sensitive time-of-flight counter for neutrons and charged particles.


Fig. 2: Light attenuation and propagation characteristics:
a) Measured light attenuation over 200 cm of a main counter rod. The mean pulse heights, $\sqrt{\text { PH1 } \times \text { PH2 }} \times 2$, were calculated off-1ine. b) Measured effective propagation time differences for a 200 cm long, $10 \mathrm{x} 5 \mathrm{~cm}^{2}$, counter rod and a 200 cm long, $25 \mathrm{x} 0.4 \mathrm{~cm}^{2}$, "anticounter" bar.
are separated by 8 cm in height, in order to facilitate their mounting in a common counter rack. For testing and monitoring purposes light emission diodes (LED's) were installed in the light pipes on both ends of each rod. Particle identification in terms of neutrons or charged particles is accomplished by five additional NE 102A "anticounter" scintillators, 200 cm long, 25 cm high and 4 mm thick, placed in front of each of the five counter modules. These anticounters are also viewed from both ends by XP 2230 photomultipliers to provide the necessary timing informations. A photograph showing the whole counter arrangement from the reverse side is shown in Fig. 1.


Fig. 3: Dalitz-1ike plot for protons from the $\pi^{+}{ }^{3} \mathrm{He} \rightarrow$ ppp reaction investigated in a recent experiment at SIN employing the new time-of-f1ight counter.

The specifications of the counter determined from test measurements with a ${ }^{60}$ Co and an Am-Be source are as follows: The light attenuation over the whole rod length of 200 cm amounts to a factor of 2.5 . But, if the geometrical mean of the analog signals from both photomultipliers is calculated off-line, the pulse height variation as a function of the incoming radiation vanishes (see Fig. 2a). For the anticounters the light attenuation factor over the whole length was determined as 8.2. The effective propagation time difference for light initiated by radiation at the near or the far end of the rod was determined as 24 ns . This effective time difference is expectedly the same for the main counter rods and the anticounter bars. The time differences from both photomultipliers versus
the distance from the middle of a rod are shown in Fig. 2 b . The time resolution of the position-sensitive counter was determined both with uncollimated and collimated irradiation from a ${ }^{60}$ Co source against a small fast NE 102 scintillator with a threshold corresponding to 500 keV . This gave a resolution of 850 ps (FWHM) for an uncollimated and of 700 ps for a collimated beam. Using the Monte Carlo Code of Cecil et al. (1) the detection efficiency for neutrons was calculated for a threshold of 5 MeV ee For neutron energies between 60 and 180 MeV the efficiency decreases smoothly from $\sim 21 \%$ to $\sim 15 \%$. Valuable additional checks of the above stated counter specifications are expected from evaluations of the recent data taken at SIN involving monoenergetic neutrons from the $\pi^{-3} \mathrm{He} \rightarrow \mathrm{dn}$ reaction.

The good overall performance of the new time-of-flight counter is certainly best illustrated by some first results obtained from the recent pion absorption experiments. Fig. 3 shows a Dalitz-like plot of protons from the $\pi^{+}{ }^{3}$ He $\rightarrow$ ppp reaction measured in kinematically related elements of the TOF counter and an additional total-absorption scintillation counter. It can be seen that the region outside the kinematically-allowed range is extremely backgroundfree, and that the events from quasifree absorption (covering the region of high event density in the upper part of the figure) are contained in a very narrow band.
(1) R.A. Ceci1, B.D. Anderson and R. Madey, Nuc1. Instr. and Meth. $161(1979) 439$
$+\quad$ On leave from Zagreb University, Zagreb, Yugoslavia

### 6.1.7 INVESTIGATION OF A VERTICAL DRIFT CHAMBER

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The systematic investigation of the properties of a vertical drift chamber (VDC) (1) to be usedas a focal plane detector for LEPS (Low Energy Pion Spectrometer, being presently installed at SIN) has been completed. Fig. 1 shows a cross-sectional view of the VDC perpendicular to the direction in which the $20 \mu \mathrm{~m}$ thick sense wires alternating with $45 \mu \mathrm{~m}$ shielding wires are stretched. The distances between a sense wire and the adjacent shielding wixes is 2 mm . The gap width L , i.e. the distance between the read-out plane and the high voltage planes is 10 mm . The chamber was operated with a $50 \%$ argon, $50 \%$ isobutane mixture at a high voltage of -8 kV .


Fig. 1: Geometry of the vertical drift chamber. The letters a -d denote particle trajectories which correspond to four points in Fig. 3.

As can directly be seen from Fig. 1 the main difference of the VDC as compared with a standard drift chamber consists in the mode of operation: the particle trajectories have to cross the chamber at angles of incidence $\alpha$ which typically range from 30 to $60^{\circ}$. This angular interval coincides with the angular range of particle trajectories crossing the LEPS focal plane which is tilted by $43^{\circ}$ with respect to the central trajectory. The simultaneous measurement of e.g. $N=3-4$ or $N=8-10$ drift times (corresponding to $\alpha=30^{\circ}$ or $\alpha=60^{\circ}$, respectively) allows a precise reconstruction of the trajectories' projection onto the $x-y$ plane.

The redundant time information obtained from the VDC also makes it possible to measure directly its spatial resolution as a function of $\alpha$. This is achieved by comparing the distance in $y$-direction between the trajectory and the sense wire measured in one drift cell with the same number determined from a least square fit to the $y$-values measured in all adjacent cells hit by the particle.

Fig. 2 shows experimental results of the variation $\sigma_{y}(\mathrm{~mm})$ versus $y(\mathrm{~mm})$ obtained with a parallel pion beam of $300 \mathrm{MeV} / \mathrm{c}$ from the $\pi \mathrm{M} 3$ channel of SIN and an angle of incidence $\alpha=60^{\circ}$. In addition, the various contributions to the finite spatial resolution of one drift cell are plotted as dashed curves. The curves a) to $c$ ) represent the contributions due to the primary ionisation statistics calculated by a Monte Carlo method, due to the diffusion of electrons and the electronic time resolution, respectively. The three curves were added quadratically resulting in the full curve, which is found to be in good agreement with the experimental values. The variation $\sigma_{y}$ for one cell was typically $\pm 100 \mu \mathrm{~m}$. The spatial resolution based on drift time
measurements for $N$ sense wires is given approximately by $\sigma_{x} \simeq \sigma_{y} \operatorname{tg} \alpha / \sqrt{N}$, which represents the resolution within the wire plane and equals $\pm 60 \mu \mathrm{~m}$, almost independent of angle.


Fig. 2: The variation of the spatial resolution $\sigma_{y}$ in mm versus the spatial resolution $y$ in mm for one drift cell as obtained in an experiment with a parallel pion beam of $300 \mathrm{MeV} / \mathrm{c}$ from the $\pi \mathrm{M} 3$ channel of $\operatorname{SIN}\left(\alpha=60^{\circ}\right.$ ). The dashed curves a - c represent various contributions to the finite spatial resolution as explained in the text. The full curve is the quadratic sum of the three curves $a-c$ and agrees very well with the experimental points.

Fig. 3: Correlation of the lengths $t_{1} \times v_{D}$ and $t_{2} \times v_{D}$ along the drift paths in two adjacent drift cells $\left(v_{D}=d r i f t\right.$ velocity). The points a-d correspond to particle trajec. ries indicated in Fig. 1.


Fig. 4: Experimental results for the correlation of the drift times $t_{1}$ and $t_{2}$ of two neighbouring cells measured with electrons of a ${ }^{90} \mathrm{Sr}$ source. The shadow of the $45 \mu \mathrm{~m}$ shielding wire between two sense wires is observed at $t_{1}=t_{2}$ revealing the excellent spatial resolution of the drift chamber.


The excellent spatial resolution of the chamber is also clearly revealed by the time correlation spectra of adjacent sense wires. The drift times of neighbouring cells should be located on a $u$-shaped curve as shown in Fig. 3 (see Fig. 1 for the explanation of numbers). Experimentally electrons from a ${ }^{90}$ Sr source were detected and the drift time correlation of two adjacent wires relative to a scintillation counter was registered by a twodimensional multichannel analyser. The result is shown in Fig. 4 for $\alpha \simeq 45^{\circ}$. For $t_{1}=t_{2}$ the number of events on the $U$-shaped curve is drastically reduced due to electrons, which are scattered by the shielding wire between the two sense wires. Hence, Fig. 4 demonstrates clearly that the spatial resolution is good enough to display the "shadow" of a $45 \mu \mathrm{~m}$ thick shielding wire.
(1) Annual Report on Nuclear Physics Activities 1981/82

KfK 3427 (1982) 210
6.1.8 DEVELOPMENT AND TEST OF MWPC's FOR NEUTRON INDUCED REACTIONS
A. Bischoff, F.P. Brady, P. Doll, H.O. Klages, H. Krupp

In continuation of our work to investigate the operation of multiwire proportional chambers for detecting charged particles produced by fast neutrons the MWPC described last year was tested in the neutron beam. The collimated neutron beam from POLKA exhibits a bell-shaped energy distribution extending up to a twobody reaction peak. It turned out that under normal operation conditions (- 2500 V on the Cathode foils, about $74 \%$ Argon, $1 \%$ Freon (13B1) and $25 \%$ Isobuthane at normal pressure) signals on the anode wire planes did range from $1-10 \mathrm{mV}$ on $50 \Omega$ load. (With the geometry chosen for this prototype chamber a gas multiplication factor $A \sim 10^{4}$ is garantied.) They result from charged particles like protons and $\alpha$-particles and photons produced by the neutron beam in the 1 mm CH 2 window of the chamber. Tests with an ${ }^{55} \mathrm{Fe}$ source exhibited similar signal heights. However, the energy resolution was poor, as a test run in combination with a charged particle $\Delta E / E$ telescope showed. Several cathode foil materials where investigated, to improve the performance of the chamber. The risetime of the signals on the anode wire plane were about $2-5 \mathrm{~ns}$ and similar fast on the cathode planes.

Fig. 1 shows the time-of-flight distribution between the cyclotron and the charged partic1e event detected in the MWPC. The time signal and the energy signal was deduced from the cathode foils. While the time-of-flight coordinate covers the bell-shaped flux distribution from the neutron beam nearly twice, the charged particle energy spectra are rather continuous. There is a weak indication of a pulse height separation of highly ionizing $\alpha$-particles and less ionizing protons.

Fig. 2 shows the spacial flux distribution of the narrow collimated neutron beam ( $20 \mathrm{~mm} \emptyset$ ) after 5 m flight path behind the collimator. This distribution is measured by means of the two crossed anode wire planes, which cover an area of $100 \times 100 \mathrm{~mm}^{2}$. The multiplied signal on an anode wire propagates into both directions of a well matched delay line, where the time difference between both ends is a measure of the charge coordinate. The longtime behaviour of the chamber in the neutron beam was studied. With charged particle countrates of about 100 Hz , corresponding to a detection efficiency close to one (estimated from an integral conversion rate of neutrons into charged particles), the MWPC worked fine as monitor, however, deteriorated after several days, building up discharge tracks (Ma1ter-effect), due to the lack of methylal in the gas mixture.

For future n, p experiments at the cyclotron a large area MWPC is presently constructed (Fig.3). It will be operated at normal pressure of either "magic" gas (1) or a argon- $\mathrm{CO}_{2}$ mixture.



Fig. 2

## XY-MULTIWIRE-CHAMBER



Fig. 3

This new chamber typ, however, will use the influenced charge on two perpendicular oriented cathode wire planes to determine the position of the charge produced on the thin anode wires.
(1) G. Charpak, H.G. Fischer, C.R. Cruhn, A. Minten, F. Sauli, G. Ploch and G. Flügge, NIM 99 (1972) 279

### 6.1.9 PROPERTIES OF A $\triangle E / E$ TELESCOPE FOR NEUTRON FLUX DETERMINATION

F.P. Brady, P. Doll, E. Finckh, H.O. Klages, J. Wilczynski

At POLKA most of the neutron flux is in the broad lower energy continuum due to deuteron breakup. The full energy peak at 50 MeV contains only about $5 \%$ of the total flux. (However such a continuum beam allows simultaneous measurements at neutron energies from 17 to 50 MeV .) To build a monitor which would be very stable and essentially threshold independentit was decided to use DE•E plastic scintillator telescopes ${ }^{1)}$. These were mounted near the exit of the neutron collimator with the polyethylene $\mathrm{CH}_{2}$ target at $\sim 3.4 \mathrm{~m}$ from the neutron source and two symmetrical $\Delta E \cdot E$ telescopes at small angles $\left( \pm 14^{\circ}\right.$ and later $\left.\pm 20^{\circ}\right)$ to the neutron beam as shown in Fig. 1 . These telescopes also provided accurate measurements of $A_{y}$ for $n-p$ at large scattering angles.


Fig. 1: Horizontal schematic view of the $\triangle E \cdot E$ telescopes
The E plastics are coupled to the PMT via lucite light pipes as shown. Brass collimators 5 mm thick with 40 mm opening diameter are directly in front of the E plastics which are wrapped in $3 \mu \mathrm{Al}$ foil and taped on the sides to be light tight.

The $\Delta E$ plastics were mounted on two lucite "peg-legs" (2-3 mm diameter) so that the spacing between the top of the conical light pipe and the bottom of the scintillator is 11 mm . Clear epoxy is used on the joints, top and bottom of the pegs. This stand-off distance and the use of highly reflective Al foil on the inside of the tent surrounding the scintillator gives good uniformity in light collection from different parts of the scintillator. The uniformity of response of the $\Delta E$ detectors was checked by scanning them with a collimated $\beta$-source. A uniformity of response to within $2-3 \%$ is observed.

The $\Delta E \cdot E$ coincidence gates the $E$ and $\Delta E$ signals and provides the start TOF signal which, relative to the phase of the cyclotron RF, gives a measure of the incident neutron beam energy. The $\Delta E, E$, and TOF parameter along with the telescope addresses are stored on magnetic tape for later off-1ine analysis. Gain changes are accounted for in $\triangle \mathrm{E}$ and E , and time shifts in the TOF so that all runs have the same spectrum in three parameter space. This allows the same time and/or energy bin selection to be made for each run. After chosing an equivalent TOF flight window for the left and right telescope the integrated energy spectra provide an accurate monitor for the neutron flux. In Ay measurements, e.g. in $\vec{n}-p$, the flux cancels out when symmetric left-right pairs are used. However in measurements of $A y$, where the
target spin cannot (in the case of brute-force polarization) be quickly turned over, or in cross section measurements, a stable flux monitor is important.


Fig. 2: $\Delta E$ vs. E matrix of the "left" telescope for one run of 6000 s

For typical polarized deuteron beam currents (25nA) and geometries $\left(\Omega_{\mathrm{n}}=10^{-4} \mathrm{sr}, \Omega_{\mathrm{p}} \sim 10^{-3} \mathrm{sr}\right.$ ) and 3 mm CH 2 target, the count rate in each telescope is $\sim 3 \mathrm{cps}$. Combined, the two telescopes give a statistical uncertainty of $\sim 0.5 \%$ for each run of 6000 secs. The spin "up-down" flux ratio can be monitored to better than $1 \%$.
(1) T.S. Subramanian, J.L. Romero and F.P. Brady, NIM 174 (1980) 475
6.1.10 DETERMINATION OF ${ }^{3} \mathrm{He}-$ DETECTOR PROPERTIES USING A ${ }^{252}$ Cf NEUTRON SOURCE
A. Chalupka*, P. Do11, B. Haesner, H.O. Klages

With the availability of a low-mass ${ }^{252}$ Cf neutron source, embedded in a small ionization chamber (1) it was challenging to attempt to calibrate neutron detectors. The fission chamber was exploited to calibrate a liquid
${ }^{3}$ He-scintillator detector (2). Because the flux distribution of the fission neutrons from the ${ }^{252}$ Cf source is assumed to be known fairly well (3), the absolute cross section for a specific reaction in the ${ }^{3} \mathrm{He}$-scintillator: $3^{H e}(n, p)^{3}$, can be measured. However, this measurement requires besides a careful knowledge of all background and attenuation effects from surrounding materials, a detailed understanding of the pulse height spectrum produced in the liquid scintillator. For a more quantitative understanding of pulse height variations in liquid ${ }^{3}$ He, a standard Monte-Carlo code (4) was modified to calculate multiple scattering effects in a ${ }^{3}$ He-scintillator of 7.2 cm diameter and 7 cm height exposed vertically to a neutron field of 6.4 cm diameter like in the real experiment. Cross sections and angular distributions for elastic scattering and all reaction channels were taken from the literature. For neutron energies between 0.3 and 4 MeV the knowledge of the total, elastic and the partial ( $n, p$ ) cross section is poor (5). This statement holds even more for the differential cross sections. For higher energies, cross sections were taken from ref. 6 and 7. Hardly anything is known about the light output of low energy charged particles in liquid ${ }^{3}$ He. For scattering processes below 3 MeV the calculated pulse height spectra are very sensitive to the angular distributions and the light output for ${ }^{3}$ He particles, protons and tritons. (Further investigations should be carried out in these energy range.) Fig. 1 shows typical pulse height spectra at 2.1 and 4.2 MeV incident neutron energy. While the description of the measured spectra is satisfactory at 4.2 MeV , the above mentioned problems are evident at 2.1 MeV . The low pulse height spectrum is dominated by elastic scattering processes. The structure with the highest pulse height is due to the combined light output for the proton and triton from the ${ }^{3} \mathrm{He}(\mathrm{n}, \mathrm{p})^{3} \mathrm{~T}$ reaction. The steep fall-off (short dashed line) indicates the boundary for events which experienced in the 1 . scattering elastic scattering from ${ }^{3}$ He only. Consequently, the pulse-height range between elastic and $p+t$ scattering events seems to be contaminated by multiple scattering processes. By improving the description of the spectra, we will be able to correct the experimentally evaluated count rates in the ${ }^{3} \mathrm{He}(\mathrm{n}, \mathrm{p})^{3} \mathrm{~T}$ reaction peak taking into account the finite energy resolution. This will allow us to determine the absolute cross section for this monitor reaction.
(1) A. Chalupka, NIM 164 (1979) 105
(2) R. van Staa, J. Reher, B. Zeitnitz, NIM 136 (1976) 241
(3) J.A. Grund1 and C.U. Eisenhauer, Symp. on Neutron Standards and Applic. Gaithersburg 28-31, 1977, NBS 493
(4) G. Kanisch, Diplomarbeit, University Hamburg (1976)
(5) R. Batchelor and K. Parkor, AWRE-Report 0-78/64
(6) B. Haesner, Ph.D. thesis, University Karlsruhe (1982)
(7) M. Drosg, Los Alamos LA-7268-MS 1978 UC-34C


Fig. 1: Pulse height distributions in a liquid ${ }^{3}$ He scintillator detector
6.1.11 SHIELDING OF PMs IN STRONG MAGNETIC FIELDS

F.P. Brady ${ }^{*}$, P. Do11, H.O. Klages, P. Jany

Here we discuss the specific problems of shielding the $\Delta E-E$ telescopes and the neutron detectors from the magnetic field of the polarized proton target.

The magnetic field variation in the median plane (at 9 Tesla central field) is shown in the figure below. The detector arrangement for the $\vec{n}-\vec{p}$ Ay measurement was the following: The E-detectores from the telescope were 125 cm from the target center ( 70 Gauss), the $\Delta E$-detectors were 140 cm from the target ( 55 Gauss); a pair of small neutron detectors were at 55 cm ( 900 Gauss) and at 75 cm ( 338 Gauss) and a pair of large neutron detectors were at 125 cm ( 70 Gauss) from the target center.


Fig. 1: Radial dependence of the external field of the polarized proton target

The magnetic field inside the detector should be $\ll 0.5$ gauss for a gain shift of < $10 \%$ in XP 2020 and 2040 tubes. ( 0.1 G gives $<\underset{\sim}{\alpha} 1 \%$ gain shift.) The main contributions to $B_{i}$, the field inside the shields, from the external field, $B_{e x t}$, come from through field, $B_{i}^{t}$, and both end fields, $B_{i}^{e}$ (1). These contributions for single shields of inside diameter $D$, wall thickness, $d$, length, L, and (high) permeability, $\mu$, have the form

$$
\begin{aligned}
& B_{i}^{t} \simeq B_{e x t} \cdot D / \mu d \\
& B_{i}^{e}(x) \simeq B_{o p} e^{-2 k \cdot x /} D
\end{aligned}
$$

$x$ is the acial distance from the opening, $B_{o p}$ is the field on axis at the opening. These depend on whether $B_{\text {ext }}$ is parallel or perpendicular to the cylinder axis, and are respectively (static field):

$$
\begin{array}{lll}
\mathrm{B}_{\mathrm{op}} \simeq 1.3 \sqrt{\mathrm{~L} / \mathrm{D}} \mathrm{~B}_{\text {ext }} & \mathrm{k}=2.25 & \text { parallel } \\
\mathrm{B}_{\mathrm{op}}=0.33 \cdot \mathrm{~B}_{\text {ext }} & k=3.5 & \text { perpendicular }
\end{array}
$$

Quenching factors $Q \equiv B_{i} / B_{e x t}$ or shielding factors, $S=Q^{-1}$, are defined in this way. Because of the smaller $k$-parallel it was determined for the case of $B_{\text {ext }}$ parallel to the axis of the cylindrical shield that it was not possible to shield the $\Delta E$ detectors adequately in this geometry. So the scintillators were rotated to have $B_{\text {ext }}$ perpendicular to the cylinder axis.

Due to saturation effects it was necessary to use (at least) two shields, the outer one of iron which saturates around 20 kG (in the iron).

In the case of a $\mu$-metal and a Fe shield the total shielding factor $S_{\text {tot }}$ is somewhat less than $S_{\mu} \cdot \mathrm{S}_{\mathrm{Fe}}$. In the case of through fields, $\mathrm{S}_{\text {tot }}$ is further reducted by approximately $4 \Delta / D$ (Fe) where $\Delta$ is the air gap between the $\mu$ and Fe shields.

Similarly for end fields the shielding factor for double shields is taken to be

$$
B_{i}^{(e)}(x)=B_{i}^{(e)}\left(x_{o}\right) \cdot e^{-2 k x / D}
$$

where $B_{i}^{(e)}\left(x_{o}\right)$ is the end field for the inner cylinder which is $x_{o}$ inside the outer cylinder.

For the $\Delta E$ and $E$ detectors it was necessary to shield the anodes with $\mu$ metal shields $(D=60, L=110 \mathrm{~mm})$ which fitted tightly around the bottom of the tube. A second $\mu$ shield $D \approx 76 \mathrm{~mm}$ was used to shield the photocathodes. This overlapped the anode shield and enclosed the scintillator and light pipe of the $E$ and went nearly to the top of the light pipe in the $\Delta E$ detector. Iron shields of 3 mm wall enclosed the $E$ from base to the end of the collimator, and the $\Delta E$ from base to the bottom of the scintillator. These arrangements gave shielding factors of $10^{3}$ for both "end" and "through" fields in the $\Delta \mathrm{E}$ and E detectors.

For the small neutron detectors, two separate $\mu$ shields were necessary. The outer Fe shield was 5 mm thick and ran from base to the end of the scintillator ( 50 mm long on a 50 mm long light pipe) except for a slot at the end for the expansion tip. It was estimated that the field in the Fe cylinder's wall was on the verge of saturating. In tests with the magnetic field it was shown that the extra layer of $\mu(d=1 \mathrm{~mm})$ was necessary to avoid a large gain decrease in the field.

The large neutron detectors had one large $\mu$ metal shield ( $d=1 \mathrm{~mm}$ ) and outside that a soft iron shield also of $d=1 \mathrm{~mm}$. The field in the Fe wall is close to saturation. A thickness of $2 \rightarrow 3 \mathrm{~mm}$ would be safer. (But this makes the detectors heavier.) Assuming the field in the walls is < 10 kG (as will be the case for most of the length then $\mu_{\mathrm{Fe}}$ at 1 to 10 kG is $\simeq 2000$. For $\mu$ metal assume $\mu=50 \mathrm{~K}$. Then through field quenching is $Q_{\text {tot }} \sim 0.003$. In the actual experimental arrangement the gain changes for the telescope detectors due to $B$ field off and on were reasonably small, of the order of $10 \%$. While for the neutron detectors at the largest distance no changes where observed the closest detectors exhibited changes of the same order of magnitude.
(1) Wirbe1ströme und Schirmung in der Nachrichtentechnik, H. Kadem, Springer-Ver1ag (1959)

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### 6.2 INSTRUMENTATION

### 6.2.1 POLARIZATION OF PROTONS IN TiH2 SAMPLES

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A brute-force polarized proton target has been developed for scattering experiments with fast polarized neutrons. The target material is pressed TiH 2 powder which has been pressed into copper cylinders. A method to determine the proton polarization is described here, in which the heat capacity of the proton spins in the polarizing magnetic field is measured.

THe schematic set-up is shown in Fig. 1. The sample employed here has a length of 3.7 cm and a diameter of 2.5 cm . It has in its center a small copper rod on which a cobalt crystal (with a small amount of ${ }^{60}$ Co) and a


Fig. 1: Schematic view of the set-up inside the cryostat
small heating coil are mounted. The anisotropy of the ${ }^{60}$ Co $\gamma$-radiation is used for the temperature determination. With the heating coil accurate amounts of heat can be supplied to the sample. A second copper rod connects the sample
to the mixing chamber of a ${ }^{3} \mathrm{He}-{ }^{4} \mathrm{He}$ dilution refrigerator. Shortly below the mixing chamber this copper rod has been interrupted to insert a thermal switch. This is a piece of high-purity tin, which can be made superconducting or normally conducting by changing the magnetic field of a small coil around it. In the superconducting state the thermal conductivity is very low (switch open), in the normal state it is very high (switch closed).

The measurements were carried out as follows. First, the sample was polarized in a field of 4.2 T with the switch closed. It lasted about 30 h to reach 10 mK . Then the switch was opened to isolate the sample. Small amounts of heat were given to the target until it had reached about 30 mK . After each heat input we waited $0.5-1 \mathrm{~h}$ for the temperature to stabilize at its new higher value. These warming-up measurements were carried out 4 times with a $\mathrm{TiH}_{2}$ sample and 2 times with a Ti dummy sample with the same dimensions and the same amount of Ti as the $\mathrm{TiH}_{2}$ sample.

The dummy measurements are necessary to determine the heat capacity, which is not due to the protons. This background heat capacity is mainly caused by the copper nuc1ei in the cylinder and in the rods, which polarize to some extent. The background heat was determined to be about $30 \%$ of that of the protons. This agrees with the calculations.


Fig.2: The theoretical polarization of protons in a field of 4.2 T (full line) compared with the experimental results

The amount of heat $\Delta Q$, that a number of $N$ protons in a magnetic field $B$ can absorb is equal to $\Delta Q=N \mu B \quad \Delta p$, where $\Delta p$ is the change in polarization. With this relation the amount of heat $\Delta Q$ that was supplied to heat the sample over a certain temperature interval can be converted into a polarization decrease $\Delta p$ over that interval. These polarization steps are shown in Fig. 2 . The warming-up could not be followed until zero polarization because the temperature measurements with ${ }^{60}$ Co is too inaccurate above 30 mK .

Thus from these measurements an absolute value for the polarization is not obtained, but they allow to verify whether the theoretical curve is followed property between 10 and 30 mK . The data of the 4 runs were normalized to the theoretical curve at the highest temperature, where deviations are supposed to be smallest. Although there are large differences between the results of the different runs, the data follow the theoretical curve quite well on the average. The data indicate that the polarization at 10 mK has been at least $35 \%$. It may be expected that in a larger field larger polarizations are achievable.

### 6.2.2 DEMAGNETIZATION OF PROTONS IN TiH 2 SAMPLES

R. Aures, W. Heeringa, R. Maschuw, F.K. Schmidt

Samples of pressed titanium hydride powder have been prepared to be used as polarized proton targets in scattering experiments with fast polarized neutrons. The protons are polarized with the brute-force method. Results of polarization measurements are given elsewhere in this report.

A sample of polarized protons is a potential nuclear refrigerant: it is a cooling medium, with which it should be possible to reach very low temperatures by adiabatic demagnetization. In fact titanium hydride shows some favourable properties compared to other nuclear refrigerants. It has a high nuclear density and the magnetic moment of the proton is large. Hence a large entropy reduction per unit of volume can be obtained. This property is expressed in a high value of the molar Curie constant $\lambda$ divided by the molar volume $V$. In Table 1 the values of $\lambda / V$ are compared for several possible materials for nuclear refrigeration. Another advantage of $\mathrm{TiH}_{2}$ is that it does not become superconducting. For superconductors it is not possible to continue the demagnetization to below the critical field $B_{c}$. We report here on the
first adiabatic demagnetization experiments carried out with protons.

| Material | $\lambda / \mathrm{V}(\mu \mathrm{K})$ | $\mathrm{B}_{\mathrm{c}}(\mathrm{mT})$ |  |
| :--- | :---: | :--- | :--- |
| Nb | 1.99 | 250 | Table $1:$ |
| V | 1.91 | 13 | Molar Curie constant $\lambda$ di- |
| TiH | 2 | 1.62 | - |
| In | 1.11 | 30 | vided by the molar volume V |
| Al | 0.87 | 10 | for some materials. The |
| Cu | 0.57 | - | critical field $\mathrm{B}_{\mathrm{c}}$ is also |
| T 1 | 0.21 | 20 | listed |

The set-up was basically the same as that shown in the previous contribution. The titanium hydride is contained in a copper cylinder, which is connected to the mixing chamber of the dilution refrigerator by a copper rod. Because the Cu -atoms in the sample also polarize and demagnetize, we have made a dummy sample with the same dimensions containing pressed titanium powder instead of pressed titanium hydride powder. The results of two demagnetization runs are shown in Fig.1. The samples were magnetized at 4.2 T and precooled to somewhat below 10 mK . The demagnetization was performed in 3 h to a final field of $B_{f}=0.22 \mathrm{~T}$. Ideally the temperature of the $\mathrm{TiH}_{2}$ sample should have followed the lower straight line ending at 0.55 mK . The measured temperature


Fig. 1: Results of the demagnetization experiments
did however not fall below 2.0 mK . This difference can not be explained by eddy currents, because a calculation allowing for very large eddy currents still comes down to 0.75 mK (upper straight line). We suppose that our thermometer ( ${ }^{60}$ Co $\gamma$-anisotropy) was not reliable below about 3 mK and that the actual temperature has come below 2 mK . This is confirmed by the following observation. After 5 h we heated the samples with an amount of 0.64 mJ . The nuclei in the dummy sample could not absorb this amount, hence its temperature remained high, whereas the $\mathrm{TiH}_{2}$ sample came back to 3.2 mK after having been heated from 2.6 mK . Theoretically the $\mathrm{TiH}_{2}$ sample could only absorb 0.11 mJ between 2.6 and 3.2 mK . If we suppose that the 3.2 mK are correct, than the starting temperature should have been 1.3 mK instead of 2.6 mK to absorb 0.64 mJ .

We conclude that demagnetized protons in $\mathrm{TiH}_{2}$ can effectively be used as a heat sink at least down to 2 mK . More precise measurements, falling outside the scope of our research program, are necessary to establish the properties at lower temperatures.

### 6.2.3 INJECTION STUDIES WITH A CYCLOTRON TRAP

P. Blium, D. Gotta, W. Kunold, M. Schneỉder, L.M. Sịmons

The cyclotron trap (1) is a device designed to achieve high stopping densities of exotic atoms in dilute gases with pressures even lower than 1 Torr. The first problem is to inject the charged particle beam into the Trap in a way, that experiments can be done under best conditions. The whole injection process, which is similar to the injection of particles into a storage ring, is divided into two parts: 1) the transport of the beam through an entrance moderator into the focusing region of the Trap, 2) keeping the injected beam inside the Trap and avoiding that the beam hits the entrance moderator again after one or more revolutions.

To inject a beam, whose initial momentum is larger than the 'acceptance momentum' (about $110 \mathrm{MeV} / \mathrm{c}$ ), at a desired radius, requires a moderator block with a given thickness to slow down the particles of the beam. The injection radius must be less than 140 mm , because of the radial focusing properties of the field. For radii less than 140 mm the field index $n$ is less than 1 , which means that particles whose momenta are essentially equal to the equilibrium momentum are focused to a therefore
given equilibrium radius, which must be smaller than the injection radius. The difference between injection radius and equilibrium radius should not be too large, to keep the betatron amplitudes small. From injection studies with the computer code PATRAC (2) and first experimental injection studies in air as well as studies with alpha-particles im Helium, it turned out, that the injection radius has to be about 125 mm , whereas the equilibrium radius is then about 115 mm .

In order to keep the geometrical extension of the beam passing through the entrance moderator as small as possible, and therefore the amount of accepted particles of the beam as high as possible, it is necessary to overlap the acceptance of the Trap and the emittance of the initial beam. A maximum overlap of the acceptance of the Trap and the emittance of the beam requires essentially, that an injected beam has to have a given phase space ellipse. In addition the acceptance of the Trap as well as the emittance of the beam depend on the plane, where the beam 'sees' the magnetic field. To realise a suitable overlap of acceptance and emittance in the experimenc, it is necessary to shield the incoming beam with an iron tube from the magnetic field up to a certain plane, which should be tangential to the rotationally symmetric field to avoid an azimuthal disturbance of the field. The measured small radial disturbance was incorporated into PATRAC, allowing more realistic conditions for the calculations of the now slighly different acceptance of the Trap.

Fig. 1 shows the calculated vertical acceptance of the Trap with the actual configuration of the iron shielding compared with the $\overline{\mathrm{p}}$-beam at LEAR with a distance between the Trap and the last quadrupole of about 3.5 m . To keep the once accepted particles inside the radial focusing region of the Trap, the beam coming out of the entrance moderator (injection point) must be restricted to a limited range of radial and axial momenta. At the injection point the beam starts with betatron oscillations, corresponding to the radial and momentum mismatch and the radial momentum itself.

Because the energy loss, and therefore the decreasing of the equilibrium radius, for muons and pions is too small in a gaseous medium, the beam would hit again the entrance moderator after one of the following revolutions (Poincaré's Theorem). To avoid this, an inner soft moderator system (polyethylene foils, some hundred microns thick) was installed. To achieve then a well concentrated stopping distribution, the absolute value of the radial momentum must be kept as low as possible. With a suitable

Trap acceptance / LEAR- $\overline{\mathrm{p}}$ beam ernitiance

$\mathrm{p}=300 \mathrm{MeV} / \mathrm{c}$
Fig. 1: Vertical acceptance of the cyclotron trap for $300 \mathrm{MeV} / \mathrm{c}$ antiprotons, compared with vertical emittance of the $\overline{\mathrm{p}}$-beam at LEAR at a distance of about 3.5 m from the window
arrangement of the foils this objective can be reached, thus ensuring that the absolute value of the radial momentum at the injection point is not exceeded during the whole slowing down process. Values from about $-5 \mathrm{MeV} / \mathrm{c}$ to about $-10 \mathrm{MeV} / \mathrm{c}$ in combination with the foil are sufficient for the particles to pass the entrance moderator, and to get a concentrated stopping distribution.
(1) P. Blüm, E. Borie, D. Gotta, H. Koch, W. Kunold, M. Schneider, L.M. Simons, SIN proposal R8-02.1, and
P. Blüm, D. Gotta, R. Guigas, H. Koch, W. Kunold, M. Schneider, L.M. Simons, CERN proposal PS 175/1980
(2) W. Kunold, Annual Report on Nuclear Physics Activities 1981/82, KfK 3427 (1982)

### 6.2.4 SLQWING DQWN OF $\alpha$-PARTICLES IN A CYCLQTRQN TRAP

P. Blüm, D. Gotta, W. Kunold, K. Meissner, M. Schneider, L, M. Cimons

To slow down particles in a "cyclotron trap" (1) the beam must be injected into the region of field index $n$ between 0 and 1 , where the
focusing properties of the magnetic field force the particles to perform betatren oscillations around equilibrium orbits in the symmetry plane. In a first measurement the incoming particle beam has been replaced by a (collimated) ${ }^{241}$ Am $\alpha$-source with an angular distribution of 150 mrad FWHM and $p_{\alpha}=184 \mathrm{MeV} / \mathrm{c}$ momentum. The momentum $\mathrm{p}_{\alpha}$ corresponds to $92 \mathrm{MeV} / \mathrm{c}$ momentum for single charged particles and an equilibrium orbit of 11.6 cm for a 3.7 T field in the central region. The $\alpha$-particles were detected by a cylindrical plastic scintillator rod of 5 mm diameter and 80 mm extension perpendicular to the symmetry plane. The $\alpha$-source and the detector were mounted opposite in a vacuum chamber. The radial positions of source and detector could be varied countinously from outside in the range of $r=0$ to 18 cm (Fig. 1).

Fig. 1: Experimental set-up



Fig. 2: Radial distribution of the $\alpha$-particles in the center of the He-gas chamber. 0 o: image of the $\alpha$-source after $1 / 2$ revolution (with diaphragm), $*:$ radial distribution corrected for detector efficiency (without diaphragm). Statistical errors only, - - : Prediction by the scaling factor $\sqrt{1-n}$

First the image of the $\alpha$-source after $1 / 2$ revolution in vacuum was recorded (diaphragm behind the source from $r=7$ to 18 cm ). Second the motion in a He atmosphere of different pressures was examined. Due to energy loss the particles spiral toward the center. A minimum pressure of 8 mbar is required to protect the $\alpha$-particles from hitting the back of the source.

Fig. 3: Comparison between predicted and measured radial distribution of the $\alpha$-particles. $*$ : measurement (statistical errors only), : simulation without decharging of the $\alpha$-particles, $o$ : simulation decharging taken into account


The maximum rate was achieved by injecting $12 \%$ outside of the equilibrium orbit ( $r=13.1 \mathrm{~cm}$ ) and the minimum pressure ( 8 mbar ). The intensity increases by the predicted factor $\sqrt{1-n}$ (Fig. 2). The rates are corrected for the detection efficiency of the detector.

The results were compared to a computer simulation. The measured radial distribution at 8 mbar and the simulation with and without decharging of the $\alpha$-particles taken into account is shown in Fig. 3.
(1) P. Blüm, E. Borie, D. Gotta, R. Guigas, H. Koch, W. Kunold,
M. Schneider, L.M. Simons, SIN proposal R-81-02.1, and
P. Blüm, D. Gotta, R. Guigas, H. Koch, W. Kunold, M. Schneider, L.M. Simons, CERN proposal PS 175/1980

### 6.2.5 FIRST EXPERIENCE WITH THE MAGNETIC SPECTROGRAPH "LITTLE JOHN"

H.J. Gils, J. Buschmann, M. Heinz, J. Krisch, H. Rebe1, S. Zagromski, H. Schlösser ${ }^{+}$

First test measurements have been performed with the magnetic spectrograph "Little John" partly using provisional equipment since some parts of the vacuum system (pumps, acceptance slits, target sluice) and of the control system (1) have not been completed. The main purposes of the tests were to study vacuum properties, reproducibility and constancy of magnetic fields and to obtain rough estimates of the ion optical properties of the spectrograph. The results of these tests are summarized in the following:

- Magnets: The settings of the magnets were found to be close to the design values and the reproducibility was sufficient within the so-far reduced requirements of the tests. No instabilities have been observed. The computer
controlled operation (software not yet realized) is necessarily needed in particular for calibration purposes during the experiments.
- Vacuum system: The design goal of the vacuum system of a pressure of $10^{-6}$ mbar anticipating a leakage rate of less than $10^{-5} \mathrm{mbar} 1 / \mathrm{sec}$ has been reached only when the focal plane detector was not mounted. The leakage rate of the detector foil ( $25 \mu \mathrm{~m}$ thickness, $50 \times 6 \mathrm{~cm}^{2}$ area) separating the vacuum system from the gas volume of the detector was found to be of the order of $10^{-3}$ mbar $1 / \mathrm{sec}$ making the installation of an additional pump unit (already considered in the control system (1)) just in front of it necessary in order to reach the vacuum design goal under realistic conditions. The use of a cryotrap (2) between beam-1ine and target chamber was successful in bridging the difference in pressure of more than one order of magnitude between these parts.
- Ion optical properties: The ion optical properties have roughly been studied for one focal plane position FP8 (3). The preliminary values found are for the energy range $E_{\min } / E_{\max }=0.8$ and for the momentum resolution (beam spot $\emptyset 1 \mathrm{~mm}$, position resolution $\Delta x=0.8 \mathrm{~mm}$ ) $\Delta \mathrm{p} / \mathrm{p}=8 \times 10^{-4}$.

These values indicate that already in the present status the design values (3) are closely approached.
(1) J. Buschmann, F. Deutsch, H.J. Gils, G. Ludwig, D. Manger, H. Rebe1, K.J. Rist, W. Seith, T. Thouw, O. Walz, contr. 6.2.3
(2) J. Buschmann, G. Ludwig, unpublished results, Kernforschungszentrum Kar1sruhe (1981)
(3) H.J. Gils, KfK 2972 (1980)
H.J. Gils, J. Buschmann, H. Rebel, S. Zagromski, unpub1ished results (1980)

+ Tandem Labor der Universität Erlangen


### 6.2.6 CHARACTERISTIC FEATURES OF THE DETECTION SYSTEM OF THE MAGNETIC SPECTROGRAPH "LITTLE JOHN"

S. Zagromski, M. Heinz

The focal plane (FP) detector (1) and its periphery were mounted and tested including operation in the spectrograph. The gas supply (argon/methan, $90 / 10 \%, \mathrm{p}=970 \pm 0.2 \mathrm{mbar}$ ) and the high voltage were optimized for best resolution of the two position sensitive (PS) detectors and the energy loss ( $\triangle E$ ) detector, respectively. Up to now a position resolution of 0.8 mm (FWHM)
was achieved with an alpha source $\left({ }^{241} \mathrm{Am},{ }^{244} \mathrm{Cm},{ }^{239} \mathrm{Pu}, 0.1 \mu \mathrm{Ci}\right)$ and a collimator of $0.2 \times 10 \mathrm{~mm}^{2}$ in front of it. The energy resolution of the $\Delta E-$ detector was determined with the same source to be 67 keV .

The window foils ( $25 \mu \mathrm{~m}$ hostaphan) of the FP-detectors burst two times in the spectrograph since they were stretched by gas pressure in the opposite direction than in laboratory tests by air pressure. This can be avoided when taking care that the foils are stretched only in one direction.

Preliminary data acquisition programs were developed which include charge division, display of spectra, particle track projections from the PS detectors and two dimensional displays for the $\Delta E$ - and the scintillation (E) detector. A position spectrum of the first PS detector on the focal plane obtained from alpha particles (mixed Am, Cm, Pu source) is shown in Fig. 1.

In addition to the $F P$-detector the spectrograph is equipped with an acceptance detector (2) placed between the target chamber and the mechanical acceptance slit. The detector and its supply have been collected and tested in the laboratory with an $\alpha$-source. Optimal values for gas pressure (isobutane) and anode voltage were found to be 16 mbar and 650 V , respectively. The position resolution was deduced to be $0,6 \mathrm{~mm}$ (FWHM), which should be improved furtheron. The efficiency of the detector does not yet fulfill the requirements. Changes of the geometry of the anode and cathode are promising ways to reach the planned detector specifications (2).


Fig. 1 Position spectrum from the first position sensitive detector obtained with an alpha source $\left({ }^{241} \mathrm{Am},{ }^{244} \mathrm{Cm},{ }^{239} \mathrm{Pu}, 0,1 \mu \mathrm{Ci}\right)$ collimated with a slit of 0.2 mm width just in front of the detector at three positions of 170 mm distance.
(1) S. Zagromski, unpublished results (1980)
S. Zagromski, H.J. Gi1s, H. Rebel, KfK 3068 (1980) p. 111
(2) M. Heinz, H.J. Gils, H. Rebel, KfK 3427 (1982) p. 170
M. Heinz, Diplomarbeit Universität Karlsruhe (1983)

### 6.2.7 THE CONTROL OF THE MAGNETIC SPECTROGRAPH "LITTLE JOHN"

J. Buschmann, F. Deutsch, H.J. Gils, G. Ludwig, D. Manger, H. Rebel, K.J. Rist, W. Seith, T. Thouw, O. Walz

The computer control of the magnetic spectrograph performs the following functions:

1. Setting and control of the magnetic fields
2. Setting of the acceptance slits
3. Setting of the target position and angle
4. Setting of the detector support
5. Control of the vacuum system.

Most of the control and driving functions can be done either manual$1 y$ or computer controlled according to the choice of the experimenter. The tentative inclusion of a computer offers useful possibilities like automatic control or documentation of relevant experimental parameters.

The computer control is provided by a CAMAC system connected to a NOVA 2 computer (Fig.1) which is available exclusively for this purpose. The choice between manual and computer control is done by key switches on the contro1 boards.

The magnetic field control and regulating units are the same as for the beam transport magnets of the Karlsruhe isochronous cyclotron (1).

For the mechanical setting of the acceptance slits, the target position, the target angle and the detector support angle the control concepts are similar: The desired positions will be given either by manually operated multi-switches or by the computer via CAMAC. The actual positions will be read by 36 turns absolute angular digitizers with 100 digits per turn. When the difference between the preselected and the actual position values is nonzero, the stepping motors are engaged in the required sense, otherwise they are stopped. In case of manual control, a start signal is needed to activate the comparator.


Fig. 1
Block diagram of the computer control system of the magnetic spectrograph

For the control of the vacuum system, two $19^{\prime \prime}$ switch boards are provided, one of which is installed in the experimental hall and the other in the electronic room for remote operation. It allows for simultaneous control of all important functions and to indicate the success of an action. In case of manual control, each manipulation on the push buttons directly drives the corresponding aggregates. In the computer control mode, manipulations of the vacuum system still have to be carried out manually via the switch boards. The corresponding switching signals are generated by CAMAC only if the control program (at present under development, (2)) accepts them to be allowed. In addition, in case of unforeseen events (vacuum leak etc.) the computer response protects the vacuum system against consequential damages.
(1) W. Kneis, KfK-report 2835 (1979)
(2) D. Manger, Diplomarbeit 1982

### 6.2.8 STATUS OF THE LOW ENERGY PION SPECTROMETER PROJECT

H. Matthäy, A. Höhne, K. Kärcher, W. Kluge

The mechanical design and the construction of the magnetic spectrometer LEPS (Low Energy Pion Spectrometer) have been completed during the period of the present report.

A cross-sectional median plane view of LEPS is shown in Fig. 1. The design parameters have been given elsewhere (1,2). Fig. 1 displays the spectrometer in the so called HHH-mode of operation which will be used for test runs at the $\pi M 3$ channel (HHH stands for Horizontal dispersion, Horizontal scattering, Horizontal analysis). Finally LEPS will be operated in the VHVmode for which it is optimised (V for Vertical).

The main components of LEPS are a quadrupole triplet $Q_{1}, Q_{2}, Q_{3}$ and two dipoles $D_{1}$ and $D_{2}$, the latter being mounted in a splitpole assembly. The quadrupoles refocus the target spot onto an intermediate image detector (ac-


Fig. 1: Sketch of the Low Energy Pion Spectrometer LEPS. SC = Scattering Chamber, Q1-Q3 = Quadrupoles, $I D=$ Intermediate image detector, D1, D2 = Dipoles of the Splitpole assembly, FD = Focal plane detector. The dashed-dotted curve corresponds to the unphysical lay-out ray of the code RAYTRACE of $H$. Enge and S. Kowalski.
tually two multiwire proportional chambers) determining the primary momentum of the incoming pions. The momentum of the scattered pions is analysed by the splitpole using the focal plane detector (a vertical drift chamber (3)) and the intermediate image detector which both measure coordinates and angles of the particle trajectories. This allows detailed software corrections indispensable for a broad range ( $\Delta \mathrm{p} / \mathrm{p}=-15,+20 \%$ ) and large acceptance ( $\Delta \Omega \approx 20 \mathrm{msr}$ ) spectrometer.

So far all parts except for the vacuum chamber around the dipoles have been delivered. The pole plates of the two dipoles, as delivered by the manufacturer, had, however, intolerable mechanical deviations from the specifications and had to be remachined.

After the reassembly of the splitpole detailed measurements of field maps will be started. Higher multipole field components of the quadrupoles have proved to be sufficiently low after some shimming.

In the meantime considerable efforts have been paid, together with the SIN beam line group, to the design of the magnetic pion channel for LEPS. The basic requirements to be met by this channel in order to match the design parameters of the spectrometer are as follows:

```
energy range 30-100 MeV
vertical dispersion D D 5 cm/(% in }\Delta\textrm{p}/\textrm{p}
momentum bite }\Delta\textrm{p}/\textrm{p}=\pm1
momentum resolution < 10-3 +)
flux }\quad=5\cdot1\mp@subsup{0}{}{6}\mp@subsup{\textrm{m}}{}{+}/\textrm{s}+\mathrm{ +),++)
+) for the full solid angle,
++) at 50 MeV for a proton current I}\mp@subsup{I}{p}{}=100\mu\textrm{A}\mathrm{ and full momentum
        bite.
```

After thorough studies of numerous possible variants for pion beams the most promising solution was found for the existing $\pi E 3$ channel by adding a. $90^{\circ}$ bending magnet, four quadrupoles and one sextupole. Before finally freezing the beam parameters careful raytrace calculations including both the beam and the spectrometer have stili to be carried out.
(1) H. Matthäy, W. Kluge, H.A. Thiessen

Annual Report on Nuclear Physics Activities
KfK 3427(1982)208
(2) W. Kluge, H. Matthäy

SIN Jahresbericht 1982, p. JB23
(3) J. Jaki, H. Matthäy, A. Höhne, K. Kärcher, U. Klein, W. Kluge Contribution to the present Annual Report on Nuclear Physics Activities 1982/83, p. 149

### 6.2.9 A FAST $\mu$ P-CONTROLLED DATA ACQUISITION FOR SPECTROSCOPY EXPERTMENTS

T. Köhler, U. Raich ${ }^{+}$, J. Hauth

In experiment PS-176 at the Low Energy Antiproton Ring (LEAR) at CERN we had to face new problems in the readout of the experimental data due to extremely high event rates ( $10^{4}$ ) and due to the unusual beam structure of LEAR ( 15 min beam - 3 min pause). 6 semiconductor detectors are used to look for atomic X-rays from antiprotonic atoms. Each of these gives energy and time information. Since no other detectors are needed by the experiment, aside from scintillation counters used for trigger purposes, the complete information of an event may be stored in 2 dimensional spectra as long as one does not consider coincidence events. This sort of data structure is typical for spectroscopy experiments which are usually read out with multichannel analysers (MCAs). Since we are using 6 detectors covering different energy ranges, 6 MCAs including magnetic tape stations would be required. Such a set-up seems unreasonable because of the high cost. MCAs are also not able to cope with coincidence events.

The data acquisition system developed in our group uses a microcomputer system acting like a collection of 6 MCAs. The microprocessor system consists of a VME crate containing 2 Mbytes of memory and a CPU card with a MC 68000 microprocessor. The VME bus has been developed and standardized by several big microcomputer companies (Motorola, Mostek, Philips etc.). The large memory is necessary to allow the generation of 2 dimensional spectra with much higher precision than could be obtained with MCAs. The microprocessor program simulating the MCAs is a rather short assembler language program wich runs extremely fast. The readout speed obtained is comparable to MCAs ( $40 \mu \mathrm{~s} /$ event). One big advantage of the system is the capability of taking coincidence events by simply storing them into a memory region reserved for that purpose. The coincidence events may then later be evaluated offline.

The interfacing of the VME system to the experiment is done through CAMAC. CAMPORT, a CAMAC module developed in our group, allows the microprocessor to act as an auxiliary crate controller when used in conjunction with an A2 crate controller. This means that the MC 68000 may read the energy information stored in CAMAC $A D C S$ and the time information stored in

CAMAC TDCs. It may also execute any other CAMAC funtion desired. On the other hand CAMPORT may be used as a normal CAMAC station executing CAMAC commands ịssued by a minicomputer ( a PDP-11 in our case). The CAMAC commands implemented in CAMPORT allow the PDP-11 to read and write VME memory and to interrupt the MC 68000. Since there are no pripherals connected to the VME system the spectra accumulated must be transferred to the PDP-11 before they can be visualized and saved onto magtape. This transfer is done during the beam pause using CAMPORT. Once the data are on the PDP-11 disk (they fill one complete RKO5J disk!) the microprocessor may continue to readout the experiment while at the same time the PDP-11 processes the spectra accumulated during the spill before.

Fig. 1: Photograph of the data acquisition system


During this year the system has been debugged and extensivly tested. It has taken it's first antiproton data in August 1983. Fig. I shows a photograph of the system.

### 6.3 ACCELERATORS

### 6.3.1 OPERATION OF THE ISOCHRONOUS CYCLOTRON

F. Schulz, H. Schweickert

During the period of this report the machine has been in full operation (see Tables I,II). Even with a large total of available beam time of 7667 h for experiments we could not fulfill all the beam time requirements of our users. At the moment the machine is overbooked by more than $20 \%$. Table II displays the user statistics and the share of beam time between various activities.

| Cyclotron Operational | With internal <br> lon Sources |  | With external <br> lon Sources |  | Total |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |
| For Experiments <br> Beam Development <br> Testing new Components <br> Developments for Isoiope <br> Production | 4759 h | $85.6 \%$ | $2407 \mathrm{~h} *$ | $80.8 \%$ | 7166 h | $83.9 \%$ |
| Total Time of Operation <br> with the Beam on Targels | 5098 h | $61.7 \%$ | 2569 h | $86.2 \%$ | 7667 h | $89.8 \%$ |
| Scheduled shut-down for <br> maintenance, Repair and <br> Installation | 131 h | $2.4 \%$ | 87 h | $2.9 \%$ | 218 h | $3.1 \%$ |
| Unscheduled shut-down | 329 h | $5.9 \%$ | 324 h | $10.9 \%$ | 653 h | $7.6 \%$ |
| Total Shift Time | 5558 h | $100.0 \%$ | 2980 h | $100.0 \%$ | $8538 \mathrm{~h} *$ | $100.0 \%$ |

* Polarized Douteron 1440 h ; $6 \mathrm{~L} \mathrm{l}^{3+}-\mathrm{Ion}$ ( 156 MeV ) 967 h
** The real time of 8760 h is achieved by adding a total of 9 days shut down 24.12.82-2.1.83


## Table I Statistics of the cyclotron from July 1982 to June 1983

The machine operation in the preceding year was rather strenuous because of the increased demand for beam time to produce medical isotopes and the high interest in injected particles ( ${ }^{6} \mathrm{Li}^{3+}, \mathrm{d} \uparrow$ ). To combine both activities, the machine has to be adjusted for high currents from Monday to Thursday, whereas over the weekends the axial injection system has to be implemented. These boundary conditions not only impose a lot of time critical work on the operating group, but sometimes also lead to ineffective beam times for


KiK-Karlsruhe Users

| Instilut für Kernphysik I <br> Instliut tür Kernphysik III <br> Instilut für Nukleare Fesikörper-Physik <br> Institut für Radiochemie <br> Labor für Isotopentechnik <br> Institut für Heibe Chemie <br> Technologie Transier |
| :---: |
|  |  |
|  |  |
|  |  |
|  |  |
|  |  |
|  |  |


| 1436 h | $20.1 \%$ |
| ---: | ---: |
| 768 h | $10.8 \%$ |
| 520 h | $7.2 \%$ |
| 81 h | $1.1 \%$ |
| 72 h | $1.0 \%$ |
| 47 h | $0.6 \%$ |
| 36 h | $0.5 \%$ |
| 2960 h | $\mathbf{4 1 . 3 \%}$ |

External Users

| Freie Universilät Berlin | 849 h | 11.7 \% |
| :---: | :---: | :---: |
| Technische Universiiä̀ München | 665 h | 9.2 \% |
| Max Planck Instilut tür Kernphysik Heidelberg | 589 h | 8.2 \% |
| Universität Erlangen | 435 h | 6.3 \% |
| Universität Bonn | 117 h | 1.5 \% |
| Universitäl Ulm | 83 h | 1.1 \% |
| GSI Darmstadt | 82 h | 1.1 \% |
| IRC Ispra | 71 h | 1.0\% |
| Universität Konstanz | 71 h | 1.0\% |
| Universität Mainz | 36 h | 0.5 \% |
| Medizinische Hochschule Hannover | 31 h | 0.5 \% |
| Universität des Saarlandes | 8 h | 0.1 \% |
| Universität Hamburg | 6 h | 0.1 \% |
| Transurane Karlsruhe | 3 h | 0.1 \% |
| Universität Münster | 3 h | 0.1 \% |
|  | 3049 h | 42.5 \% |
| Akitivierung von Maschinenteilen | 571 h | 8.0 \% |
| Commercial Iodine-123-Production | 392 h | 5.5 \% |
| Commercial Rb-81-Production | 194 h | 2.7 \% |
|  | 7166 h | $100.0 \%$ |

the nuclear physics group because of the short available times for the machine optimization (beam quality and current). The situation will improve at the end of 1983, when the new compact cyclotron takes over the isotope production.

At the end of 1982 it was decided to replace the existing self-excited $r f-s y s t e m$ by a modern driven amplifier system. A 100 kW amplifier for a frequency range of $20-40 \mathrm{MHz}$ was ordered from ZARAT (Warsaw) and will be delivered in late 1983. The amplifier will be connected to the cyclotron by a. 30 m long $50 \Omega$ transmission line。

### 6.3.2 STATUS OF THE KARLSRUHE CP-42H ${ }^{-}$CYCLOTRON (KAZ)

J. Möllenbeck, H. Schweickert

The cyclotron was delivered with a delay of 3 months in August 1982 to Karlsruhe. In the beginning the installation by the TCC-group was very effective but at the end of 1982 it had fizzled out, as TCC found itself in financial problems and could not finish the project. In April 1983 we decided, after a hard battle with the bonding company, to complete the cyclotron installation by ourselves. A11 hard- and software are now operating rather well and since the end of July the beam can be extracted in an energy range from 30 to 42 MeV on the D-side. In this energy range the beam is slightly bent downard in the fringing field, resulting in an angle of $0.37^{\circ}$ with respect to the axis of the beam guiding system. This small angle has been corrected by introducing a 50 cm long electrostatic deflector ( $10 \mathrm{kV} /$ cm) inside the cyclotron vacuum chamber. Figs. 1 to 6 illustrate the activities at the compact cyclotron during the period of report.


Fig. 1 Rolling in the cyclotron magnet into the vault (August 1982)

Fig. 2 View of the $\mathrm{CP}-42 \mathrm{H}$ from the side of the ion source


Fig. 3 First part of the beam guiding system manufactured by Bruker


Fig. 4 View into the vacuum chamber on the $D$-side, showing part of the $D$ system, stripping foil mechanism, and ค1 actrostatic corrector.


Fig. 6 Beam at the position of the iodine production target.

### 6.3.3 ION SOURCE DEVELOPMENTS

V. Bechtold, L. Friedrich, H.P. Ehret, J. Ka1tenbaek, L. Wiss,
P. Ziegler

Three external ion sources are operated at the Karlsruhe cyclotron. The Penning source for $\mathrm{Li}^{3+}$ ions and the Lambshift source are extensively used, delivering some tens of nA of $26 \mathrm{MeV} / \mathrm{A}$ particles for nuclear physics experiments. The ECR-source HISKA is designed to deliver highly charged light ions, although at present only fully stripped light ions can be accelerated
by the cyclotron. A sudden rise of the liquid He evaporation rate of the superconducting coils allowed only short times for testing, but nevertheless progress was made concerning quality and charge state distributions of the extracted beam. In order to extract the ions along the axis of the longitudi nal magnetic mirror field the extraction system was made movable in the radial direction and adjusted observing the electron flow coming from the plasma. In the new construction, more care was taken over the pumping speed in the extraction region as well as in the charge state analysing system, for which a new Wien filter with permanent magnets inside a vacuum chamber was built. This resulted in a remarkable rise of the mean charge of the extracted $N$-ions from 2.8 to 3.8 . Charge state spectra (Fig.1) show that now most of the extracted ions have only two electrons.


Fig. 1 Charge state spectra of the extracted beam from HISKA operated with $\mathrm{N}_{2}$ and $\mathrm{CO}_{2}$

The ECR-technique delivers high plasma density with low neutral background. This allows the extraction of low velocity deuteron beams of about 10 mm diameter. This and the magnetic field might be favourable to overcome the space charge problems in the Lambshift source. First experiments in this field led one to expect a substantial gain of intensity, but it is not believed that one can reach the intensity of the atomic beam source. So it was decided to replace the Lambshift source by an atomic beam source.

Experiments with Li-vapour in the ECR source p-HISKA demonstrated that at least $10 \%$ of the extracted Li-ions were in the $3+$ state. A special ECR
source for Li-ions has been constructed and will be assembled in the near future.


Fig. 2 LISKA, the ECR source for $\mathrm{Li}^{3+}$ ions uses an external hexapole to enable heating of the walls to achieve the right vapour pressure of Li . At 1 east $30 \mu \mathrm{~A}$ of $\mathrm{Li}^{3+}$ are expected.

### 6.3.4 ON THE POSSIBILITY TO CONVERT THE KARLSRUHE ISOCHRONOUS CYCLOTRON TO AN ENERGY VARIABLE MACHINE

J. Schwabe, H. Schweickert

In context with the discussions of the future use of the Karlsruhe Isochronous Cyclotron for nuclear physics it turned out, that for any postaccelerator project $(1,2)$ a better flexibility of the injector is desirable. The actual machine is a simple FFAG-cyclotron for $Z / A=1 / 2$ particles with a

## 

| Particle | Energy in MeV/Nucleon |  |
| :--- | :--- | :--- |
| $p, p^{\dagger}$ | $40-84$ | $h=2$ |
| $p, d^{1}, Z / A=1 / 2$ | $18-35$ | $h=3$ |
| ${ }^{3} \mathrm{H}^{+}$ | $10-15$ | $h=3$ |
| ${ }^{3} \mathrm{He}^{2+}$ | $32-60$ | $h=2$ or 3 |
| ${ }^{6} \mathrm{Li}^{2+}$ | $11-15$ | $h=4$ |
| ${ }^{12} \mathrm{C}^{5+}$ | $18-24$ | $h=3$ |
| ${ }^{14} \mathrm{~N}^{6+}$ | $18-24$ | $h=3$ |
| ${ }^{16} \mathrm{O}^{7+}$ | $19-26$ | $\mathrm{~h}=3$ |
| ${ }^{20} \mathrm{Ne}^{9+}$ | $19-28$ | $\mathrm{~h}=3$ |


| Particle | I $[\mu A]$ <br> extracted | Ion Source |
| :--- | :--- | :--- |
| $p, d$ |  |  |
| $\alpha,{ }^{3} \mathrm{He}^{2+}$ | $80-100$ <br> $30-50$ | Internal PIG |
| $Z / A=1 / 2$ | $0,01-3$ | ExternalHISKA |
| $Z / A=<1 / 2$ | $0,2-2$ | Atomic Beam <br> (ordered) |
| $d \uparrow, p \uparrow$ | $1-5$ |  |



Fig. 1: Schematic diagram of the planned modification of the cyclotron, showing the new pole- tips and the resonators. All other components can be taken over from the "old" machine.
maximum energy of $26 \mathrm{MeV} / \mathrm{Nuc} 1 \mathrm{eon}$. The main design goals for the modified machine are:

1. Acceleration of protons up to an energy of $\geqq 70 \mathrm{MeV}$.
2. Possibility for the acceleration of only partly stripped light ions, giving us at least one order of magnitude larger extracted currents for Lithium up to Neon ions.
3. Energy variability of a factor of two.

In the last monthsthe theoretical basis for such a modification has been worked out. It could be shown, that by only exchanging the pole tips and the rf-system the above mentioned goals can be fulfilled.

The original straight Thomas sectors have to be replaced by spiraled ones. The spiral starts at 50 cm and is increased slowly, according to the beam dynamics calculations, up to $32^{\circ}$. The three-fold symmetry is maintained.

The rf-system has to be changed from the self-excited configuration, using a vertical transformation line to couple the 3 dees in the center, to a configuration where each dee is separately coupled with a special feeder to a common rf-amplifier. The maximum power level is 90 kW for 3 dees including the feeders and the coupling system for a D-voltage of 50 kV . For the second harmonic operation the second dee gets a phase shift of $180^{\circ}$. The frequency range of $25-40 \mathrm{MHz}$ is adjusted by two vertical trimmer plates with a maximum lift of 15 cm .
(1) E. Poth, this Annual Report, p. 192
(2) H. Schweickert, Ninth international conference on cyclotrons and their applications, Caen University, France, 1981

### 6.3.5 ISOTOPE PRODUCTION FOR MEDICAL APPLICATIONS

P. Kernert, J. Peters, H. Schweickert

Since the isotope production at the new compact cyclotron (KAZ) will start in autumn 1983, the installation of the isotope production targets at this facility were started while continuing the routine production with the Isochronous Cyclotron. The status of the various activities is as follows: ${ }^{123}$ I The demand for ${ }^{123}$ I again increased during the period of this report. In total, 846 batches were produced with a total activity of 38 Ci . They were sent to 10 hospitals and to 3 pharmaceutical firms where the labelling of more complex radiopharmaceuticals like hippuran, amphetamine, fatty acids etc. is carried out.

An irradiation facility for ${ }^{123}$ I has been built at the KAZ. The target holder and the transportation system are fully compatible with the existing system. So, the already existing hot cell for dry-distillation and distribution needs no modification.


Fig. $1{ }^{123}$ I irradiation facility at the compact cyclotron
${ }^{81} \mathrm{Rb} /{ }^{81 \mathrm{~m}_{\mathrm{Kr}} \text {-Generator }}{ }^{81} \mathrm{Rb} /{ }^{81 \mathrm{~m}_{\mathrm{Kr}}-\text { generators arebeing produced two times per }}$ weak since November 1982. In total, 216 generators ( $\cong 6.5 \mathrm{Ci}{ }^{81} \mathrm{Rb}$ ) have been delivered for medical use. Additionally, 4 generators were produced for technical tracer experiments in nuclear reactor safety studies, each of them containing about 150 mCi of ${ }^{81} \mathrm{Rb}$.

Some changes in the cooling and target handing system improved the stability of the entrance foil. The target assembly for the compact cyclotron
is completed and installation will start in August 1983.
${ }^{201}$ T1 To improve the chemical processing and to assure the radiochemical purity 12 Tl targets were irradiated and sent to Hoechst. One of the se was irradiated at the JULIC-cyclotron of the KfA Juilich. Furthermore, seven batches were produced to develop an independent chemical process at the cyclotron laboratory in order not to sell the ${ }^{201} \mathrm{~Pb}$ but the final product 201

T1. Thus, only the pharmaceutical control, the distribution and the marketing is done by the industrial partner.

The production target system is being installed at the KAZ facility. In order to ensure the reliability, two identical targets can be connected to the beam-1ine by remote control.


Fig. 2
${ }^{201}{ }^{\mathrm{Pb}}$ production target and standby target

### 6.3.6 STATUS OF THE ELECTRON COOLING DEVICE FOR LEAR

L. Hütten, H. Poth, A. Wolf, H. Haseroth ${ }^{+}$, C.E. Hi11 ${ }^{+}$, (1)

The construction of the LEAR electron cooler has advanced considerably in the past year. A description of the system has been given in the last annual report.

The present set-up is shown in Fig. 1. The high voltage platform is installed next to the cooler magnet. A subassembly of the vacuum system has been set-up for the determination of the achievable vacuum conditions with cold and hot electron gun and for the study of the performance of a


Fig. 1 LEAR electron cooler
new type of NEG-pumps (2). The measurements of the main magnet field properties are completed and evaluated (3).

The LEAR electron cooler will be completely controled by an LSI 11 computer, CAMAC data acquisition and Fiber optics transmission between the high voltage platform and the ground level. This system is to be completed in the near future. During the next months an electron beam will be generated and its properties will be studied using the subassembly. The full system will be assembled by the end of 1983.
(1) contributed paper to the $12^{\text {th }}$ Int. Conf. on High-Energy Accelerators Fermilab (1983)
(2) This Annual Report, p. 191
(3) This Annual Report, p. 189
$+\quad$ CERN, PS Division

### 6.3.7 THE MAGNETIC FIELD PROPERTIES OF THE LEAR ELECTRON COOLER

L. Huitten, H. Poth, A. Wolf

The field of the magnet assembly to be used for electron cooling at LEAR (solenoids and toroids from Initial Cooling Experiment ICE) has been investigated by careful field mapping. The properties of the longitudinal field guiding the electron beam deserve attention under the following aspects:

- transverse electron temperature: electron motion through an inhomogenous field excites cyclotron rotation in the transverse degrees of freedom;
- electron-proton alignment: the benefit that in practice can be gained from the strong magnetic cooling force depends on the straightness of the field lines in the cooling region: minimal rms variation of the field angle along the common trajectory of electrons and protons is crucial.

The guiding field was measured in a volume with a cross section of $3 \times 3 \mathrm{~cm}^{2}$ around the electron beam axis which is 4.5 m long and bent twice, in the toroids. A Hall plate field scanner designed for the mapping of dipole magnets at the CERN EP division (1) was slightly modified and used to obtain separate field maps of all cartesian components, in ten rectangular boxes covering the forementioned measuring region. In total, the field components were measured at appr. 26000 points. For merging these maps into a consistent overall field table, a special procedure has been developed in this work that includes careful survey and the correction of errors inherent to Hall-probe measurements (2).

Fig. 1 shows the field components longitudinal (a) and transverse ( $b, c$ ) to the electron beam along the beam axis, and on trajectories displaced from the beam axis diagonally by appr. 3 cm in the transverse plane. The shaded areas in Fig. 1 illustrate the non-uniform transverse field shape due to variations of the longitudinal field strength. It demonstrates the improvement we achieved by mounting additional coils at both ends of the cooling solenoid where, otherwise, a field shape more similar to that measured at points 2 and 3 would appear.

Next steps in employing the results of the magnetic measurements are:


Fig. 1: Magnetic field components along the electron beam. S = path length along the beam.

- design of correction coils remoying uniform transyerse components in the cooling solenoid,
- calculation of electron trajectories for a drift motion in the measured field.
(1) O. Runolfsson: Magnet Measurement in High-Energy Physics Experiments, Inivited paper, $6^{\text {th }}$ Int. Conf. on Magnet Technology, Bratislava (29.8.-2.9.1977)
(2) L. Hütten, H. Poth, A. Wolf, The magnetic field properties of the LEAR Electron Cooler, KfK report (in preparation)


### 6.3.8 ULTRA HIGH VACUUM TEST MEASUREMENTS ON THE LEAR ELECTRON COOLER

L. Hütten, H. Poth, A. Wolf

As the $\bar{e}$-cooler will be a highly gas producing device special effort has to be made in order to meet the extraordinary vacuum conditions of the LEAR, that is an average pressure of about $3 \times 10^{-12} \mathrm{mbar}$ in the beam area. This and the fact that there will be no access for a vacuum pipe into the $e^{-}$-gun part where the greatest part of the gas produced by the cathode has to be trapped focussed the vacuum-pumping on the application of getters. A special device was made with some $51 / \mathrm{sec}$ getter pump units of the same Al-Zr getter material that is to be used in LEP (non evaporable getter NEG). The NEG getter traps the gas molecules of the ultra high vacuum selectively on its surface until saturation which stops the pumping process; that means increase of the partial pressure of the gas. The surface has to be activated again by heat treatment which provokes the diffusion of the gas molecule into the bulk or releases it into the space depending on the sort of gas trapped. Thus the physics of gettering implies that getters have to be activated and regenerated from time to time in a strictly observed process that has to be optimized by experience.

The construction of the experimental set up is shown in Fig. 1. The other pumping facilities are a $1001 /$ sec Varian ion getter pump and a $4001 / \mathrm{sec}$ turbomolecular pump that was separated from the recipient by a UHV-valve after preevacuation and bake-out cleaning.

The first part of the measurements was to provide the fundamental data for the calculation of the conductance and effective pumping speed in various points as well as the degas-rates of the cathode. They were made without NEG integrated. The degas-rate of the cathode at $1060^{\circ} \mathrm{C}$ was found to be about $2 \times 10^{-6} \mathrm{mbar} 1 / \mathrm{sec}$, the greatest part of it being CO . In the second part of the measurements NEG was integrated to study its influence upon the vacuum.
After activation a final pressure of $6 \times 10^{-12} \mathrm{mbar}$ was reached with ion pump in operation and cold cathode. A long term run with cathode heated was made with various pumping conditions. The results for the final pressure are:

- with cold cathode, NEG, and ion pump $6 \times 10^{-12} \mathrm{mbar}$
- with hot cathode, NEG, and ion pump $2,6 \times 10^{-10} \mathrm{mbar}$
- with hot cathode and NEG $2 \times 10^{-9}$ mbar

Mass spectra were taken to measure the components of the degas flow and the saturation behaviour of the NEG. Further measurements with e beam under final operation conditions are being prepared and will concern the optimal NEG operation as well as final vacuum equipment to match LEAR.


### 6.3.9 CONSIDERATIONS ABOUT AN ELECTRON COOLED STORAGE RING FOR MEDTUM ENERGY (POLARIZED) PROTQNS AND LIGGT IONS

## H. Poth

The recent progress in accelerator technology (particularly in beam cooling and the application of internal cluster beam targets), the evidence for dibaryon resonances and the need for high precision experiments in medium energy and nuclear physics initiated a first study on a cooler and storage ring for polarized protons, deuterons and for light ions. We consider the basis for such a project at the KfK very favourable for the
following reasons:

- The KfK-cyclotron could be used after slight modifications as an injector. The developments of different types of ion sources of the cyclotron group are recognized. Applying this experience high flux beams of polarized protons, deuterons and light ions could be anticipated. Moreover there exists a profound knowledge in accelerator physics at the KfK.
- We have collaborated with CERN during the ICE-electron cooling experiments. The gained experience was deepened during the development of LEAR at CERN. Presently we construct in collaboration with CERN the electron cooler for LEAR. The technical and physical boundaries in ICE and LEAR are very similar to those arising from the scheme discussed here.
- The cluster beam technology was pioneered in IKVT at KfK and served as a model for many internal cluster beam targets applied in various accelerators today. A configuration of an internal target for LEAR on the basis of such a cluster beam was worked out. The implications of the operation of this internal target were studied in detail in close collaboration with CERN. A similar target could be used in the machine discussed here. It would have to operate under similar but less restrictive conditions as in LEAR.
- There is a large community of present cyclotron users from the KfK and outside institutes interested in such a facility. There exists a strong activity in nuclear and medium energy physics at the KfK.

Such a machine would open the door to a wide field of high precision physics and would enable unique experiments.

The concept for a cooler and storage ring had previously been discussed under the hypothesis that the ICE-ring could be used (1) and recently in the context of a more general ring configuration (2). The basic features of the scheme are summarized in the following,

The ring has the task to collect (polarized) protons, deuterons or light ions at the cyclotron end-energy ( 80 MeV for protons, $35 \mathrm{MeV} / \mathrm{N}$ for $q / M=1 / 2$ ions). During collection the circulating beam is cooled continuously by electrons allowing to reach the space charge limit. Thereafter the beam is accelerated to the desired working energy, the internal targets are activated and the experimental phase is started. The beam is continuously cooled. From physics arguments an upper operation energy of 1.3 GeV would be desirable. The experimental phase would range from a fraction of a minute to a few hours depending on stored projectile and the thickness of the internal targets. At the end the magnetic field will
be decreased and the filling is restarted. The collection and acceleration phase would last only a few seconds.

As projectiles (polarized) protons, deuterons and light ions could be collected, accelerated and stored. A high degree of vertical polarization for protons could be expected. An external neutron beam emerging tangentially out of the ring from an internal target could be set-up. The (polarized) neutrons would be produced through ( $p, n$ ) and $(d, n)$ reactions. By detection of the reaction partner they could be tagged.

Molecular clusters, (polarized) atomic and jet beams would be used as internal targets yielding a target density in the range between a few $1 \mu \mathrm{~g} / \mathrm{cm}^{2}$ and $5 \mathrm{pg} / \mathrm{cm}^{2}$ or less. A tentative ring parameter 1 ist is given in Table 1. Similar concepts are presently realized at the Indiana cyclotron in Bloomington and in Uppsala and had been discussed in the KfK, Jiilich. Some of the most important virtues of the scheme discussed here are again summarized in the following:

- The use of electron cooling in combination with accumulation would allow to built up dense stored particle beams even for low current cyclotron beams (e.g. $p_{\uparrow}, d_{\uparrow}, t$, light ions).
- The combination of electron cooling and internal target operation would allow for the most efficient use of the accumulated beam and enable high resolution experiments with high luminosity. The beam target interaction volume could be made very small.
- The application of thin internal targets would allow the detection of heavy and slow reaction products and would kender target windows unnecessary, Secondary interactions would not occur.
- A high degree of beam and target polarization could be achieved, with the possibility of a rapid change of the target polarization direction and a change of beam polarization from cycle to cycle.
- A high energy neutron beam could be set-up. Through polarization transfer reaction the neutrons could be polarized. By detecting the reaction product the neutron could be tagged.
- Rare isotopes could be used as internal targets since only little material is needed for atomic beams.

Table 1: Tentatiye parameters of an electron cooled storage ring

(1) A. Citron, J. Gspann, W. Kluge, H. Poth, G. Rebe1, G. Schatz, H. Schweickert, Ch. Weddigen, unpublished notes (1981/83)
(2) KfK-Cyclotron users meeting, Karlsruhe (June 1983)

### 6.4 APPLICATIONS <br> 6.4.1 ELEMENTAL ANALYSIS OF SINGLE CELLS OF ORGANIC MATTER BY PROTON INDUCED X-RAYS

D. Heck, KfK-report 3399 (1982)

Trace element analyses in single cells of organic matter may be performed with ion microbeams. The method uses proton induced X-ray emission, the principles of which are presented; the Karlsruhe micro beam setup is described. The requirements for sample preparation are discussed in detail. The mass of the irradiated sample volume is determined by Rutherford backscattering. The calibration of the elemental concentrations is performed by comparison with a reference standard of known composition. The efficiency of the described method is demonstrated at the example of erythrocytes (red blood cells).

### 6.4.2 ELEMENTAL ANALYSES PERFORMED WITH THE KARLSRUHE NUCLEAR MICROPROBE

D. Heck, E. Rokita ${ }^{+}$, IEEE Trans. Nuc1. Sci. NS-30 (1983) 1220

Proton induced X-rays have been used in the Karlsruhe nuclear microprobe to determine the content of some trace elements in normal and cancerously degenerated human stomach mucousa tissue. A 3 MeV proton beam produced in a single stage Van de Graaff accelerator is focussed down to $3 \mu \mathrm{~m}$ at 600 pA and swept across the specimen in a meandrous pattern. The local mass density of the specimens is determined by the yield of elastically scattered protons, wich are registered in a surface barrier detector. Erythrocytes served as test objects to demonstrate the feasibility of the procedure. Specimens of stomach mucousa were prepared by two different methods of fixation. After irradiation the specimens were stained to identify the histological structure. All specimens show rather low structure in elemental concentrations except for iron. No significant shifts in the trace element contents between normal and concerous stomach mucousa could be observed.
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Jage11onian University, Institute of Physics, Cracow, Poland

### 6.4.3 DOSE DISTRIBUTION AND ENERGY DEPOSITION SPECTRA OF $\pi^{-}$MESON BEAMS FOR BIOMEDICAL APPLICATIONS

G. Büche

Distributions of absorbed dose in depth and in lineal energy have been worked out using our Monte Carlo program named PIONDOSE. This computer code is based on a large set of experimentally approved input data and uses only a few model approximations. It has been the aim to get results which cannot be obtained from conventional dosimetry or even microdosimetry and which are essential for therapy planning procedures.

Fig. 1 shows a series of depth dose curves for a pencil beam in a water phantom. The absorbed dose as a function of depth has been integrated over $\pm 10 \mathrm{~g} \mathrm{~cm}^{-2}$ in both directions perpendicular to the beam axis. In particular, the contributions of pions, (secondary) protons and ${ }^{4}$ He-ions from reactions in flight as well as from nuclear absorption of stopped pions are given. According to their short ranges the peak of He-ions is equivalent to the distribution of stopped pions and shows a shift existing between the maxima of absorbed doses from pions and from densely ionizing secondaries.

Fig. 2 shows energy deposition spectra calculated for various depths along the beam axis. Shortly behind the phantom surface ( 2.5 cm ) the spectrum shows a sharp peak which is generated from fast pions crossing the sensitive volume ( $2 \mu \mathrm{~m}$ in diameter) of the spectrometer and a shoulder from $\delta$-electrons. Above $10 \mathrm{keV} \mu \mathrm{m}^{-1}$ the contributions from pion reaction products are maximum. With increasing depth (up to 16.5 cm ) the peak from pions is shifted to the right and the contributions from $\delta-e l e c t r o n s$ as well as from secondaries decrease. The following pictures of Fig. 2 cover the stopping region of pions. The Bragg peak is at 19.5 cm ; the spectra at 18.3 and 20.9 cm result for $\pm 1 / 2$ FWHM and the positions 16.5 and 22.5 cm are related to $1 / 10$ of the maximum value of the stopping distribution. This sequence of $y$-spectra shows clearly the variation of ionization density over the Bragg peak which is correlated with a variation of radiobiological effectiveness (RBE) of absorbed dose. Together with the decomposition into contributions from various secondaries ( $p, d, t,{ }^{3} \mathrm{He}$, ..., n, recoils; not shown in this context) a series of spectra like this one is the physical part of the estimate of radiation quality in therapy planning in the case of negatively charged pions.

A third aspect of the calculations can be read from Table 1. The total absorbed dose and the part from the kinetic energy of the pions is given as a function of the phantom material integrated over the stopping region of

Table 1: The absorbed dose from secondaries related to the dose from pions in the Bragg peak for various phantom materials

| Phantom material $(i=0, H, T E, C)$ | $\mathrm{H}_{2} \mathrm{O}$ | pig muscle | re-plastic | $\left(\mathrm{CH}_{2}\right)_{n}$ |
| :---: | :---: | :---: | :---: | :---: |
| Number of transported ${ }^{-}$ | $10^{4}$ | $2 \cdot 10^{3}$ | $2 \cdot 10^{3}$ | $10^{4}$ |
| $\begin{aligned} & \text { Chemical composition } \\ & H: C: 0: N \end{aligned}$ | 2:0:1:0 | $\begin{array}{c:c} 34,24: \\ : 12,94 & 37 \\ : 1 \end{array}$ | $\begin{gathered} 39,23: 24,62 \\ : 1,42: 1 \end{gathered}$ | 2:1:0:0 |
| Capture ratio 0: (C4N) | 1:0 | 0,836:0,164 | 0,080: 0,920 | 0:1 |
| Total dose $D_{\text {tot }} \underbrace{\text { integrated }}_{\text {over }}$ | 27,571 $\frac{\mathrm{HeV}}{\mathrm{cm}}$ | $29.398 \frac{\mathrm{MeV}}{\mathrm{cm}}$ | $32,590 \frac{\mathrm{MeV}}{\mathrm{Cm}}$ | $31,394 \frac{\mathrm{MeV}}{\mathrm{cm}}$ |
| Pion dose $\mathrm{D}_{\pi} \int^{\text {stopping }} \begin{aligned} & \text { region } \\ & \text { res }\end{aligned}$ | 11,267 $\frac{\mathrm{HeV}}{\mathrm{cma}}$ | 11,915 $\frac{\mathrm{MeV}}{\mathrm{Cm}}$ | $11.724 \frac{\mathrm{MeV}}{\mathrm{cm}}$ | $11.076 \frac{\mathrm{MeV}}{\mathrm{CmI}}$ |
| $\left(D_{\text {tot }}-D_{n}\right) / D_{\pi}=D_{\text {sek }} / D_{\pi}=D_{i}$ | 1,447 | 1,467 | 1,780 | 1,834 |
| $\underset{\substack{\text { overestimation } \\ \text { of absorbed }}}{ } \frac{1+\mathrm{D}_{\text {IE }}}{1+\mathrm{D}_{\mathrm{H}}}=$ dose | 0,992 | 1 | 1,127 | 1,149 |



Fig. 1: Distribution of absorbed dose in depth for a pencil beam of $\pi^{-}$in a water phantom. The contributions from the kinematic energy of pions and from two secondaries ( $p,{ }^{4} \mathrm{He}$ ) are given separately.


Fig. 2: Distributions of absorbed dose in lineal energy for various positions along the axis of a pencil beam of negatively charged pions.
pions. It is clearly revealed that the absorbed dose from heavy secondaries $D_{\text {sek }} / D_{\pi}=\left(D_{\text {tot }}-D_{\pi}\right) / D_{\pi}$ increases with increasing carbon content of the phantom material. In cases of application an absorbed dose that is expected within e.g. muscle tissue is measured using a dosimeter which is made usually out of tissue equivalent (TE) plastic material. According to our calculation the absorbed dose in pig muscle tissue is overestimated by about $13 \%$. A correction factor like this for differing target and dosimeter materials can not be obtained by other methods than a calculation.

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6th Internat. Balaton Meeting on High Energy Nuclear Physics, Balatonfuered, (Hungary), June 6-11, 1983

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The status of the $\sigma N$ - curve and the $s$-process neutron density
 detector

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[^1]:    * DESY 83-034, submitted to Zeitschrift für Physik C

[^2]:    * DESY 83-019, submitted to Phys.Lett. B

[^3]:    *It should be noted here that the $1 \sigma$ limit for $\nu_{\tau}$ derived from $e^{+} e^{-} \rightarrow \tau^{+} \tau^{-}$ total cross section measurements (2) is as large as $\pm 6.5$, mainly due to the systematic errors ( $\sim 8 \%$ ) involved.

